



University of Granada
Department of Mineralogy and Petrology

Pre-Variscan magmatism of the Central Iberian Zone:
Chemical and isotope composition, geochronology and
geodynamic significance



Cristina Talavera Rodríguez

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Que. Dña. **Cristina Talavera Rodríguez** ha realizado bajo mi dirección la Tesis Doctoral que lleva por título: ***Pre-Variscan magmatism of the Central Iberian Zone: Chemical and isotope composition, geochronology and geodynamic significance.***

Y para que así conste, lo firmo en Granada a 13 de Enero de 2009

Fdo. Pilar G. Montero

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INDEX

RESUMEN	<i>p. 3</i>
1. INTRODUCTION	<i>p. 9</i>
1.1. Problem definition and objectives	<i>p. 9</i>
1.1.1. The problem	<i>p. 9</i>
1.1.2. Objectives	<i>p. 11</i>
1.2. Geological setting	<i>p. 13</i>
1.2.1. The Iberian Massif	<i>p. 13</i>
1.2.2. The Central Iberian Zone	<i>p. 14</i>
1.2.2.1. <i>The Ollo de Sapo Domain</i>	<i>p. 15</i>
1.2.2.2. <i>The Schist-Graywacke Complex Domain</i>	<i>p. 19</i>
1.2.2.3. <i>Deformation in the Central Iberian Zone</i>	<i>p. 25</i>
1.3. Previous work	<i>p. 28</i>
1.3.1. The Ollo de Sapo Formation and associated metagranites	<i>p. 28</i>
1.3.1.1. <i>General geology and lithology</i>	<i>p. 28</i>
1.3.1.2. <i>Stratigraphic position</i>	<i>p. 30</i>
1.3.1.3. <i>Structural geology</i>	<i>p. 30</i>
1.3.1.4. <i>Radiometric dating and isotope geology</i>	<i>p. 31</i>
1.3.2. The pre-Middle Ordovician metagranites of the Schist-Graywacke Complex Domain	<i>p. 34</i>
1.3.2.1. <i>General geology and composition</i>	<i>p. 34</i>
1.3.2.2. <i>Stratigraphic position</i>	<i>p. 36</i>
1.3.2.3. <i>Structural geology</i>	<i>p. 36</i>
1.3.2.4. <i>Radiometric dating and isotope geology</i>	<i>p. 37</i>
2. SAMPLES AND METHODS	<i>p. 41</i>
2.1. Samples	<i>p. 41</i>
2.2. Analytical methods	<i>p. 46</i>
2.2.1. Whole-rock major and trace element analyses	<i>p. 46</i>
2.2.2. Sr and Nd isotope analyses	<i>p. 48</i>
2.2.3. Zircon separation and dating	<i>p. 48</i>
2.2.3.1. <i>Zircon U-Pb microprobe analyses</i>	<i>p. 49</i>
2.2.3.2. <i>Zircon U-Th-Pb, Zr/Hf and Ti ablation ICP-MS analyses</i>	<i>p. 49</i>
2.2.3.3. <i>Zircon Pb-Pb stepwise evaporation analyses</i>	<i>p. 50</i>
3. PETROGRAPHY AND MINERALOGY	<i>p. 53</i>

3.1. Ollo de Sapo Domain	p. 53
3.1.1. The Ollo de Sapo Formation gneisses	p. 53
3.1.1.1. The Ollo de Sapo coarse-grained gneisses	p. 53
3.1.1.2. The Ollo de Sapo fine-grained gneisses	p. 56
3.1.2. The Ollo de Sapo Formation metagranites	p. 57
3.2. The Schist-Graywacke Complex Domain	p. 61
3.2.1. The Tormes Dome	p. 61
3.2.2. The Sierra de Gredos	p. 59
3.2.3. The Sierra de Guadarrama	p. 65
3.2.3.1. The metagranites	p. 65
3.2.3.2. The metavolcanic rocks	p. 71
3.2.4. The Anatectic Complex of Toledo	p. 73
4. GEOCHRONOLOGY	p. 77
4.1. The Ollo de Sapo Domain	p. 77
4.1.1. The Ollo de Sapo Formation	p. 77
4.1.1.1. Vivero	p. 78
4.1.1.2. Trives	p. 80
4.1.1.3. Sanabria	p. 81
4.1.1.3.1. Lower metavolcanic rocks	p. 81
4.1.1.3.2. Upper metavolcanic rocks	p. 83
4.1.1.4. Villadepera	p. 84
4.1.1.5. Hiendelaencina	p. 86
4.1.1.5.1. Lower metavolcanic rocks	p. 86
4.1.1.5.2. Upper metavolcanic rocks	p. 87
4.1.2. The metagranites of the Ollo de Sapo Formation	p. 90
4.1.2.1. Viana do Bolo	p. 90
4.1.2.2. San Sebastián	p. 92
4.1.2.3. Miranda do Douro	p. 93
4.1.2.4. Antoñita	p. 94
4.2. The Schist-Graywacke Complex Domain	p. 96
4.2.1. The Tormes Dome	p. 97
4.2.1.1. Fermoselle	p. 97
4.2.1.2. Vitigudino	p. 98
4.2.1.3. Ledesma	p. 99

4.2.2. The Gredos sector	p. 100
4.2.2.1. <i>San Pelayo</i>	p. 100
4.2.2.2. <i>Castellanos</i>	p. 102
4.2.2.3. <i>Bercimuelle</i>	p. 103
4.2.3. The Guadarrama sector	p. 104
4.2.3.1. <i>La Cañada</i>	p. 104
4.2.3.2. <i>La Hoya</i>	p. 106
4.2.3.3. <i>La Estación I</i>	p. 107
4.2.3.4. <i>La Estación II</i>	p. 108
4.2.3.5. <i>Vegas de Matute (Melanocratic metagranites)</i>	p. 109
4.2.3.6. <i>Vegas de Matute (Leucocratic metagranites: El Caloco)</i>	p. 110
4.2.3.7. <i>La Morcuera</i>	p. 111
4.2.3.8. <i>Buitrago de Lozoya</i>	p. 112
4.2.3.9. <i>La Berzosa</i>	p. 113
4.2.3.10. <i>El Cardoso</i>	p. 114
4.2.3.11. <i>Riaza</i>	p. 116
4.2.4. The Anatectic Complex of Toledo	p. 117
4.2.4.1. <i>Polán</i>	p. 117
4.2.4.2. <i>Mohares</i>	p. 118
4.3. Summary	p. 119
4.3.1. The Olo de Sapo Domain	p. 119
4.3.2. The Schist-Graywacke Complex Domain	p. 121
4.3.3. The significance of the younger ages	p. 122
5. CHEMICAL COMPOSITION AND Sr-Nd ISOTOPE GEOLOGY .	p. 125
6. DISCUSSION	p. 133
6.1. The nature of the magmatic sources	p. 133
6.2. The reasons for the abnormally elevated zircon inheritance	p. 136
6.2.1. Zircon saturation temperatures	p. 137
6.2.2. Ti-in-zircon temperatures	p. 140
6.2.3. Zircon dissolution kinetics in the Cambro-Ordovician magmas of Iberia	p. 141
6.2.3.1. <i>Watson's equation</i>	p. 142
6.2.3.2. <i>Choice of parameters</i>	p. 142

6.2.3.3. <i>Estimation of heating rates</i>	p. 143
6.2.3.4. <i>Estimation of cooling rates</i>	p. 144
6.2.3.5. <i>Discussion of results</i>	p. 147
6.3. Origin and geodynamic significance of the Cambro-Ordovician magmas of the Central Iberian Zone	p. 149
6.4. The age of the Cambro-Ordovician rifting in Iberia	p. 151
7. SUMMARY AND CONCLUSIONS	p. 157
7.1. Summary and conclusions	p. 157
7.2. Conclusiones	p. 163
REFERENCES	p. 171
APPENDIX	p. 195

1. RESUMEN

RESUMEN

Los estudios geocronológicos de los últimos años indican que el magmatismo Cambro-Ordovícico en Iberia y en el resto de Europa central es mucho más abundante de lo que previamente se creía. En Iberia, y en concreto en la Zona Centroibérica, el basamento pre-Varisco contiene tres alineamientos de rocas ígneas de probable edad Cambro-Ordovícica: las rocas metavolcánicas y metagraníticas de la Formación Olló de Sapo, los metagranitos y las rocas metavolcánicas de la zona norte del Dominio del Complejo Esquisto-Grauváquico, y más al sur, en el límite con la Zona de Ossa Morena, los granitoides de Carrascal Portalegre y las rocas metavolcánicas de la Formación Urrea, mucho menos voluminosos y situados ya en Portugal. La presente Tesis Doctoral trata específicamente de este magmatismo Pre-Varisco y en ella se abordan dos cuestiones fundamentales: por un lado, la datación y el estudio geoquímico e isotópico de los cuerpos ígneos de las bandas situadas más al norte, y por otro lado, la integración de todos los datos en un modelo de generación de magmas durante el Cambro-Ordovícico compatible y extensible al resto de rocas de la misma edad en Europa.

La dificultad que históricamente han tenido estas rocas desde el punto de vista geocronológico radica en que la orogenia Varisca sobreimpuesta ha modificado los sistemas isotópicos en roca total, de forma que la aplicación de métodos de datación como Rb-Sr y Sm-Nd no han dado buenos resultados. Por otra parte, los métodos de datación basados en concentrados de minerales (circones), tampoco han sido eficaces debido a la elevadísima fracción de circones que contienen núcleos pre-magmáticos, algo que solamente se ha puesto de manifiesto en estudios recientes realizados por nuestro grupo de trabajo.

