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# Relationship between the sub-micron fraction (SMF) and fine-mode fraction (FMF) in the context of AERONET retrievals

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Abstract. The sub-micron (SM) aerosol optical depth (AOD) is an optical separation based on the fraction of particles below a specified cutoff radius of the particle size distribution (PSD) at a given particle radius. It is fundamentally different from spectrally separated FM (fine-mode) AOD. We present a simple (AOD-normalized) SM fraction versus FM fraction (SMF vs. FMF) linear equation that explains the well-recognized empirical result of SMF generally being greater than the FMF. The AERONET inversion (AERinv) products (combined inputs of spectral AOD and sky radiance) and the spectral deconvolution algorithm (SDA) products (input of AOD spectra) enable, respectively, an empirical SMF vs. FMF comparison at similar (columnar) remote sensing scales across a variety of aerosol types.

SMF (AERinv-derived) vs. FMF (SDA-derived) behavior is primarily dependent on the relative truncated portion ( $\varepsilon_c$ ) of the coarse-mode (CM) AOD associated with the cutoff portion of the CM PSD and, to a second order, the cutoff FM PSD and FM AOD ( $\varepsilon_f$ ). The SMF vs. FMF equation largely explains the SMF vs. FMF behavior of the AERinv vs. SDA products as a function of PSD cutoff radius ("inflection point") across an ensemble of AERONET sites and aerosol types (urban-industrial, biomass burning, dust, maritime and a mixed class of Arctic aerosols). The overarching dynamic was that the linear SMF vs. FMF relation pivots clockwise about the approximate (SMF, FMF) singularity of

(1, 1) in a "linearly inverse" fashion (slope and intercept of approximately  $1 - \varepsilon_c$  and  $\varepsilon_c$ ) with increasing cutoff radius. SMF vs. FMF slopes and intercepts derived from AERinv and SDA retrievals confirmed the general domination of  $\varepsilon_c$  over  $\varepsilon_f$  in controlling that dynamic. A more general conclusion is the apparent confirmation that the optical impact of truncating modal (whole) PSD features can be detected by an SMF vs. FMF analysis.

### 1 Introduction

Anderson et al. (2005) noted the "decades-old observation that aerosol mass generally consists of two modes: (1) a mechanically produced coarse mode (CM) and (2) a fine mode (FM) produced by combustion and/or gas-to-particle conversion". Typical CM examples include wind-eroded desert dust and sea salt, while FM aerosols tend to be dominated by biomass-burning smoke and anthropogenic and biogenic (ABF) fine particles (the latter classification being proposed, for example, by Lynch et al., 2016). Given this speciated physical reality (and notwithstanding the known chemical and microphysical internally mixed changes that occur in aerosol properties as they are transported through the atmosphere), there is a certain level of justification in treating different species of aerosols as independent FM and CM par-

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ticle size distributions (PSDs). The degree of rigor in this modal paradigm is, in fact, the subject of model analyses that compare prescribed bulk aerosol PSDs with sectional (binned) PSDs (see, for example, Mann et al., 2012, for the case of size-referenced PSDs and Kodros and Pierce, 2017, for the case of mass-referenced PSDs).

AERONET inversions (AERinv), derived from input spectral AODs and solar almucantar radiances (Dubovik and King, 2000), provide a comprehensive suite of microphysical and optical products (at approximately hourly temporal resolutions). The general tendency towards (particle-volume) PSD bimodality is notably evidenced in retrieved AERinv PSDs (see, for example, Fig. 1 of Dubovik et al., 2002; Fig. 11a of Eck et al., 2009; and Fig. 3 of AboEl-Fetouh et al., 2020). A bimodal optical representation of that PSD bimodality (without the requirement of having to specify the PSD shape of each mode) can largely determine the measurable (low spectral order) optical behavior of remote sensing data (O'Neill et al., 2001, and references cited therein). O'Neill et al. (2003) employed spectral AODs (sampling intervals  $\sim$  3 min) as an input to their spectral deconvolution algorithm (SDA) to separate aerosols into (a) extensive (quantity dependent) FM and CM AOD components ( $\tau_f$  and  $\tau_c$ ) of the total AOD  $(\tau_a)$ , (b) intensive (quantity-independent) spectral derivatives of  $\tau_f$  ( $\alpha_f$ ,  $\alpha_f'$ ,  $\alpha_f''$ , etc.) and  $\tau_c$  ( $\alpha_c$ ,  $\alpha_c'$ ,  $\alpha_{\rm c}^{\prime\prime}$ , etc.) as components of the total AOD spectral derivatives  $(\alpha, \alpha', \alpha'', \text{ etc.})$  and (c) semi-intensive FM and CM fractions (FMF and CMF represented, respectively, by  $\eta = \tau_f/\tau_a$ and  $1 - \eta = \tau_c/\tau_a$ ). If aerosols can be viewed as independent coarse and fine modal features, then this purely spectral technique acts to separate those modal features in an optical sense.

The SMF (sub-micron fraction) is a microphysically determined alternative to the optically determined FMF: the division into fine and coarse components is effected by an explicit separation of the PSD at a cutoff radius ( $r_0$ ) that typically ranges from  $\sim 0.4$  to 1.1 µm across different types of aerosol volume sampling instruments.<sup>2</sup> SMF separation represents a moderate but significant difference relative to the optically based FMF separation. Anderson et al. (2005) describe the SMF as "an operational definition ... to distinguish it from the theoretical concept of fine-mode fraction, FMF". Because the AERinv algorithm incorporates a cutoff radius division of the retrieved PSD into sub- and supermicron parameters (Dubovik and King, 2000), it is actually an SMF approach. Comparisons with the AERONET SDA product provide a unique opportunity to empirically ana-

lyze the SMF vs. FMF approaches at similar remote sensing (columnar) scales and across a global variety of aerosol types.

The literature on the direct comparison of the SMF vs. FMF retrievals is sparse and oftentimes at the margins of significance. The significance problem relates to the fact that one may be forced to extract relatively subtle microphysical and optical changes in the face of SMF and FMF variations that are often limited in range (lack of extensive-parameter aerosol variation and/or types for example). The sparsity of SMF vs. FMF investigations relates to limitations such as the rarity of experiments designed for such a comparison or the sparsity of data sets whose modes can be readily separated (by source information and/or chemical identification for example).

In the most controlled experiments, volume-sampled SMF estimates are compared with FMF retrievals derived from applying the SDA to volume-sampled, three-channel CRD (cavity ring-down) or three-channel nephelometer "spectra" (Atkinson et al., 2010; Kaku et al., 2014 (Ka); Atkinson et al., 2018). In general we find, as expected, the SMF to be ≥ FMF in Atkinson et al. (2010) and Ka: this is notably true at the near-IR (700 nm) wavelength in Ka (ACE-Asia data of their Fig. 2).³ The cases where this is not true (the T1 "ext\_sum" results of Atkinson et al., 2018) are likely attributable to the spectral sensitivity of the FMF: this is most severe in the case of just three channels (for example, Ka's VOCALS 5% calibration correction of one 450 nm nephelometer channel transformed a case of SMF being substantially < FMF to a corrected case of SMF ~ FMF).

Lesser constrained experiments involve multi-altitude, volume-sampled SMF estimates (from nephelometer scattering coefficient and absorption coefficient devices) compared with layer and column SDA-derived FMF estimates from multi-altitude AOD and  $\triangle$ AOD spectra acquired with an airborne sun photometer (for example, Gassó and O'Neill, 2006, and Shinozuka et al., 2011, respectively, represented by the acronyms GO and SHN). In these cases, the lesser degree of experimental control (associated with multi-altitude flights) was somewhat offset by a more generous number of sun photometer spectral bands (four bands from the UV to the NIR in the GO case and five bands near the five bands employed in the AERONET SDA in the SHN case). The GO results were fairly coherent with SMF/FMF expectations (especially when the layered estimates were added to obtain columnar estimates) and quite marginal for SHN with a large spread of near-unity points about the SMF=FMF line (a

<sup>&</sup>lt;sup>1</sup>The examples in these papers can include the apparent presence of more than two modes (notably what often appear to be two CM sub-modes). Such cases are discussed below.

<sup>&</sup>lt;sup>2</sup>The expression "sub-micron" is generally "defocused" (relative to a literally exact value of  $1.0 \, \mu m$ ) to encompass this approximate range of  $r_0$  values. It is worth noting that the in situ community almost universally refers to diameter rather than radius.

