- 1 M. Sicard, M.J. Granados-Muñoz, L. Alados-Arboledas, R. Barragán,
- 2 A.E. Bedoya-Velásquez, J.A. Benavent-Oltra, D. Bortoli, A. Comerón,
- ³ C. Córdoba-Jabonero, M.J. Costa, A. del Águila, A.J. Fernández, J.L.
- 4 Guerrero-Rascado, O. Jorba, F. Molero, C. Muñoz-Porcar, P. Ortiz-
- 5 Amezcua, N. Papagiannopoulos, M. Potes, M. Pujadas, F.
- 6 Rocadenbosch, A. Rodríguez-Gómez, R. Román, R. Salgado, V.
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Ground/space, passive/active remote sensing observations coupled with particle dispersion modelling to understand the inter-continental transport of wildfire smoke plumes

- 20 M. Sicard^{1,2}, M. J. Granados-Muñoz¹, L. Alados-Arboledas^{3,4}, R. Barragán^{1,2}, A.E. Bedoya-
- 21 Velásquez^{3,4}, J.A. Benavent-Oltra^{3,4}, D. Bortoli⁵, A. Comerón¹, C. Córdoba-Jabonero⁶, M. J.
- 22 Costa⁵, A. del Águila⁶, A. J. Fernández⁷, J.L. Guerrero-Rascado^{3,4}, O. Jorba⁸, F. Molero⁷, C.
- 23 Muñoz-Porcar¹, P. Ortiz-Amezcua^{3,4}, N. Papagiannopoulos^{1,9}, M. Potes⁵, M. Pujadas⁷, F.
- 24 Rocadenbosch^{1,2}, A. Rodríguez-Gómez¹, R. Román¹⁰, R. Salgado⁵, V. Salgueiro⁵, Y. Sola¹¹, M.
- 25 Yela⁶
- 26 ¹Remote Sensing Laboratory / CommSensLab, Universitat Politècnica de Catalunya, Barcelona, 08034, Spain
- 27 ²Ciències i Tecnologies de l'Espai Centre de Recerca de l'Aeronàutica i de l'Espai / Institut d'Estudis Espacials de
- 28 Catalunya (CTE-CRAE / IEEC), Universitat Politècnica de Catalunya, Barcelona, 08034, Spain
- ²⁹ ³Department of Applied Physics, University of Granada, Granada, 18071, Spain
- 30 ⁴Andalusian Institute for Earth System Research (IISTA-CEAMA), Granada, 18006, Spain
- 31 ⁵Institute of Earth Sciences and Dept. of Physics, Universidade de Évora, Évora, 7000-671, Portugal
- 32 ⁶Instituto Nacional de Técnica Aeroespacial (INTA), Atmospheric Research and Instrumentation Branch, Torrejón de
- 33 Ardoz (Madrid), 28850, Spain
- 34 ⁷Dept. of Environment, Research Centre for Energy, Environment and Technology (CIEMAT), Madrid, 28040, Spain
- 35 ⁸Dept. of Earth Sciences, Barcelona Supercomputing Center (BSC), Barcelona, 08034, Spain
- 36 ⁹Consiglio Nazionale delle Ricerche, Istituto di Metodologie per l'Analisi Ambientale (CNR-IMAA), Tito Scalo,
- 37 85050, Italy
- 38 ¹⁰Atmospheric Optics Group (GOA), University of Valladolid, 47002, Spain
- ³⁹ ¹¹Dept. of Astronomy and Meteorology, Universitat de Barcelona, Barcelona, 08028, Spain
- 40 Correspondence to: Michaël Sicard (msicard@tsc.upc.edu)

41 Abstract. During the 2017 record-breaking burning season in Canada / United States, intense wild fires raged during 42 the first week of September in the Pacific northwestern region (British Columbia, Alberta, Washington, Oregon, Idaho, Montana and northern California) burning mostly temperate coniferous forests. The heavy loads of smoke particles 43 44 emitted in the atmosphere reached the Iberian Peninsula (IP) a few days later on 7 and 8 September. Satellite imagery allows to identify two main smoke clouds emitted during two different periods that were injected and transported in 45 the atmosphere at several altitude levels. Columnar properties on 7 and 8 September at two Aerosol Robotic Network 46 47 (AERONET) mid-altitude, background sites in northern and southern Spain are: aerosol optical depth (AOD) at 440 48 nm up to 0.62, Ångström exponent of 1.6-1.7, large dominance of small particles (fine mode fraction > 0.88), low 49 absorption AOD at 440 nm (<0.008) and large single scattering albedo at 440 nm (>0.98). Profiles from the Cloud-50 Aerosol Lidar with Orthogonal Polarization (CALIOP) show the presence of smoke particles in the stratosphere during 51 the transport, whereas the smoke is only observed in the troposphere at its arrival over the IP. Portuguese and Spanish 52 ground lidar stations from the European Aerosol Research Lidar Network / Aerosols, Clouds, and Trace gases Research InfraStructure Network (EARLINET/ACTRIS) and the Micro-Pulse Lidar NETwork (MPLNET) reveal 53 54 smoke plumes with different properties: particle depolarization ratio and color ratio, respectively, of 0.05 and 2.5 in the mid troposphere (5 - 9 km) and of 0.10 and 3.0 in the upper troposphere (10 - 13 km). In the mid troposphere the 55 56 particle depolarization ratio does not seem time-dependent during the transport whereas the color ratio seems to 57 increase (larger particles sediment first). To analyze the horizontal and vertical transport of the smoke from its origin to the IP, particle dispersion modelling is performed with the Hybrid Single Particle Lagrangian Integrated Trajectory 58 Model (HYSPLIT) parameterized with satellite-derived biomass burning emission estimates from the Global Fire 59 60 Assimilation System (GFAS) of the Copernicus Atmosphere Monitoring Service (CAMS). Three compounds are simulated: carbon monoxide, black carbon and organic carbon. The results show that the first smoke plume which 61 62 travels slowly reaches rapidly (~1 day) the upper troposphere and lower stratosphere (UTLS) but also shows evidence 63 of large scale horizontal dispersion, while the second plume, entrained by strong subtropical jets, reaches the upper 64 troposphere much slower (~2.5 days). Observations and dispersion modelling all together suggest that particle depolarization properties are enhanced during their vertical transport from the mid to the upper troposphere. 65

Keywords. Time-space monitoring, ground-based and space-borne lidars, long-range transport of smoke plume,
injection of particles up to the upper troposphere, particle dispersion model, smoke particle absorption and
depolarization properties.

69 1 Introduction

70 It is well established that atmospheric biomass burning from either prescribed fires or natural wildfires have effects 71 on air quality, atmospheric circulation and climate (Stocks et al., 2003). Wildfires have recently become a focus of 72 growing interest and attention because of their capabilities to inject smoke particles at high altitude levels. The 73 mechanisms leading to the vertical transport of smoke particles are either direct injection by pyroconvection (Fromm 74 et al., 2000; Fromm and Servranckx, 2003), a combination of pyroconvection and radiatively driven uplift forces (de Laat et al., 2012) or a combination of pyroconvection and gravito-photophoresis (Rohatschek, 1996; Pueschel et al., 75 76 2000). The first mechanism, pyroconvection, materializes through the formation of pyrocumulus (pyroCu) and their 77 most extreme form, namely pyrocumulonimbus (pyroCb; Fomm et al., 2005). The characteristic injection height of pyroCu and pyroCb emissions is the upper troposphere (UT) and less frequently the lower stratosphere (LS) (Fromm 78 79 et al., 2010). The second mechanism, called self-lifting, is based on the absorption of incoming solar radiation by soot 80 and smoke particles which may cause sufficient warming for air masses to provide buoyancy and subsequent lofting of the injected plume (Boers et al., 2010; de Laat et al., 2012). The third mechanism, gravito-photophoresis, is due to 81 82 "a sunlight-induced force acting on particles which are geometrically asymmetric and which have uneven surface 83 distribution of thermal accommodation coefficients" (Pueschel et al., 2000). It is strongly altitude-dependent because 84 of the weak lifting forces involved and it is most effective above 10 km. Renard et al. (2008) suggested that soot from biomass burning could reach the stratosphere owing to the gravito-photophoresis effect. The last two mechanisms, 85 self-lifting and gravito-photophoresis, can only act on particles which are already settled in the free troposphere or in 86 the stratosphere, and thus require a prior injection of the particles usually produced by pyroconvection. 87

Once in the UT, the tropopause acts as a dynamic barrier to the upward transport of smoke particles from the 88 89 troposphere because of the steep gradient in the temperature lapse rate, and in most cases the particles stay in the 90 troposphere. The conditions (burnt matter, fire characteristics, latitude range, local meteorology, synoptic conditions, 91 dynamics, etc.) allowing for the penetration of smoke particles through the tropppause are still not yet entirely clear, 92 and many conclusions of the recent literature on the subject call for more investigation on the topic. The transport of 93 particles in the upper troposphere and lower stratosphere (UTLS) has several effects: (i) in this altitude range, the 94 particles can persist for long durations (Robock, 2000), allowing for gradual spread over hemispheric or global scales; 95 (ii) the long-lived aerosol radiative effects, especially marked for smoke which is a warming agent, may cause differential regional heating patterns that affect regional circulation (Lau et al., 2008; Son et al., 2009); (iii) complex 96 97 interactions with clouds due to their capability to serve as cloud condensation nuclei, producing in the end a reduction of precipitation (see details in Rosenfeld et al., 2007); (iv) effects on UTLS ozone chemistry (Crutzen and Andreae,
1990; Forster et al., 2001; Real et al., 2008). An increasing number of recent studies report on the observation of the
presence of smoke particles in the UTLS: Nédélec et al. (2005), Damoah et al. (2006), Rosenfeld et al. (2007), Fromm
et al. (2010) (and references therein), Siddaway and Petelina (2011), de Laat et al. (2012), Khaykin et al. (2018),
Ansmann et al. (2018), Haarig et al. (2018), and Hu et al. (2018), among others. The number of modelling studies
dealing with the injection of smoke into the UTLS is more reduced: Trentmann et al. (2006), Luderer et al. (2006),
Cunningham and Reeder (2009), Cammas et al. (2009), and Peterson et al. (2017).

105 During summer 2017, North America lived one of its worst burning season on record. On 16 August an aerosol index 106 (AI), a qualitative index indicating the presence of elevated layers of aerosols with significant absorption, of 55.4 was 107 recorded over Canada by the Ozone Mapping and Profiling Suite (OMPS) on board Suomi National Polar-orbiting 108 Partnership satellite (Seftor, 2017a). It breaks the record of AI values by far, the previous record being 31.2 registered 109 by the Total Ozone Mapping Spectrometer (TOMS) on 29 May 2001 during the Canadian Chisholm fires (Fromm et al., 2008). The cluster of the most intense fires of August 2017 was located in Canada near the intersection border of 110 111 Saskatchewan, Alberta and the Northern Territories at latitude 60 °N. These intense fires produced strong pyroCb 112 which injected smoke particles in the LS which travelled eastward, entrained and dispersed zonally by polar jet streams 113 (Khaykin et al., 2018). Smoke layers at 14 - 16 km with an aerosol optical depth (AOD) at 532 nm of 0.6 were 114 observed in Germany (Ansmann et al., 2018) on 22 August. The event is already documented by a series of papers: Khaykin et al. (2018), Ansmann et al. (2018), Haarig et al. (2018), Hu et al. (2018) and Baars et al. (2019). Fifteen 115 116 days later, on 30 August, AI from OMPS peaked again at 23 in a smoke plume detected over the southern parts of 117 Alberta and Saskatchewan and the upper Great Plains of the United States (US) (Seftor, 2017b). Most of the fires of 118 this new burning period were in the Pacific northwestern region (British Columbia, Alberta, Washington, Oregon, 119 Idaho, Montana and northern California). In the US severe air quality issues were reported in Washington and Oregon 120 at least until 6 September (NYT, 2018). Prevailing winds and the presence of a frontal boundary across the North 121 American continent created the conditions for the formation of a long, wide, arching ribbon of smoke that stretched thousands of kilometers from the source region all the way to Newfoundland, location from where it was further 122 123 transported towards Europe. The smoke hit the Iberian Peninsula (IP) in southwestern Europe on 7 and 8 September 124 (Sicard et al., 2018). Although the smoke plume was detected in the LS at some points during its transport, it was 125 only detected in the UT over the IP.