En esta memoria hemos estudiado un total de 26 cuerpos ígneos (volcánicos y plutónicos) de la Formación Olló de Sapo, los sectores de Guadarrama y Gredos, el domo del Tormes y el Complejo Anatéctico de Toledo. Sobre estos cuerpos hemos llevado a cabo un programa sistemático de datación de circones, complementado con un estudio geoquímico e isotópico (Sr y Nd) de roca total. En primer lugar, hemos estudiado más de 2.000 circones mediante catodoluminiscencia, lo que ha puesto de manifiesto la elevadísima tasa de herencia de estas rocas. En segundo lugar, hemos llevado a cabo la datación sistemática de los circones mediante U-Pb por microsonda iónica (SIMS) y por ICP-MS Ablación Láser, y mediante Pb-Pb por TIMS. Localmente se ha utilizado también el método Rb-Sr en roca

total y una isocrona interna realizada en un megacrystal de feldespatos potásicos. Los resultados geocronológicos se detallan en el capítulo 4 y se resumen en la Tabla 4.1. Además de la edad de la cristalización, mayoritariamente en torno a 480-490 Ma, hemos podido establecer las distintas edades de poblaciones heredadas, que van desde Ediacaran (la más importante en torno a 605-625 Ma), a Toniense, Orosiriense y Mesoarcaica.

Geoquímica e isotópicamente las rocas Cambro-Ordovícicas (tanto las volcánicas como las plutónicas) de los dos dominios de la Zona Centroibérica son muy similares. Geoquímicamente son rocas marcadamente peraluminosas, magnesianas y alcalino-cálcicas (en el sentido de Frost et al., 2001), caracterizadas por el enriquecimiento en los elementos más incompatibles, con una anomalía positiva de Li y negativa de Sr, características típicas de granitos derivados de rocas corticales antiguas. Isotópicamente, estas rocas tienen $(^{87}\text{Sr}/^{86}\text{Sr})_{490\text{Ma}} \approx 0.707 - 0.713$, y $\epsilon\text{Nd}_{490\text{Ma}} \approx -5.8$ a -3.3 , lo que apunta a que proceden de materiales continentales antiguos con poca o ninguna aportación mantélica excepto, quizá, los metagranitos de Miranda do Douro y San Sebastian, en el dominio del Olló de Sapo, ligeramente más primitivos ($(^{87}\text{Sr}/^{86}\text{Sr})_{490\text{Ma}} \approx 0.704$, $\epsilon\text{Nd}_{490\text{Ma}}$ entre -3.40 y 2.1).

A partir de estos datos, y a partir de las edades de las poblaciones heredadas y de la información extraída de los estudios morfológicos y de catodoluminiscencia de los circones se ha establecido la naturaleza de la posible fuente de este magmatismo Cambro-Ordovícico, que ha debido consistir en rocas ígneas félsicas Panafricanas y situadas en el margen norte de Gondwana, posiblemente asociadas espacialmente a las que forman el actual Anti-Atlas marroquí.

La elevadísima proporción de herencia encontrada en los circones de estas rocas es un hecho tan anómalo que ha requerido un estudio específico. Para ello, hemos calculado la temperatura de saturación de circón (≈ 815 °C), la temperatura de cristalización de los circones con el termómetro Ti-en-circón (≈ 826 °C) y modelado la cinética de disolución de circón en los magmas Cambro-Ordovícicos usando las ecuaciones de Watson (1996). A partir de estos datos hemos podido extraer información muy valiosa acerca de los mecanismos de fusión que han dado origen a los magmas Cambro-Ordovícicos y su posterior emplazamiento. Se concluye que la fusión del protolito cortical tuvo lugar en la base de la corteza, y fue desencadenada por la afluencia de pulsos de magmas máficos muy calientes en un ambiente geodinámico extensional.

Finalmente, con todos los datos geocronológicos del magmatismo Cambro-Ordovícico de las Zonas Centroibérica, Galicia-Tras-Os-Montes y Ossa-Morena, ha sido posible establecer la edad del rifting Cambro-Ordovícico en Iberia: en las las Zonas Centroibérica y Galicia-Tras-Os-Montes empezó aproximadamente a 500 Ma, alcanzó un máximo a 495-483 y finalizó a 470 Ma. En la Zona de Ossa-Morena la situación es menos clara, la duración del rifting parece haberse extendido desde 515 hasta 460 Ma, sin que exista un máximo bien definido.

1. INTRODUCTION

1. INTRODUCTION

1.1. Problem definition and objectives

1.1.1. The problem

Growing geochronological evidence indicates that the Cambro-Ordovician magmatism of South and Central Europe was more voluminous than earlier believed (e.g., Delaperriere and Respaut, 1995; Schaltegger and Gebauer, 1999; Valverde-Vaquero and Dunning, 2000; Barbey et al., 2001; Deloule et al., 2002; Schaltegger et al., 2003; Friedl et al., 2004; Mazur et al., 2004; Roger et al., 2004; Helbing and Tiepolo, 2005). Igneous rocks of this age are well represented in the Central Iberian Zone where pre-Middle Ordovician felsic volcanic rocks and granitoids, strongly deformed and variably metamorphosed during Varis-

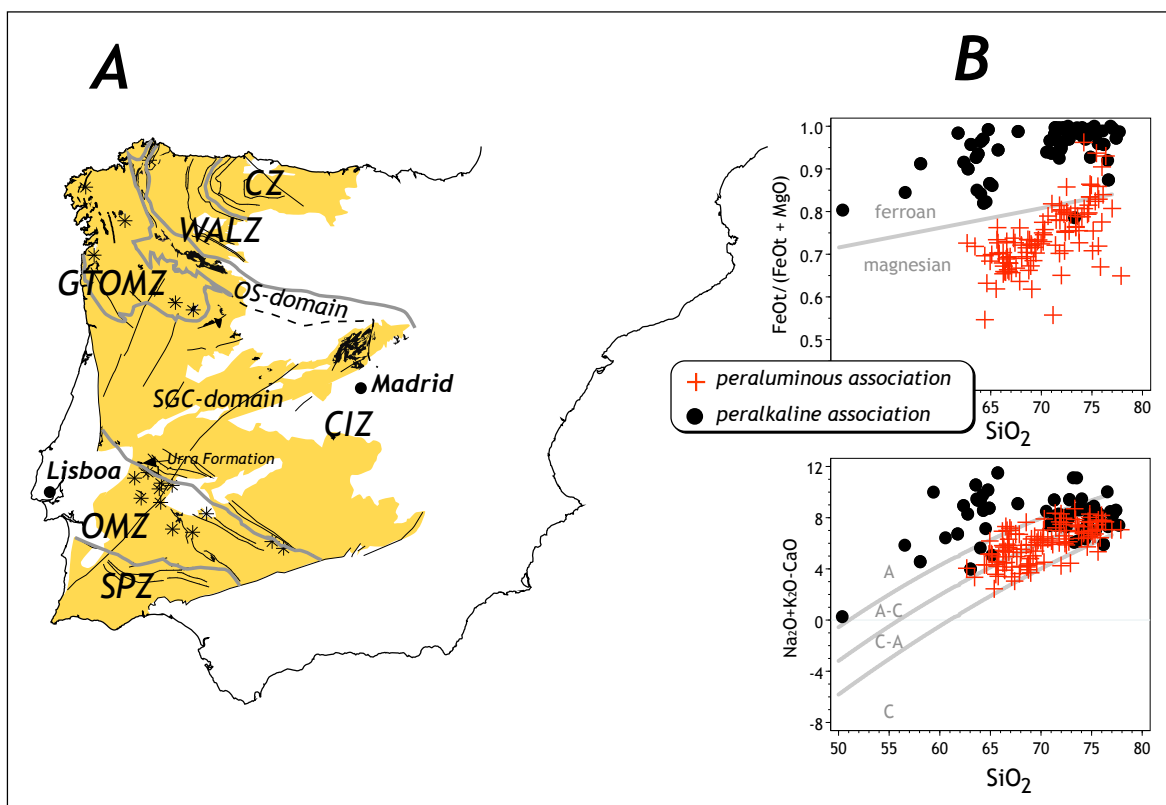


Figure 1.1: A) Exposures of the Cambro-Ordovician igneous rocks in the Iberian Massif. Black spots represent the peraluminous association; stars represent bodies of the peralkaline Association. CZ: Cantabrian Zone; WALZ: West Asturian Leonese Zone; CIZ: Central Iberian Zone, OS and SGC: Ollo de Sapo and Schist-Graywacke Complex domains; GTOMZ: Galicia Tras-os-Montes Zone; OMZ: Ossa-Morena Zone; SPZ: South Portuguese Zone. B) Frost et al. (2001) plots of the Cambro-Ordovician igneous rocks of Iberia revealing the contrasting composition of the peraluminous and peralkaline associations (see text for discussion).

can times, crop out following three NW-SE roughly parallel curved lineaments (Fig. 1.1A): The first, and most voluminous, comprises the Ollo de Sapo Formation, located near the boundary with the West Asturian Leonese Zone; the second comprises the orthogneisses of the northern part of the Schist-Graywacke Complex Domain, and the third, and less voluminous, includes the Urra Formation (Portugal), and is located at the boundary with the Ossa-Morena Zone.

All these rocks are augen-gneisses mostly derived from igneous rocks, either volcanic or plutonic, so that all of them can be properly named orthogneisses. To avoid confusion and distinguish between those derived from volcanic rocks from those derived from plutonic rocks, we shall use the terms “metavolcanic rocks” and “metagranites”. Obviously, this last term does not mean “a granite of metamorphic origin” but a “metamorphosed granite”.

The study of the Cambro-Ordovician igneous rocks of the Central Iberian Zone has historically raised much interest because of the following two main reasons:

- First, they represent the largest pre-Variscan magmatic episode in the Central Iberian Zone, the significance of which is still poorly understood. Whereas there is a general agreement that the Cambro-Ordovician peralkaline rocks and associated gabbros of the Ossa Morena Zone originated in a rifting environment (e.g., Ribeiro, 1987; Ribeiro and Floor, 1987; Santos Zalduegui et al., 1995; Montero and Bea, 1998; Sánchez García et al., 2003; Montero and Floor, 2004a), the Cambro-Ordovician peraluminous to calc-alkaline rocks of Central Iberia, solely by virtue of their chemical signature, have been interpreted by several authors as evidence for an active margin setting (e.g., Gebauer et al., 1993; Valverde-Vaquero and Dunning, 2000; von Raumer et al., 2003).
- Second, they are an important component of the source layer of the Iberian Variscan granites (Villaseca et al., 1998; Castro et al., 1999; Castro et al., 2000; Dias et al., 2002; Bea et al., 2003).

However, the study of the Iberian Cambro-Ordovician magmatism has been plagued with difficulties, most of them arising from the lack of precisely determined ages, which in the literature have been variously estimated, from Early Ediacaran to Late Ordovician.