<sup>&</sup>lt;sup>3</sup>For which the fixed cutoff radius for the optically smaller particles (and thus more optically active in the sense of more strongly attenuating) results in a greater relative contribution to the FM AOD (by "optically smaller", we mean that the ratio of particle size to wavelength has decreased because the wavelength has increased). The optically smaller particles are more optically active as they are ascending the right side of the anomalous diffraction peak (see, for example, O'Neill et al., 2005).

situation with practically no significant SMF or FMF range and hence little relevant testing of their relationship). Analyses of the FMF vs. SMF relationship comparing satellite estimates of FMF<sup>4</sup> with, for example, airborne estimates of SMF (Anderson et al., 2005) or coastal/island AERinv estimates of SMF (Kleidman et al., 2005) have proven difficult given the spectrally sensitive nature of the few bands employed in satellite-derived estimates of FMF and all the possible sources of incoherencies between the satellite retrievals and the ground-based or airborne estimates of SMF (differences in spatiotemporal sampling volumes for example).

A simple (approximate) relationship between the SMF and FMF will be presented below. We seek to demonstrate that the AERinv-derived value of SMF and the SDA-derived value of FMF are largely linked by that simple relationship and that fitting parameters extracted from their empirical comparison yield insight into their fundamental optophysical dynamics. We argue that the similar columnar scales as well as the diversity of AERONET aerosol types shared by the two retrievals facilitates the analysis of their second-order intensive-parameter relationship. It is emphasized that the AERinv and SDA algorithms were developed independently and share no explicit algorithmic links: what they do share is that the four spectral-AOD inputs to the AERinv represent four of the five spectral AODs employed as input to the SDA. What they do not share is the almucantar (angularly variable) radiance input employed in the AERinv retrieval. It is notable that the AERONET AOD is accurately measured with an estimated uncertainty of  $\sim 0.01$  and 0.02 in the visible/NIR and the UV, respectively (Eck et al., 1999), for an overhead sun (solar air mass of M=1). This data quality enables both algorithms to, for example, retrieve notably consistent intensive parameter information such as particle size. The consistency is also attributable to the AERinv requirement that the retrieved AODs be within 0.01 of the measured AODs, while the second-order spectral fit to the AODs employed in the SDA has similar constraints (see, for example, Fig. 4 of O'Neill et al., 2001).

### 2 Theoretical considerations

### 2.1 Size-cutoff integration versus modal integration

Figure 1 (generated from a simple lognormal fit to a sample AERinv particle-volume PSD retrieval<sup>5</sup>) illustrates the theoretical mechanical/optical framework of this paper. That fit was applied to lend an air of empirical relevance to this the-

oretical section: the discussion presented here is otherwise of a general nature and is not intended to represent an algorithmic step of the AERinv retrievals (or of the SDA for that matter). The FM and CM PSDs are represented by the red and blue lognormal curves, respectively. The SMF-type cutoff radius of  $r_0$  is represented by the black dashed vertical line where the associated optical and microphysical bimodal quantities are computed from both FM and CM PSDs to obtain sub-micron parameters to the left of  $r_0$  and super-micron parameters to the right of  $r_0$ . The rest of the parameters in Fig. 1 are defined immediately below as we develop the theoretical framework.

Letting prime variables refer to integrations which are carried out over size regimes of  $(0, r_0)$  and  $(r_0, \infty)$  and unprimed variables refer to integrations carried out over entire modal features (over the entire FM or the entire CM PSDs), we can write the total AOD  $(\tau_a)$  at some given reference wavelength as the sum of the fine and coarse total-modal AODs:

$$\tau_{\rm a} = \tau_{\rm f} + \tau_{\rm c},\tag{1}$$

where (letting  $\tau_x$  ( $r_1$ ,  $r_2$ ) refer to the optical integration from  $r_1$  to  $r_2$ )

$$\tau_{\rm f} = \tau_{\rm f}(0, r_0) + \tau_{\rm f}(r_0, \infty), \qquad (2a)$$

$$\tau_{\rm c} = \tau_{\rm c} (0, r_0) + \tau_{\rm c} (r_0, \infty).$$
 (2b)

The analogous relationships in the domain of the cutoff size regimes are written as

$$\tau_{\rm a} = \tau_{\rm f}' + \tau_{\rm c}',\tag{3}$$

$$\tau_{\rm f}' = \tau_{\rm f}(0, r_0) + \tau_{\rm c}(0, r_0),$$
 (4a)

$$\tau_{\rm c}' = \tau_{\rm f}(r_0, \infty) + \tau_{\rm c}(r_0, \infty), \tag{4b}$$

with an emphasis on the "conservation of  $\tau_a$ " ( $\tau_f + \tau_c = \tau_f' + \tau_{fc}' = \tau_a$ ). The differences in the FM and CM (cutoff vs. modal) integrations are, respectively,

$$\begin{split} \tau_{\rm f}' - \tau_{\rm f} = & \Delta \tau_{\rm f} \\ = & \tau_{\rm f}(0, r_0) + \tau_{\rm c}(0, r_0) - [\tau_{\rm f}(0, r_0) + \tau_{\rm f}(r_0, \infty)] \\ = & \tau_{\rm c}(0, r_0) - \tau_{\rm f}(r_0, \infty) \,, \end{split} \tag{5a}$$
 
$$\tau_{\rm c}' - \tau_{\rm c} = & \Delta \tau_{\rm c} \\ = & \tau_{\rm f}(r_0, \infty) + \tau_{\rm c}(r_0, \infty) \\ - & [\tau_{\rm c}(0, r_0) + \tau_{\rm c}(r_0, \infty)] \\ = & \tau_{\rm f}(r_0, \infty) - \tau_{\rm c}(0, r_0) \,, \end{split} \tag{5b}$$

from which

$$\Delta \tau_{\rm c} = -\Delta \tau_{\rm f}. \tag{5c}$$

The last expression is nothing more than confirmation of the "conservation of  $\tau_a$ ".

<sup>&</sup>lt;sup>4</sup>These are retrievals that generally employ prescribed and speciated bulk (modal) FM and CM PSDs (of constant shape and position as a function of radius) as a basis for fitting spectral AODs at a few wavelengths.

<sup>&</sup>lt;sup>5</sup>The AERinv PSD particle-volume PSD  $(dV/d \ln r)$  is readily transformed to the particle-surface PSD of Fig. 1  $(dS/d \ln r)$  =  $3/(4r)dV/d \ln r$ ). See the Fig. 1 caption for details.

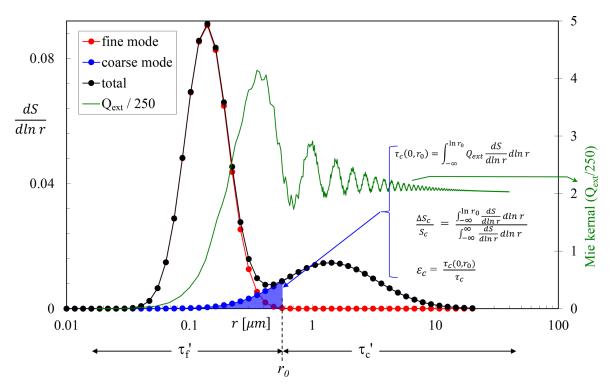


Figure 1. Illustrative fine- and coarse-mode (red and blue) lognormal particle-surface size distributions  $(dS/d\ln r)$  and the Mie kernel (extinction efficiency  $Q_{\rm ext}$ ) for a wavelength of 0.5 µm and a refractive index of 1.44-0.0035i (by particle-surface we mean the projected particle surface area of  $\pi r^2$ ). The cutoff radius ("inflection point" as it is called in the AERONET documentation)  $(r_0)$  applies to an SMF type of computation. The total  $dS/d\ln r$  (black curve) is (in order to have a level of grounding in reality) the sum of the FM and CM lognormal  $dS/d\ln r$  curves that were individually manipulated to fit the mean of a series of AERONET and ground-based  $(dV/d\ln r)$  FM and CM sea-salt curves reported in Fig. 8f of Reid et al. (2006). We base this illustration on  $dS/d\ln r$  as it is more optically fundamental than the AERinv particle-volume PSD (the projected particle surface area is employed to normalize the extinction cross section to obtain the fundamental extinction efficiency). An analogous explanation (and a graph similar to that of Fig. 2) would apply to the particle-volume PSD.