126 This paper investigates the time-space evolution of the smoke plume detected at its arrival over the IP on 7 and 8 127 September with ground-based multi-wavelength lidars and backward in time with the CALIOP (Cloud-Aerosol Lidar 128 with Orthogonal Polarization) spaceborne lidar, in terms of optical properties and vertical distribution. Sun-sky 129 photometers at mid-altitude, background sites with no local sources are used to monitor the smoke columnar properties 130 over the IP. A dispersion model parameterized with satellite-derived fire products simulates the vertical and horizontal 131 transport of 3 smoke-related compounds: carbon monoxide, black carbon and organic carbon. Simulations, and 132 especially the injection heights computed by the dispersion model, are qualitatively evaluated against observations 133 and used to understand the atmospheric causal mechanisms yielding to the differences observed in the optical 134 properties over the IP in the mid and upper troposphere.

135 2 Instrumentation and tools

136 The tools used in our methodology include passive/active, ground-based and spaceborne observations, as well as a 137 particle dispersion model. The observations, listed in For each fire simulated, a series of common parameters (in 138 brackets we indicate HYSPLIT denomination) are necessary: location (Release location), start time (Release start 139 time), duration (Release duration) and heat release (Heat release for plume rise). And for each chemical compound 140 simulated (gas or particle), the emission rate (Emission rate) is necessary. Such information is extracted from the 141 biomass burning emission estimates from GFAS (Global Fire Assimilation System; Kaiser et al., 2012) data from 142 CAMS (Copernicus Atmosphere Monitoring Service). GFAS Table 1, are used to follow the transport of the smoke 143 plumes from the source to the IP. In addition, passive spaceborne observations are also used to parameterize the 144 emission of the particle dispersion model.

145 2.1 Ground observations

The ground-based observations include lidars and sun-sky photometers in the IP. A total of five lidar systems are used: three from the EARLINET/ACTRIS (European Aerosol Research Lidar Network / Aerosols, Clouds, and Trace Gases Research Infrastructure Network; https://www.actris.eu/default.aspx; Pappalardo et al., 2014) network in Évora (EV), Granada (GR) and Madrid (MA), and two from MPLNET (Micro-Pulse Lidar Network; https://mplnet.gsfc.nasa.gov/; Welton et al., 2001) in El Arenosillo/Huelva (AR) and Barcelona (BA; see For each fire simulated, a series of common parameters (in brackets we indicate HYSPLIT denomination) are necessary: location (Release location), start time (Release start time), duration (Release duration) and heat release (Heat release for plume rise). And for each chemical 153 compound simulated (gas or particle), the emission rate (Emission rate) is necessary. Such information is extracted 154 from the biomass burning emission estimates from GFAS (Global Fire Assimilation System; Kaiser et al., 2012) data 155 from CAMS (Copernicus Atmosphere Monitoring Service). GFAS Table 1 for more details and Figure 1 for the 156 geographical position of the stations). The EARLINET lidars are multi-wavelength systems measuring at least at three 157 elastic wavelengths. In addition, EV and GR have Raman and depolarization-sensitive channels. The MPLNET 158 systems have one wavelength at 532 nm and an additional polarization-sensitive channel. A review of the lidar 159 techniques using elastic, Raman and depolarization-sensitive channels, among others, for the remote sensing of 160 aerosols can be found in Comerón et al. (2017). For the characterization of the smoke plume, we use the particle depolarization ratio, δ_p , in EV, AR, GR and BA and the pair (color ratio, depolarization ratio) in EV and GR. The 161 162 particle depolarization ratio and the color ratio provide significant information on the particle shape and dominant size 163 (Burton et al., 2012), respectively. The particle depolarization ratio is defined as (Freudenthaler et al., 2009):

164
$$\delta_p = \frac{\beta^{\square}}{\beta^{\square}}$$
(1)

165 where β^{\Box} and β^{\bot} are the particle parallel and perpendicular backscatter coefficients, respectively. The color ratio, 166 *CR*, is defined as a function of the particle backscatter coefficient at 532 nm, β_{532} , and at 1064 nm, β_{1064} , as:

167
$$CR = \frac{\beta_{532}}{\beta_{1064}}$$
(2)

The reason for using this definition of the color ratio between the wavelengths of 532 and 1064 nm is that it allows direct comparison with the space-borne lidar (see Section 2.2) and it is a common parameter used in aerosol classification (Burton et al., 2012; Groß et al., 2013). To understand the reasons of the differences and similarities found in the upcoming discussion, we also define the extinction-related Ångström exponent (AE) between the wavelengths of 355 and 532 nm:

173
$$\alpha - AE = -\ln\left(\frac{\alpha_{355}}{\alpha_{532}}\right) / \ln\left(\frac{355}{532}\right)$$
(3)

174 where α_{355} and α_{532} are the extinction coefficient at 355 and 532 nm, respectively. This quantity is calculated only 175 at EV which is the only stations where Raman inversions were successfully performed. Similarly to the color ratio, 176 $\alpha - AE$ provides information on the particle dominant size. The advantage of $\alpha - AE$ is that it can be directly 177 compared to the Ångström exponent retrieved by AERONET and defined in the next paragraph. While the MPLNET 178 and the EV systems work continuously 24/7, GR and MA measurements are discontinuous.

179 Due to the high vertical extension of the smoke plume (up to 14 km) and the high AOD values at 440 nm reached at 180 peak (0.6), no Raman inversions could be performed satisfactorily in Granada. Raman inversions performed in Évora 181 yielded a lidar ratio (the extinction-to-backscatter ratio), LR, at 532 nm, LR_{532} , in the mid troposphere smoke plume 182 on the order of 55 steradian (sr) (see Section 5.2). This value of 55 sr is used in the elastic inversions performed for 183 the other systems at both 532 and 1064 nm. To maximize the signal-to-noise ratio and thus minimize the retrieval uncertainties, all ground-based lidar measurements presented in this work are nighttime measurements. For the 184 185 EARLINET systems, the Raman-inverted extinction coefficient has an accuracy of 10 - 30 %, the backscatter coefficient of 5 - 10 % and the lidar ratio of 20 - 35 % (Ansmann et al., 2002). As far as elastic inversions are 186 concerned, the uncertainty of the backscatter coefficient is 10-20 % according to Ansmann et al. (2002) and the one 187 188 of the extinction coefficient is almost directly proportional to the uncertainty of the lidar ratio assumed. Thus, a 25% 189 uncertainty in the lidar ratio input parameter (assuming variations of 14 sr around 55 sr, see Section 5.2) of the elastic 190 inversion leads to a relative uncertainty of about 25% in the extinction coefficient. The particle depolarization ratio 191 uncertainty can reach up to 50 % in the UTLS (Rodríguez-Gómez et al., 2017). For the MPLNET systems, according 192 to Córdoba-Jabonero et al. (2018) the backscactter coeffcient and the particle depolarization ratio retrieved from MPL 193 data have a relative uncertainty of 5 to 20 % and of 10 to 60 %, respectively.

194 In order to monitor the event over the IP from columnar optical properties we looked at mid-altitude AERONET 195 (Aerosol Robotic Network; Holben et al., 1998) sites with no local sources so as to maximize the signature of the 196 smoke long-range transport. Such sites are Montsec in northeastern Spain and Cerro Poyos in south Spain (see For 197 each fire simulated, a series of common parameters (in brackets we indicate HYSPLIT denomination) are necessary: 198 location (Release location), start time (Release start time), duration (Release duration) and heat release (Heat release 199 for plume rise). And for each chemical compound simulated (gas or particle), the emission rate (Emission rate) is 200 necessary. Such information is extracted from the biomass burning emission estimates from GFAS (Global Fire 201 Assimilation System; Kaiser et al., 2012) data from CAMS (Copernicus Atmosphere Monitoring Service). GFAS 202 **Table 1** for more details and Figure 1 for the geographical position of the stations). We considered AERONET 203 Version 3 products: AOD and SDA (Spectral Deconvolution Algorithm; O'Neill et al., 2001; 2003) inversions data 204 level 1.5 in Montsec (level 2.0 is not available yet) and 2.0 in Cerro Poyos; and aerosol inversions data level 1.5 at 205 both sites. The AERONET products used in our work are:

• The AOD at 440 nm, AOD_{440} , which has an estimated accuracy of ± 0.02 (Eck et al., 1999).

- The Ångström exponent calculated between the wavelengths of 440 and 870 nm, $AE_{440-870}$, which has an accuracy of ±0.25 for $AOD_{440} \ge 0.1$ (Toledano et al., 2007).
- The fine mode fraction, *FMF*, which has an uncertainty of ~25% for an AOD at 500 nm greater than 0.3 (O'Neill et al., 2003). *FMF* represents the ratio of the fine-mode AOD to the total AOD.
- The absorption aerosol optical depth, AAOD, which has an accuracy of ±0.01 for wavelengths greater than 440
 nm (Sicard et al., 2016). AAOD represents the AOD due to absorption.
- The single scattering albedo, SSA, which has an accuracy of ± 0.03 for $AOD_{440} \ge 0.5$ for biomass burning (Sicard
- et al., 2016). SSA represents the fraction of the AOD due to scattering (i.e. AOD AAOD) to the total AOD.
- The asymmetry factor, g, which has an accuracy in the range $[\pm 0.03, \pm 0.08]$ for biomass burning (Sicard et al.,
- 216 2016). g represents a measure of the preferred scattering direction and varies between -1 (only backward-
- scattering, i.e., at 180° relative to the incident direction) and +1 (only forward-scattering at 0°).

218 2.2 Spaceborne observations

219 Several types of satellite sensors are used to fulfill the objectives of the study. The Atmospheric Infrared Sounder 220 (AIRS; Chahine et al., 2006), on board the Aqua satellite, is a hyperspectral instrument with 2378 infrared channels 221 and 4 visible/near-infrared channels. AIRS, together with the Advanced Microwave Sounding Unit (AMSU-A) and 222 the Humidity Sounder for Brazil (HSB), form the AIRS instrument suite which is designed to measure the Earth's 223 atmospheric water vapor and temperature profiles on a global scale. The physical product from AIRS used in our study 224 is the Carbon Monoxide (CO) Total Column science parameter which is a parameter of the AIRS Level 2 standard 225 retrieval product using AIRS only (AIRS2RET_NRT). It indicates the amount of CO in the vertical column of the 226 atmosphere and is measured in parts per billion by volume (ppbv). The spatial resolution of the AIRS2RET_NRT 227 product is 45 km at nadir. The temporal resolution is twice daily (day and night).