Stratigraphically, the position of the metavolcanic rocks is poorly constrained. In the case of the Ollo de Sapo Formation, they are located below a siliciclastic series of probable Arenig age, the Armorican Quartzite, and overlie a non-fossiliferous series of unknown age which has been variously interpreted as Early Paleozoic or Late Proterozoic. Radiometrically, no attempt to date the Ollo de Sapo metavolcanic rocks by Rb-Sr, Sm-Nd or conventional U-Pb on zircon concentrates has succeeded so far. Only the El Cardoso gneiss in the eastern Guadarrama mountains (to the south of Riaza, Fig. 1.3), which although it does not belong to the Ollo de Sapo Formation bears many similarities to it, has been dated at 480 ± 2 Ma, U-Pb on zircons (Valverde-Vaquero and Dunning, 2000). Equally problematic is the age of the metagranites, especially those of the Ollo de Sapo Domain, in which U-Pb dating of zircon concentrates has characteristically yielded highly discordant and very imprecise ages, from 618 Ma (Lancelot et al., 1985) to 440 Ma (Fernández-Suárez et al., 1999).

The reasons for these difficulties may be summarized as follows:

- The lack of initial isotopic homogeneity and the disturbances produced during the Variscan metamorphism cause many difficulties when applying Rb-Sr and Sm-Nd methods.
- The zircons of these rocks are very complex, as revealed by our previous studies. Cathodoluminescence imaging revealed that a significant fraction of grains contain inherited cores which record different ages, from Ediacaran up to Archean (Bea et al., 2006; Montero et al., 2007; Bea et al., 2007; Sola et al., 2008; Talavera et al., in press). The simultaneous occurrence of two or more different age components make dating methods like conventional U-Pb in zircon concentrates useless, single crystal methods are the only way to properly date these rocks.

1.1.2. Objectives

The main goal of this work is to make a systematic study of the geochronology, Sr and Nd isotope geology, and geochemistry of the Lower-than-Middle Ordovician magmatism of the Central Iberian Zone. In particular, we aimed at the following objectives:

- To date, using zircon U-Pb ion microprobe and LA-ICPMS analyses, single-grain Pb-Pb stepwise evaporation analyses and, occasionally, whole rock and internal Rb-Sr isochrons, (i) the metavolcanic and metagranitic rocks of the Ollo de Sapo

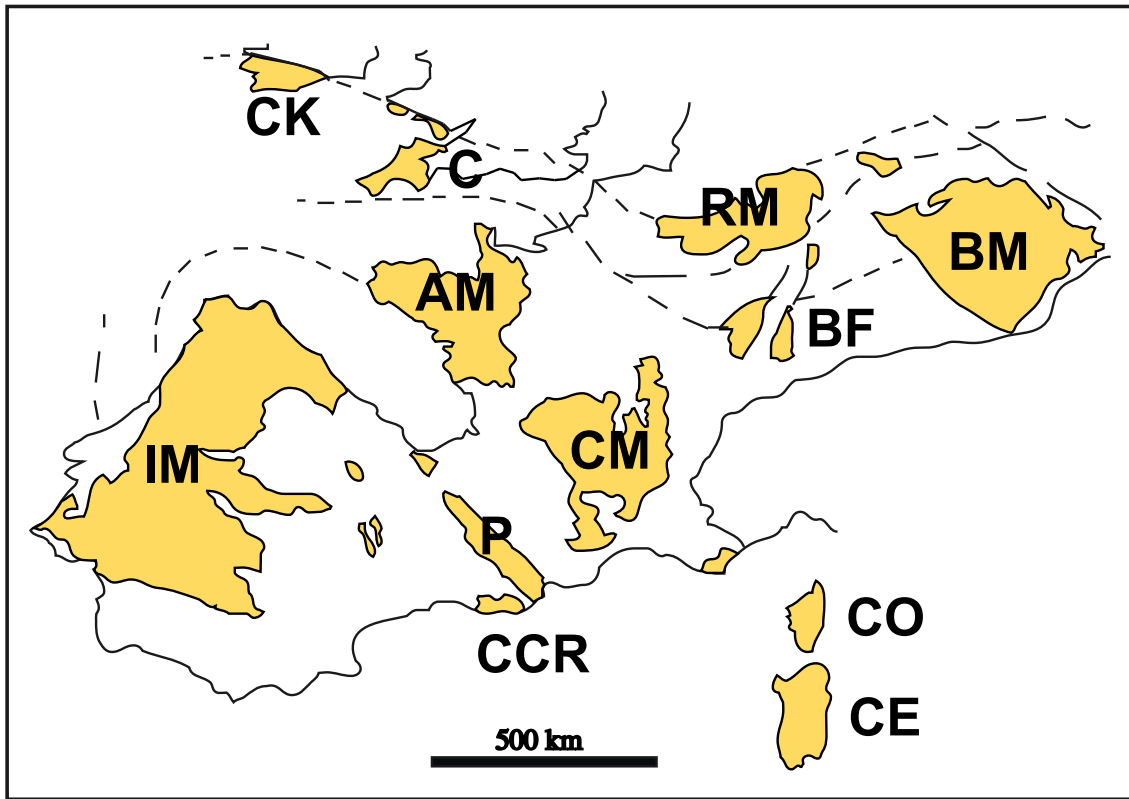


Figure 1.2. Main segments of the Variscides during the Early Carboniferous. IM, Iberian Massif; CCR, Catalan Coastal Ranges; P, Pyrenees; CM, Central Massif; AM, Armorican Massif; C, Cornubia; CK, Cork; RM, Renish Massif; BF, Black Forest; BM, Bohemian Massif. (Taken from <http://www.geofys.uu.se/eprobe/Projects/iberia/SWIberia.htm>).

Domain, and (ii) the pre-Middle Ordovician metagranites and metavolcanic rocks of the Schist-Graywacke Complex Domain.

- To study the major and trace element geochemistry and the Sr and Nd isotope geology of all these materials.
- To understand the reasons for the abnormally elevated zircon inheritance of these rocks, which largely exceeds what is to be expected for common felsic volcanic rocks and granites (e.g., Miller et al., 2003). To do this, we developed a numerical model based on zircon solubility and solution kinetics that constrains the minimum heating and cooling rates of the melting event that caused the Cambro-Ordovician anatexis.
- Finally, to propose a geodynamic model compatible with these data that may explain the origin of the Cambro-Ordovician magmatism of Central Iberia.

1.2 Geological setting

1.2.1. The Iberian Massif

The Iberian Massif is one of the largest exposures of the European Variscan Orogen (Fig. 1.2) occupying the west and northwest areas of the Iberian Peninsula (Fig. 1.1A). Lotze (1945), based on stratigraphic, structural, metamorphic and magmatic features, divided it into six paleogeographic zones which from the north to the south are: Cantabrian, West Asturian Leonese, Castellán-Galician, Alcuadian-Lusitanian, Ossa Morena and South Portuguese Zones. Afterwards, Julivert et al. (1972) joined the Castellán-Galician Zone and the Alcuadian-Lusitanian Zone together to form the Central Iberian Zone. Some years later, Fariás et al. (1987) and Arenas et al. (1988) identified a new zone within the Central Iberian Zone which they called the Galicia Tras-os-Montes Zone.

The internal regions of the Iberian Massif: the Ossa Morena, Galicia-Tras-os-Montes and Central Iberian Zones stand out because of the abundance of pre-Middle Ordovician magmatic rocks. These form two different geochemical associations. The most abundant is composed of intermediate to silicic peraluminous rocks ($\text{mol. Al}_2\text{O}_3/(\text{CaO} + \text{Na}_2\text{O} + \text{K}_2\text{O}) \approx 1.34 \pm 0.23$) which according to the classification scheme of Frost et al. (2001) are magnesian, and alkaline-calcic to calc-alkaline (Fig. 1.1B). The less abundant association is composed of ferroan, alkaline to calc-alkaline, and predominantly high-silica rocks (Fig. 1.1B) which often contain Na pyroxenes and/or amphiboles. Whereas the latter unequivocally form an A-type association, the former, despite having most features similar to S-type rocks including the occasional presence of cordierite and garnet, cannot be correctly termed so because, according to zircon evidence, they do not represent melts of sedimentary rocks, which strongly suggests that their magmatic source was predominantly composed of Pan-African igneous materials (Montero et al., 2007; Bea et al., 2007). Therefore, for the sake of simplicity, in this work we shall refer to these two different rock associations as “peraluminous” and “peralkaline”, noting that a few rocks of the first group have $\text{mol. Al}_2\text{O}_3/(\text{CaO} + \text{Na}_2\text{O} + \text{K}_2\text{O}) < 1.05$ and about 50-70% of the rocks of the second group, depending on the massif, have $\text{mol. (Na}_2\text{O} + \text{K}_2\text{O})/\text{Al}_2\text{O}_2 \leq 1$, i.e., they are not strictly peralkaline. Whereas in Ossa-Morena and Galicia-Tras-os Montes Zones, the two associations are well represented, in the Central Iberian Zone, despite of hosting the most voluminous pre-Middle Ordovician magmatism, only the peraluminous association is found.

