1	ESR and OSL dating of fossil-bearing deposits from Naracoorte Cave Complex
2	palaeontological sites, South Australia
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34 setting; (ii) highlight the merits of applying combined ESR-OSL analyses in tandem, and; (iii) provide

one of the first reliable MAAD evaluations of quartz ESR MC dating for samples with natural dose
 ranges as low as only a few tens of Gy. These results show that the Whale Bone, Specimen and

37 Alexandra cave sites are temporally related and can be used to derive multi-site reconstructions of

38 faunal assemblages and palaeoenvironmental history.

39

40 **Keywords:** Quartz ESR dating; Multi-centre approach, single-grain OSL dating; Naracoorte cave

- 41 **1.** Introduction
- 42

43 Combined application of electron spin resonance (ESR) and optically stimulated luminescence (OSL)

- 44 dating of quartz grains the two most widely used quartz dating techniques in Quaternary studies –
- 45 can provide improved insights into methodological reliability over Middle to Late Pleistocene
- 46 timescales. Comparisons of optically bleached quartz dating methods using shared sedimentary
- 47 samples offer optimal assessments of dating reliability by enabling stratigraphically constrained
- 48 evaluations of dose rate or palaeodose complications that may otherwise be difficult to unravel in
- 49 certain depositional settings (Demuro et al., 2020a). In spite of these advantages, stratigraphically
- 50 paired evaluations of ESR and OSL dating remain relatively uncommon (Demuro et al., 2020b;
- 51 Tanaka et al., 1997; Tissoux et al., 2010). In Australia, relatively little work has been undertaken on
- 52 ESR dating of sedimentary quartz (Beerten et al., 2008; Rink et al., 2007; Tanaka et al., 1995;
- 53 Yoshida, 1996) and there have been even fewer comparative studies of multiple centre (MC) ESR
- 54 and OSL dating techniques (Beerten et al., 2006; Rittner, 2013).
- 55 The Naracoorte Cave Complex (NCC) of South Australia comprises >150 cave sites (White and Webb,
- 56 2015), many of which contain extensive infill sequences with rich megafauna fossil assemblages

57 spanning the last 550 thousand years. This karst system is considered a key fossil locality for

- 58 understanding the drivers of Australia-wide Late Pleistocene megafaunal extinction given the
- 59 presence of numerous, overlapping fossil sites in direct association with well-stratified
- 60 palaeoenvironmental proxies. An extensive radiometric dating program is currently in progress to
- 61 improve chronological constraints on a large number of NCC cave sites (e.g. Arnold et al., this
- 62 volume). Establishing reliable chronologies on the NCC fossil-bearing deposits using quartz trapped
- 63 charge dating techniques is critical since most NCC sites lie close to, or well-beyond, the radiocarbon
- 64 dating range (Grün et al., 2001; Moriarty et al., 2000; Prideaux et al., 2007).
- 65 In this study, we focus on three of these Middle-Late Pleistocene NCC sites that share similar types
- of deposits (sediment cones associated with solution pipe entrances) and potentially
- 67 contemporaneous faunal assemblages currently lacking robust chronostratigraphic frameworks:
- 68 Whale Bone, Specimen and Alexandra Caves. Palaeontological excavations are currently being
- 69 undertake in Whale Bone and Specimen caves, hence establishing a firm dating framework is
- 70 important for guiding ongoing research programs.
- 71 The main aim of the study is to assess the suitability of a combined ESR and OSL dating approach for
- 72 establishing improved chronologies at NCC sites. This is achieved by applying complementary ESR
- 73 MC and single-grain OSL techniques to splits of the same samples to minimise stratigraphic
- 74 uncertainties. U/Th dating of a closely associated speleothem is also undertaken at Specimen Cave
- to provide independent age control. These comparative dating assessments are used to: (1) evaluate
- 76 characteristics of radiation-induced ESR signals associated with quartz  $[AIO_4]^0$  and  $[TiO_4/M^+]^0$  (M<sup>+</sup>= Li<sup>+</sup>
- 77 and H<sup>+</sup>) paramagnetic centres (simplified as Al and Ti centres below); (2) examine ESR signal bleaching
- 78 adequacy, dose response properties and dating reliability over relatively low natural dose ranges; (3)
- 79 provide insights into potential multi-grain OSL averaging effects for NCC deposits; (4) Establish
- 80 improved chronological constraints for sedimentary and fossil deposition in the three cave sites,
- 81 thereby expanding the existing NCC chronostratigraphic framework.
- 82