To track back the vertical distribution of the smoke plume before its arrival in the IP, we use the spaceborne lidar Cloud-Aerosol Lidar with Orthogonal Polarization (CALIOP; Winker et al., 2007), on board the Cloud-Aerosol Lidar and Infrared Pathfinder Satellite Observation (CALIPSO) satellite. CALIOP is a two-wavelength polarizationsensitive lidar that provides high-resolution vertical profiles of aerosols and clouds. It utilizes three receiver channels: one measuring the 1064 nm backscatter intensity and two channels measuring orthogonally polarized components of the 532 nm backscattered signal. The data used in our study are the CALIOP Aerosol Profile Lidar Level 2 data, 234 version 4.10. Profiles of extinction and backscatter coefficients at 532 and 1064 nm, as well as particle depolarization 235 ratio at 532 nm, are given at a horizontal resolution of 5 km and a vertical resolution of 60 m. The uncertainty in the 236 aerosol extinction coefficient is 40 % (assumed a 30-% lidar ratio uncertainty) and the one in the aerosol backscatter 237 coefficient is 20 – 30 % at 532 nm (Young et al., 2009). The CALIOP Level 2, version 4.10 data products used in this 238 study contain substantial changes over the earlier releases, among which the most significant is the updated lidar ratio 239 assignment (Young et al., 2018). Information on CALIOP aerosol sub-typing algorithm and assigned lidar ratios can 240 be found in Omar et al. (2018) and Kim at al. (2018). CALIOP observations have been used for the study of long-241 range transport of fire smoke locally (Kar et al., 2018) and also globally (Mehta and Singh, 2018).

The Moderate Resolution Imaging Spectroradiometer (MODIS; Kaufman et al., 2003), on board Aqua and Terra 242 243 satellites, is used for various purposes: 1) to quantify and monitor the smoke AOD at the global scale, 2) to confirm 244 the fires position and active period, and 3) to parameterize the smoke emission in the dispersion model. For the AOD 245 we use the near real-time value-added MODIS AOD level 3 gridded product (MCDAODHD) based on MODIS level 246 2 aerosol products combined from Aqua and Terra satellites. The sensor resolution is 0.5°, imagery resolution is 2 km, 247 and the temporal resolution is daily. For the fire information (position and active period), MODIS Fire and Thermal 248 Anomalies products, either from Terra (MOD14), Aqua (MYD14) or a combination of them (MCD14), are used. Each 249 MODIS active fire location represents the center of a 1-km pixel that is flagged by the algorithm as containing one or 250 more fires within the pixel.

251 2.3 Particle dispersion modeling

252 2.3.1 Model overview

253 Back-trajectory and dispersion calculations are performed with the Hybrid Single Particle Lagrangian Integrated 254 Trajectory Model (HYSPLIT; Stein et al., 2015; Rolph et al., 2017). HYSPLIT is developed at NOAA's Air Resources 255 Laboratory and is one of the most widely used models for atmospheric trajectory and dispersion calculations. It is a 256 complete system for computing simple air parcel trajectories as well as complex transport, dispersion, chemical 257 transformation, and deposition simulations. The model calculation method is a hybrid between the Lagrangian 258 approach and the Eulerian methodology. Apart from calculating back-trajectories, HYSPLIT is mostly used in this 259 study to calculate the transport, dispersion, and deposition of emitted CO (used as a tracer of the transport) and 260 particulate matter (black carbon, BC, and organic carbon, OC). The specificity of our HYSPLIT runs is that the heat release from the fires is used to estimate the smoke release height, i.e. no release heights were a priori set. The initial 261

262 particle height is assumed equal to the final buoyant rise height as computed using the method of Briggs (1969) with 263 the fire heat release given in input, implying that the final rise is a function of the estimated fire heat release rate, the 264 atmospheric stability, and the wind speed. Stein et al. (2009) tested the sensitivity of HYSPLIT to fixed and variable 265 release heights by comparing PM2.5 levels modelled and measured at the surface of northwestern US fires in 266 September 2006. They found that the case when the heat release from the fire was used to estimate the release height 267 showed the best performance, although they also concluded that the model is highly sensitive to variations in the 268 smoke release height and to whether the smoke injection actually occurred below or above the planetary boundary 269 layer. Rolph et al. (2009) also used HYSPLIT plume rise computation from the fire heat release.

270 For each fire simulated, a series of common parameters (in brackets we indicate HYSPLIT denomination) are 271 necessary: location (Release location), start time (Release start time), duration (Release duration) and heat release 272 (Heat release for plume rise). And for each chemical compound simulated (gas or particle), the emission rate (Emission 273 rate) is necessary. Such information is extracted from the biomass burning emission estimates from GFAS (Global Fire Assimilation System; Kaiser et al., 2012) data from CAMS (Copernicus Atmosphere Monitoring Service). GFAS 274 275 **Table 1:** Instruments used in this study. The nomenclature $3\beta+2\alpha+1\delta$ stands for 3 elastic channels (here, 355, 532, 276 1064 nm), 2 Raman channels and one depolarization channel; $1\beta+1\delta$ stands for 1 elastic channel (here, 532 nm) and 277 one depolarization channel; 3 β stands for 3 elastic channels (here, 355, 532, 1064 nm). 8- λ refers to the number (8) of

Ground	-based						
	Station /	Latitude, longitude, altitude	Instrument type				
	Network						
Active	EV /	38.57N, 7.91W, 293 m asl	$3\beta+2\alpha+1\delta$ lidar				
	EARLINET						
	AR /	37.10N, 6.73W, 59 m asl	1β+1δ lidar				
	MPLNET						
	GR /	37.16N, 3.61W, 680 m asl	$3\beta+2\alpha+1\delta$ lidar				
	EARLINET						
	MA /	40.45N, 3.72W, 669 m asl	3β lidar				
	EARLINET						
	BA /	41.39N, 2.11E, 115 m asl	1β+1δ lidar				
	MPLNET						
Passive	Cerro Poyos /	37.11N, 3.49W, 1830 m asl	8-λ sun-sky photometer				
	AERONET						
	Montsec /	42.05N, 0.73E, 1574 m asl	8-λ sun-sky photometer				
	AERONET						
Spacehorna							
Spacebo	Instrument	Satellite	Instrument type				
	msuument	Satemite	instrument type				

278	wavelengths	of the	photometers
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Active	CALIOP	CALIPSO	2β +1 δ lidar			
Passive	MODIS	Aqua and Terra	Moderate resolution imaging			
			radiometer			
	AIRS	Aqua	High-spectral resolution,			
			multispectral infrared sounder			

279

assimilates fire radiative power (FRP) observations from satellite-based sensors (Freeborn et al., 2014), namely MODIS/Aqua and Terra and SEVIRI (Spinning Enhanced Visible and InfraRed Imager), to produce daily estimates of biomass burning emissions. GFAS data (in brackets we indicate GFAS denomination) used in our work include daily information of the fire location and heat release (Wildfire radiative power), and for each chemical compound the emission rate (Wildfire flux). Data are available globally on a regular latitude-longitude grid with horizontal resolution of 0.125° x 0.125°. We used the current version of GFAS, i.e. GFAS v1.2. This work contains modified Copernicus Atmosphere Monitoring Service Information (CAMS, 2018).

The quantification of the HYSPLIT dispersion model uncertainties is not straightforward and it is usually performed through complex sensitivity studies (Mosca et al., 1998; Pielke and Uliasz, 1998; Straume, 2001; Warner et al., 2002). In general, the performance of dispersion models is largely attributed to uncertainty in the input fields (Challa et al., 2008). For our case, the GFAS data used for estimating the magnitude and timing of fire emissions have a typical uncertainty around 30% (Andela et al., 2013).



292

Figure 1: MODIS/Aqua corrected reflectance (true color) map centered over Spain on 8 September. Green bullets indicate lidar

stations (EV: Évora, AR: El Arenosillo/Huelva, GR: Granada, MA: Madrid, BA: Barcelona) and red bullets indicate AERONET

295 sites. Map created from https://firms.modaps.eosdis.nasa.gov/map/.

296 2.3.2 Model parametrization

297 The dispersion of CO, BC and OC is simulated in the forward direction, with a time resolution of 6 hours and at 15 298 altitude levels: one between 0 and 2.5 km and then 14 adjacent 1-km thick layers up to 16.5 km. The meteorology is 299 taken from GDAS (Global Data Assimilation System) data with a horizontal resolution of 0.5° x 0.5°. Noteworthy is 300 the fact that the first simulations with GDAS 1° x 1° meteorological data (not shown) simulated the dispersion of the 301 smoke plume too far north reaching France and Germany, instead of the IP. The use of the finer resolution of 0.5° x 302 0.5° improved significantly the arrival location of the plume, and put forward the importance of the horizontal 303 resolution of the meteorological data upon the correct dispersion of the emitted plume studied. The vertical limit of 304 the internal meteorological grid of HYSPLIT was set to 20 km, which, according to the following sections, is well 305 above the maximum height at which the smoke particles were observed. The daily number of active fires, the FRP 306 per fire and the emission rate per fire and chemical compound are from GFAS 0.125° x 0.125° data. Since the time 307 resolution of GFAS data is daily, the emission rate and FRP are assumed constant during the day the fires are active. 308 In all simulations 2500 particles were released to calculate the transport.

- 309 In the case of CO, dry deposition is neglected and wet removal is parameterized with a Henry's law constant of 9.9×10^{-1}
- 310 ⁴ mol atm⁻¹. BC (OC) is parameterized with the following values (Chin et al., 2002):
- 311 Particle radius: 0.0118 (0.0212) μm.
- 312 Particle density: 1.0 (1.8) g cm⁻³.
- Gravitational settling velocity (for dry deposition): 0.5 cm s⁻¹ for both types.
- Scavenging coefficient in- and below-cloud (for wet deposition): 8x10⁻⁵ s⁻¹ for both types.

315 3 Methodology

The proposed methodology is a two-way process, posterior to an initial phase (step 0) consisting in visualizing the "big picture" of the event at global scale with satellite images and back-trajectories. A flowchart of the methodology is shown in Figure 2. The first step of the methodology (step 1) consists in monitoring the smoke optical properties observed over the IP and their backward evolution back to the source with CALIOP retrievals. The second step (step 2) consists in parameterizing the smoke emission and run HYSPLIT forward simulations to obtain 4D (space and time) dispersion maps of the concentration of smoke-related compounds such as carbon monoxide, black carbon and

- 322 organic carbon. The main contribution of the modelling in Section 6 is to support the possible hypothesis made along
- 323 the discussion in Section 5.2.



324

325 **Figure 2:** Flowchart of the methodology.