### 2.2 SMF versus FMF

Using the definitions and relationships given above, we can now define FMF and SMF, respectively, as

$$\eta = \frac{\iota_1}{\tau_a}$$

and

$$\eta' = \frac{\tau_{\rm f}'}{\tau_{\rm a}}.\tag{6}$$

Given that  $\tau_f + \tau_c = \tau_f' + \tau_c'$ , we divide by  $\tau_a$  to obtain

$$\eta + \frac{\tau_{\rm c}}{\tau_{\rm a}} = \eta' + \frac{\tau_{\rm c}'}{\tau_{\rm a}},$$

and with a little manipulation,

$$\eta' = (1 - \varepsilon_{\rm c} - \varepsilon_{\rm f}) \, \eta + \varepsilon_{\rm c},\tag{7a}$$

where  $\varepsilon_f$  and  $\varepsilon_c$  represent the pure truncation errors of the FM and CM PSDs (the second terms of Eq. 5a and 5b normalized by  $\tau_a$ ):

$$\varepsilon_{\rm f} = \frac{\tau_{\rm f}(r_0, \infty)}{\tau_{\rm f}}$$

and

$$\varepsilon_{\rm c} = \frac{\tau_{\rm c}(0, r_0)}{\tau_{\rm c}}.\tag{7b}$$

These two quantities represent intensive parameter (largely quantity independent)<sup>6</sup> attributes. Typically however,  $\varepsilon_c \gg \varepsilon_f$  (the major part of the FM PSD is well displaced to the left of  $r_0$  as per the illustration of Fig. 1). This has much more of a cutoff impact on the blue-colored CM PSD of Fig. 1 than on the red-colored FM PSD: the solid blue cutoff portion ( $\Delta S_c$ ) is a significant portion of the CM particle-surface density ( $S_c$ ), and  $\tau_c(0,r_0)$  is a significant optical depth portion of  $\tau_c$  (i.e., both  $\Delta S_c/S_c$  and  $\varepsilon_c$  are typically significant), while the analogous FM fractions above  $r_0$  are relatively insignificant. This affirmation, which we will empirically demonstrate in the multi-station analysis below, is, in part, related to the fact that the FM PSD is approximately half the width of the CM PSD (Dubovik et al., 2002; in their paper "width" specifically refers to the  $\sigma$  value of fitted lognormal distributions).

<sup>&</sup>lt;sup>6</sup>By "largely quantity independent", we mean, from an empirical standpoint, that  $\tau_{\rm f}(r_0,\infty)$  and  $\tau_{\rm c}(0,r_0)$  are typically well correlated with  $\tau_{\rm f}$  and  $\tau_{\rm c}$ , respectively.

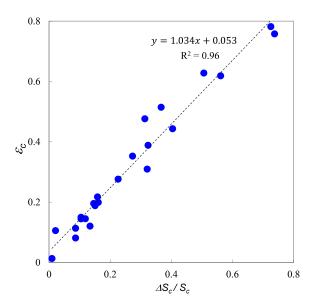
Sites	Latitude	Longitude	Elev.	Region	Aerosol class/	Sampling period	$N^{a}$	Sample citations
			(111)		type	period		
GSFC <sup>b</sup>	38°59′33″ N	76°50′23″ W	87	Maryland, USA	Urban-industrial	1994-2020	13 070	Holben et al. (2001)
Mongu	15°15′13″ S	23°9′3″ E	1047	Zambia	Biomass burning	1997-2009	4751	Eck et al. (2001)
Hamim	22°58′0″ N	54°18′0″ E	105	United Arab Emirates	Dust	2004-2007	2295	Eck et al. (2008)
Solar Village	24°54′25″ N	46°23′50″ E	764	Saudi Arabia	Dust	1999-2012	14 171	Hamill et al. (2016)
FORTH, Crete	35°19′58″ N	25°16′57″ E	20	Greece	Maritime and dust	2003-2017	4039	Hamill et al. (2016)
Midway Island	28°12′35″ N	177°22′42″ W	20	Midway Atoll	Maritime	2001-2014	644	Smirnov et al. (2003)
Lanai	20°44′6″ N	156°55′18″ W	20	Hawaii	Maritime	1997-2004	1175	Reid et al. (2006)
PEARL <sup>c</sup>	80°3′13″ N	86°25′1″ W	615	Nunavut, Canada	Arctic	2007-2019	270	Ranjbar et al. (2019)
Barrow <sup>d</sup>	71°18′44″ N	156°39′54″ W	8	Alaska, USA	Arctic	1997-2020	351	Stone et al. (2014)
Thule	76°30′58″ N	68°46′8″ W	225	Northern Greenland	Arctic	2007-2019	500	AboEl-Fetouh et al. (2020

**Table 1.** Site coordinates, predominant aerosol type, sampling periods and retrieval numbers for the AERONET inversions and SDA retrievals employed in this study.

Figure 2 is a plot of  $\varepsilon_c$  vs.  $\Delta S_c/S_c$  for a variety of retrievals from the sites listed in Table 1. All these cases involved, as per the Fig. 1 illustration, the fitting of FM and CM lognormal curves to AERinv particle-volume PSDs (and a subsequent transformation to particle-surface parameters). These lognormal fits permit the explicit (Mie-based) calculations of the Eq. (7b) fractions. They represent a (AERONETgrounded) theoretical illustration of an expected strong correlation between cutoff optics and cutoff mechanics (the correlation is less than monotonic because of variations in refractive index and the width of the lognormal curves for the different cases chosen in this illustration). We emphasize that automated lognormal fits to AERinv PSDs are not part of the empirical analyses presented in this paper (nor do they have any role in the purely spectral SDA retrieval); rather, the purpose of Fig. 2 is to confirm the expected strong correlation between optical and microphysical CM cutoff fractions and thereby facilitate an understanding of  $\varepsilon_c$ 's role in the dynamics of Eq. (7a).

The linear form of Eq. (7a) with its sub-unity slope of  $(1-\varepsilon_c-\varepsilon_f)$  and positive  $\varepsilon_c$  intercept is coherent with empirical results of  $\eta'$  being generally  $\gtrsim \eta$  (the well-recognized empirical result, supported by the results presented below, of SMF being generally  $\gtrsim$  FMF). Because  $\varepsilon_c$  generally dominates  $\varepsilon_f$ , the linear  $\eta'$  vs.  $\eta$  relation pivots clockwise about the point  $(\eta,\eta')=(1,1)$  in a "linearly inverse" fashion (slope and intercept of approximately  $1-\varepsilon_c$  and  $\varepsilon_c$ ). Figure 3 shows a set of Eq. (7a) straight lines for representative ranges of different  $\varepsilon_c$  and  $\varepsilon_f$  values that were inspired by the ranges encountered in the empirical results that follow.

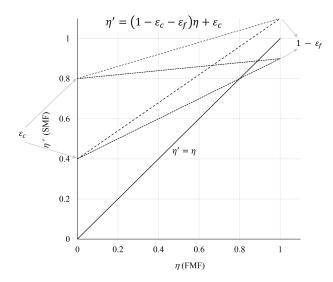
Cases of large  $\varepsilon_f$  are relatively rare but would occur during intense FM events characterized by large-amplitude FM PSDs that are shifted towards larger radii: illustrations include strong (large FM AOD) smoke events (see, for example, the sub-Arctic (Alaskan) smoke cases of AOD(440 nm)  $\gtrsim 1$  in Fig. 9a of Eck et al., 2009) and strong FM pollution events enhanced by high relative humidity (see, for example,



**Figure 2.** Plot of  $\varepsilon_{\rm C}$  vs.  $\Delta S_{\rm C}/S_{\rm C}$  for a variety of simulated optical depth retrievals. The lognormal FM and CM PSDs employed in these simulations were obtained by fitting their sum to retrieved AERinv PSDs over a variety of sites whose predominant aerosol type was urban, dust or marine (and employing the refractive indices provided by the retrieval product). The AERinv PSD retrievals were Version 2, for which  $r_0$  was fixed at  $0.6 \, \mu \rm m$  and, accordingly, for which the  $\varepsilon_{\rm C}$  vs.  $\Delta S_{\rm C}/S_{\rm C}$  regression would be only dependent on  $\varepsilon_{\rm C}$  and second-order parameters such as refractive index (i.e., the plot is independent of  $r_0$ ).

the PSDs corresponding to AOD(440 nm)  $\gtrsim 2$  in Fig. 4 of Eck et al., 2020). However an increasing  $\varepsilon_{\rm f}$  presupposes that  $r_0$  is fixed (as for volumetric surface sampling devices): the AERinv technique of setting the  $r_0$  value to the minimum of the PSD (see below) results in a minimization of  $\varepsilon_{\rm f}$ .

<sup>&</sup>lt;sup>a</sup> N = the number of retrievals. See the Methodology section of the text for details on how the AERONET inversion and SDA retrievals were matched. <sup>b</sup> Goddard Space Flight Center. <sup>c</sup> Polar Environment Arctic Research Lab. The acronym represents the total atmospheric research infrastructure at Eureka, Nunavut, Canada. The AERONET/AEROCAN site is more accurately referred to as the PEARL Ridge Lab. <sup>d</sup> Barrow Atmospheric Baseline Observatory.