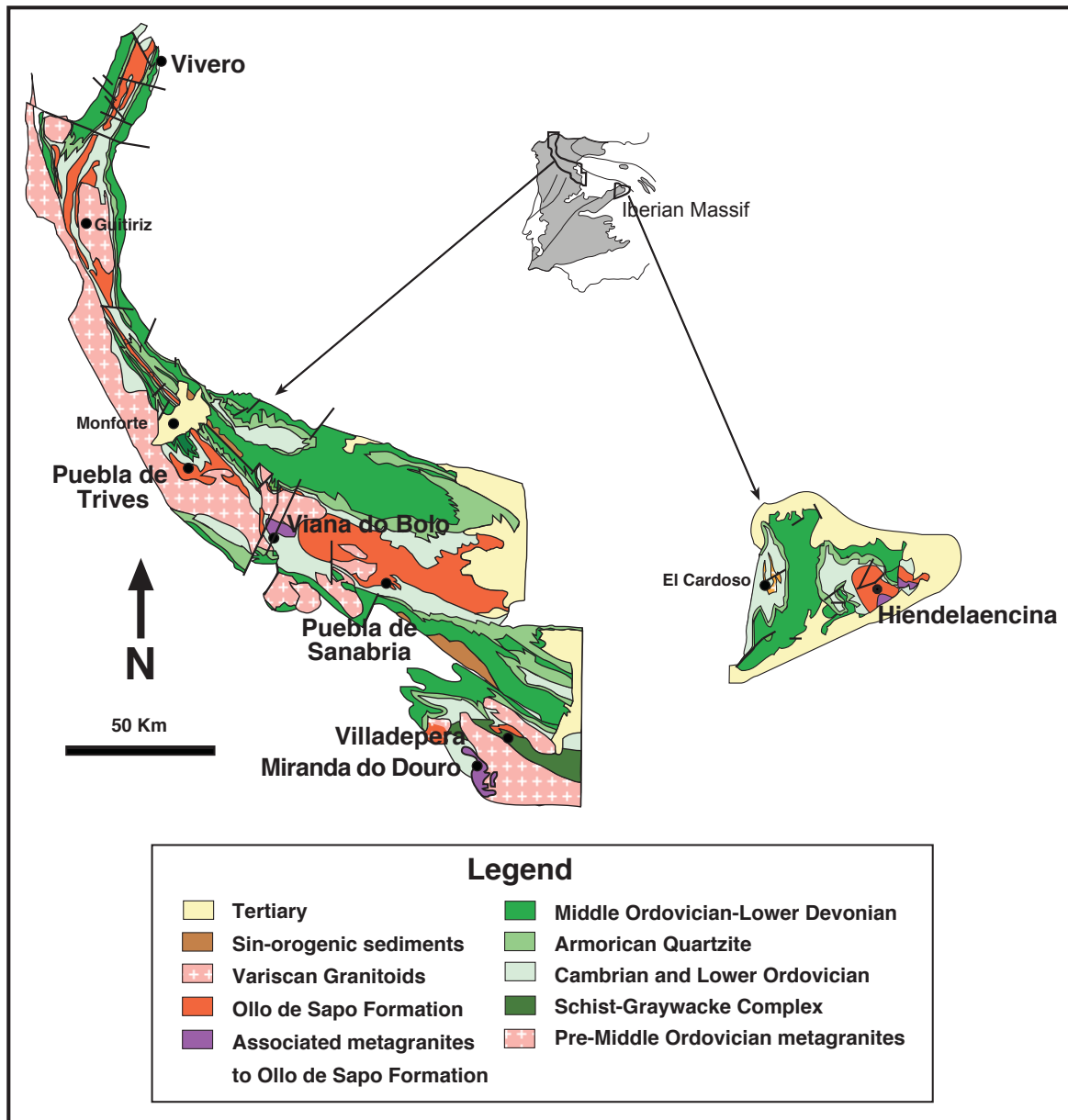


Figure 1.3: Geological scheme of the Ollo de Sapo Domain (Martínez Catalán et al., 2004).

The following section describes the features of the Central Iberian Zone which are relevant to the present work.

1.2.2. The Central Iberian Zone

The Central Iberian Zone is located in the central part of the Iberian Massif. Its boundaries are the Vivero fault and the Courel and Peñalba synclines in the north (Martínez Catalán, 1985), and the Badajoz-Córdoba shear zone in the south (Díez Balda et al., 1990; Azor et al., 1994). Rock formations in this area mostly consist of Late Proterozoic to Early

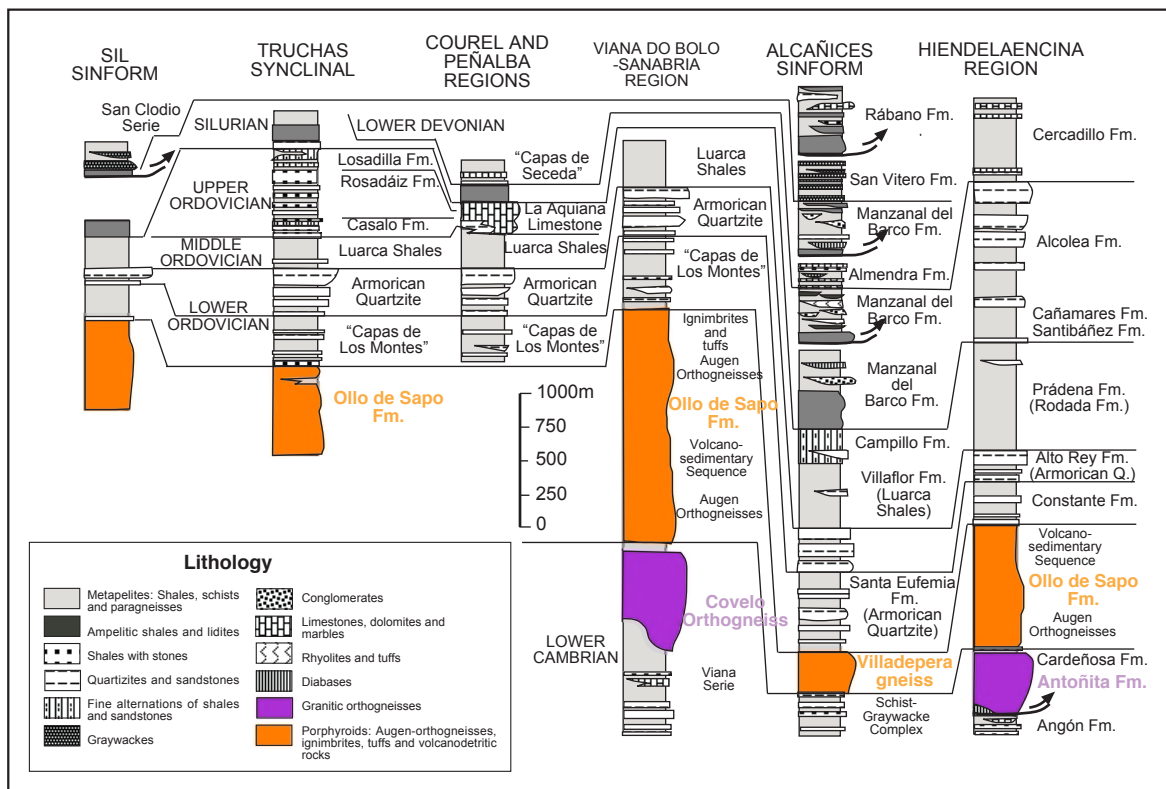


Figure 1.4: The most representative stratigraphical columns of the Oлло de Sapo Domain (Martínez Catalán et al., 2004)

Paleozoic metasediments and minor pre-Middle Ordovician metagranites intruded by numerous syn- or late-kinematic Variscan granitoids. Discordant Carboniferous deposits are common in the cores of large synclines.

The Central Iberian Zone can be subdivided into two domains: the Oлло de Sapo Domain to the north and the Schist-Graywacke Complex Domain to the south (Pérez-Estaún et al., 2004). The most important features of these areas are as follows:

1.2.2.1. The Oлло de Sapo Domain

The Oлло de Sapo Domain is a narrow zone located in the north of the Central Iberian Zone (Fig. 1.3), limited by the Vivero fault and Peñalba and Courel synclines to the east and northeast, which mark the boundary with the West Asturian Leonese Zone (Fig. 1.1A) (Martínez Catalán, 1985).

The Oлло de Sapo Domain is defined by the presence of the Oлло de Sapo Formation, which extends for about 570 kilometers, from the northeast Galician coast to the Sanabria region. From this region to the southeast, it is covered by the Tertiary deposits of the Duero

Basin to reappear again in the Hiendelaencina region (Fig. 1.3). This domain consists of: (i) a pre-Ordovician sequence, (ii) the Ollo de Sapo Formation, (iii) an Ordovician-Lower Devonian sequence, (iv) an Upper Devonian-Upper Carboniferous sequence (Fig. 1.4).

- **The pre-Ordovician sequence** crops out below the Ollo de Sapo Formation (Fig. 1.4), in the cores of large antiforms in the Hiendelaencina and Viana do Bolo region. In the last region this series has been called the Porto-Villavieja Series (Martínez García and Corretgé, 1970), Porto Series (Martínez García, 1973) or Viana Series (Ferragne, 1968). From bottom to top the series consists of: (i) quartzites and quartzitic micro-conglomerates, (ii) limestones associated with quartzites, (iii) schists and migmatitic gneisses in which Covelo and San Sebastián metagranites are intercalated, and (iv) limestones and calcic-silicate rocks. In Hiendelaencina, the sequence is very similar and from bottom to top it consists of: (i) a sequence of schists, quartzites, micro-conglomerates, marbles and amphibolites (the Angón Formation: González Lodeiro, 1980; Fernández Rodríguez, 1991), (ii) the sill-like Antoñita metagranite (Schäfer, 1969), and (iii) a sequence of calcic-silicate rocks and marbles with a few meters of feldspathic quartzites and schists (the Cardeñosa Formation). In Galicia the base of the Ollo de Sapo Formation is not exposed.
- **The Ollo de Sapo Formation** is intercalated between a pre-Ordovician sequence and the Lower Ordovician shales and quartzites, called “Capas de los Montes” in Galicia (Riemer, 1963), Puebla de Sanabria Formation in the Sanabria region (Martínez García, 1971, 1973), Santa Eufemia Formation in the Villadepera region (Vacas and Martínez Catalán, 1987) and Constante Formation in the Hiendelaencina region (Schäfer, 1969; González Lodeiro, 1980, 1981a, 1981b) (Fig. 1.4). The lower contact is always very deformed and is metamorphosed to medium or high grade. The upper contact is locally discordant. Two facies of the Ollo de Sapo Formation are often distinguished, one coarse-grained and the other fine-grained, these form alternating layers of variable thickness (Parga-Pondal et al., 1964; Navidad, 1978) :
 - *Coarse-grained facies.* The coarse-grained facies is preferentially, but not exclusively, in the lower part of the formation. It is composed of K-feldspar megacrysts ($\varnothing \approx 5-7$ cm, locally up to 15 cm) and phenocrysts of plagioclase



Figure 1.5: Facies of the Ollo de Sapo Formation: A) Coarse-grained facies in the Ria del Barquero (Lugo); B) Coarse-grained facies in the Hiendelaencina region; C) Fine-grained facies in the Ria del Barquero (Lugo); D) Fine-grained facies in Mañón (A Coruña).

and quartz ($\varnothing \approx 0.5\text{-}3\text{ cm}$) in a gray to greenish fine-grained groundmass (Fig. 1.5, A and B). K-feldspar megacrysts are usually euhedral or subhedral, locally rapakivi, and contain abundant inclusions of plagioclase, quartz, biotite and zircon. Plagioclase phenocrysts (An_{15-25}) are euhedral or subhedral. Quartz phenocrysts tend to be subhedral, locally rounded, and often show embayed contours. The groundmass is composed of quartz, K-feldspar, muscovite, biotite, rare albite, chlorite and sericite. The accessory mineral assemblage is often composed of apatite, zircon, Fe-Ti-oxides, monazite, xenotime, ilmenite and Fe-Cu-sulfides.