326 4 Canadian/United States fires and general overview

BA

327 The first hint of the arrival and the presence of the smoke plume over the IP is given by the temporal evolution of a 328 combination of AERONET parameters in Montsec and Cerro Poyos, namely the AOD at 440 nm, $AE_{440-870}$, and FMF (Figure 3). According to Sola et al. (2014) the mean AOD at 500 nm ($AE_{440-870}$) in Montsec during the month of 329 September is ~0.1 (~1.5) which corresponds to an AOD at 440 nm of 0.12. In Montsec AOD₄₄₀ starts to exceed this 330 331 value on 4 September, day from which the AOD increases continuously until it reaches its peak value of 0.55 (0.54) 332 on 7 (8) September. These peak values of AOD are associated with values of $AE_{440-870}$ of 1.7 (1.6) on 7 (8) September. 333 On both days the fine mode fraction is higher than 0.98, leaving basically no room for the presence of coarse mode 334 EV AR 335 GR 336 MA



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Figure 3: AOD₄₄₀ (black), FMF (blue) and AE₄₄₀₋₈₇₀ (red) in (top) Montsec, northeastern Spain, and (bottom) Cerro Poyos, south
Spain. The gray areas in the bars on top of the figures indicate coincident lidar measurements.

341 smoke particles (diameter $> 1 \mu m$). In Cerro Poyos, the background AOD at 440 nm is even lower than in Montsec, 342 being smaller than 0.1 (AERONET, 2018). This value is exceeded from 5 September on, and the AOD increases until 7 September when it reaches its peak value of 0.62 with an associated $AE_{440-870}$ of 1.6. The fine mode fraction is 343 344 higher than 0.88 on both 7 and 8 September. The main difference between Montsec and Cerro Poyos is their proximity 345 to anthropogenic emissions: while Montsec is a remote site, far away from any industrial or large metropolitan area, 346 Cerro Poyos, although higher in altitude, is only 12 km SE of the city of Granada (~600,000 inhabitants including 347 metropolitan area). Due to its position with respect to Granada and the prevailing winds during the period under study, 348 Cerro Poyos was downwind of the city. This has several implications: the AOD in Cerro Poyos shows a diurnal cycle 349 related to the anthropogenic emissions of Granada, and the fine mode fraction is lower than in Montsec due to the 350 same emissions. However one can appreciate from Figure 3 that during the night of 7-8 September FMF in Cerro 351 Poyos is nearly 1 like in Montsec. In terms of AOD, the biomass burning contribution at both sites is roughly five times higher than the background values, and the peak values (~0.6) are considered very large for biomass burning long-range transport. In comparison the North American biomass burning event detected 15 days earlier (~22 August 2017) in northern Europe produced AODs at 500 nm near 1, $AE_{440-870} \sim 1.1$ and $FMF \sim 1$ (Ansmann et al., 2018).

355 To track back the plume transport in the atmosphere from the source to the IP, we use maps of columnar CO (AIRS) 356 and AOD (MODIS; Figure 4) as well as CALIOP curtains and HYSPLIT back-trajectories (Figure 5). The combined 357 day/night columnar CO maps are reported for the period 30 August - 8 September and a threshold of 95 ppb was 358 applied in order to highlight strong concentrations. The active fires are indicated by a red star (Figure 4) centered in 359 a region defined by the orange square visible in Figure 5 (bottom plot). This square covers the provinces of British 360 Columbia and Alberta (Canada) and the states of Washington, Oregon, Idaho, Montana and northern California (US) 361 where more than 90% of the active fires in North America are present during the period considered. The habitat type, 362 a little more south than the Canadian boreal forests, corresponds to temperate coniferous forests (Ricketts et al., 1999). 363 Some important forests in this region of North America are the National Forests of Wenatchee, Flathead, Nez Perce-364 Clearwater or Payette, among others, which, under the influence of both continental and maritime climates, produce 365 a large variety of ecosystems ranging from wet, western redcedar bottoms to high alpine peaks, and forests of alpine 366 larch and whitebark pine. So, from this region, a first plume (Plume 1) is released from the source region on 30 August, 367 travels E-NE on 31 August and then eastwards on 1 and 2 September. On 3 September a second plume (Plume 2) is 368 released from the same source region and starts travelling east, slightly SE. On 4, 5, and 6 September Plume 2 is 369 carried by the jet stream and travels rapidly towards the east, while at the same time Plume 1 travels slowly eastwards 370 above the Atlantic. The column concentration of Plume 2 is stronger than the one of Plume 1. On 7 September both 371 plumes merge into one and reach the IP. The high MODIS AOD values on 8 September over the IP confirm that the 372 high level of column CO is accompanied with high aerosol loads. These aerosols are also clearly visible especially in 373 the eastern part of the IP as a gray/brownish smoke shroud on MODIS true color image of 8 September (Figure 1). 374 The 10-day back-trajectories at selected heights are shown in Figure 5. Although we computed back-trajectories at

375 all lidar stations, only Madrid is shown as point of arrival for the sake of clarity of the figure and because Madrid is



376



Figure 4: Total column carbon monoxide (day/night) from AIRS/AQUA from 30 August until 8 September. The extra plot at the bottom to the right represents the MODIS combined (Aqua and Terra) value-added AOD at 550 nm on 8 September. The red star indicates the position of the active fires. On the plots of 3 and 4 September the descending, nighttime orbits of CALIPSO are reported. Maps created from https://worldview.earthdata.nasa.gov/.



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Figure 5: (top) 10-day back-trajectories, 1-hour resolution, arriving in Madrid, in the center of Spain, on 8 September at 00UT at heights of 3 (red), 6 (green) and 11 (blue) km; (bottom) Same back-trajectories, different viewing angle and superposition of CALIOP curtains on 4 September at 05:10UT (D-4, day-4 before arrival) and on 3 September at 09:23UT (D-5) where the smoke plumes, clearly visible, match very well in space and time with the back-trajectories. Pink crosses indicate active fires in the period 30 August – 5 September. The red rectangle of corner coordinates (125W, 40N; 93W, 58N) is the area in which the fires were taken into account in the dispersion modelling analysis (see Section 6). The orange rectangle simply highlights the region containing most of the fires. Maps created with Google Earth.

395 located in the center of the IP. The selected heights (3, 6 and 11 km asl) have been chosen by looking at the smoke 396 vertical distribution from the lidar data (see next section). All three trajectories pass over the region containing most 397 of the active fires (orange square): the trajectory arriving in Madrid at 3 km passes over this region on 31 August at 398 08UT (7-8 days of transport) and the trajectories arriving at 6 and 11 km, very similar in path and speed, pass over the 399 region of the fires on 3 September between 17 and 22UT (4-5 days of transport). These results suggest that the 400 airmasses arriving above the IP at 3 km on 8 September picked up smoke most likely from Plume 1, while those 401 arriving at 6 and 11 km most likely from Plume 2. During the transport CALIPSO orbits intersect the back-trajectories 402 in space and time in two occasions: once southeast of Greenland on 4 September at 05:10 UT (D-4, 4 days before 403 arrival in Madrid, intersects with Plume 1 which is 5 days old) at 3 km height, and another time on 3 September at 404 09:23 UT (D-5, intersects with Plume 2 which is less than 1 day old) at 6 and 11 km heights. On D-5 the shortest distance between CALIPSO curtain and the center of the region of the fires (orange square) is 700 km. On both occasions the cloud-free CALIOP curtains show clearly the large spatial extension of the smoke: 1700 km (below orbit) x 15 km (height) on D-4 and 1100 km (below orbit) x 8-9 km (height) on D-5. The attenuated backscatter of CALIOP on D-5 is clearly much stronger than on D-4 because of the proximity of the orbit to the source region. On D-5, in the southernmost part of the plume, most of the smoke between 4 and 7 km height is optically so thick that it attenuates the lidar signal below it.

411 **5** Optical properties of the smoke particles

412 Many papers, most of them listed in the literature overview of Ortiz-Amezcua et al. (2017) or of Haarig et al. (2018), 413 deal with the optical properties of long-range transport smoke particles derived from observations of photometers, 414 lidars or a combination of them. More general aerosol-typing literature based on lidar remote sensing and including 415 biomass burning are available in Burton et al. (2012), Groß et al. (2013), Illingworth et al. (2015) and Baars et al. 416 (2016; 2017).

417 5.1 Column-averaged properties

418 Figure 6 shows the spectral AAOD, SSA and asymmetry factor retrieved from AERONET sun-sky photometer measurements at Montsec and Cerro Poyos on 7 and 8 September. Several aspects are noteworthy. AAOD, similar 419 420 at both sites in absolute values, is surprisingly very low. Compared to the climatological AAOD representative of 421 boreal forests from US and Canada (Russell et al., 2010), recalculated from Dubovik et al. (2002), our AAOD values 422 are 2 to 3 times lower. As a consequence of the small AAOD observed in Montsec and Cerro Poyos, SSA is large 423 (~0.98 at 440 nm) and indeed much larger than the climatological values for boreal forest biomass burning (0.94 at 424 440 nm) from Dubovik et al. (2002). However it is in the range of values of SSA at 355 nm obtained in Europe by 425 Markowicz et al. (2016), 0.91 - 0.99, and Ortiz-Amezcua et al. (2017), 0.965 - 0.991, in smoke plumes originating from North America in July 2013. In particular Markowicz et al. (2016) attribute these high SSA values to "a 426 427 transformation of [biomass burning] during long-range transport [...] and mixing of the [biomass burning] with non-428 absorbing aerosol species". The high transport altitude of the fire smoke observed over the IP in summer 2017 makes 429 the second hypothesis (mixing with non-absorbing aerosol species) highly improbable. If these low AAOD and high





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433 Figure 6: AERONET daily mean spectral (top) AAOD, (center) SSA, and (bottom) asymmetry factor at Montsec and Cerro Poyos on 7 and 8 September.

435 SSA were due to low BC emission at the source, the following rationale can be made. According to Radke et al. 436 (1991) and more recently to Russell et al. (2014) the absorption properties of biomass burning in its smoldering 437 combustion phase are lower than during its flaming phase, the reason being a larger production of black carbon in the 438 flaming phase relative to the smoldering phase (Radke et al., 1991). In addition, smoldering combustion occurs over

439 a much longer period of time relative to the comparatively short lives flaming phase of tree-crown fires of, e.g., pines, 440 cedar or cypress that commonly populate temperate coniferous forests. These results suggest that the smoke particles observed over the IP might be the product essentially of smoldering combustion at the source. We also recall that 441 442 level 1.5 AERONET data are not totally quality assured and that the values of AAOD and SSA should be taken with certain caution. The values and spectral behavior of g in Montsec and Cerro Poyos are in good agreement with 443 444 results for biomass burning aerosols from other studies (Dubovik et al., 2002; Sayer et al., 2014; Nikonovas et al., 445 2015), with g presenting a sharp decrease with increasing wavelength. Nikonovas et al. (2015) distinguished the 446 behaviour of fresh (within the first 24 h) and aged (more than 72 h) smoke and values reported therein for aged smoke 447 agree quite well with the mean values obtained at Montsec and Cerros Poyos on 7 and 8 September 2017 (~0.70 at 448 440 nm and ~0.53 at 1020 nm). This spectral behavior is typical of the dominance of fine particles that are scatterers 449 of solar radiation more efficient at lower wavelengths, with the forward scattering decreasing with increasing 450 wavelength.