- 83 2. Study context and sample details
- 84

85 The three NCC fossil deposits at Alexandra, Specimen and Whale Bone Caves have all accumulated 86 via narrow solution pipe entrances that acted as pitfall traps for vertebrates and sources for 87 sediment cone deposition. Alexandra Cave has an exposed sedimentary section that is laterally 88 continuous for 30 m in the entrance chamber (McCluskey, 2012). Previous study of the cave deposit 89 attributed the sediment source to be alluvial and colluvial, with chronology from radiocarbon ages 90 on charcoal constraining the top 1.9 m of the sediment sequence to 17.5 - 30.8 thousand years ago 91 (ka) (Table S6; McCluskey, 2012). Vertebrate fossils recovered during excavations indicate a diverse 92 mammal fauna including megafaunal species *Thylacoleo carnifex* (an extinct marsupial carnivore), 93 and two extinct short-faced kangaroos, Procoptodon goliah and Simosthenurus occidentalis (Reed 94 and Bourne, 2000). The presence of extinct megafauna taxa provides a terminus ante quem age 95 estimate of 36.7 – 48.1 ka (Saltré et al., 2016) for the part of the sequence located beneath a depth 96 of 1.9 m (Table S6). Four sediment samples were collected from Alexandra Cave to provide sufficient 97 coverage of the stratigraphic sequence (Fig. 1) located below the first 1.9 m previously dated by 98 radiocarbon. The lowermost sample (ALEX19-10) was collected from the base of the exposed 99 sequence to constraint the initial phase of sedimentary infilling. Samples ALEX19-3 and ALEX19-6 100 were collected to bracket the lower and intermediate palaeontological levels.

101 Specimen Cave consists of a solution pipe leading to a large chamber containing a >1 m thick clastic

102 fossil deposit in the southwest corner, which is capped by an extensive (30-40 cm-thick) flowstone

103 layer. A small excavation was undertaken in the cave in 1908 and yielded megafauna fossils from

104 *Thylacoleo carnifex, Protemnodon sp.* and *Simosthenurus* sp. (Reed and Bourne, 2000). New

105 excavations are currently underway in the cave. A single ESR-OSL dating sample (SPEC18-2) was

106 collected from the uppermost layer of the current site excavation to provide terminal age constraint

107 for the fossil deposit. In addition, two U-Th samples (SP-18 and SP-17) were extracted from the top-

108 and mid-section of the closely related speleothem formation. Sample SPEC18-2 was taken 55 cm

109 beneath the base of this speleothem and ~1.5 m to the side of the U-Th sampling positions. Further

110 details of the U-Th methods and results obtained for these two samples are provided in S.M.5.

111 Whale Bone Cave contains a smaller sediment cone deposit that is currently classified as a single

112 stratigraphic unit. The ~1 m-thick fossil sequence contains abundant extinct Pleistocene vertebrate

113 remains, including Phascolarctos cinereus, Thylacoleo carnifex, Macropus fuliginosus, 'Procoptodon'

browneorum and '*Procoptodon' gilli* (Reed and Bourne, 2009). There is currently no available

115 numerical age control for the fossil deposit, though the presence of extinct megafauna indicates an

116 expected age of >36.7 ka (Saltré et al., 2016). A single ESR-OSL dating sample (WBC19-1) was

117 collected from the uppermost exposed layers of the main excavation (10 cm beneath the surface) to

118 constrain the last appearance of megafauna at the site.

119

## 120 **3. Experimental procedures**

121 Samples were prepared under safe-light conditions according to standard laboratory procedures

- 122 outlined in Aitken (1985) (see Fig. S2 for details), and the resulting purified quartz grains of each
- 123 sample were split for luminescence and ESR analyses.
- 124

- 125 ESR dose evaluation was performed on varying grain size fractions of quartz (Table 1) at the
- 126 Australian Research Centre for Human Evolution, Griffith University, using the standard multi-grain
- 127 multiple aliquot additive dose (MAAD) method. ESR measurements were performed using an Elexsys
- 128 E 500 Bruker X-band ESR spectrometer at low temperature (77 K) using a finger dewar filled with
- 129 liquid nitrogen (see details of the experimental setup and performance in Guilarte et al., submitted).
- 130 In accordance with the MC approach (Toyoda et al., 2000), the radiation-induced ESR signals
- associated with the Al, Ti (option D) and Ti-H centres were acquired in a single spectrum and
- 132 measured as per Bartz et al. (2020) (see an example in Fig. S3). The full analytical procedure is
- 133 provided in the Supplementary Information (*S.M.2. ESR dating*).
- 134
- 135 Single-grain (SG) OSL analyses were undertaken on the 212 250  $\mu m$  quartz fraction (90 250  $\mu m$
- 136 quartz fraction for WBC19-1) in the Prescott Environmental Luminescence Laboratory at the
- 137 University of Adelaide. Equivalent dose (D<sub>e</sub>) measurements were conducted on 500 1500
- 138 grains/sample for SG OSL analysis, while multi-grain (MG) OSL analysis of the Alexandra Cave
- 139 samples was undertaken using a 5 mm mask size (approximately 500 grains/aliquot). De values were
- 140 determined using the SAR protocol shown in Table S2 and the analytical procedures of Arnold et al.,
- 141 2016, 2012a), which are further detailed in the Supplementary information (*S.M.3. OSL dating*).
- 142