- *Fine-grained facies.* The fine-grained facies is typically located in the upper part of the Ollo de Sapo Formation. It consists of small phenocrysts of K-feldspar (up to 2 cm) and quartz (up to 1.5-2 cm) inside a groundmass with a modal composition identical to that of the coarse-grained facies (Fig. 1.5, C and D). Often the fine-grained facies is intercalated with layers of metasediments, mostly schists, quartzites and quartz-feldspatic schists. In some areas, such as the Sanabria and Hiendelaencina regions, both facies are finely intercalated, so forming the “mixed facies”.
- **The Ordovician-Lower Devonian sequence** overlies the Ollo de Sapo Formation. It is composed of pre-orogenic deposits from the Lower Ordovician to the Lower Devonian (Fig. 1.4). The Ordovician deposit consists of: (i) alternating quartzites and black shales with minor intercalations of micro-conglomerates (Matte, 1968; Pérez Estaún, 1978), and a quartzitic formation, both of which are Lower Ordovician, (ii) a formation of gray-bluish shales with minor quartzitic intercalations which is Middle Ordovician (Gutiérrez-Marco et al., 1999), and (iii) gray and gray-greenish shales with intercalations of sandstones and quartzites, and sometimes beds of limestones and dolomites, which are Upper Ordovician.

According to Schäfer (1969), Bultynck and Soers (1971), Soers (1972), Carls (1977) and Gutiérrez-Marco and Robardet (1991), the Silurian rocks are composed of: (i) few quartzitic beds, (ii) black carbonaceous shales (ampelites), and (iii) gray shales, thin beds of gray and white quartzites, lidites, black and very fine-grained silicious and carbonaceous rocks and thin beds of limestones.

Finally, as described by Drot and Matte (1967), Carls (1969) and Bultynck and Soers (1971), the Lower Devonian is made up of shales and limestones.

- **The Upper Devonian-Upper Carboniferous sequence** is formed of a succession of syn-orogenic flysch facies, mostly turbiditic deposits of argillites, graywackes and micro-conglomerates (Martínez Catalán and González Clavijo, 2004).

Spatially associated with the metavolcanic rocks of the Ollo de Sapo Formation there are some augen-gneisses that undoubtedly represent intrusive granite bodies. They appear as sills or lacoliths up to 300 m thick (González Lodeiro, 1981b; Iglesias Ponce de Leon and Ribeiro, 1981) that intrude the metasedimentary sequence underlying the Ollo de Sapo metavolcanic rocks. The best known are: (i) the Antoñita metagranite in the Hiendelaencina region, (ii) the Miranda do Douro metagranite in the Villadepera region, and (iii) the Covelo and San Sebastián metagranites in the Viana do Bolo area. These rocks are coarse-grained augen-gneisses that differ from those of the Ollo de Sapo in having a coarser groundmass, a lack of blue quartz crystals and the presence of intrusive contacts, aplopegmatitic dikes and metapelitic enclaves. They are composed of large K-feldspar megacrysts ($\varnothing \approx 5\text{-}12$ cm) within a coarse-grained foliated groundmass of quartz, oligoclase, K-feldspar, biotite, muscovite, and rare tourmaline, cordierite and garnet. The K-feldspar megacrysts frequently have multiple tiny inclusions of plagioclase and biotite aligned following crystallographic contours (Fral's inclusions), and also often have rapakivi structures. The accessory minerals are apatite, ilmenite \pm magnetite, zircon, monazite and xenotime.

1.2.2.2. The Schist-Graywacke Complex Domain

The Schist-Graywacke Complex Domain occupies the major part of the Central Iberian Zone (Fig. 1.6). It is bounded by the Badajoz-Córdoba shear zone in the south, marking the boundary with the Ossa Morena Zone. At the northeast boundary of this domain, to the west of Zamora, there are recumbent folds with northeast vergence which define a transitional zone to the Ollo de Sapo Domain. The main sedimentary units of this domain are: (i) The Schist-Graywacke Complex, (ii) the Cambrian transgressive deposits, and (iii) the post-Cambrian-Paleozoic sequence.

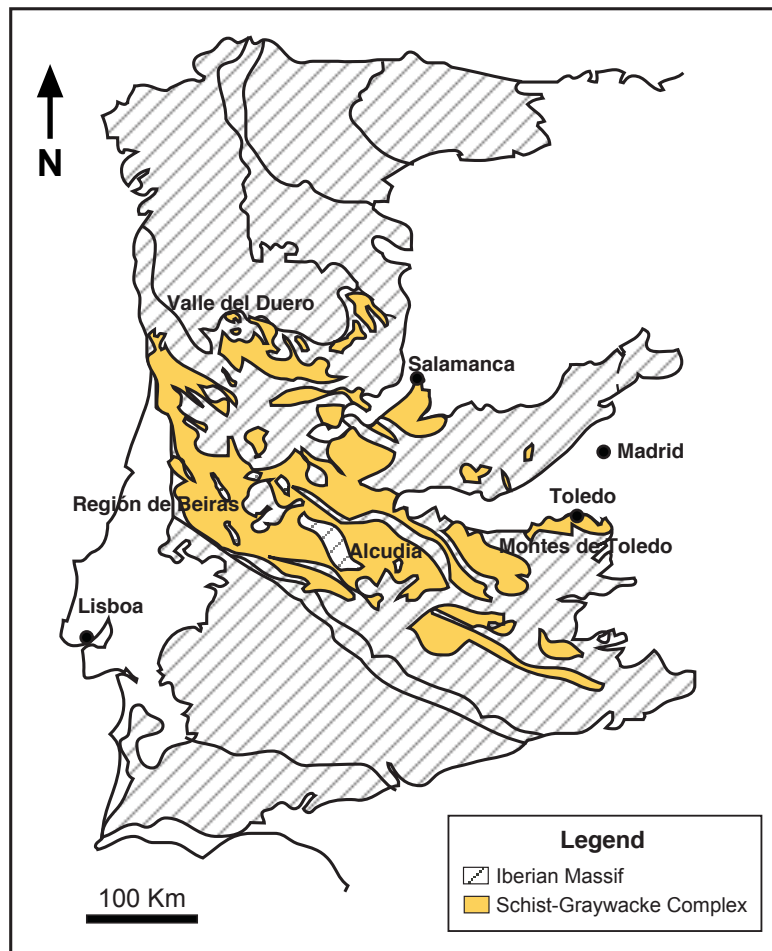


Figure 1.6: Outcrops of the Schist-Graywacke Complex in the Central Iberian Zone (Bernardo de Sousa, 1984).

- **The Schist-Graywacke Complex.** This term refers to important Ediacaran to Cambrian monotonous non-fossiliferous pelitic and psammitic successions (Car-rington da Costa, 1950) which crop out in the core of large antiforms in central Portugal (Valle del Duero and região de Beiras), Salamanca, Extremadura, Mon-tes de Toledo and Alcudia (Fig. 1.6). The Schist-Graywacke Complex can be separated into two units which have different names depending on the places where they have been defined (Fig. 1.7). According to Rodríguez Alonso et al. (2004) and Diez Balda (1986) these units can be described as follows:
 - The lower unit, often called the Monterrubio unit. This is a monotonous suc-cession consisting of lutites and sandstones with minor intercalations of con-glomerates, chaotic beds and few volcanoclastic beds, which appear in the up-per part of the series. It crops out throughout all the Central Iberian Zone ex-

Age	Portugal (Bernardo de Sousa, 1984)	South of Salamanca-Ciudad Rodrigo-Hurdes-Sierra de Gata (Rodríguez Alonso, 1985); (Diez Balda, 1986); (Valladares et al., 1998, 2002a, 2002b)	Montes de Toledo-Alcudia-Extremadura (Herranz et al., 1977);(Álvarez Nava et al., 1988); (Robles, 1988);(Palacios, 1989); (San José et al., 1990); (Vidal et al., 1994);(García-Hidalgo, 1985); (Pieren, 2000)
Ordovician			
Upper Vendian	Lower Cambrian	Endrinal Shales	Los Cortijos S.
		Tamames Limestones	Solares Fm.
		Tamames Sandstones	Los Navalucillos Limestones
		Monterrubio Fm. Aldeatejada Fm.	Azorejo Sandstones
		Upper Unit	Pusa Shales
		Unit XII to Unit VI	Fuentes Bed/ El Membrillar
		Unit V El Bodón Breccia/ Conglomerates and vulcanites	Tañeijo shales and conglomerates/ Cubilar Fm./ Estomiza Fm./ Valcasa S.?
		Unit IV Pastores Limestones/ Fuente-guinaldo / Agueda	Villarta Limestones/ Peraleda/ San Román/Ibor
		Lower Unit	Nª Sª de La Laguna Congl.
		Unit III	Lower Alcudian/ Guadiana Shales/ Extremadura Dome Group
		Unit II	
		Unit I	
		Grupo do Douro	Pusa Group
		Grupo das Beiras	Valdecaza G./Rio Huso G.
			Ibor-Navalpino Group
			Shales
			Conglomerates
			Upper Shales with Anabarella
			San Lorenzo Conglomerates
			Cabezarrubias shales and limestones
			Tamujar Fm.
			Talarrubias Dolomites
			Orellanita Congl. Unit
			Valdemanco Dolomites
			Collado de la libre Quartzites

Figure 1.7: Scheme of the chronostratigraphy and the correlation of the units, groups and formations which have been defined in different areas of the Schist-Graywacke Complex Domain for pre-Ordovician sequence to the north of Los Pedroches batholith (Rodríguez Alonso et al., 2004).

cept in the northwest of the domain (in the areas of Salamanca, Fregeneda and Douro). Its thickness is unknown because the lower boundary is not exposed.