451 We analyze the wavelength dependence of AAOD by means of the absorption Ångström exponent, AAE, calculated 452 between the wavelength of 440 and 870 nm. Although a clear difference is observed between both sites in terms of AAE: 1.24 < AAE < 1.35 in Montsec and 0.99 < AAE < 1.03 in Cerro Poyos, conclusions are not straightforward. 453 454 We rely our discussion on the results of Lack and Cappa (2010). According to these authors, the AAE for pure BC 455 cores varies in the range [-0.2, +1.3], for BC cores coated in non-absorbing matter (i.e. coated with a purely scattering 456 shell) it can be as high as 1.6 -1.7, and for BC cores coated in absorbing matter, namely brown carbon (mildly 457 absorbing organic matter; Andrea and Gelencsér, 2006), it is usually greater than unity even if for certain combinations 458 of core/shell size pairs and values of the imaginary part of the refractive index, it can be close to unity. Thus the absolute attribution of BC or brown carbon is hampered when AAE < 1.6. However Lack and Cappa (2010) also 459 460 showed that high SSA values (> 0.9) could only be achieved for BC cores coated in absorbing matter. The results 461 allow us to conclude (i) that, without any doubt, the AAE values in Montsec (~1.3) are representative of brown carbon 462 (or BC coated in brown carbon), likely contained in the long-range transport smoke plume detected, and (ii) that the 463 AAE ~ 1 in Cerros Poyos is most probably caused by brown carbon from biomass burning origin and maybe pure BC 464 from the anthropogenic fossil fuel emissions of the nearby city of Granada. Another possible reason for AAE ~ 1 in 465 Cerros Poyos may be the presence of nearby persistent local fires in Sierra Morena, approximately 150 km northwest of Cerro Poyos. For comparison, Bergstrom et al. (2007) measured an AAE of 1.45 in the range 325 – 1000 nm in a 466

plume of South Africa biomass burning with data from the SAFARI (Southern Africa Regional Science Initiative)campaign.

469 5.2 Vertically-resolved properties

470 To relate smoke optical properties and their vertical distribution, we use ground- and space-borne lidar profiles. The 471 availability of lidar measurements in the period 3-9 September is indicated by the gray areas in the bars of Figure 3. 472 Because of the high aerosol load and the high vertical extension of the plumes (> 10 km) on the night of 7 to 8 473 September and their implication on the signal-to-noise ratio of the lidar signals and thus on the quality of the 474 inversions, Raman inversions were performed only the night of 6 to 7 September. In Figure 7 we show the result of a 475 Raman inversion in Évora on 7 September between 04 and 06UT. Although this measurement time is a few hours 476 before the arrival time fixed for the back-trajectory simulations (8 September at 00UT) our back-trajectory analysis 477 (not shown) confirms that air mass paths were very similar during the 48 hours of both days 7 and 8 September. A series of quality checks have been applied to Évora lidar profiles: negative optical properties are not considered and 478 intensive properties ($\alpha - AE$, $\beta - AE$, CR, LR and δ^{p} , see caption of Figure 7 for symbol definition) are 479 480 calculated only for optical properties greater than a minimum threshold in order to guarantee the presence of aerosols 481 and to avoid physically meaningless retrievals. In addition to the profiles, smoke layer-mean values are given in two 482 altitude ranges corresponding to the mid and upper troposphere. These layer mean values are also reported in Table 483 2. We find smoke particles up to 12.7 km, below the tropopause height, with a clear plume extending from 2.3 to 8.1 484 km and a very shallow one from 11.5 to 12.7 km. The AOD at 532 nm is 0.24. The color ratio is significantly different 485 in the two altitude levels considered (~ 2.49 in the mid troposphere and ~ 3.31 in the upper troposphere). For 486 comparison Haarig et al. (2018) found color ratios of 1.8 and 2.3 in the troposphere and the stratosphere, respectively, 487 for the North American biomass burning detected in northern Europe 15 days earlier (22 August). Our higher values 488 indicate particles of smaller size. This finding is corroborated by the columnar effective fine mode radius measured 489 by AERONET in Montsec and Cerro Poyos on 7 September which vary in the range $0.14 - 0.18 \mu m$, while values larger than 0.23 μ m were found during the 22 August event (Ansmann et al., 2018). $\alpha - AE$, only retrieved in the 490 491 mid troposphere, is 1.51. For comparison Haarig et al. (2018) found a $\alpha - AE$ of 0.9 in the troposphere. The difference with our $\alpha - AE$ is probably due to different absorption properties: low in our case and rather large on 22 492 August (Ansmann et al., 2018). Low absorption properties yield to $\alpha - AE$ very similar to the scattering Ångström 493

494 exponent. According to Valenzuela et al. (2015) scattering AE larger than 1.5 indicates that submicron particles

495 dominate the aerosol size distribution, which is in agreement with our findings.



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Figure 7: Nighttime multi-wavelength lidar inversion in Évora on 7 September between 04 and 06UT. The first plot represents the quicklook of range-square corrected signal at 1064 nm in arbitrary units. β is the particle backscatter coefficient, α the particle extinction coefficient, $\alpha - AE$ the extinction-related AE, *CR* the color ratio, *LR* the lidar ratio and δ_p the particle depolarization ratio. Mean values in the mid troposphere and stratosphere (as depicted by the gray rectangles) for $\alpha - AE$, *CR* 501 , *LR* and δ_p are reported in the plots. The horizontal dash lines at 13.6 km indicate the tropopause height calculated with 1° x 1° GDAS data.

503 The lidar ratio at 532 nm is 55.1 \pm 14.2 and 34.3 \pm 10.5 sr in the mid and upper troposphere, respectively. Our LR_{532} 504 values are slightly lower than those of Haarig et al. (2018), 65 - 80 sr, if we take into account the standard deviations 505 associated to our retrievals, but are definitely in the range of literature values (26 - 80 sr) for North American biomass 506 burning detected in Europe (Ortiz-Amezcua et al., 2017). A lower lidar ratio in the upper troposphere compared to 507 the mid troposphere might indicate less absorbing particles in higher altitude (Ortiz-Amezcua et al., 2017), maybe 508 related to a lesser amount of BC with respect to organic carbon, or to a lesser degree of coating on BC since coatings 509 on BC enhance scattering and absorption properties (Cheng et al., 2014). These hypotheses are further investigated 510 in the next section. At 355 nm we find in Évora $LR_{355} = 41.8 \pm 6.8 \, sr$ in the mid troposphere, which is in good agreement with the range of 40 - 45 sr found by Haarig et al. (2018) and with the range of literature values (26 - 80511 512 sr) from Ortiz-Amezcua et al. (2017).

513 Table 2: Layer mean values of the color ratio, the Ångström exponent, the lidar ratios at 355 and 532 nm and the particle

		UTLS		Mid troposphere			
Parameter	Évora	CALIOP D-4	CALIOP D-5	Évora	CALIOP D-4	CALIOP D-5	
	7/9 at 04UT	Plume 1	Plume 2	7/9 at 04UT	Plume 1	Plume 2	
Color ratio	3.31±0.27	3.06±1.19	-	2.49±0.24	2.17±0.28	1.86±0.36	
α-AE	-	NA	NA	1.51±0.76	NA	NA	
LR_{355} (sr)	-	NA	NA	41.8±6.8	NA	NA	
<i>LR</i> ₅₃₂ (sr)	34.3±10.5	NA	NA	55.1±14.2	NA	NA	
δ_{p}	0.12±0.02	0.12±0.02 0.12±0.03		0.04±0.01	0.05±0.01	0.05±0.02	

514 depolarization ratio at 532 nm in Évora, and from CALIOP on D-4 and D-5.

515

516 Last but not least, the analysis of the profile of the particle depolarization ratio at 532 nm also reveals interesting results. The layer mean values of δ_p are 0.04 and 0.12 in the mid- and upper troposphere, respectively. While the 517 518 mid troposphere value falls in the range of literature values (Ortiz-Amezcua et al., 2017) and indicate spherical or 519 almost spherical smoke particles, the value of 0.12 in the upper troposphere is rather unusual. Some works 520 investigating the inter-continental transport of North American fire smoke to Europe from August 2017 also report 521 unusually high depolarization ratios (Khaykin et al., 2018; Haarig et al., 2018; Hu et al., 2018; Sicard et al., 2018) up 522 to 0.20 at 532 nm in the stratosphere. The causes of such high depolarizing capabilities of smoke particles are still 523 not well understood. Recently Burton et al. (2015) made a nice discussion based on literature to explain the high 524 values of three-wavelength depolarization ratios and their spectral dependence that they observed for smoke particles 525 from North American fires retrieved by high-spectral resolution lidar. They proposed two possible explanations of 526 the depolarization by smoke: the "lifting and entrainment of surface soil into the smoke plume and asymmetry of smoke particles themselves". Haarig et al. (2018) hypothesized that high δ_p values may be the result of dried out 527 528 smoke particles (relative humidity ~0 %) with a non-spherical shape. This hypothesis, however, is probably unlikely 529 in the range of altitude considered here (< 13 km over the IP) as radiosoundings in Barcelona (not available in Évora) 530 on 8 September at 00 UT indicate a relative humidity in the range 20 - 30 % in the upper troposphere. At this stage of the paper, our intention is not to give a single explanation of our high δ_p values, as we believe that the main 531 532 features observed over the IP (injection in the upper troposphere, low absorption and high depolarization properties) 533 are somehow connected, but to list some fire characteristics and physical/chemical mechanisms which could lead to 534 such features: the burnt material at the source (BC and OC contents), flaming versus smoldering phases, fire power, 535 BC aging processes (coagulation, condensation, and heterogeneous reactions) during transport resulting in changes in

536 its morphology and mixing state, relative humidity. Literature on these issues can be found in Fromm et al. (2003; 537 2008), Zhang et al. (2008), Lack and Cappa (2010), Adachi et al. (2010), Cheng et al. (2014), China et al. (2015). 538 Forrister et al. (2015), Burton et al. (2015), among many others, and will be used in the discussion of the next section. To analyze the spatio-temporal evolution of the smoke transport, we compare the smoke CR and δ_p profiles from 539 540 Évora (7 September at 04 UT) with CALIOP retrievals in Plume 1 (5 days old, D-4) and in Plume 2 (fresh < 1 day, 541 D-5). CALIOP retrievals are shown in Figure 8. CALIOP quicklooks of the total attenuated backscatter at 532 nm 542 show a spatial extension clearly larger for Plume 1 than for Plume 2 both horizontally and vertically. Plume 1 extends 543 up to ~15 km and into the stratosphere while Plume 2 stays in the troposphere below 9 km. This result suggests that 544 the UTLS injection of smoke particles does not occur immediately a few hours after fire ignition but during the 545 transport. Indeed Cammas et al. (2009) simulated with the anelastic non-hydrostatic mesoscale model Meso-NH the time needed for a boundary layer tracer to reach the tropppause to be about 7.5 hours. Logically, the particle 546 backscatter coefficient at 532 nm is much stronger in Plume 2 (> 5 Mm⁻¹sr⁻¹ below 4 km at latitudes of $48 - 50^{\circ}$ N) 547 548 resulting in a high AOD at 532 nm of 1.20 (versus 0.78 for Plume 1). The color ratios in the troposphere are 2.17 and 549 1.86 for Plume 1 (5 days old) and Plume 2 (< 1 day), respectively, indicating a decrease of the particle size as the 550 plume gets older. In Évora CR is 2.49. In the stratosphere the color ratio of Plume 1 is 3.06, while it is 3.31 in the 551 upper troposphere in Évora. Given the large standard deviation of CALIOP CR retrieval in the stratosphere (Table 552 2), the relatively small difference between both values (3.06 and 3.31) cannot be interpreted as a decrease of particle 553 size. In fact, once in the UTLS the smallest particles (with radii $< 0.5 \mu m$), tend to maintain at their altitude level or 554 to ascend. Rohatschek (1996) and Pueschel et al. (2000) explained the self-lofting of UTLS-level BC with the gravito-555 photophoresis mechanism consisting in sunlight-induced upward forcing. It is interesting to note a significant increase 556 of CR (> 3) close to the ground in Plume 2 which probably reflects freshly emitted, small soot particles, before they 557 undergo any of the various aging processes that lead to their size increase. The CALIOP particle depolarization ratio at 532 nm in the troposphere is 0.05 in both plumes, a value similar to $\delta_p = 0.04$ found in Évora in the mid 558 559 troposphere. It is an indication that the smoke particle depolarizing capabilities, and subsequently also their shape, in the troposphere are stable during transport. In the stratosphere δ_p in Plume 1 increases from 0.04 to a peak value of 560 561 0.16, the mean value being 0.12. In Évora the same value of 0.12 is found in the upper troposphere between 11.5 and 562 12.7 km. As far as Plume 1 is concerned, the smoke particles reached the UTLS in less than 5 days after their release 563 in the atmosphere and it seems that the smoke particle depolarizing capabilities (and thus their shape) at UTLS level are also stable during transport. The quasi-linear increase of δ_p with height may be an indication of the heightdependence of the ongoing aging processes leading to the transformation of the smoke particle depolarization properties from low- (0.04 at 12.25 km) to moderately-depolarizing (0.16 at 14.95 km). In Évora δ_p in the upper troposphere does not seem height-dependent as the particles must have already undergone these aging processes.