143 Environmental dose rate assessments were conducted in tandem for the ESR and OSL sample splits.

- As such, the only differences between the ESR and OSL dose rates for a given sample relate to grain-
- 145 size attenuation effects and alpha efficiency values. Field gamma dose rate measurements were
- 146 made using a portable gamma spectrometer NaI:Tl detector connected to an InSpector 1000 digital
- 147 handheld multichannel analyser. The 'energy windows' approach described in Arnold et al. (2012b)
- and Duval and Arnold (2012) was used to derive individual estimates of U, Th and K concentrations
- 149 from the field gamma-ray spectra. External beta dose rates were calculated from measurements
- 150 made on a Risø GM-25-5 beta counter, using homogenised sediment sub-samples collected from the
- 151 main ESR-OSL dating sample positions. Radionuclide concentrations have been converted to dose
- rates using the conversion factors given in Guérin et al. (2011), making allowance for beta-dose
- 153 attenuation (Mejdahl, 1979; Brennan, 2003). Cosmic-ray dose rates were calculated using the
- approach described in Prescott and Hutton (1994), taking into consideration the site altitude,
- 155 geomagnetic latitude as well as the density, thickness and geometry of sediment and bedrock above
- 156 the sample. The gamma, beta and cosmic-ray dose rates are corrected for long-term sediment water
- 157 content (Aitken, 1985). Further details of dose rate evaluations are given in Table 1.
- 158 **4. Results**
- 159 3.1. ESR dating
- 160 3.1.1. ESR dose determination

All ESR dose response curves obtained for the Al, Ti (D) and Ti-H signals are displayed in Fig. S4 to S9, while numerical fitting results are given in Table S1. Bleaching coefficient values obtained from the Al signal (expressed as the relative difference between the ESR intensities of the natural and the bleached aliquots), range from 39 % to 70 %, suggesting highly variable bleaching conditions for all samples, while no apparent stratigraphic trend is observed for the samples from Alexandra Cave. Measurement repeatability for the Al signal ranges from 1.1% to 11 %, resulting in D<sub>e</sub> variability of between 11% and 14 %, which is within the standards usually obtained for this centre. The exception

- 168 is WBC 19-1, which shows significantly enhanced  $D_e$  precision (5 %). The Ti centre (option D) has a
- 169 measurement repeatability of 1–6 %, and  $D_e$  repeatability of 3–13 % (Table S1), which is slightly better
- 170 than the Al centre. In particular, four out of six samples show excellent  $D_e$  repeatability (<10%). The
- 171 Ti-H centre has a measurement repeatability of 1–6 %, and  $D_e$  repeatability of 6–25 % (Table S1). In 172 comparison to the AI and Ti (D) centres, the average intensity repeatability of the Ti-H centre is
- superior (2.6%), though its  $D_e$  repeatability is inferior (14.7%) as a result of weaker signal intensity
- and higher signal-to-noise ratio (see example in Fig. S3).
- 175 The Al centre D<sub>e</sub> values were obtained by fitting an exponential + linear (EXP+LIN) fitting function (data
- 176 weighted by inverse of the squared intensities, 1/l<sup>2</sup>) through the pooled repeated ESR intensities, with
- 177 a maximum applied irradiation dose (D<sub>max</sub>) of 12000 Gy. The calculated D<sub>e</sub> values are in the range of
- 178  $27 \pm 1.9 141 \pm 13.4$  Gy (Table S1). Only SPEC 18-2 had an adjusted r<sup>2</sup> value <0.98, suggesting overall
- 179 reliable fitting results for the combined dataset.
- 180 The Ti (D) and Ti-H signal D<sub>e</sub> values were obtained using both Ti-2 and single saturating exponential 181 (SSE) functions as described in Duval and Guilarte (2015). The Ti (D) ESR intensities do not exhibit the 182 decreasing dose-dependency behaviour typically observed at high dose values (Fig. S6, S7); therefore, 183 the Ti-2 fitting function cannot be employed for these samples. Similar atypical behaviour has been 184 observed in previous studies (e.g., Duval et al., 2017) and is likely a result of the limited D<sub>max</sub> value 185 employed here (12000 Gy). The SSE fitting function is more appropriate to describe the experimental 186 data points showing apparent dose saturation. Since this function can potentially return 187 overestimated D<sub>e</sub> values when adopting inappropriate D<sub>max</sub> values (e.g., Duval et al., 2009), the 188 evolution of both parameters (De and Dmax) for Ti (D) is plotted in Fig. S10. The impact of Dmax was 189 assessed by performing the SSE fitting repeatedly after removing successive data points from the high 190 dose range. Interestingly, all samples show a similar trend, with De values remaining relatively constant 191 (or within error) up to D<sub>max</sub> = 1000-3000 Gy, followed by a significant increase at higher doses. The 192 strong correlation between  $D_e$  and  $D_{max}$  values shows that the SSE function is not appropriate to 193 describe the behaviour of the ESR intensities for  $D_{max}$  > 3000 Gy. In contrast, the resulting  $D_e$  value 194 reaches a plateau in the < 1000-3000 Gy dose range (note the large  $D_e$  errors for  $D_{max}$  < 1000 Gy as a 195 result of the limited number of data points available for the fitting), indicating the absence of 196 significant correlation between De and Dmax. Consequently, a Dmax value of 1522 Gy was used for most 197 samples, as this produced the lowest inter-sample variability, while D<sub>max</sub> values of 2978 Gy and 997 Gy 198 were used for Alex 19-8 and WBC 19-1, respectively.
- 199 Since the Ti-H signal typically reaches apparent saturation much earlier than the Ti (D) signal (usually 200 in the 3000-6000 Gy dose range; see Duval and Guilarte, 2015), slight decreasing dose-dependency 201 behaviour is observed in the DRCs (Figs. S8, S9) for this signal. Consequently, the standard Ti-2 fitting 202 function can be used for the Ti-H data obtained here (Duval and Guilarte, 2015); although the fitting 203 of the decreasing dose component can only be achieved with a very limited number of points (1 or 2), 204 which may result in significant uncertainty in the fitting results. As such, all Ti-H DRCs were fitted with 205 both SSE (with increasing D<sub>max</sub> from 403 to 11995 Gy) and Ti-2 functions (over the full dose range, with 206  $D_{max}$  =11995 Gy) for comparison (Fig. 2). Excluding the scattered  $D_e$  values derived using  $D_{max}$  = 403 207 and 606 Gy, which is the result of having very few experimental data points, produces relatively 208 constant De values for all samples in spite of the increase in Dmax from 998 to 11995 Gy (which is unlike 209 the Ti (D) signal). The D<sub>max</sub> value therefore has limited influence on D<sub>e</sub> values for the SSE fitting results, 210 suggesting that the use of this function is appropriate to describe the data set. These Ti-H D<sub>e</sub> values