- The upper unit, often called the Aldeatejada unit. This occurs on top of the Lower unit and it is mainly composed of pelites with some beds of black lutites, conglomerates and sandstones, together with limestones and discontinuous olistostromic beds and intercalations of phosphates and volcanic and volcanoclastic rocks. This unit has been divided into a several lithostratigraphic units which mainly crop out in the anticlines of Alcudia, Ibor, Navalpino and Valdelacasa, at the edge of the Ciudad Rodrigo-Hurdes-Sierra de Gata area and in the Douro-Fregeneda-Salamanca area and its eastern prolongation.

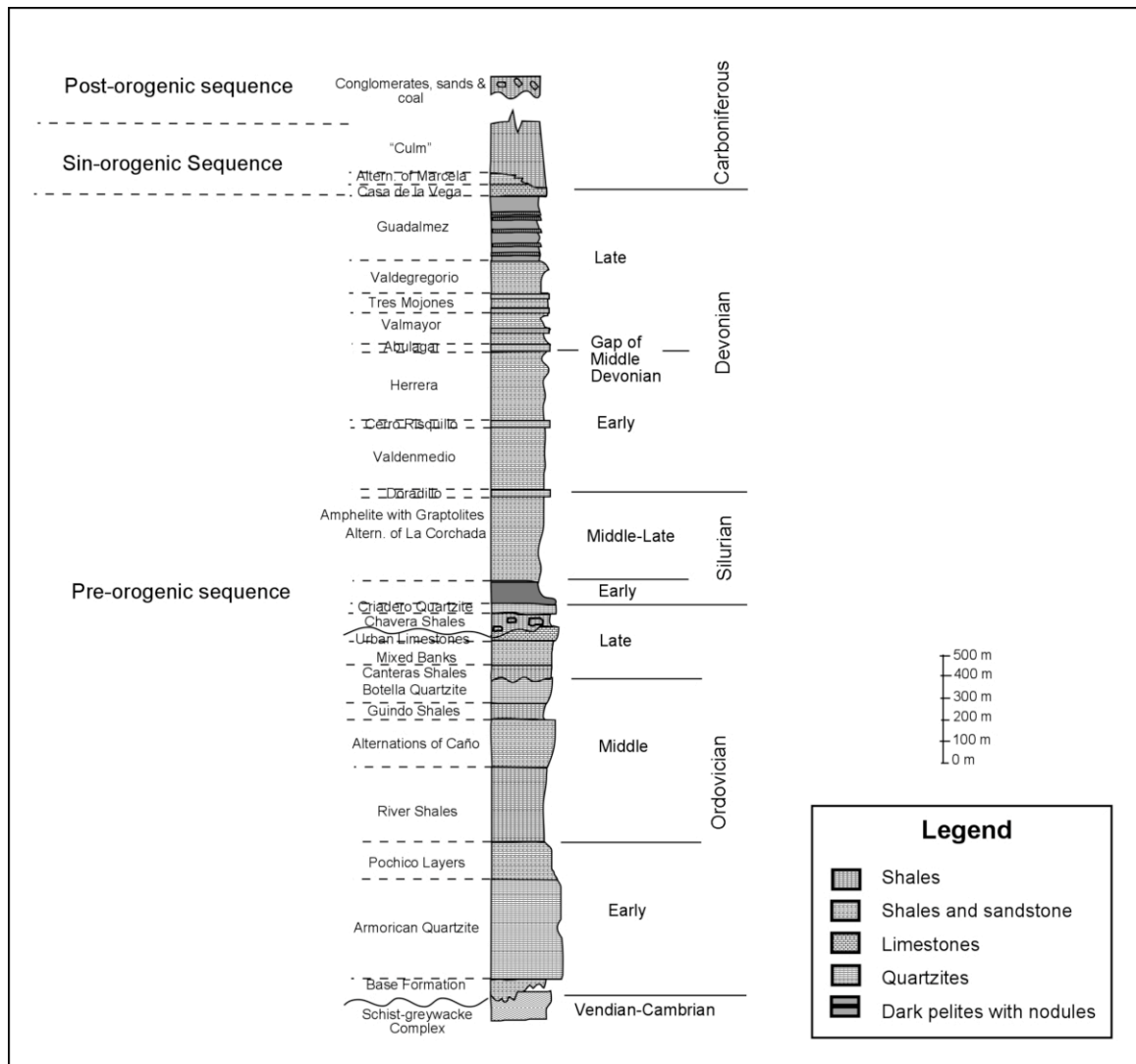


Figure 1.8: Stratigraphic column of the Ordovician-Carboniferous rocks of the Schist-Graywacke Complex Domain to the north of Los Pedroches Batholith (Martínez Poyatos et al., 2004).

These units can be recognized throughout all the Central Iberian Zone except to the south of the Pedroches Batholith, where the Cambrian-Ediacaran sedimentary sequence bears more similarities to the sequences of the same age found in the Ossa Morena Zone to the south (Fig. 1.1A).

- **The Cambrian transgressive deposits** crop out discontinuously throughout the Schist-Graywacke Complex Domain, except in the southeast and the central-eastern areas of the Central Iberian Zone, and the Portuguese region of the Douro. In the central and eastern areas of the Central Iberian Zone, this sequence consists of shallow marine sandy deposits overlain by platform carbonates. In the

Fregeneda and the Portuguese region of the Douro, the sequence is composed of a pelitic-psammitic succession with calcic-silicate and conglomeratic beds.

- **The post-Cambrian Paleozoic sequence.** The post-Cambrian Paleozoic sequence crops out in the cores of synclines with northwest-southeast orientation and in the southern part of the Schist-Graywacke Complex Domain. According to Martínez Poyatos et al. (2004), it can be divided in three sequences (Fig. 1.8): pre-orogenic, sin-orogenic and post-orogenic.

- **Pre-orogenic sequence.** This mainly consists of a succession of detritic formations deposited in a shallow-marine siliciclastic platform environment. These sediments are discordantly deposited over the Cambrian deposits (Lotze, 1956; Gutierrez Marco et al., 2002). In detail, the sequence is composed of: (i) conglomerates, sandstones and silts, Armorican Quartzite, and alternating quartzites and shales which are Lower Ordovician, (ii) dark shales, alternating sandstones and shales, gray shales, and quartzites which are Middle Ordovician, (iii) shales (at the base of which there is a conglomeratic bed with phosphates), alternating shales and sandstones, discontinuous calcareous bed, a disconformity related to an erosive event (Robardet, 1981), and diamictites which are Upper Ordovician, (iv) orthoquartzites at the top of which is the Ordovician-Silurian boundary, (v) black shales with volcanic beds and alternating Silurian shales, sandstones, and orthoquartzites; (vi) alternating shales and sandstones, orthoquartzites and ferruginous sandstones, alternating shales and sandstones with intercalated bioclastic limestones in the lower part which are Lower Devonian; (vii) a stratigraphic gap (Puschmann, 1967) which covers almost all the Middle Devonian; (viii) a quartzitic unit, orthoquartzites inter-layered within two important units of alternating shales and sandstones, dark gray shales with nodules, limestones, and, finally, shales and sandstones, which are Upper Devonian-Lower Carboniferous.

To the south of the Albuquerque-Pedroches magmatic lineation, the Devonian pre-orogenic sequence is different. It has disconformities between the Devonian series and the previous series, and the extension of the stratigraphic gap is bigger.