571 Figure 8: CALIOP images and products on (left) 4 September at 05:10UT (D-4, Plume 1 released 5 days earlier) and (right) 3 572 September at 09:23UT (D-5, Plume 2, fresh < 1 day). (top) CALIOP quicklooks of the total attenuated backscatter signal at 532 573 nm; (center) CALIOP quicklooks of the retrieved backscatter coefficient at 532 nm restricted to the smoke plume (red squares); 574 (bottom) CALIOP mean profiles of backscatter coefficient at 532 and 1064 nm, the color ratio and the particle depolarization ratio 575 at 532 nm. The horizontal black dash lines indicate the tropopause height calculated with 1° x 1° GDAS data.

576 To close this section, we compare the time-height evolution of mid and upper tropospheric particle depolarization 577 ratio at all Iberian lidar stations (capable of measuring particle depolarization) plus CALIOP, and the dependency of δ_p versus CR in Évora and Granada plus CALIOP. The results are shown in Figure 9. The reason for choosing to 578 plot δ_n versus *CR* is twofold: 1) they are the two intensive parameters provided by CALIOP, and 2) low-level aerosol 579 580 typing is possible with these parameters (Groß et al., 2013), although the classification they propose also includes the 581 lidar ratio. The profiles in the IP were selected during the night of 7 to 8 September, close to the back-trajectory 582 arrival time (8 September at 00UT), according to measurement availability and clear-sky conditions. They all fall 583 around the back-trajectory arrival time -3/+1 hour. Contrarily to the observations of North American smoke in the 584 stratosphere in France (Khaykin et al., 2018; Hu et al., 2018) and Germany (Ansmann et al., 2018; Haarig et al., 2018) 585 earlier in August, 2017, over the IP no aerosols are observed in the stratosphere in the period considered. At all 586 stations of the IP, a continuum of aerosols is observed up to the upper troposphere: aerosols are present in the whole 587 troposphere. In order to identify representative layers and give layer mean values, we selected in both the mid and 588 upper troposphere the layers centered around the backscatter coefficient peak value in each altitude range. This 589 methodology guarantees a higher representativeness of the smoke particle properties, but in certain cases the selected 590 layer may be spatially thin which may bias the interpretation of the top plot of Figure 9. Hence the layer height and 591 its thickness represented in this plot has to be interpreted as the layer of maximum intensity, i.e. of maximum aerosol 592 load. Before entering in the discussion, it is worth noting the important difference between the AOD at 532 nm in 593 Barcelona (0.65) and the rest of the stations of the IP (0.27 - 0.34). This difference is indeed not that surprising if we 594 look back at MODIS AOD on 8 September (Figure 4) which clearly shows a decreasing AOD tendency along the axis NE-SW. In the mid troposphere, with the exception of Granada, all measurements including CALIOP give a particle 595 596 depolarization ratio of 0.05 - 0.06. This result reflects again that in the mid troposphere the smoke particle 597 depolarization ratio was neither time- nor plume-dependent. The particle depolarization in Granada at ~7 km is 0.01 598 and clearly indicates non-depolarizing particles, slightly different from what is observed over the other stations of the 599 IP in the mid troposphere. The layers of maximum intensity are rather high (between 6 and 8.3 km) and are higher in 600 Évora, Granada and El Arenosillo/Huelva than in Barcelona. The center of the youngest plumes (CALIOP) are slightly below the layers of maximum intensity detected over the IP. In the upper troposphere δ_p varies between 0.09 and 601 0.10 over the IP, while in the low stratosphere $\delta_p = 0.12$ is found for Plume 1 (CALIOP, D-4). Given the large 602 standard deviations associated to the retrieval of δ_p in the UTLS, these findings are here again not sufficient to point 603

out a clear difference between the plume over the IP on D-0 and CALIOP on D-4. In case this difference is real, at
this stage our findings only allow to give hypothetical explanations to be taken with care, listed in order of likelihood:
1) the stations over the IP are representative of Plume 1 + Plume 2 while CALIOP D-4 is representative of Plume 1;
on D-0 the plume is in the upper troposphere while on D-4 it is in the stratosphere; and 3) a transformation of the
smoke depolarizing capabilities between D-4 and D-0.



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Figure 9: (top) Mid and upper tropospheric layer mean particle depolarization ratios at 532 nm at all Iberian lidar stations on the night of 7 to 8 September. Cyan and Purple bullets represent CALIOP measurements. The vertical bars indicate the vertical extension of the smoke layers of maximum intensity (base to top height). The horizontal bars indicate the standard deviation associated to δ_p in these layers. (bottom) Layer mean particle depolarization ratios at 532 nm vs. layer mean color ratio. The bullet color code is the same as in the top plot. We have reported four aerosols classes adapted from Groß et al. (2013). The vertical and horizontal bars indicate the standard deviation associated to δ_p and *CR*, respectively.

The bottom plot of Figure 9 represents pairs of (δ_p , *CR*) mean layer values in Évora and Granada (D-0) and in Plume 1 (D-4) and Plume 2 (D-5). The color-coded shaded areas, representative of different aerosol classes, are adapted from Groß et al. (2013). The results can be summarized as follows:

620 UTLS-level smoke particles have large color ratios ($\sim 2.5 - 3$) and moderate particle depolarization ratios (~ 0.10). • 621 Between D-4 and D-0 a small decrease of both parameters, within the statistical variability of one another, is noted. In the mid troposphere, $\delta_p = 0.05$ is stable with time, except in Granada. CALIOP *CR* values are smaller than in 622 Évora (and also Madrid (not shown) as $CR \square 2.95$ in the smoke layer centered around 6 km on 7 September at 623 624 21UT; in Madrid no smoke layer was observed in the UTLS), indicating that as the smoke gets closer to its arrival 625 in the IP, the particles get smaller. The difference between CALIOP CR values cannot be evaluated since they 626 correspond to two different smoke plumes which may initially have different morphology and thus different optical properties. The results obtained in Granada ($\delta_p = 0.01, CR = 4.30$) are an indication of ultrafine, non-depolarizing 627 628 particles and reveals a clear difference in the smoke properties with the rest of observations in the mid troposphere. 629 The back-trajectories in all three southern stations (EV, AR and GR) are very similar and do not allow to give an 630 explanation related with long-range transport. In turn, locally, Granada may have been exposed to nearby persistent fires in Sierra Morena, approximately 150 km northwest of the city. Fresh smoke produces low δ_p and large CR 631 632 , but this is only an hypothesis at this stage.

• The pairs of (δ_p, CR) fall in the Canadian biomass burning type, but often on the edges. In the mid troposphere, except in Granada, the pairs of (δ_p, CR) actually overlap between the classes of Canadian and African biomass burning and marine aerosols. It is worth recalling that the fires studied are not exactly "Canadian biomass burning", which stands for boreal forest fires in the literature, but fires from temperate coniferous forests. This result calls for further investigation on biomass burning properties in relation to their origin which goes beyond the usual Amazonian, African and North American classes.

Note *en passant* that the mid and upper tropospheric values of δ_p and *CR* in Évora on 7 September at 21UT (close in time to the back-trajectory arrival time of 8 September at 00UT, so called D-0) are not significantly different from the values found on 7 September at 04UT, for which the Raman inversion was performed (see Figure 7 and associated text).

643 6 UTLS injection and inter-continental transport

644 In order to investigate the role of each of the plumes identified in Figure 4, each plume is simulated separately and

- 645 then together. The emission of Plume 1 is set to 30/8 1/9 (3 days) and the emission of Plume 2 to 2/9 5/9 (4 days).
- 646 Only fires falling inside the red square defined in Figure 4 are considered. For each chemical compound (CO, BC and
- 647 OC), three simulations are run, corresponding to:
- Plume 1 (noted P1 from now on) from 30/8 until 8/9 with emissions limited to the period 30/8 1/9.
- Plume 2 (P2) from 2/9 until 8/9 with emissions limited to the period 2/9 5/9.
- Plume 1 + Plume 2 (P1+2) from 30/8 until 8/9 with emissions from 30/8 until 5/9.

From the results of Section 4, the hypothesis is implicitly made that the emitted matter before 30/8 and after 5/9 is not

- affecting the IP on 7 and 8 September. The fire characteristics are summarized in Table 3.
- Before entering in the discussion, we recall the questions raised in Section 5.2 and left opened: 1) injection mechanisms

responsible of the injection in the upper troposphere, 2) smoke particles with low-absorbing properties and decrease

655 of the absorption properties with height, 3) high depolarization properties, and 4) differences observed between the

656 smoke plume observed in the IP (D-0) and the younger plume (D-4) observed by CALIOP.

- **Table 3:** Characteristics of the fires at the origin of the emission of Plume 1 (emission: 30/8 1/9), Plume 2 (emission: 2/9 5/9)
- 658 and for the whole period (emission: 30/8 5/9). The data are from GFAS daily estimates of biomass burning emissions.