are in the range of  $26.1 \pm 1.3 - 128.1 \pm 9.5$  Gy and  $25.3 \pm 1.3 - 138.1 \pm 7.7$  Gy for the SSE and Ti-2 functions, respectively (Table S1). Interestingly, both the SSE ( $D_{max} = 1522$  Gy) and Ti-2 functions yield consistent  $D_e$  values at 1 $\sigma$  for three out of six samples, while all replicate results are in agreement at  $2\sigma$ . The magnitude of the  $D_e$  estimates for these samples are all within the range of values (<300-400

- 215 Gy) for which the Ti-H signal is known to provide reliable dose estimates (Bartz et al., 2020).
- 216

## 217 3.1.2. <u>Multiple centre approach</u>

The Ti-H centre signal consistently produces the lowest D<sub>e</sub> values of the ESR MC dataset, while the Ti 218 219 (D) signal provides, somewhat unexpectedly, the highest  $D_e$  values; although similar trends have 220 been observed in other studies (e.g. Asagoe et al., 2011; Beerten et al., 2006; Beerten and Stesmans, 221 2007; Rittner, 2013; Tissoux et al., 2007, 2008). Importantly, however, the D<sub>e</sub> derived from the Ti (D) 222 and AI signals are consistent at 2 $\sigma$  for all six samples (Fig. S11), which suggests that these systematic 223 differences may not be significant. This observed D<sub>e</sub> pattern is consistent with the known respective 224 bleaching rates of each signal, with the Ti-H signal achieving full resetting much faster than any other 225 signal (Rink et al., 2007; Tissoux et al., 2007). It is also worth noting that other sources of uncertainty 226 might have partially impacted the MC fitting results in this study. In particular, the low natural dose 227 ranges under consideration here (<150 Gy) are especially challenging from an ESR methodological 228 perspective. This could result in curve fitting artefacts and an inability to accurately determine De 229 values for the Al and Ti (D) centres, given the existing uncertainty on the ESR intensities (typically, a 230 few %) and the lower radiation sensitivity of these two signals compared with that of Ti-H (Duval and 231 Guilarte, 2015).

#### 232 3.2. OSL dating

233 The SG-OSL De distributions show that five of the six samples (ALEX19-3, ALEX19-8, ALEX19-10, SPEC 234 18-2, WBC 19-1) are well bleached and have not been affected by syn-depositional mixing 235 complications (Fig. S13). The overdispersion values for these samples range between  $26 \pm 2$  and  $33 \pm 2$ 236 2% (Table 1), and are in agreement with values commonly reported for ideal (well-bleached and unmixed) sediment samples (e.g. Arnold et al., 2019), including those from other NCC sites (Arnold 237 238 et al., this volume). Sample ALEX19-6 is an exception to this trend, exhibiting a notably higher 239 overdispersion of 41 ± 2% (Table 1), a significant positively skewed D<sub>e</sub> distribution according to the 240 test of Bailey and Arnold (2006) and Arnold et al. (2011), and a maximum log likelihood (L<sub>max</sub>) test 241 score (Arnold et al., 2009) favouring use of the minimum age model (MAM; Galbraith et al., 1999) 242 over the weighted mean (central age model; CAM) for De evaluation. These characteristics suggest 243 that ALEX19-6 may have been affected by more significant syn-depositional mixing (i.e., mixing of 244 pre-existing cave sediments with externally derived quartz grains during their initial transportation 245 within the cavity; e.g. (Arnold et al., 2019) or bleaching complications compared to other samples. 246 This interpretation is supported by the scattered replicate D<sub>e</sub> distributions obtained for this sample 247 using single-grain thermally transferred OSL (TT-OSL) and post-infrared infrared stimulated 248 luminescence (pIR-IRSL) (Arnold et al., this volume). It is also consistent with the sedimentological 249 characteristics of the ALEX 19-6 host deposits, which (unlike the other sampled deposits) lack 250 bedding structures (laminations, charcoal bands) and likely indicate a higher-energy depositional 251 environment that could have contributed to syn-depositional mixing.