**Figure 3.**  $\eta'$  (SMF) vs.  $\eta$  (FMF) lines of Eq. (7a) for three  $\varepsilon_c$  values and three  $\varepsilon_f$  values (0, 0.4 and 0.8 and -0.1, 0 and 0.1, respectively). The  $\varepsilon_c$  values were inspired by the range of empirical values seen in Fig. 6 (and, in the case of  $\varepsilon_f$ , by the range of values, whether real or artifactual, seen in Fig. 7).

### 3 Methodology

Hesaraki et al. (2017) give an overview of the AERinv and SDA with an emphasis on the fact that the former is a significantly more comprehensive algorithm whose low frequency sampling rate is appropriate for detailed climatological-scale analyses (see, for example, Dubovik et al., 2002; AboEl-Fetouh et al., 2020), while the high frequency sampling rate of the SDA is better suited to the detailed analysis of diurnal events (see, for example, Fig. S5 of Saha et al., 2010). The SDA is readily applied to AOD spectra generated by star photometers and moon photometers (see, for example, Baibakov et al., 2015), while analogous AERinv products (requiring an almucantar scan) do not exist for nighttime conditions.

In this investigation, we "matched" SDA to AERinv retrievals by employing averages of Version 3 Level 2.0 AODs (Giles et al., 2019) as inputs to the SDA if those AODs were within a time window of  $\pm 16$  min about the nominal AERinv times. Version 3 AERinv products (Sinyuk et al., 2020) were employed to derive estimates of SMF,  $\tau_f'$  and  $\tau_c'$ . The cutoff radius for those products is actually defined as the minimum of the AERinv output particle-volume PSD ( $dV/d \ln r$ ) with a restriction that the radius bin centers of that minimum must be one of four AERinv bins ( $r_0$ , which is referred to as an "inflection point" in AERONET documentation is allocated a AERinv bin-center value of 0.439, 0.576, 0.756 or 0.992  $\mu$ m). The AERinv retrievals of  $\tau_f'$  and  $\tau_c'$  are interpolated to the SDA reference wavelength of 500 nm using a second-order

(log-log space) spectral polynomial regressed to the  $\tau'_f$  and  $\tau'_c$  values at the four AERinv wavelengths of 440, 675, 870 and 1020 nm (the same technique employed in the SDA).

A variety of AERONET sites (Fig. 4), representing different types (classes) of aerosols, were chosen to investigate the SMF versus FMF relationship. These aerosol classes (Table 1) included sites known for FM urban-industrial aerosols (GSFC in Greenbelt, MD, USA), FM biomass burning (Mongu, Zambia), CM dust (Crete, Hamim and Solar Village), CM maritime aerosols (Midway Island and Lanai), a mixture of dust and marine aerosols (FORTH, Crete) and a mixture of high-Arctic aerosols (PEARL and Thule) as well as low-Arctic aerosols (Barrow Atmospheric Baseline Observatory, referred to as Barrow hereafter). Our choice of aerosol types was not intended to be comprehensive from the standpoint of investigating variations in, say, different types of smoke aerosols or different types of dust aerosols; rather, we sought to properly exercise the SMF versus FMF relationship by largely filling the admissible portion of the SMF versus FMF scattergram (see Fig. 3).

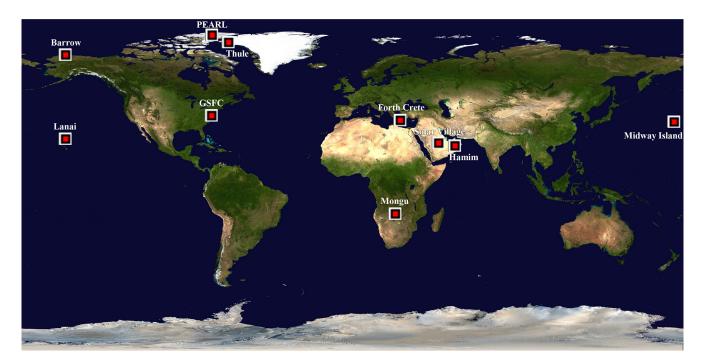
The Arctic category is a mixed class of aerosols: its inclusion in our table of aerosol types is more in terms of it representing an important comparative test (relative to the applicability of southern latitude findings) in a region where the aerosol signal is notably weak. Arctic illustrations of FMF vs. SMF principles feature more in the analysis presented below precisely because the signal is weak: the observation of an independently verified FMF vs. SMF trend or characteristic in the Arctic is an indicator of the robustness of that observation. This weak-signal robustness is something we often see for Arctic retrievals (see, for example, the FM and CM results of AboEl-Fetouh et al., 2020): it may well be (at least in part) attributable to the large solar zenith angles (large solar air mass, M) and attendant 1/M decrease in optical depth errors (a typical finding over 20 years of AERONET Mauna Loa calibrations according to co-author Thomas F. Eck; see also Fig. 2 of Karanikolas et al. (2022) for an empirical validation of this 1/M error dependency).

### 4 Results

### 4.1 Frequency of occurrence of inflection $(r_0)$ points

Figure 5 shows the relative frequency of occurrence (FO) of the four different AERinv inflection  $(r_0)$  points for the different Table 1 classes. The relative importance of CM PSDs vs. FM PSDs for the dust class sites of Solar Village and Hamim and the marine sites of Midway Island and Lanai push the AERinv minimum to smaller  $r_0$  values (resulting in an FO dominance at the 0.439  $\mu$ m inflection point). The marine sites are moderately less asymmetric (less pushed towards 0.439  $\mu$ m) than the dust sites, because the CM PSDs for the dust sites tend to be of larger amplitude. There are, however, other factors at play that could have some impact on

<sup>&</sup>lt;sup>7</sup>The number of AODs required to define a match in the time window could be as small as one (i.e., there was no minimum number).



**Figure 4.** AERONET stations employed for generating the statistics employed in this paper. These stations were chosen to represent a regional variety of aerosols (pollution, biomass burning, dust mixtures, sea-salt mixtures and Arctic aerosols). See Table 1 for details.

the FO distributions; for example, large FM sulfatic particles from Kīlauea eruptions that might push back on the CM PSD dominance of marine particles at Lanai and/or non-sphericity effects of dust that, when corrected (Dubovik et al., 2006), would produce a significantly larger FM PSD amplitude (a complication that does not impact the predominantly spherical marine particles).

The two sites that are heavily influenced by strong FM PSDs (the Mongu biomass-burning site and the GSFC urbanindustrial site) tend (as suggested above in an  $\varepsilon_f$  context) to "push" the PSD minimum towards larger  $r_0$  values relative to the dust and marine sites (resulting in a more balanced FO curve in Fig. 5). Eck et al. (2001) attributed this FM particle-size increase to coagulative effects for Mongu (see their Fig. 6), while Eck et al. (2012) attributed the increase to the effects of hygroscopically induced FM particle growth at the GSFC site (see their Fig. 17). Other smoke-impacted and urban-industrial sites show similar coagulative particle growth effects (see Fig. 10b of Eck et al., 2019, for specific cases recorded in the Southeast Asian tropical forest) and hygroscopic particle growth impacts (see, for example, Fig. 13 of Eck et al., 2005, for Beijing). We argue below that these large-amplitude increases in FM particle size are, given the AERinv technique of variable inflection points, relatively minor in terms of producing significant  $\varepsilon_f$  values.

The Arctic sites show a FM PSD domination which produces an FO distribution that is not unlike the Mongu distribution in the singular case of Barrow, while being much more skewed towards large  $r_0$  for PEARL and Thule. We illustrate

below (as part of a discussion on the variation of  $\varepsilon_c$ ) that a systematic (seasonal) spring-to-summer inflection point increase can be attributed to the Arctic sites.

### 4.2 $\eta'$ vs. $\eta$ (SMF vs. FMF) scattergrams

Figure 6 shows  $\eta'$  vs.  $\eta$  (SMF vs. FMF) scattegrams representing four key aerosol types of Table 1 (scattergrams for the rest of the aerosol types and sites can be seen in Fig. S1 of the Supplement). The theoretical solid black lines of Fig. 6 represent various values of  $\varepsilon_c$  in Eq. (7a); we chose to set  $\varepsilon_f$  to zero in tracing those lines because its value is, as indicated above, generally small (and to not obscure the graphs with second-order detail). We note (as per the previous discussion of the FO distributions) that as the AERinv PSD minimum radius (inflection point) increases, the cutoff portion of the CM PSD increases (resulting in the transition from red to blue curves). The slopes tend to decrease (swing clockwise) with an attendant increase in the intercepts.