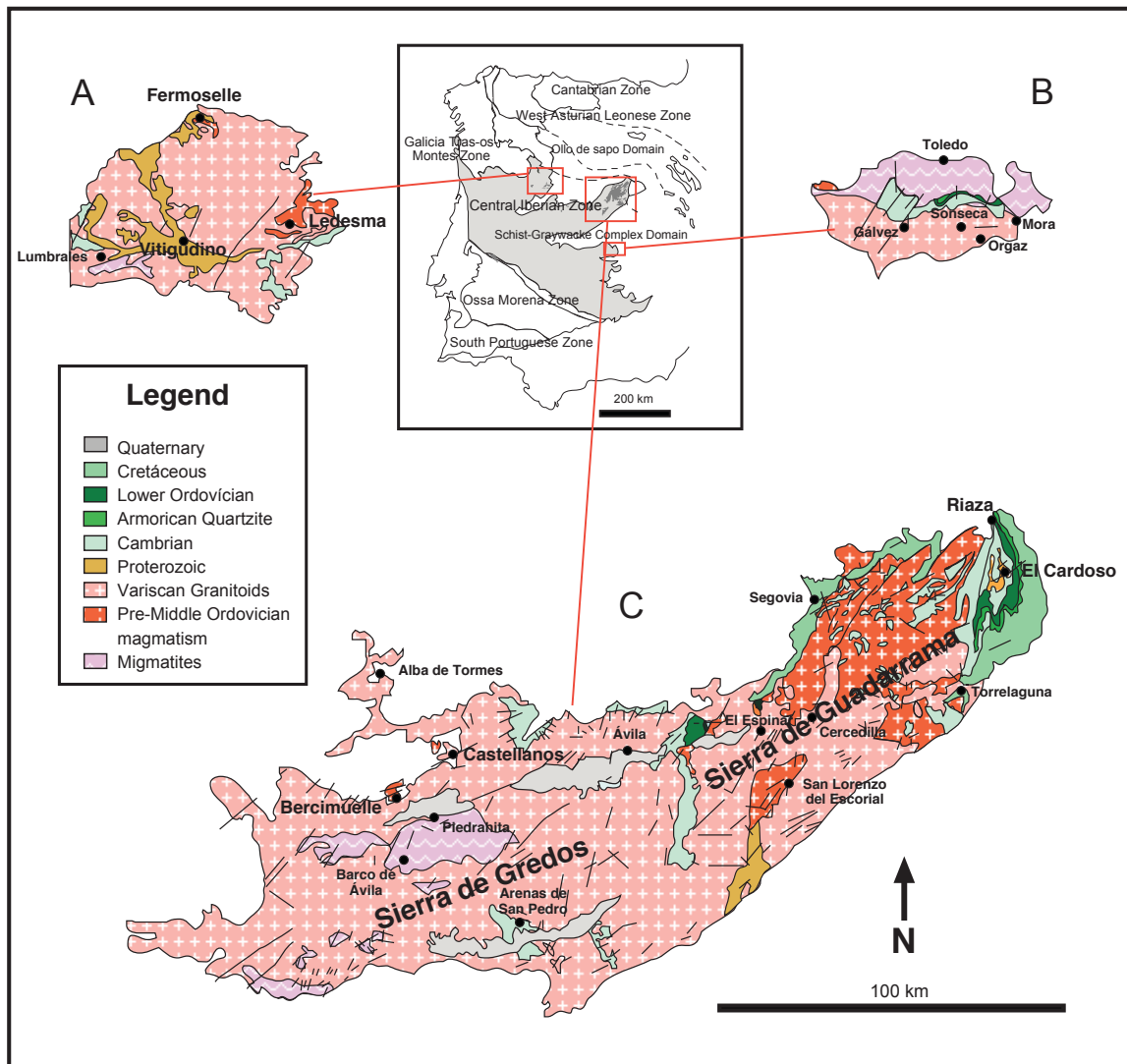


Figure 1.9: Geological Scheme of areas of the Schist-Graywacke Complex Domain where pre-Middle Ordovician metagranites and metavolcanic rocks crop out. The geological scheme represents a section of the geological map of Spain (Álvaro et al., 1994).

- **Sin-orogenic sequence.** It crops out in the south of the Schist-Graywacke Complex Domain and is arranged in northwest-southeast bands. It consists of alternating Lower-Middle Carboniferous shales and graywackes in the Culm facies with intercalations of conglomerates, acid and basic volcanic rocks and minor limestones, which are preferentially found at the bottom of this sequence (García-Alcalde et al., 1984; Rodríguez et al., 1990). At the southern edge, there are two outcrops of sin-orogenic sequences (Mamet and Martínez, 1981; Andreis and Wagner, 1983; Cózar and Rodríguez, 1999), one of them

has paralic sedimentation and the another limnic sedimentation and cyclic sedimentation of conglomerates, sands, silts and coal.

- **Post-orogenic sequence.** This is a sub-horizontal cyclical discordant sequence composed of Upper Carboniferous conglomerates, sands, silts and coal beds (Wagner and Utting, 1967) which crops out in few isolated intra-mountain basins.

The pre-Middle Ordovician magmatism in the Schist-Graywacke Complex Domain is mostly represented by (i) intrusive granitoids and metavolcanic rocks cropping out in the northern sector of the domain, in the areas of Salamanca, Ávila, Toledo and, especially, Guadarrama (Fig. 1.9), and (ii) the Urra Formation, totally analogous to the Ollo de Sapo Formation albeit much less voluminous, which crops out in the south, just at the boundary with the Ossa Morena Zone.

1.2.2.3. Deformation in the Central Iberian Zone

The pre-Variscan and Variscan deformation in the Central Iberian Zone can be summarized as follows:

- **The pre-Variscan deformation** is revealed by three discordancies in the Schist-Graywacke Complex Domain, named Alcludian, Cadomian and Toledanic. The first two formed during the Cadomian or Pan-African orogeny, and the last one during a Middle-Upper Cambrian low intensity tectonic event (González Lodeiro et al., 2004).
- The **Alcludian discordance** is found in the southern part of the Schist-Graywacke Complex Domain between the Lower and Upper Alcludian. It is an angular discordance which was identified by the higher dip of the intersection lineation in the Lower Alcludian than in the Upper Alcludian (Ortega Gironés and González Lodeiro, 1986).
- The **Cadomian discordance** has only been observed in the Peraleda Anticline of the southern Schist-Graywacke Complex Domain, (Llopis, 1970; Capdevila et al., 1971), where a schistosity, dated at 550 ± 10 Ma (Blatrix and Burg, 1981) is found in the “Serie Negra”. This schistosity does not affect the overlying Ordovician succession.

- The **Toledanic discordance** is at the bottom of the Ordovician succession. It is affected by minor folds without associated foliation and metamorphism (Diez Balda et al., 1990), it is recognizable because the Ordovician succession lies on top of different formations of the underlying succession. This discordance has been called "Sardic" (Lotze, 1945) and was attributed to the effects of pre-Ordovician movements in a block-tectonic regime. The Toledanic discordance is always between the Armorican Quartzite and the Ollo de Sapo Formation, and affects the Central Iberian and Ossa-Morena Zones.
- **The Variscan deformation** took place through the Late Devonian-Late Carboniferous and was developed in three phases. The most remarkable features of these phases are:
 - **First phase (D1)**. The first phase is characterized by the formation of recumbent folds with east and northeast vergence and an S1 cleavage in the Ollo de Sapo Domain (Matte, 1968) and the north and southeast area of the Schist-Graywacke Complex Domain (Macaya et al., 1991; Azor et al., 1994). In the central area of the Schist-Graywacke Complex Domain, folds have sub-vertical axial planes (Diez Balda et al., 1990).
 - **Second phase (D2)**. The second phase is distinguished by the development of the thrust faults which cut the aforementioned recumbent folds. In the Ollo de Sapo Domain, these thrust faults generally form within Silurian metapelites, with abundant graphite-rich layers (González Clavijo and Martínez Catalán, 2002).
 - **Third phase (D3)**. The third phase is recognized by the formation of back-folding structures whose superposition formed several large folds in the Ollo de Sapo Domain, for instance the El Cardoso Antiform and the Majaelrayo Synform (González Lodeiro et al., 1988; Macaya et al., 1991). In the central and south areas of the Schist-Graywacke Complex Domain the previous recumbent folds and thrust faults were bent by folds with a vertical axial plane or northeast vergence and northwest-southeast direction.
- Moreover, **some late-orogenic extensional structures** have been identified in both domains. These structures are generally described as extensional shear

zones and crop out in the deeper structural levels of antiforms and domes (Escuder Viruete et al., 1994; Escuder Viruete et al., 2000; Diez Balda et al., 1995).

1.3. Previous work

Numerous researchers have studied the stratigraphy, petrology, structural geology and geochronology of the Cambro-Ordovician magmatism of the Central Iberian Zone. What follows is a brief summary of the most important studies.

1.3.1. The Ollo de Sapo Formation and associated metagranites

1.3.1.1. General geology and lithology

The term “Ollo de Sapo” derives from the presence of quartz phenocrysts that, when the metamorphic grade is medium or low, have a deep-blue color, probably due to inclusions of sagenitic rutile. The study of the rocks currently grouped under the name of Ollo de Sapo started at the end of the nineteenth century with Cortázar (1874) and Macpherson (1883). The first author considered the Ollo de Sapo of the Mombuey region (Zamora province) as a poligenetic psefite. The second author also proposed a detritic origin for the Ollo de Sapo from Ria del Barquero (Lugo) and suggested that certain fragments of this detritic formation may have come from granitic rocks.

In the first part of the twentieth century, Hernández Sampelayo (1922) described a micaceous nodular gneiss in the Colleira island (Lugo province) which was called “Ollo de Sapo” by the villagers of this region. This gneiss had abundant feldspar crystals from 2 millimeters to 8 centimeters and bluish quartz phenocrysts. Schroeder (1930) described the upper part of the Hiendelaencina formations and considered these rocks as conglomeratic gneisses. Finally, Lotze (1945) studied the porphyric formation of the Sanabria region and considered it to be of volcanic origin, similar to volcanic gneisses of Azuaga, in the Ossa Morena Zone.

In the second part of the twentieth century, numerous researchers have studied the Ollo de Sapo Formation. San Miguel de la Cámara and Lobato Díez (1955) working in the area surrounding the Sanabria lake described the presence of an augen-gneiss. Riemer (1963) in the south of the Lugo province described a volcanic acid association of lavas, tuffs and variably metamorphosed sub-volcanic rocks and named it the Vilachá complex, nowadays known as part of the Ollo de Sapo. Parga-Pondal et al. (1964) provided the first description of the Ollo de Sapo as a whole. These authors defined the Ollo de Sapo as a complex for-

mation consisting of two types of rocks: (i) Detritic rocks consisting of arkosic conglomerates, schistose arkoses, arkoses s.s. with some intercalations of schists and quartzites, which were probably derived from granites and gneisses, and (ii) Tuffaceous volcanic rocks of predominantly rhyolitic composition.

Anthonioz and Ferragne (1967), on the other hand, reached a contrasting conclusion. These authors suggested that in the Sanabria area the Ollo de Sapo showed features of a sedimentary complex: the fact that the coarse-grained facies are located at the bottom of the complex and the fine-grained facies on top was considered to be due to stratigraphic polarity. Furthermore shear structures that often affect these rocks were considered to reflect sedimentary characteristics such as grain selection and crossed stratification. The authors also concluded that most of the detritic material of the Ollo de Sapo came from Galician orthogneisses. Finally, Ferragne (1968) studied the pre-Cambrian basement of the Viana do Bolo region (Southern Galicia) noting the presence of orthogneisses and paragneisses.

In the 1970s, Martínez García and Corretgé (1970) listed the most distinctive features of the different petrographic units of the Porto-Villavieja Series, including the San Sebastián and Porto gneisses. Martínez García (1973) recognized two facies in the Ollo de Sapo Formation of the Sanabria region and concluded that although the sedimentary contribution was important in some areas, the origin was basically volcanic. Navidad (1978), in agreement with Parga-Pondal et al. (1964) and Martínez García (1973), differentiated two facies: the first one was composed of metavulcanites and metaarkoses preferentially located in the lower part of the formation, and the second one consisted of graywackes and pelites preferentially located in the upper part of the formation. Navidad (1978) suggested an igneous-sedimentary origin for the Ollo de Sapo Formation based on geochemical data.