252 The replicate MG-OSL ages obtained for four of the Alexandra Cave samples provide useful insights 253 into multi-grain averaging effects for NCC deposits. The MG-OSL ages for ALEX 19-3 and ALEX 19-6 254 are in agreement with their SG-OSL counterparts, confirming the suitability of both approaches for 255 D<sub>e</sub> determination. However, the MG-OSL ages for ALEX 19-8 and ALEX 19-10 are significantly (~50%) 256 higher than the corresponding SG-OSL ages. The reason for this upward shift in MG-OSL ages can be 257 gleaned by examining the characteristics of the rejected grains populations (Table S4). In particular, 258 a significantly higher proportion of measured grains for samples ALEX 19-8 and ALEX 19-10 are 259 rejected for exhibiting aberrant luminescence properties associated with high sensitivity-corrected 260 natural signals; namely, non-intersecting, extrapolated and saturated grains (e.g. Fig. S12). The 261 number of rejected grains falling into this category amounts to ~50-62% of the accepted grain populations for samples ALEX19-8 and ALEX19-10, compared to only 1-3% of the accepted grain 262 263 populations for samples ALEX 19-3 and ALEX 19-6 (Table S4). While these unreliable grains can be 264 routinely rejected during SG-OSL analysis, they cannot be excluded from MG OSL measurements and 265 thus contribute to a significant upward shift in the cumulative sensitivity-corrected natural signals of 266 the final MG-OSL D<sub>e</sub> datasets for ALEX 19-3 and ALEX 19-6. The same type of biasing MG averaging 267 effects have been reported elsewhere for samples that contain significant populations of non-

268 intersecting, extrapolated and saturated grain types (Arnold et al., 2012a, 2013).

### 269 Discussion

#### 270 4.1. ESR-OSL comparisons

There is good overall agreement between the ESR (Ti-H) and SG OSL ages at the three study sites (Table 1; Fig. 3). For SPEC 18-2, the ESR (Ti-H) and OSL results are additionally compatible with Useries ages from a calcite flowstone directly above the sample (Table 1), whereas both the AI and Ti (D) ESR ages overestimate the independent age control. This confirms that these two ESR signals have most likely not been fully reset during sediment transport.

276

277 The SG OSL analysis similarly show that the sampled sediments are generally well-bleached and 278 unlikely to have been affected by syn-depositional mixing, with the exception of ALEX 19-6 (Table 1; 279 Fig S13). Given the depositional context and related sedimentological evidence, the D<sub>e</sub> scatter 280 observed for ALEX 19-6 is likely the result of mixing with previously deposited sediments during 281 transportation through the cave or partial bleaching of sediments prior to deposition. Interestingly, 282 the corresponding ESR MC ages for sample ALEX19-6 do not appear to exhibit any overestimation 283 compared to the SG-OSL age, which might otherwise be expected if they have experienced the same 284 depositional history (especially given MG averaging effects). It is possible that the coarser grain

- 285 fraction used for the SG-OSL analysis (212 250 μm) is more susceptible to progressive, short-
- 286 distance mobilisation through the cave system during successive high-energy transportation events 287 compared to the finer grain fraction used for the ESR analysis ( $125 - 212 \mu m$ ). The latter would
- certainly have lower entrainment thresholds and thus may be transported more directly through the
- 289 cave system during single flow events. Such sediment routing grain-size dependencies remain largely
- 290 unexplored in karst settings, but should be investigated further using replicate OSL or ESR
- 291 measurements performed on various grain size fractions.

The results of the SG-OSL versus MG-OSL comparisons reveal that some NCC deposits may suffer
 from adverse multi-grain averaging effects that could affect the reliability of conventional OSL dating
 in this setting. Complex MG-OSL averaging effects related to the inclusion of aberrant grain types