The FO dominance of the 0.439 µm  $r_0$  bin is most evident for the CM-dominated dust class (Hamim) with, practically speaking, a single (small  $\varepsilon_c$ ) red-colored grouping of points being observable, while the more balanced distributions of FO curves for GSFC and Mongu provide four clear point groupings. In the extreme of CM cutoff, virtually all of the optically significant contributions of the CM PSD are eliminated, and an asymptote of  $\eta' \to 1$  with  $\varepsilon_c \to 1$  for all values of  $\eta$  is approached (the slope of Eq. 7a approaches zero). That extreme CM cutoff condition is most evident for

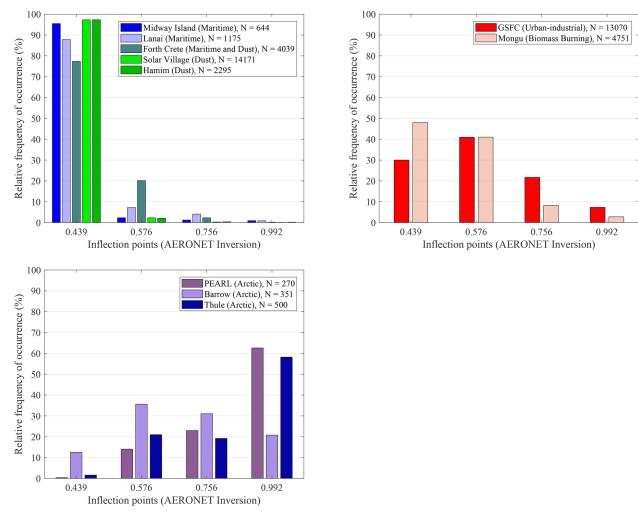


Figure 5. Relative frequency occurrence (FO) distributions for the four different AERONET inflection points ( $r_0$  values with units of  $\mu$ m) for all Table 1 AERONET sites.

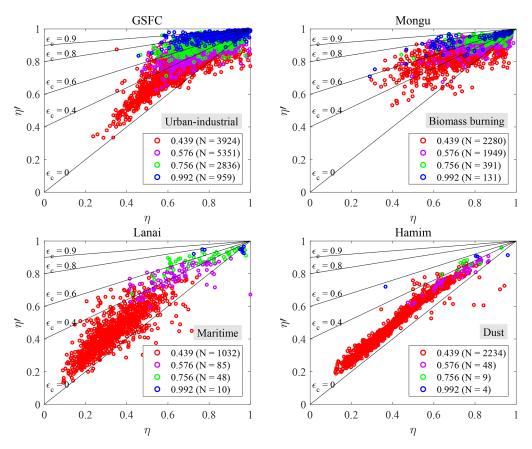
the blue, large- $r_0$  points of the GSFC scattergram), where a regression slope could be almost parallel to the  $\eta'$  axis. We did indeed find  $r_0 = 0.992$  cases associated with near-unity  $\eta'$  values for which most of the optically significant portion of the CM PSD was cut off, while the FM and CM PSDs were not inordinately unbalanced in terms of amplitude ( $\eta$  was some sub-unity value of significance for which there were no apparent problems associated with the AOD spectra).

The small to large  $r_0$  (red to blue) transformation translates, for example, into a classical seasonal pattern for the Arctic scattergrams of Fig. S1; the springtime amplitude of the small-sized CM PSD (that AboEl-Fetouh et al., 2020, associated with Asian dust) decreases progressively from spring to summer. This weakening of CM influence induces a spring-to-summer increase in the value of  $r_0$  and, by extension,  $\varepsilon_c$  (see the seasonal  $r_0$  histograms and derived table for Barrow in Fig. S2). This is an effect that is more notice-

able in the Level-1.5 AERinv products of AboEl-Fetouh et al. (2020) than the Level-2 products of this paper.<sup>8</sup>

Within the theoretical context of Eq. (7a), the scattergrams of Fig. 6 show (predominantly for the red-colored 0.439  $\mu$ m inflection value of  $r_0$ ), a minority of unphysical points below the  $\eta' = \eta$  line (points for which  $\eta' < \eta$ ). An investigation into the most extreme cases of this inequality indicated that the wayward points were more inclined to be associated with abnormally large values of  $\eta$  (rather than abnormally small values of  $\eta'$ ) and small AODs ( $\lesssim$  0.05). Figure 10 of O'Neill et al. (2003) shows the noise sensitivity of  $\eta$  values to such small AODs if one assumes an rms AOD error of

<sup>&</sup>lt;sup>8</sup>The Level-2 processing tends to eliminate springtime retrievals completely: the majority of eliminations are due to excessive residuals in the retrieved vs. measured sky radiances that are, in turn, largely incited by the strong reflectance uncertainty of springtime snow (and its attendant impact on computed sky radiance) as well as the AERinv protocol of eliminating Level-2 retrievals if any snow is detected (by MODIS) within a 5 km radius of the site.



**Figure 6.**  $\eta'$  vs.  $\eta$  (AERinv-derived SMF vs. SDA-derived FMF) scattergrams for the sites of GSFC, Mongu, Lanai and Hamim (aerosol types of urban-industrial, biomass burning, marine and dust, respectively, as per Table 1). The four colors represent the four different AERinv inflection points ( $r_0$  values with units of  $\mu$ m). The remaining scattergrams for the other sites are shown in Fig. S1. The solid black lines are those of Eq. (7a), with  $\varepsilon_f$  set to zero (for the sake of the simplicity of the presentation coupled with the fact that  $\varepsilon_f$  plays a relatively minor role).

0.01 in all bands (0.01 being  $\lesssim$  errors of AERONET field instruments). In general, this type of small-AOD sensitivity was reported empirically by, for example, Fig. 7 of Eck et al. (1999) and Figs. 8 and 9 of O'Neill et al. (2001). The fundamental dynamic is that a band-to-band AOD discontinuity  $\lesssim$  0.01 can generate important perturbations in small-AOD curvature spectra and produce significant outliers in the spectrally sensitive  $\eta$  values.

## 4.3 Analysis of the slopes and intercepts (derivation of $\varepsilon_c$ and $\varepsilon_f$ )

The  $\eta'$  vs.  $\eta$  scattergrams support the hypothesis that there is a physical/optical interpretation that can be given to the slope and intercept of Eq. (7a) (that Eq. 7a and 7b are theoretically relevant approximations). Figure 7a and b show the regression slopes vs. the intercepts derived for, respectively, all sites and all  $r_0$  values of the Fig. 6 and Supplement scattergrams, <sup>9</sup> while Table 2 lists the associated regression statistics

for the 0.439 µm inflection point. The large range of the intercept seen in Fig. 7 confirms that the cutoff portion of the CM PSD ( $\varepsilon_c$  of Eq. 7b) is much more determinant in affecting the interplay of  $\eta'$  vs.  $\eta$ . The position of the colored circles relative to the dashed or dotted black lines (derived from Eq. 7a and 7b) visually support the Table 2 results of small  $\varepsilon_f$  ( $\varepsilon_f$  values being  $\lesssim 0.07$  according to their positions between the  $\varepsilon_f = 0$  and  $\varepsilon_f = 0.1$  grid lines).

The values corresponding to the five red-colored (0.439 µm inflection point) circles for the desert and marine sites (bottom right-hand corner of Fig. 7b) are, however, outside the physically coherent range of Eq. (7). The negative values of  $\varepsilon_{\rm f}$  (see Table 2) can be inferred from the  $\varepsilon_{\rm f}=0$ , 0.1, 0.2 grid of Fig. 7b or, more visually, from the unphysical (super-unity) slopes that arise from a regression through the red points of the desert and marine scattergrams of Fig. 6. This incoherency is very likely the result of those regression lines representing a non-unique ensemble of  $\varepsilon_{\rm c}$  and  $\varepsilon_{\rm f}$  values. From an optophysical standpoint, they vary as a function of the diversity of dust and marine CM PSDs that are predominantly associated with the 0.439 µm inflection point (they

<sup>&</sup>lt;sup>9</sup> with a coherent color scheme between the scattergrams and Fig. 7b

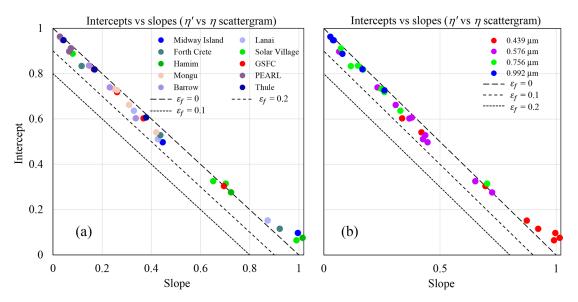


Figure 7. Intercept vs. slope plots for the  $\eta'$  vs.  $\eta$  scattergrams as a function of (a) all the sites of Table 1 and (b) the cutoff  $r_0$  value (AERinv inflection point radius). The black dashed or pointed lines represent the family of straight lines generated by Eq. (7a): slope =  $(1 - \varepsilon_c - \varepsilon_f)$  and intercept =  $\varepsilon_c$  for values of  $\varepsilon_c = 0$  to 1 and three values of  $\varepsilon_f$  (0.0, 0.1 and 0.2).