		P1			P2			P1+2			
Simulation period		30/8 - 8/9 (10 days)			2/9 – 8/9 (7 days)			30/8 - 8/9 (10 days)			
Emission period		30/8 – 1/9 (3 days)			2/9 – 5/9 (4 days)			30/8 – 5/9 (7 days)			
Number of active		836		772			1073				
fires											
Number of active		1843			2123	2123		3966			
fires x day											
FRP per fire (MW)	Min	0.1			0.1			0.1			
	Mean	95.1		137.0		117.5					
Max		5405.7		7162.2		7162.2					
Number of fires 232			190		277						
with $FRP > 50 MW$											
		P1		P2		P1+2					
		CO	BC	OC	CO	BC	OC	CO	BC	OC	
Emission rate per	Min	0	0	0	0	0	0	0	0	0	
fire	Mean	26.63	0.15	2.23	32.39	0.17	2.79	29.71	0.16	2.53	
(T h ⁻¹)	Max	2031.15	10.80	175.79	2561.98	13.63	221.73	2561.98	13.63	221.73	

During the emission period 30 August – 5 September a total of 1073 fires were detected in the domain considered (red square, Figure 5), which in terms of fires per day, defined as the sum of all fires multiplied by the number of days they were active, represents nearly 4000 emitting fires. From these numbers one can deduce that, on average, each fire had an emission duration of approximately 4 days. The mean FRP is 117.5 MW and it is approximately 50 % higher during P2 than during P1. Maxima are also higher during P2 (7162.2 MW) than during P1 (5405.7 MW). Two hundred and seventy seven fires had a FRP larger than 50 MW. For comparison, Ansmann et al. (2018) reports a number of 10000 fires with FRP larger than 50 MW in Canada for the month of August 2017.

667 Before presenting the results of the dispersion analysis, we will make a point about the geographic location of the fires 668 of this study. A notable difference between the fires of August (Khaykin et al., 2018; Ansmann et al., 2018; Haarig 669 et al., 2018; Hu et al., 2018) vs. September 2017 (this study) is the latitude at which they occurred. In August the most intense fires were located in Canada, 49 < latitudes < 67 °N (see Ansmann et al., 2018), while the emission region 670 671 considered in our study goes from British Columbia down to northern California, 40 < latitudes < 53 °N. In this lower part of the mid-latitude region, air masses can be under the influence of either the polar or the subtropical jet streams, 672 and therefore be entrained either north- or south-ward, respectively. The latitude difference, 49 - 67 °N vs. 40 - 53673 674 "N, also results in a higher tropopause height in September than during the August event, and also a thicker 675 troposphere-stratosphere transition layer due to the vicinity of the subtropical jet in September (Pan et al., 2004). It 676 has also another important implication: the material burnt, and consequently the content of emitted CO, BC and OC, 677 are different. While the fires in August were from boreal forests, in September the fires occurred in a region of 678 temperate coniferous forests. According to Lavoué et al. (2000) the main difference between boreal and temperate 679 forests does not rely on the canopy itself, but in the shrubland and the grassland which are more abundant in temperate 680 forests. McMeeking et al. (2009) who made controlled laboratory burns of Alaskan spruce and forest floor (duff), 681 among other fuels, found that the forest floor has a strong contribution from smoldering combustion, but a lower 682 carbon monoxide emission factor than most of the other fuels because it contains less carbon per mass unit. This 683 result supports low BC and OC contents at the source and thus low absorption properties as discussed in Section 5.1. Results are shown in Figure 10 and Supplements (S) 1-12. Figure 10 shows the dispersion maps of CO, BC and OC 684 over the IP at time of arrival on 8 September at 00UT in terms of column density, i.e. the concentration integrated 685 686 along the vertical axis. S1-S4, S5-S8 and S9-S12 are 6-hour time resolution animated gif images of the dispersion of, 687 respectively, CO, BC and OC for the total column density, and the concentration at 3, 6 and 11 km. The heights of 6 688 and 11 km are representative of the mid- and upper troposphere, where smoke particles were detected in the IP. The 689 images of the total column density (S1, S5 and S9) are the same as in Figure 10, i.e. a dispersion map on top and a 690 longitudinal cross-section below at the latitude of Madrid, taken as a central point in the IP. On all dispersion maps, 691 the color bar for P1 goes from red (low) to yellow (high) and for P2 from blue (low) to green (high). On a horizontal 692 scale all three compounds have similar dispersion patterns. For this reason, the interpretation of the dispersion maps 693 is made independently of the compound. Partly because the simulation time of P1 is longer relative to P2, its dilution 694 in the northern hemisphere is wider and circulations around the globe start to be visible, although in low concentration 695 level, in the final dispersion maps of all compounds. On 8 September CO/BC/OC from P1 are present around the pole 696 and also in eastern Russia. Interpretations of S1, S5 and S9 confirm that:

P1 is transported northeast-ward the first four days of the simulation. On 3 September a large swath of the US is
 covered by P1. This feature is confirmed by OMPS images (Seftor, 2017c). Later, as a large smoke tongue travels
 slowly over the Atlantic towards the IP (this transport coincides with the interpretation of the satellite images in
 Figure 4, see Section 4), large scale jets make the plume start meandering anti-clockwise around a point centered
 initially above Iceland which drifts slowly with time towards Ireland.

P1 reaches the IP (Madrid) on 6 September at 12UT with column density levels of CO of 0.02 mg m⁻². For
 comparison on 8 September at 00UT the CO column density of P1 is on the order of 0.83 mg m⁻².

P2 is travelling eastward since the first day of emission. On 4 September at 18UT one can already observe the
 beginning of the stretching of P2 located on the cyclonic-shear side of a strong jet, probably of subtropical origin
 since it ends up in northern Africa. Later a relatively thin smoke tongue travels rapidly along the large scale jet
 towards the IP. Residual smoke from P1 is also marginally carried with this flow.

P2 reaches the IP (Madrid) on 7 September at 12UT with column density levels of CO below 1.20 mg m⁻². For comparison on 8 September at 00UT the CO column density of P2 is on the order of 18.60 mg m⁻². This result indicates that at the peak of the event the CO level observed over the IP and emitted by P2 is roughly 20 times larger than the one emitted by P1. For comparison, Yurganov et al. (2001) reports values of total column CO in Moscow, Russia, in the vicinity of strong wildfires ~50 times larger than the our values of ~20 mg m⁻².

As far as the vertical transport is concerned, we first analyze S3, S7 and S11 to identify where and when the injection at 6 km (mid troposphere) occurs, and then S4, S8 and S12 for the injection at 11 km (upper troposphere). CO/P1 (S3) appears for the first time at 6 km 18 hours after the first emission of the fires and close to the source region. CO/P2 appears for the first time at 6 km much later, ~30 hours after the emission, but also much farther, ~2000-3000 km from the source. In the upper troposphere, CO/P1 (S4) appears for the first time at 11 km 36 hours after the first 718 emission of the fires and at ~2000 km from the source region, while CO/P2 appears for the first time at 11 km 60 719 hours after the emission and about 4000 km east of the source region. The time difference between injections from 6 720 to 11 km is 18 hours for P1 and 30 hours for P2, indicating a much faster ascending rate for P1 than for P2, despite 721 higher FRPs during P2 relative to P1 (Table 3). It seems there is a tradeoff between vertical and horizontal transport: 722 slow horizontal transport is favorable to vertical motion whereas strong horizontal transport reduces it. In the case of 723 P2, it is highly probable that the strong jets leading to its fast transport towards the IP contributed at the same time to 724 limit its vertical transport. In addition, we believe that the injection at higher altitudes of P1 is favored by its transport 725 above the region of Lake Winnipeg (a large lake visible on the Supplements to the southwest of Hudson Bay) where 726 wildfires are active, especially in the northern part of the lake. The hot region of the active fires is prone to increase 727 the convection of upper air masses travelling above it. The analysis of S7 and S11 (BC and OC at 6 km) and S8 and 728 S12 (same compounds at 11 km) indicates that the ascending rate of BC and OC is slower than for CO. BC and OC 729 reach the altitude of 6 km approximately 6 hours after CO does, and they reach the altitude of 11 km approximately 730 12 hours later than CO. No significant difference is observed between the two types of particles in terms of ascending 731 rate.

732 As far as the maximum injection height is concerned, interestingly none of the altitude levels is empty, indicating that 733 the dispersion model injects smoke at all altitude levels considered, i.e. up to 16 km. However, above 10 km the 734 number of pixels with non-zero concentration significantly decreases. We investigate the maximum injection height 735 calculated by the model by defining a threshold of significant aerosol load at a given height when the probability of 736 occurrence is greater than 2 %, i.e. when more than 2 % of the pixels at a given height are filled with non-zero values. 737 With such a criterion, we find that CO/P1 (CO/P2) generally stays below 13 (9) km, BC/P1 (BC/P2) below 12 (9) km 738 and OC/P1 (OC/P2) below 12 (11) km. These approximations of the maximum injection heights are in good 739 agreement with the profiles of both the ground-based lidar stations and CALIOP. One sees clearly that the injection 740 at high levels is much less efficient for P2 than for P1. One singular feature is the small difference (1 km) between 741 the maximum injection heights of OC/P1 and OC/P2. It may be related to the emission rate increase between P1 and P2 which is the strongest for OC compared to CO or BC. 742

To the right of the longitudinal cross-sections of Figure 10 we plot the vertical distribution of both P1 and P2 above Madrid on 8 September at 00UT and superimpose the particle backscatter coefficient measured in Madrid on 7 September at 21 UT. In order not to rely on a single profile of the simulations, P1 and P2 profiles are averaged over a square of 9 pixels centered around the coordinates of Madrid. At time of arrival over the IP, HYSPLIT results for all 747 compounds contribute to assign Plume 2 as the main source of smoke particles, representing more than 90 % of the 748 column density. At arrival over the IP CO is present up to 13 km, BC up to 11 km and OC up to 12 km (Figure 10, 749 right plots). The concentration of all three compounds is low in the first height interval between 0 and 2.5 km, except 750 for CO/P1. For CO and BC the concentration levels of P1 are higher than for P2 which supports a former suggestion 751 (see Section 4) that the airmasses arriving above the IP at 3 km on 8 September picked up smoke most likely from 752 Plume 1, while those arriving at 6 and 11 km most likely from Plume 2. The peak of CO near 5-km height is very 753 well reproduced by the model: it matches exactly the peak of the backscatter coefficient. For BC and OC the 754 concentration peak (at 4 km) is 1 km lower than the peak of the backscatter coefficient (at 5 km). For both types of 755 particles the gradual decrease of the concentration with increasing height above the peak reflects well the behavior of 756 the backscatter coefficient. Given the poor model vertical resolution and the long distance of the horizontal transport 757 (~10000 km), the particle transport is indeed very well simulated at its arrival in the IP. HYSPLIT simulates the 758 presence of a layer of BC at ~ 11 km in the upper troposphere and a layer of OC at 10 - 12 km just below the tropopause, 759 whereas the observation indicates that the smoke plume is not present above 10 km. Interestingly enough if the fact 760 that a tiny layer of BC is simulated by HYSPLIT above the troppopuse at 14 km. With respect to the literature, the 761 concentrations simulated by HYSPLIT correspond to relatively small amount of what is usually measured at ground 762 level. In a city like Barcelona, where BC is abundantly produced, the background BC concentration is usually higher 763 than 1000 ng m⁻³ (Pérez et al., 2010), i.e. much higher than the values simulated after long-range transport which peak 764 at 10 ng m⁻³.