- 295 have been widely reported (Arnold et al., 2012a, 2013, 2016, 2019; Demuro et al., 2013, 2008;
- 296 Jacobs et al., 2006; Russell and Armitage, 2012; Stone and Bailey, 2012). Our results indicate the
- 297 potential for systematic biases associated with large populations of non-intersecting, extrapolated
- 298 and saturated grain types, and suggest it may be prudent to focus on SG-OSL rather than MG-OSL
- 299 analysis for NCC deposits. It is plausible that the same upward shift in MG-OSL ages observed for
- 300 samples ALEX19-8 and ALEX19-10 may explain the 10 – 25 ka overestimation of MG-OSL ages
- 301 reported previously at the nearby NCC site of Blanche Cave (Darrénougué et al., 2009).
- 302 Unfortunately, it remains unclear whether multi-grain ESR signals are similarly affected by adverse 303 averaging effects related to unsuitable grain types. This is conceivable (e.g., Beerten and Stesmans,
- 304
- 2006a, 2006b), and could contribute to some of the systematic differences observed with ESR MC 305 ages (assuming such averaging effects affect the Ti-H, Ti (D) and AI ESR signals to differing extents).
- 306 However, this would need to be specifically investigated using ESR single-grain analysis before
- 307 drawing any further conclusions.
- 308 A potential explanation for the systematically older Ti (D) ages observed in this study (Fig. 3), is that 309
- this particular ESR signal is not fully bleached, and retains a residual (unbleached) dose prior to 310 burial (Tissoux et al., 2007; Toyoda et al., 2000). Beerten et al. (2006) showed Ti-Li residual signals of
- 311 up to 70% of the natural (option E from Duval & Guilarte 2015). Richter and Tsukamoto (2021)
- 312 showed residual doses in all centres (86 – 2039 Gy), including for the Ti-H signal (86 – 263 Gy).
- 313 Measurements of residual dose in modern analogue samples have also showed residual doses of 61
- 314 - 466 Gy for the Ti-Li signal (option D from Duval & Guilarte 2015) (Tsukamoto et al., 2017).
- 315 However, if the Ti (D) ESR signal had not been fully reset prior to sediment deposition, it is expected
- 316 that incomplete bleaching would also apply to the Al signal, which has much slower bleaching
- 317 kinetics, and would have a larger De value than the Ti (D) signal, contrary to the results observed in
- 318 this study. While incomplete bleaching of the Ti (D) signal is unlikely, a study on modern analogue
- 319 sediment samples from NCC would enable improved evaluations of the ESR Ti (D) signal.
- 320 Another possible explanation for the systematically older Ti (D) ages observed in this study could 321 relate to the relatively large intensity of Ti-H and Ti-Li signals, which may impact the accuracy of the 322 dose estimates derived from each centre. Beerten and Stesmans (2006b) observed that for samples 323 with both Ti-H and Ti-Li signals, the latter provided overestimated dose estimates, while no 324 significant bias was observed in samples that did not have a Ti-H signal. The Naracoorte samples are 325 characterised by exceptionally high Ti-H intensities, despite their relatively young ages, which may 326 impact Ti-Li centre signal evaluation. In particular, the C/D intensity ratio (Duval and Guilarte, 2015) 327 calculated for the NCC natural samples are exceptionally high (0.9 to 1.22) compared to those 328 observed in quartz from other localities (Duval and Guilarte, 2015; Demuro et al., 2020b). This 329 variability in C/D intensity ratio likely relates to the nature and origin of the quartz being dated (see 330 comparative data in Bartz et al., 2020; Demuro et al., 2020b), and suggests that the NCC samples 331 considered in this study might be more susceptible to the complications suggested by Beerten and 332 Stesman (2006a).
- 333 The measurement of Ti ESR intensities can vary between studies (see Duval et al., 2020), making
- 334 comparisons across sites difficult. Depending on how the ESR intensities are measured (see options
- 335 A – E Duval and Guilarte, 2015) and the source of quartz (Demuro et al., 2020b), the relative
- 336 contribution of Ti-H and Ti-Li signals may vary. Standardising the evaluation of ESR intensities within

- the dating community could help to improve our understanding of the nature and composition ofESR signals related to Ti centres.
- 339 Taking these various factors into consideration, it seems likely that the Ti (D) D<sub>e</sub> overestimations
- 340 observed for the NCC samples relate to their high Ti-H signal intensities and related adverse impacts
- 341 on Ti (D) ESR evaluation, rather than incomplete signal resetting during transport. Irrespective of the
- 342 cause of the Ti (D) offset, the Al and Ti (D) signals ultimately produce age estimates that are within
- $2\sigma$  of each other. Hence, any apparent systematic differences observed with these samples are not
- 344 statistically significant.
- The Ti-H centre ages presented in this study are, to our knowledge, the youngest obtained so far using ESR dating with the MAAD method, providing reliable D<sub>e</sub> estimates over natural dose ranges as low as < 30 Gy. These results confirm that the quartz preserved at NCC sites is particularly well suited for ESR dating, and that the Ti-H centre has great potential for dating Late Pleistocene NCC fossil deposits with corresponding burial doses < 100 Gy.
- 350

## 351 4.2. Implications for NCC site histories

- 352 The OSL and ESR ages provided in this study have extended the chronology of the Alexandra Cave 353 entrance chamber sediment sequence by 130 ka; building on the initial radiocarbon ages of 17.5 – 354 30.8 ka (Table S6; McCluskey, 2012), we have established a basal age of 157.8 ± 9.8 ka (ALEX 19-10) 355 for the lowermost sample in the stratigraphic section. This indicates that 6 m of sediment infill 356 occurred at the site between marine isotope stage (MIS) 6 and MIS 2 (Lisiecki and Raymo, 2005), and 357 confirms that the Alexandra Cave sequence spans multiple glacial-interglacial cycles. These 358 chronological results are significant because they reveal that the deposit overlaps with the 359 megafauna extinction period (36.7 – 48.1 ka) and >100 ka leading up to the event, flagging this 360 sedimentary section as one of the few sites in Australia to enable detailed assessment of megafauna
- 361 extinction dynamics over long timescales.
- The Specimen Cave deposit was deposited during late MIS 6 or early MIS 5 (134.7 ± 8.7 ka), and can be chronologically correlated to the base of the Alexandra Cave entrance chamber deposit (Alex 19-10). Examining palaeoenvironment proxies from the infill sequence at Specimen Cave and Alexandra Cave would be useful to reconstruct regional climatic conditions during the MIS 6 to MIS 5 transition, as this time period is currently not represented by any other well-dated sedimentary deposits at NCC sites.
- 368 The Whale Bone Cave deposit is most likely constrained to the final stage of the last interglacial 369 complex or the start of MIS 4, with mean ages centred on MIS 5a to 5b (71 - 87 ka). These ages 370 overlap with the Alexandra Cave entrance chamber deposit presented here, and also Grant Hall 371 chamber in Victoria Fossil Cave (Macken et al., 2011). Previous palaeoenvironmental and faunal 372 studies at Grant Hall chamber suggested well-forested, dense woodlands (Fraser and Wells, 2006; 373 Macken et al., 2011), with increased effective moisture during this period supported by speleothem 374 deposition (Ayliffe et al., 1998). The extent of moisture availability during the last interglacial 375 complex can be further assessed from the two newly identified MIS 5 cave sites presented in this 376 study, Whale Bone and Alexandra caves, which share similar aged deposits.