**Table 2.**  $\eta$ ' vs.  $\eta$  regression stats for  $r_0 = 0.439 \,\mu\text{m}^a$ .

	m	b	$R^2$	$\sigma(\eta'_{\rm res})$	$\sigma(m)$	$\sigma(b)$	N	$\varepsilon_{ m c}$	$arepsilon_{ m f}$	$\sigma(\varepsilon_{ m c})$	$\sigma(\varepsilon_{ m f})$
GSFC	0.695	0.304	0.749	0.043	0.409	0.288	3924	0.304	$9 \times 10^{-4}$	0.288	0.29
Mongu	0.42	0.541	0.641	0.038	0.314	0.256	2280	0.541	0.039	0.256	0.182
Hamim	1.015	0.076	0.969	0.025	0.172	0.069	2234	0.076	-0.091	0.069	0.158
Solar Village	0.99	0.065	0.912	0.043	0.314	0.12	13 790	0.065	-0.055	0.12	0.291
FORTH, Crete	0.922	0.116	0.941	0.044	0.226	0.139	3125	0.116	-0.038	0.139	0.178
Midway	0.996	0.097	0.832	0.059	0.39	0.15	615	0.097	-0.093	0.15	0.36
Lanai	0.873	0.152	0.674	0.072	0.427	0.154	1032	0.152	-0.025	0.154	0.398
PEARL		$n/a^b (N=1)$									
Barrow	0.34	0.604	0.447	0.053	0.462	0.341	44	0.604	0.057	0.341	0.311
Thule	0.259	0.66	0.196	0.038	0.482	0.405	8	0.66	0.081	0.405	0.262

a See, for example, Taylor (1997) for typical regression relationships. The  $\varepsilon_{\rm C}$  and  $\varepsilon_{\rm f}$  values as well as their standard deviation ( $\sigma\varepsilon_{\rm C}$ ) and  $\sigma(\varepsilon_{\rm f})$  were derived from the regressed slope and intercept ("m" and "b") using the slope and intercept expressions of Eq. (7a). The values of ( $\sigma(\varepsilon_{\rm C})$  and  $\sigma(\varepsilon_{\rm f})$ ) are computed from effective standard deviations of m and b. By this, we mean that the standard error (from sources such as Taylor, 1997) is multiplied by  $\sqrt{N}$  to yield  $\sigma(m)$  and  $\sigma(\varepsilon_{\rm f})$  and  $\sigma(\varepsilon_{\rm f})$  is multiplied by  $\sigma(\varepsilon_{\rm f})$  and  $\sigma(\varepsilon_{\rm f})$  and  $\sigma(\varepsilon_{\rm f})$  is multiplied by  $\sigma(\varepsilon_{\rm f})$  and  $\sigma(\varepsilon_{\rm f})$  and  $\sigma(\varepsilon_{\rm f})$  is change is coherent with the notion that those variations are more likely due to  $\varepsilon_{\rm C}$  and  $\varepsilon_{\rm f}$  being characterized by a systematic range of values for any given inflection point ( $r_{\rm O}$ ) value. Note that  $\sigma(\eta'_{\rm res})$  represents the standard deviation of the regression residuals.

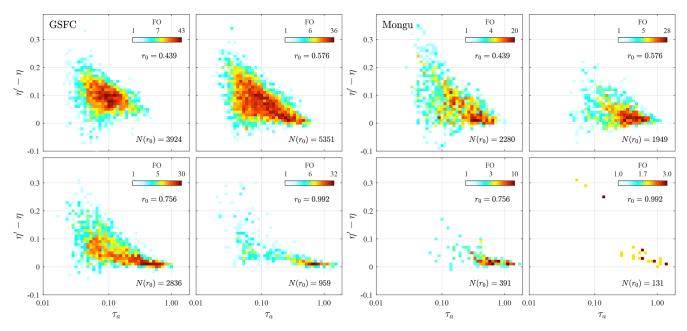
roughly lie between the  $\varepsilon_c = 0$  and  $\varepsilon_c = 0.4$  boundaries of Fig. 3; the  $\eta = 1$  intercept<sup>10</sup> will (in consequence of a superunity slope) yield negative values above the  $\eta' = \eta$  intercept).

Non-systematic noise will also impact the derived  $\varepsilon_c$  and  $\varepsilon_f$  values. Scattergrams of  $\tau_f'$  vs.  $\tau_f$  for the two sites that are largely dominated by strong FM variations (GSFC and Mongu) and which showed the most balanced (Fig. 5) FO distributions are shown in Fig. A1. One can observe, in the first instance, the point-dispersion reduction and the convergence towards the  $\tau_f' = \tau_f$  line as both increase (as  $\tau_a$  increases). The extensive (quantity-dependent) nature of

those scattergrams complements the analogous  $\eta'$  vs.  $\eta$  semiintensive scattergrams of Fig. 6 by explicitly displaying the noise-like influences of the amount of aerosol ( $\tau'_f$  or  $\tau_f$ ) as well as  $r_0$ . The  $\tau'_f$  vs.  $\tau_f$  dispersion of Fig. A1 is nonetheless generally small; this underscores a hypothesis that the  $\eta'$  vs.  $\eta$  results presented above are robust second-order findings. The FO distributions of  $\eta'$  vs.  $\eta$  in Fig. A2 effectively eliminate the extensive variations of Fig. A1 with an attendant enhancement of those second-order influences.

The Fig. 8 FO distributions of  $\eta' - \eta$  vs.  $\tau_a$  more readily show the decreasing dispersion and the convergence towards the  $\eta' = \eta$  ( $\tau_f' = \tau_f$ ) line with increasing  $\tau_a$  (increasing  $\tau_f$  val-

 $<sup>^{10}</sup>$  for which  $\eta' = 1 - \varepsilon_{\mathrm{f}}$ 



**Figure 8.**  $\eta' - \eta$  vs.  $\tau_a$  FO distributions for GSFC and Mongu.

ues for GSFC and Mongu). The broad left-to-right movement of the (red) FO peak values with increasing  $r_0$  (notably in the case of GSFC) also clarifies an aspect that is not easily discernable in the highly correlated scattergrams of Fig. A1: that an increase in  $r_0$  is coarsely associated with  $\eta'$  values that approach  $\eta$  or, hence,  $\tau_f'$  values that approach  $\tau_f$  (a trend that effectively drives the  $\eta' = \eta$  ( $\tau_f' = \tau_f$ ) convergence induced by increasing  $\tau_a$ ).

Figure 9 shows the regression-derived  $\varepsilon_c$  and  $\varepsilon_f$  variation as a function of an artificial minimum ( $\tau_{a,\,min}$ ) in the lower bound of the  $\tau_a$  regression range (see the Fig. 9 caption for more details). The result shows no strong  $\varepsilon_c$  dependency on  $\tau_{a,\,min}$  and more variable but consistently small amplitudes of  $\varepsilon_f$ . 11 Qualitatively, these observations are not unexpected given the fairly persistent slope of the high FO (red-colored ellipses of Fig. A2).

The choice of a variable  $r_0$  in the AERinv retrievals clearly means that the affirmation of increasing  $\varepsilon_f$  in the presence of strong FM events (i.e., what would be measured by a traditional fixed- $r_0$  device) cannot readily be observed with the AERinv retrievals. Strong FM events basically push the FM PSD to larger radii (while the CM PSD remains relatively inactive  $^{12}$ ). The AERinv approach of selecting  $r_0$  as the minimum of the PSD basically neuters the cutoff of significant optical portions of the FM PSD. The general tendency for  $r_0$ 

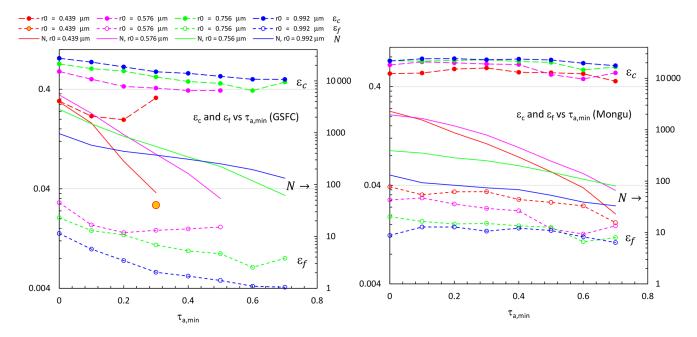
to increase with increasing  $\tau_f$  in Fig. 8 is consistent with the concept of the FM PSD being pushed to the right as  $\tau_f'$  and  $\tau_f$  increase.