765 Finally, we now come back to some of the hypotheses made in Section 5.2 and look for supporting arguments with 766 the results of the dispersion modelling. With respect to the material burnt possibly containing low carbon content or 767 the dominance of smoldering combustion, the dispersion modelling is of no help. The lower lidar ratio at 532 nm in 768 the upper troposphere compared to the mid troposphere reflects less absorbing particles and possibly a lesser amount 769 of BC with respect to OC. This tendency is actually confirmed by the vertical distributions of BC and OC at their 770 arrival over the IP: BC and OC peak at 4 km and then gradually decrease up to 12 km, and the relative decrease of BC 771 is stronger than for OC. About the increase of the depolarization ratio with height, the dispersion modelling is of little 772 help, the result the most useful being that in the upper troposphere only P2 is present. Overall the observations and the dispersion modelling point out to the following: near the source the smoke particles slightly depolarize ($\delta_p = 0.05$ 773 774 at ~6 km height, CALIOP, D-5) and at the arrival over the IP after a 5-day transport the particles have gained altitude and δ_p has increased ($\delta_p = 0.10$ at ~12 km, Évora, D-0). The particles arriving over the IP at ~6 km have unchanged 775





780 on 8 September at 00UT; (center) the same for BC; (bottom) the same of OC. Note the different scales. The emission and dispersion

of P1 (red-yellow color bar) and P2 (blue-green color bar) are separated. To the right of the longitudinal cross-sections we report the vertical profile of each chemical compound at the coordinates of Madrid for P1 and P2, as well as the backscatter coefficient at 532 nm retrieved in Madrid on 7 September at 21UT. The yellow star indicates the fire source region.

depolarization properties ($\delta_p = 0.05$) with respect to CALIOP, D-5. These findings enlighten the enhancement of the smoke depolarization properties with vertical transport. As smoke particles are relatively effective cloud condensation nuclei (Reid and Hobbs, 1998; Warner and Twomey, 1967), we finally hypothesize that smoke particles at non-dry altitude levels such as the upper troposphere (relative humidity ~ 20-30 %, see Section 5.2) may suffer freezing which may accentuate their asymmetric form and thus their depolarization properties.

789 7 Conclusions

790 This paper documents the time-space evolution of a smoke plume detected at its arrival over the Iberian Peninsula on 791 7 and 8 September, 2017. The smoke was emitted by strong and powerful wildfires in the Pacific northwestern region 792 of North America, a region mostly composed of temperate coniferous forests. The column properties retrieved at two 793 mid-altitude, background AERONET sites in northern and southern Spain reveal AOD₄₄₀ as high as 0.62, exceeding the background AOD by a factor larger than 6, $AE_{440-870}$ of 1.6-1.7, a large dominance of small particles (FMF > 0.88794), low $AAOD_{440}$ (<0.008) and large SSA_{440} (>0.98). The low absorption properties are attributed either (i) to the 795 796 burning of low carbon content fuels such as forest floors, particularly abundant in temperate forests, (ii) the dominance 797 of smoldering vs. flaming combustion, and/or (iii) a transformation (coating processes) of the smoke particles during 798 transport. AAE ~ 1.3, together with large SSA_{440} , in northern Spain is representative of brown carbon, while AAE ~ 799 1.0, also associated with large SSA₄₄₀, in southern Spain is representative of brown carbon probably mixed with pure 800 BC from the anthropogenic fossil fuel emissions of the nearby city of Granada, or from local fires approximately 150 801 km northwest of the site.

Satellite images of total column CO allows to identify two strong periods of emission that gave birth to two different plumes reaching the IP almost simultaneously: Plume 1 is emitted from 30 August until 1 September and Plume 2 from 2 to 5 September. The vertical distribution of the smoke plumes was monitored by ground-based lidars from both EARLINET and MPLNET networks, and from space by CALIOP. Over the IP a continuum of aerosols is observed up to the upper troposphere: aerosols are present in the whole troposphere. No particles are observed in the low stratosphere. Results are given for the mid (5 - 9 km) and upper (10 - 13 km) troposphere. The analysis of the ground-based lidars indicates a color ratio of 2.5 (3.0), LR_{532} of 55 (34) sr, and δ_p of 0.05 (0.10) in the mid (upper) troposphere, which points out to smaller, less absorbing and more depolarizing particles in the upper troposphere than in the mid troposphere. Rewinding in time with CALIOP, one observes that the older the smoke plume, the larger the color ratio, i.e. that the particle size gets smaller during transport. As far as the particle depolarization ratio is concerned, no changes related to the transport are observed in the mid troposphere. The unusual values of δ_p in the upper troposphere (0.10) are further analyzed with dispersion modelling.

814 To analyze the horizontal and vertical transport of the smoke from its origin to the IP, particle dispersion modelling is 815 performed with HYSPLIT parameterized with satellite-derived biomass burning emission estimates from 816 GFAS/CAMS. We simulated CO, BC and OC, for separately P1 and P2, with a time resolution of 6 hours, at 15 817 altitude levels and using meteorology data from GDAS with a horizontal resolution of 0.5 x 0.5°. The smoke release 818 height was not artificially fixed, but calculated internally by the model and assumed to be equal to the final buoyant 819 plume rise height as computed using Briggs (1969), implying that the final rise is a function of the input fire radiative 820 power and the meteorology. The results show that the dispersion of both plumes is quite different: P1 travels slowly 821 and disperses over a large area of the northern hemisphere, while P2 is entrained by a strong subtropical jet and travels 822 quickly towards the IP. The ascending rate of CO is nearly twice larger for P1 than for P2: CO/P1 reaches the height 823 of 11 km in 36 hours, while CO/P2 needs 60 hours. There is undeniably a tradeoff between vertical and horizontal 824 transport: slow horizontal transport is favorable to vertical motion whereas strong horizontal transport reduces it. At 825 time of arrival over the IP, both BC and OC profiles over the IP are similar in shape to the lidar-derived backscatter 826 coefficient profile: they both peak at 4 km and then gradually decrease up to 12 km, and the relative decrease of BC 827 is stronger than for OC, which corroborates one of the former hypothesis, namely that particles in the upper 828 troposphere are less absorbing than in the mid troposphere because of a smaller ratio of BC to OC. HYSPLIT results 829 for all compounds contribute to assign P2 as the main source of smoke particles over the IP, representing more than 830 90 % of the column density. These findings, all together, show that δ_p increase from 0.05 to 0.10 occurs during the 831 vertical transport from the mid to the upper troposphere, and stress the influence of the vertical transport on the smoke depolarization properties. As smoke particles are relatively effective cloud condensation nuclei, we finally 832 833 hypothesize that smoke particles at non-dry altitude levels such as the upper troposphere may suffer freezing which 834 may accentuate their asymmetric form and thus their depolarization properties.

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1191

1192 List of Figure Captions

1193

1194 Figure 1: MODIS/Aqua corrected reflectance (true color) map centered over Spain on 8 September. Green bullets

- 1195 indicate lidar stations (EV: Évora, AR: El Arenosillo/Huelva, GR: Granada, MA: Madrid, BA: Barcelona) and red
- 1196 bullets indicate AERONET sites. Map created from https://firms.modaps.eosdis.nasa.gov/map/.
- 1197

1198 Figure 2: Flowchart of the methodology.

- 1199
- 1200 Figure 3: AOD440 (black), FMF (blue) and AE440-870 (red) in (top) Montsec, northeastern Spain, and (bottom) Cerro
- 1201 Poyos, south Spain. The gray areas in the bars on top of the figures indicate coincident lidar measurements.
- 1202

1203 Figure 4: Total column carbon monoxide (day/night) from AIRS/AQUA from 30 August until 8 September. The extra

1204 plot at the bottom to the right represents the MODIS combined (Aqua and Terra) value-added AOD at 550 nm on 8

1205 September. The red star indicates the position of the active fires. On the plots of 3 and 4 September the descending,

- 1206 nighttime orbits of CALIPSO are reported. Maps created from https://worldview.earthdata.nasa.gov/.
- 1207

Figure 5: (top) 10-day back-trajectories, 1-hour resolution, arriving in Madrid, in the center of Spain, on 8 September at 00UT at heights of 3 (red), 6 (green) and 11 (blue) km; (bottom) Same back-trajectories, different viewing angle and superposition of CALIOP curtains on 4 September at 05:10UT (D-4, day-4 before arrival) and on 3 September at 09:23UT (D-5) where the smoke plumes, clearly visible, match very well in space and time with the back-trajectories. Pink crosses indicate active fires in the period 30 August – 5 September. The red rectangle of corner coordinates (125W, 40N; 93W, 58N) is the area in which the fires were taken into account in the dispersion modelling analysis (see Section 6). The orange rectangle simply highlights the region

- 1214 containing most of the fires. Maps created with Google Earth.
- 1215

Figure 6: AERONET daily mean spectral (top) AAOD, (center) SSA, and (bottom) asymmetry factor at Montsec and Cerro Poyoson 7 and 8 September.

1218

- Figure 7: Nighttime multi-wavelength lidar inversion in Évora on 7 September between 04 and 06UT. The first plot represents the quicklook of range-square corrected signal at 1064 nm in arbitrary units. β is the particle backscatter coefficient, α the particle extinction coefficient, $\alpha - AE$ the extinction-related AE, *CR* the color ratio, *LR* the lidar ratio and δ_p the particle depolarization ratio. Mean values in the mid troposphere and stratosphere (as depicted by the gray rectangles) for $\alpha - AE$, *CR*, *LR* and δ_p are reported in the plots. The horizontal dash lines at 13.6 km
- 1224 indicate the tropopause height calculated with 1° x 1° GDAS data.
- 1225
- 1226 Figure 8: CALIOP images and products on (left) 4 September at 05:10UT (D-4, Plume 1 released 5 days earlier) and
- 1227 (right) 3 September at 09:23UT (D-5, Plume 2, fresh < 1 day). (top) CALIOP quicklooks of the total attenuated
- 1228 backscatter signal at 532 nm; (center) CALIOP quicklooks of the retrieved backscatter coefficient at 532 nm restricted
- 1229 to the smoke plume (red squares); (bottom) CALIOP mean profiles of backscatter coefficient at 532 and 1064 nm, the

- 1230 color ratio and the particle depolarization ratio at 532 nm. The horizontal black dash lines indicate the tropopause
- 1231 height calculated with 1° x 1° GDAS data.
- 1232
- 1233 Figure 9: (top) Mid and upper tropospheric layer mean particle depolarization ratios at 532 nm at all Iberian lidar
- 1234 stations on the night of 7 to 8 September. Cyan and Purple bullets represent CALIOP measurements. The vertical
- 1235 bars indicate the vertical extension of the smoke layers of maximum intensity (base to top height). The horizontal
- 1236 bars indicate the standard deviation associated to δ_p in these layers. (bottom) Layer mean particle depolarization
- 1237 ratios at 532 nm vs. layer mean color ratio. The bullet color code is the same as in the top plot. We have reported
- 1238 four aerosols classes adapted from Groß et al. (2013). The vertical and horizontal bars indicate the standard deviation
- 1239 associated to δ_p and CR, respectively.
- 1240
- 1241 Figure 10: (top) Dispersion map of CO column density and longitudinal cross-section of CO concentration at the
- 1242 latitude of Madrid on 8 September at 00UT; (center) the same for BC; (bottom) the same of OC. Note the different
- 1243 scales. The emission and dispersion of P1 (red-yellow color bar) and P2 (blue-green color bar) are separated. To the
- 1244 right of the longitudinal cross-sections we report the vertical profile of each chemical compound at the coordinates of
- 1245 Madrid for P1 and P2, as well as the backscatter coefficient at 532 nm retrieved in Madrid on 7 September at 21UT.
- 1246 The yellow star indicates the fire source region.
- 1247