- 377 The ages produced in this study contribute to the collective understanding of spatial and temporal
- relationships of NCC fossil sites (eg. Darrénougué et al., 2009; Forbes et al., 2007; Grün et al., 2001;
- 379 Macken et al., 2011; St Pierre et al., 2012), opening up new possibilities for undertaking multi-site
- 380 comparisons of faunal assemblages and palaeoenvironmental reconstructions spanning MIS 6 to 2,
- and improved scope for understanding the drivers of Australia-wide late Pleistocene megafauna
- 382 extinction.

## 383 4. Conclusion

- 384 Comparative quartz ESR MC and OSL ages obtained on six NCC samples are in good agreement, with
- 385 Ti-H and SG-OSL ages providing the closest age correspondence for the three study sites. This
- 386 represents the first study to demonstrate the potential for obtaining reliable quartz Ti-H centre ages
- 387 with the MAAD method over natural dose ranges as low as < 30 Gy. These promising results most
- 388 likely reflect a combination of the experimental setup employed for cryogenic ESR measurements
- 389 (finger dewar), which enabled data acquisition at much lower temperatures (77 K) and thus ensured
- higher signal intensity/resolution than with standard variable temperature units (see comparison
- 391 study in Guilarte et al., submitted), and the favourable intrinsic ESR characteristics of the NCC quartz,
- 392 which were found to be particularly well suited for ESR analyses.
- 393 SG and MG OSL comparisons show the advantages that SG OSL dating can offer for characterising
- and circumventing potentially biasing averaging effects for some NCC deposits. Further investigation
- 395 into the potential for Ti (D) bleaching residuals and ESR multi-grain averaging effects would be
- 396 worthwhile to expand the applicability of different MC signals at NCC sites.
- 397 The new SG-OSL and Ti-H chronologies presented here for Alexandra, Specimen and Whale Bone
- 398 Cave constrain three previously undated megafauna fossil deposits to 50 150 ka, collectively
- 399 spanning an important temporal gap in the existing NCC chronology. These results demonstrate the
- advantages of applying multiple palaeodosimetric dating techniques in tandem to examine
- 401 methodological reliability and reconstruct more comprehensive sedimentary histories for late
- 402 Pleistocene NCC deposits.

## 403

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598 Figures







Figure 1. Stratigraphy of entire sedimentary infill deposit from entrance chamber of Alexandra Cave,
 Naracoorte Caves, South Australia, including photos, sample position (A1, A2, A3) and locality of
 fossilised fauna (indicated by bone symbols). B: Specimen Cave photo and sample position C: Whale
 Bone Cave sample position and photos. Yellow circles indicate combined OSL–ESR sample positions,
 and are approximately 4 cm in diameter. Geographical location of cave sites are shown in Figure S1.





Figure 2. D<sub>e</sub> values obtained for the Ti-H signal and derived from SSE (red circles) and Ti-2 (green
 squares) fitting functions, with maximum irradiation dose (D<sub>max</sub>) varying between 400 and 12000 Gy.
 D<sub>e</sub> value is considered reliable in the D<sub>max</sub> range for which D<sub>e</sub> values remain constant. All samples
 show constant D<sub>e</sub> values between D<sub>max</sub> of 600-6000 Gy. D<sub>e</sub> values used for age calculations are

610 derived from the SSE  $(1/l^2)$  with  $D_{max} = 1500$  Gy (highlighted in yellow).

611



Figure 3. Comparison of final MG OSL, SG OSL and MC ESR ages for each sample, represented as mean
± total uncertainty at 95% confidence interval. The U-series ages (SP 17 = 145 ± 3.5 ka, and SP 18 =
142.8 ± 2.1 ka) for SPEC 18-2 has been derived from the capping flowstone immediate overlying the

617 OSL-ESR sample, and therefore represents a close minimum age constraint (Table S5).

#### 619 Table

620 Table 1. Summary statistics of radionuclide concentrations, environmental dose rates, De measurements and

621 final OSL and ESR samples. OSL and ESR uncertainties are quoted at 1σ. Field water content is expressed as

percentage of dry mass mineral fraction with ± 20% uncertainty. Gamma dose rates were calculated from in
 situ measurements made at each sampling position using a NaI:Tl detector. Beta dose rates were calculated

situ measurements made at each sampling position using a NaI:Tl detector. Beta dose rates were calculated
 using a Risø GM-25-5 low-level beta counter. Individual radionuclide concentrations were derived from the

625 field gamma spectra using the 'energy windows' method (Arnold et al., 2012a). Cosmic dose rates were

626 calculated by taking into account the geomagnetic latitude, altitude and thickness of overburden (Prescott &

627 Hutton, 1994). The gamma, beta and cosmic-ray dose rates are corrected for long-term sediment water

628 content (Aitken, 1985). The total dose rate for the luminescence samples includes an empirically determined

629 internal dose rate of 0.04 Gy/ka (Arnold et al., this volume) with an uncertainty of 30% (Bowler et al., 2003).