AERinv PSDs can differ from the simple bimodal paradigm incorporated in Eqs. (1) through (7). The bimodality of the Arctic CM PSDs of AboEl-Fetouh et al. (2020) were ascribed to a small-sized CM feature in the 1.3 µm AERinv bin (a systematic feature that they attributed to Asian dust) and a larger-sized mode that might have been linked to local dust and/or sea salt. Eck et al. (2012) reported a bimodal FM PSD (their Fig. 3) that they attributed to cloud processing of FM pollution (haze) aerosols. In such cases, one can appeal to the optical equivalency of, for example, a bimodal CM PSD to a single CM PSD whose curvature parameters become averages of those of the two CM components (see Appendix A for the two specific cases of a bimodal CM PSD with a unimodal FM PSD as well as a bimodal FM PSD with a unimodal CM PSD). This means that the bimodal expressions of O'Neill et al. (2001) still apply for the AERONET SDA product and thus that the FMF to SMF expressions (Eqs. 1 to 7) can still be used.

The FMF (SDA) approach is arguably the more fundamental approach for separating FM and CM optical contributions, because it is intrinsically related to the modal nature of different types of aerosols (at least to the first order: modality becomes obscured, for example, with internal mixing of different aerosol types). The SMF (AERinv) approach is the more pragmatic approach, as it is commonly and readily applied to microphysical surface and airborne measurements. It is however, as we have seen in results like those of Fig. 6, very dependent on the selected cutoff radius.

 $<sup>^{11}</sup>$ Except in the case of the (red) 0.439  $\mu m\,r_0$  (yellow-filled) value where the  $\varepsilon_f$  variations were more substantial. At the same time, the number of regression points (N) rapidly decreases with increasing  $\tau_{a,\,min}$  so that large  $\varepsilon_f$  (and  $\varepsilon_c$ ) deviations are less statistically significant.

 $<sup>^{12}</sup>$ And it suffers from a per-particle extinction kernel (the green-colored  $Q_{\text{ext}}$  factor of Fig. 1) that is weaker in CM radius range.



**Figure 9.** GSFC and Mongu variations of regression-derived values of  $\varepsilon_c$  and  $\varepsilon_f$  as a function of a minimum value of total AOD at 500 nm ( $\tau_{a, min}$ ) employed in the  $\eta'$  vs.  $\eta$  regressions (The decreasing point dispersion (the approach of  $\tau_f'$  to  $\tau_f$ ) with increasing  $\tau_f$  indicates that the regression-derived precision of the  $\varepsilon_c$  and  $\varepsilon_f$  values should, for that reason alone, generally increase with larger values of  $\tau_f$ . However, offsetting this type of influence on an increase in precision is the decrease in the number of regression points and the squeeze in  $\tau_f$  and  $\tau_f'$  variability as the  $\eta = \eta'$  singularity is approached.) (to be clear, Fig. 9 represents an exercise in testing the sensitivity to  $\tau_{a, min}$ : the regressions of Fig. 6 and the regression-derived  $\varepsilon_c$  and  $\varepsilon_f$  values of Fig. 7 were obtained with no restrictions on the matched AERinv and SDA retrievals). Same color scheme as Figs. 6 and 7. The legend of the Mongu plot is the same as the GSFC legend. The number of matched retrievals (*N*) is also shown (right-hand axis). Result are not shown in cases where N < 20. The large orange-filled (re-rimmed) circle in the GSFC case represents the only  $\varepsilon_f$  point of the  $r_0 = 0.439$  μm regressions that both survived the N < 20 rejection criterion and did not yield a negative  $\varepsilon_f$  value.

### 5 Conclusions

We presented a simple SMF vs. FMF ( $\eta'$  vs.  $\eta$ ) equation that enabled an understanding of the well-recognized empirical result of SMF being larger than the FMF. This result has been reported for in situ, satellite and ground-based remote sensing techniques; our focus was on an SMF vs. FMF interpretation in the form of AERinv SMF vs. SDA FMF retrievals. We pointed out that these two AERONET products provide a unique opportunity to empirically compare the SMF and FMF approaches at similar (columnar) remote sensing scales and across a shared global variety of aerosol types.

The SMF vs. FMF equation largely captured the SMF vs. FMF behavior of the AERinv vs. SDA products as a function of inflection point  $(r_0)$  across an ensemble of AERONET sites and aerosol types (urban-industrial, biomass burning, dust, marine, maritime and Arctic). The SMF vs. FMF behavior was primarily dependent on the intensive parameter of relative cutoff portion of the CM PSD  $(\varepsilon_c)$  and, to a second order, the relative cutoff portion of the FM PSD  $(\varepsilon_f)$ . The overarching dynamic was that the linear SMF vs. FMF relation pivots clockwise about the point  $(\eta, \eta') = (1, 1)$  in a "linearly inverse" fashion (slope and intercept of approximately  $1 - \varepsilon_c$  and  $\varepsilon_c$ ) with increasing  $r_0$ . Derived SMF vs. FMF

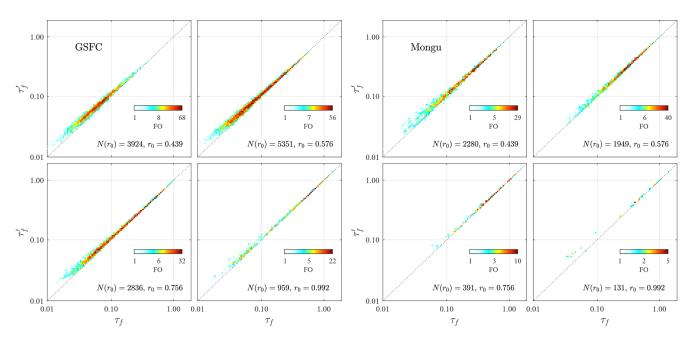
slopes and intercepts confirmed the general domination of  $\varepsilon_c$  over  $\varepsilon_f$  in controlling the "linear inverse" variation. The process of deriving and analyzing  $\varepsilon_c$  and  $\varepsilon_f$  values demonstrated an expected domination of FM optical depths for the urban pollution and biomass-burning sites of GSFC and Mongu and thus a convergence towards the  $\tau_f' = \tau_f \ (\eta' = \eta)$  line as  $\tau_a$  increased (the convergence of SDA FM AODs towards AERinv FM AODs).

The more general conclusion resulting from this analysis is the apparent empirical confirmation that the influence of PSD modal features can be detected by an indirect comparative analysis. While one would like to believe that this is true in general, a more comprehensive event-level closure experiment employing, for example, multi-altitude microphysical and optical measurements over a representative suite of AERONET instruments would do much to increase the level of confidence in such a conclusion.

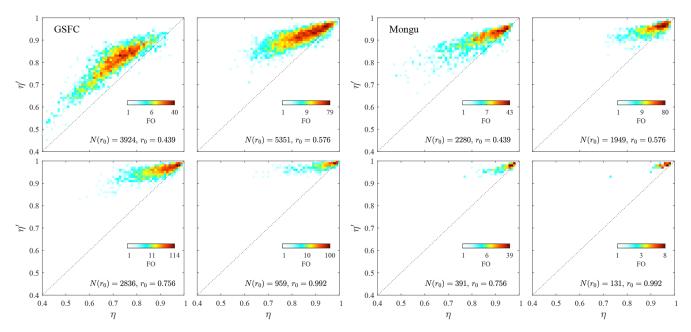
### Appendix A

### A1 Frequency of occurrence analyses

The retrieval results in this subsection are restricted to the two sites that are strongly impacted by historically large variations in FM particles (GSFC and Mongu). These sites (details in Table 1) can experience  $\tau_f$  values greater than unity.



**Figure A1.**  $\tau'_f$  vs.  $\tau_f$  frequency of occurrence (FO) distributions for GSFC and Mongu. The FO color scale is tied to variations on a logarithmic scale (with an attendant tendency to enhance the contributions of large FO values).  $N(r_0)$  is the total FO at a given  $r_0$  value.



**Figure A2.**  $\eta'$  vs.  $\eta$  FO distributions for GSFC and Mongu. See the caption of Fig. A1 for FO and  $N(r_0)$  details.

### A2 Dual-modal optical equivalency of a trimodal size distribution

### A2.1 Fine mode and two coarse modes

One often sees a bimodal coarse-mode PSD and thus a trimodal PSD. Treating the bimodal CM PSD as an optically equivalent single-mode CM PSD can be formalized as below (optical equivalent of Eq. 1a of O'Neill et al., 2003):

$$\tau_{a} = \tau_{f} + \tau_{c} = \tau_{f} + \tau_{c1} + \tau_{c2} 
= \tau_{f} + \eta_{c1}\tau_{c} + \eta_{c2}\tau_{c},$$
(A1)

where  $\eta_{c1} = \frac{\tau_{c1}}{\tau_c}$ , and  $\eta_{c2} = \frac{\tau_{c2}}{\tau_c}$ . Accordingly,

$$\tau_{a} = \tau_{f} + \eta_{c1}\tau_{c} + (1 - \eta_{c1})\tau_{c}. \tag{A2}$$

Equation (3) of O'Neill et al. (2003) can be written as

$$\alpha \tau_{a} = \alpha_{f} \tau_{f} + \alpha_{c1} \tau_{c1} + \alpha_{c2} \tau_{c2}$$

$$\alpha \tau_{a} = \alpha_{f} \tau_{f} + \tau_{c} (\alpha_{c1} \eta_{c1} + (1 - \eta_{c1}) \alpha_{c2})$$

$$= \alpha_{f} \tau_{f} + \tau_{c} \langle \alpha_{c} \rangle, \tag{A3}$$

where  $\langle \alpha_c \rangle = \alpha_{c1} \eta_{c1} + (1 - \eta_{c1}) \alpha_{c2}$  represents the optical average of  $\alpha_{c1}$  and  $\alpha_{c2}$ . Equation (A3) is exactly the analog of O'Neill et al. (2003). Differentiating and recalling that  $d\tau/d\ln\lambda = -\alpha\tau$ , one finds that

$$\begin{split} \alpha'\tau_a &= \alpha_f'\tau_f + \alpha_{c1}'\tau_{c1} + \alpha_{c2}'\tau_{c2} + \alpha^2\tau_a - \alpha_f^2\tau_f - \alpha_{c1}^2\tau_{c1} \\ &- \alpha_{c2}^2\tau_{c2} \\ &= \alpha_f'\tau_f + \left(\alpha_{c1}'\tau_{c1} + \alpha_{c2}'\tau_{c2}\right) + \left\{\alpha(\alpha\tau_a)\right\} - \alpha_f^2\tau_f \\ &- \left(\alpha_{c1}^2\frac{\tau_{c1}}{\tau_c} - \alpha_{c2}^2\frac{\tau_{c2}}{\tau_c}\right)\tau_c \\ &= \alpha_f'\tau_f + \left<\alpha_c'\right>\tau_c + \left\{\alpha\alpha_f\tau_f + \alpha\left<\alpha_c\right>\tau_c\right\} - \alpha_f^2\tau_f - \left<\alpha_c^2\right>\tau_c \\ &= \alpha_f'\tau_f + \left<\alpha_c'\right>\tau_c + \alpha_f\tau_f(\alpha - \alpha_f) + \left<\alpha_c\right>\tau_c(\alpha - \left<\alpha_c\right>) \\ &+ \left<\alpha_c\right>^2\tau_c - \left<\alpha_c^2\right>\tau_c, \end{split}$$

but since  $\alpha = \alpha_f \eta + (1 - \eta) \langle \alpha_c \rangle$ , we can replace  $(\alpha - \alpha_f)$  and  $(\alpha - \langle \alpha_c \rangle)$  to obtain

$$\begin{split} &=\alpha_f'\tau_f + \langle\alpha_c'\rangle\tau_c - \alpha_f\tau_f\left(1-\eta\right)(\alpha_f - \langle\alpha_c\rangle) \\ &+ \langle\alpha_c\rangle\tau_c\eta\left(\alpha_f - \langle\alpha_c\rangle\right) + \tau_c\left(\langle\alpha_c\rangle^2 - \langle\alpha_c^2\rangle\right) \\ &= \alpha_f'\tau_f + \langle\alpha_c'\rangle\tau_c - (1-\eta)\,\tau_f(\alpha_f - \langle\alpha_c\rangle)^2 \\ &+ \tau_c\left(\langle\alpha_c\rangle^2 - \langle\alpha_c^2\rangle\right) \end{split}$$

so that

$$\begin{split} \alpha' &= \eta \alpha_{\rm f}' + (1 - \eta) \left\{ \langle \alpha_{\rm c}' \rangle + \langle \alpha_{\rm c} \rangle^2 - \langle \alpha_{\rm c}^2 \rangle \right\} \\ &- (1 - \eta) \, \eta (\alpha_{\rm f} - \langle \alpha_{\rm c} \rangle)^2. \end{split}$$

These expressions are optically equivalent to Eq. (5) of O'Neill et al. (2003) with their Eq. (5) values being transformed according to

$$\begin{split} &\alpha_c \leftrightarrow \langle \alpha_c \rangle, \\ &\alpha_c' \leftrightarrow \langle \alpha_c' \rangle + \langle \alpha_c \rangle^2 - \langle \alpha_c^2 \rangle, \end{split}$$

where the average value of any parameter  $x_c$  is always

$$\langle x_c \rangle = x_{c1} \eta_{c1} + x_{c2} (1 - \eta_{c1}).$$

### A2.2 Coarse mode and two fine modes

In this instance one can imagine two fine-mode components; for example, (i) large-sized FM smoke and (ii) smaller sized FM urban-industrial pollution. The algebra above for the single FM PSD and two CM PSDs can be employed here: everything is perfectly reversible (interchange index c with index f) if  $\eta$  is viewed as the ratio of coarse to total AOD. Then arriving at the final equations for  $\alpha$  and  $\alpha'$ , one can return to the usual definition of  $\eta$  to obtain the same algebraic theorem. The classic equations for  $\alpha$  and  $\alpha'$  remain unchanged providing one employs the substitutions below:

$$\alpha_f \leftrightarrow \langle \alpha_f \rangle$$
,

$$\alpha_{\rm f}' \leftrightarrow \langle \alpha_{\rm f}' \rangle + \langle \alpha_{\rm f} \rangle^2 - \langle \alpha_{\rm f}^2 \rangle,$$

where  $\langle x_c \rangle = \eta_f x + (1 - \eta_f) x$ , and  $\eta_f = \tau_{f,1} / \tau_f$ . Explicitly,

$$\alpha = \alpha_{\rm c} (1 - \eta) + \eta \langle \alpha_{\rm f} \rangle$$
,

$$\alpha' = (1 - \eta)\alpha_{c}' + \eta \left\{ \langle \alpha_{f}' \rangle + \langle \alpha_{f} \rangle^{2} - \langle \alpha_{f}^{2} \rangle \right\}$$
$$- (1 - \eta) \eta (\langle \alpha_{f} \rangle - \alpha_{c})^{2}.$$

### Appendix B: Symbol and acronym glossary

AEROCAN Federated Canadian subnetwork of AERONET run by Environment and Climate Change Canada (ECCC)
AERONET AErosol RObotic NETwork: worldwide NASA network of combined sun photometer/sky-scanning ra-

diometers manufactured by CIMEL Éléctronique. See http://aeronet.gsfc.nasa.gov/ (last access: 25 Febru-

ary 2023) for documentation and data downloads.

AOD Aerosol optical depth: the community uses "AOD" to represent anything from nominal aerosol optical depth

which has not been cloud-screened to the conceptual (theoretical) interpretation of aerosol optical depth. In

this paper we use it in the latter sense and apply adjectives as required.

AERinv AERONET inversion

SDA Spectral deconvolution algorithm

CM Coarse mode

 $\varepsilon_{\rm c}$  Relative truncation error of the CM PSD (Eq. 7b)  $\varepsilon_{\rm f}$  Relative truncation error of the FM PSD (Eq. 7b)

 $\eta$  FMF (product of the SDA)  $\eta'$  SMF (product of the AERinv)

FM Fine mode

FO Frequency of occurrence

PSD Particle size distribution (notably the AERinv particle-volume size distribution,  $dV/d \ln r$ , in the context of

this paper)

FMF Fine-mode fraction (an output parameter computed from SDA products)
SMF Sub-micron fraction (an output parameter computed from AERinv products)

 $r_0$  Cutoff radius (units of  $\mu$ m) that separates the PSD into FM and CM components. The AERONET parameter

name is "inflection point" (but it would be more pragmatically labelled as the PSD minimum; the bin-center

location of the minimum of the retrieved PSD).

 $\tau_a$  Total aerosol optical depth at 500 nm (SDA)

 $\tau_{c}$  Coarse-mode aerosol optical depth at 500 nm (SDA)  $\tau_{f}$  Fine-mode aerosol optical depth at 500 nm (SDA)

 $\tau'_a$  Total aerosol optical depth at 500 nm

 $\tau'_{\rm c}$  Coarse-mode aerosol optical depth at 500 nm using a cutoff radius of  $r_0$  (derived AERinv product)  $\tau'_{\rm f}$  Fine-mode aerosol optical depth at 500 nm using a cutoff radius of  $r_0$  (derived AERinv product)

*Data availability.* All AERinv and SDA retrieval data are available from the AERONET website (https://doi.org/10.17616/R3VK9T, Lind and Gupta, 2023).

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