- 630 SG OSL and MG OSL D<sub>e</sub> datasets and overdispersion values are calculated using the central age model (CAM)
- 631 (Galbraith et al., 1999). The total dose rate for ESR samples include an empirically determined internal dose of

0.06 + -0.01 Gy/ka (Arnold et al., this volume) using an alpha efficiency of  $0.07 \pm 0.01 (1\sigma)$  for both the Al and

632 0.00 +/- 0.01 Gy/ka (Arnold et al., this volume) using an alpha enciency of 0.07 ± 0.01 (10) for both the Aran
 633 Ti centres (Bartz et al., 2019). U-series ages are presented here with 2σ uncertainty, see Table S5 for details.

Sample	Alex 19-3	Alex 19-6	Alex 19-8	Alex 19-10	SPEC 18-2	WBC 19-1
Unit	С	Е	F	G	-	-
Depth (cm)	262	411	473	585	85	20
Water content (%)	3.6 ± 0.7	$3.3 \pm 0.7$	11.1 ± 2.2	6.0 ± 1.2	16.5 ± 3.3	16.2 ± 3.2
Internal dose rate (ESR) (Gy/ka)	0.06 ± 0.02	0.06 ± 0.02	0.06 ± 0.02	0.06 ± 0.02	0.06 ± 0.02	0.06 ± 0.02
Internal dose rate (OSL) (Gy/ka)	0.04 ± 0.01	0.04 ± 0.01	0.04 ± 0.01	0.04 ± 0.01	0.04 ± 0.01	0.04 ± 0.01
Beta dose rate (ESR) (Gy/ka)	0.24 ± 0.01	0.15 ± 0.01	0.53 ± 0.03	0.27 ± 0.02	0.29 ± 0.02	0.41 ± 0.02
Grain size (µm)	125 – 212	125 – 212	125 – 212	125 – 212	212 – 250	90 – 250
Beta dose rate (OSL) (Gy/ka)	0.23 ± 0.01	0.15 ± 0.01	0.51 ± 0.03	0.26 ± 0.03	0.29 ± 0.02	0.41 ± 0.02
Grain size (µm)	212 – 250	212 – 250	212 – 250	212 – 250	212 – 250	90 – 250
Gamma dose rate (Gy/ka)	0.20 ± 0.01	0.16 ± 0.01	0.37 ± 0.02	0.24 ± 0.01	0.40 ± 0.02	0.38 ± 0.02
Cosmic dose rate (Gy/ka)	0.05 ± 0.01	0.05 ± 0.01	0.04 ± 0.00	0.04 ± 0.00	0.06 ± 0.01	0.06 ± 0.02
Total dose rate (ESR) (Gy/ka)	$0.54 \pm 0.03$	$0.42 \pm 0.02$	$1.00 \pm 0.05$	$0.61 \pm 0.03$	0.80 ± 0.05	0.90 ± 0.05
Total dose rate (OSL) (Gy/ka)	$0.51 \pm 0.03$	0.39 ± 0.02	0.96 ± 0.05	0.58 ± 0.03	0.78 ± 0.04	0.88 ± 0.05
D <sub>e</sub> (Gy) Al	27.0 ± 1.9	30.1 ± 2.0	140.7 ± 7.4	129.4 ± 7.9	141.0 ± 13.4	85.8 ± 4.0
De (Gy) Ti (D)	31.3 ± 1.4	39.7 ± 1.9	168.2 ± 9.3	136.6 ± 13.1	159.4 ± 13.1	77.0 ± 5.3
De (Gy) Ti-H	26.1 ± 1.3	28.3 ± 1.4	128.1 ± 9.5	115.0 ± 9.7	114.5 ± 12.0	50.8 ± 3.9
De (Gy) SG OSL	24.8 ± 0.6	31.3 ± 1.0	99.8 ± 2.9	90.7 ± 2.5	105.0 ± 2.6	66.7 ± 1.4
Overdispersion (%)	26.4 ± 2.1	40.6 ± 2.4	26.8 ± 2.5	26.4 ± 2.3	30.0 ± 1.9	25.9 ± 1.6
De (Gy) MG OSL	26.7 ± 0.6	34.7 ± 0.8	150.9 ± 7.0	136.0 ± 6.3	-	-
Overdispersion (%)	0	0	6.6 ± 5.6	0	-	-
Age (ka) Al	49.7 ± 4.3	72.3 ± 6.3	140.2 ± 10.3	213.4 ± 17.1	175.8 ± 19.4	95.2 ± 6.9
Age (ka) Ti (D)	57.5 ± 4.0	95.3 ± 7.2	167.5 ± 12.6	225.3 ± 24.7	198.8 ±19.8	85.5 ± 7.6

Age (ka) Ti-H	47.9 ± 3.5	67.9 ± 5.2	127.6 ± 11.5	189.6 ± 18.9	142.8 ± 17.0	56.4 ± 5.3
Age (ka) MG OSL	51.9 ± 3.0	88.9 ± 5.7	156.9 ± 11.3	236.5 ± 17.1	-	-
Age (ka) SG OSL	48.2 ± 2.9	80.4 ± 5.4	$103.8 \pm 6.4$	157.8 ± 9.8	134.7 ± 8.7	75.9 ± 4.7
Age (ka) U-series					145.0 ± 3.5	
					142.8 ± 2.1	