In the 1980s, Navidad and Peinado (1981) studied the orthogneisses and the metasediments underlying the Ollo de Sapo Formation, that is to say the Angón Formation, the Antoñita gneiss and the Cardeñosa Formation. They concluded that the Antoñita gneiss was a banded and very homogenous augen-gneiss and, based on its chemical-mineralogical features, it was classified as an S-type granitoid belonging to the very differentiated terms of the S-type granitoids with abundant primary muscovite and low-An plagioclase. This gneiss also has associated aplitic and pegmatitic leucogneisses. Iglesias Ponce de León and Ribeiro (1981) carefully mapped the Miranda do Douro and San Vitero-Carbajales regions, and described the Ollo de Sapo Formation as a monotonous series of feldspathic meta-

graywackes and rhyolites with blue quartz crystals with massive layers except at the base where they show grain selection. These authors also identified some decimetric lens of pelites and orthoquartzites placed at the bottom of this formation. González Lodeiro (1981b) studied the Lower-than-Ordovician Series of the eastern Central System (From bottom to top: the Angón Formation, the Antoñita Formation, the Cardeñosa Formation, the Hiendelaencina Formation and La Constante Formation) and their correlation. This author concluded that the Hiendelaencina Formation, which was composed of coarse-grained gneisses at the bottom and fine-grained gneisses at the top, could be correlated with the Ollo de Sapo Formation of the northwest of the Iberian Peninsula and Sanabria region. By contrast the Antoñita Formation, which is composed of quite homogenous augen-gneisses, was very difficult to correlate.

Finally, Díez Montes (2006), in detailed research about the Ollo de Sapo Formation in the Sanabria region, described this formation as a series of lithologic groups composed of augen orthogneisses and a volcanic and volcanosedimentary sequence whose textural, structural and compositional features suggested a common origin.

1.3.1.2. Stratigraphic position

Some of the first researchers who studied the stratigraphy of the Ollo de Sapo Formation were Parga-Pondal et al. (1964). These authors attributed a Lower-than-Cambrian age to this formation by analogy with similar formations of the Iberian Peninsula and France. Later, Iglesias Ponce de León and Ribeiro (1981) carefully mapped the Miranda do Douro y S. Vitero-Carbajales regions, and concluded that the Ollo de Sapo Formation is between the Lower Ordovician Puebla Formation and the Lower Cambrian Schist-Graywacke Complex. Arias et al. (2002) mapped the Viana do Bolo-A Gudiña region and deduced that the Ollo de Sapo Formation was intercalated with the Upper Cambrian-Lower Ordovician “Serie de los Cabos”. Finally, Díez Montes (2006) attributed a Lower Ordovician age to this formation placing it between the Viana do Bolo Series and “Las Capas de Los Montes”.

1.3.1.3. Structural geology

The first study of the structural geology of the Ollo de Sapo Formation was made by Matte (1963), who described the Variscan Deformation of Galicia. This author included the Anticlinorium of the Ollo de Sapo in the domain of the Recumbent and Bend Folds during

the Variscan orogeny. He also assigned names to several antiform structures in whose cores the Ollo de Sapo Formation crops out. Martínez García (1973) studied the deformation and the metamorphism of the Zamora region. In agreement with previous authors, he recognized two facies in the Ollo de Sapo Formation of the Sanabria region and concluded that the differences between both facies were not due to a different origin but a different metamorphic history. Facies with feldspar megacrysts had suffered more intense metamorphism than facies without feldspar megacrysts. González Lodeiro (1981b) analyzed the Ollo de Sapo anticline in the Hiendelaencina region (Central System) and deduced that it was a recumbent fold with east vergence produced during the first phase of the Variscan orogeny. Finally, Capote et al. (1981) looked at large ductile thrust faults of the Central System generated during the second phase of the Variscan deformation. According to these authors, this phase of thrust faults was mainly syn-metamorphic and modified the distribution of the regional metamorphism isotherms causing important thermal inversions.

1.3.1.4. Radiometric dating and isotope geology

The first radiometric studies of the Ollo de Sapo Formation and associated metagranites using U-Pb in zircon concentrates were done by Lancelot et al. (1985) in the Sanabria-Villadepera region. These authors studied the Viana do Bolo (which was actually the San Sebastián metagranite) and Miranda do Douro orthogneisses and two samples from the middle and upper facies of the Ollo de Sapo gneisses. The results were highly discordant in all cases, with upper intersections at 618 ± 9 Ma and 465 ± 10 Ma for the Miranda do Douro and Viana do Bolo orthogneisses, respectively, which they interpreted as the crystallization ages of these orthogneisses. In the Ollo de Sapo sequence, they obtained a discordia with an upper intersection at 1797 ± 30 Ma and a lower intersection at 325 ± 3 Ma. They concluded that the upper intersection represented an inherited population of zircons and the lower intersection yielded the age of the Variscan main thermal event. The authors proposed a model based on the U-Pb system of the Ollo de Sapo in which there was mixing of pre-Cambrian detrital zircons (older than 1.8 Ga) with zircons originating in sub-volcanic rocks formed at 500-600 Ma. The mixing was produced during the deposit in the Ollo de Sapo Formation. Later, during the Variscan thermal event, they suggested these zircons were affected by a substantial loss of radiogenic lead.

Wildberg et al. (1989) studied U-Pb in zircon concentrates from the Hiendelaencina series (La Constante augen-gneiss, the Bustares augen-gneiss and the Antoñita gneiss) and

obtained the following results: (i) a discordia with an upper intersection at 2055 Ma and a lower intersection at 378 ± 6 for the Antoñita gneiss, and (ii) a group of data for the Constante augen-gneiss and the Bustares augen-gneiss which did not plot a discordia. The Constante augen-gneiss data plotted above and below the Antoñita reference line and Bustares augen-gneiss data marked a curve which moved away from the Antoñita reference line.

The above-mentioned results reveal the enormous difficulties in dating these rocks with U-Pb zircon concentrates. The first spot U-Pb zircon data for the Ollo de Sapo Formation and associated metagranites were done by Gebauer et al. (1993) who, in an abstract presented at the 1993 Annual Meeting of the Geological Society of America, suggested a crystallization age of 488 Ma for Ollo de Sapo Formation, and indicated the presence of concordant ages at 2670 and 578 Ma in some inherited cores. Fernández-Suárez et al. (1999) investigated the geochronology of the northwest outcrops of the Ollo de Sapo Formation using LA-ICPMS. They detected the presence of four age groups: 440-460 Ma, 590-620 Ma, 1.2 Ga and 1.9-2.0 Ga and concluded that the youngest group (440-460 Ma) was the age of the magmatic component of the Ollo de Sapo Formation and the other age groups (590-620 Ma, 1.2 Ga y 1.9-2.0 Ga) represented recycled zircon populations.

Full understanding of the reasons for the highly discordant nature of the zircons from these rocks comes from the first studies that combined cathodoluminescence imaging and single grain or spot analysis (e.g., Bea et al., 2006a). These authors revealed that most zircon grains consist of a Cambro-Ordovician rim growth over a pre-Cambrian core (see Fig. 4.1 in Chapter 4) and obtained a concordant crystallization age of 483 ± 3 Ma for the Miranda do Douro orthogneiss. They also detected abundant inherited concordant or sub-concordant populations at 605 Ma and ca. 3.2 Ga. Zeck et al. (2007) also dated the Miranda do Douro orthogneiss using U-Pb ion microprobe and found a slightly older crystallization age of 496 ± 3 Ma and observed the presence of inherited cores with ages between c. 2700 - 550 Ma from assimilated country rock, or more probably, from anatectic source rock which was melted to produce the granitic magma that intruded to give parent rocks of the gneiss formation. Within these inherited cores, three prevailing inherited age components are 628, 580 ± 9 Ma, and c. 557 Ma.

Díez Montes (2006) dated the augen orthogneiss and the ignimbrites of the Ollo de Sapo Formation from the Sanabria region using U-Pb analysis in zircon concentrates. As

expected, the results were highly discordant. This author obtained: (i) two discordias for the augen orthogneiss which had the same lower intersection, at 472 ± 14 Ma, and different upper intersection, at 2.03 Ga and 1.53 Ga, and (ii) two discordias for the ignimbrites, one had a lower intersection at 66 ± 75 Ma and an upper intersection at 488 ± 6 Ma and the other had a lower intersection at 488 ± 6 Ma and an upper intersection at 2.14 Ga.

Recently, Montero et al. (2007) have studied the metavolcanic rocks of the Ollo de Sapo Formation in Villadepera, near the Miranda do Douro metagranite, and in Hiendelaencina, as well as the spatially-related Antoñita metagranite, using spot U-Pb ion microprobe, LA-ICPMS and Pb-Pb stepwise analysis. The results are as follows: (i) the 2000 m thick meta-volcanic rocks of the Ollo de Sapo Formation in the Hiendelaencina region show a noticeable difference in the crystallization age between top, 483 ± 3 Ma, and base, 495 ± 5 Ma. They, however, contain similar inherited populations at 603 Ma (by far the most abundant), 650-700 Ma, 850-900 Ma and ca. 2000 Ma, and (ii) the spatially-related Antoñita metagranite is younger still, with a well-defined Pb-Pb plateau-age of 474 ± 4 Ma, which is considered to be the age of crystallization.

Regarding, isotope geology, the lack of precise crystallization ages has caused the Sr and Nd isotope geology of the Ollo de Sapo materials to be poorly known.

Ortega et al. (2000) using a crystallization age of 490 Ma, calculated a $(^{87}\text{Sr}/^{86}\text{Sr})_{\text{init}}$ between 0.7084 and 0.7122 and a ϵNd between -2.0 and -2.8 for homogeneous augen-gneisses of the northern and central sectors of the Ollo de Sapo Formation. They observed that samples with the best preserved igneous features presented the lowest values of $(^{87}\text{Sr}/^{86}\text{Sr})_{\text{init}}$ and concluded that the Ollo de Sapo rocks could be late-orogenic magmatism.

Fernández-Suárez et al. (1999) studied the Nd isotopes of the volcanosedimentary rocks of the northwest sector of the Ollo de Sapo Formation and noticed that the Ollo de Sapo rocks had a narrow range of T_{DM} ages in spite of the multi-age zircon population they contained. They concluded that this reflected a relatively homogeneous mixture of crustal components (Icartian, Grenville, Cadomian and a c. 440-460 Ma component) over a large area which was consistent with the proposed proximal nature of the crustal sources.

Montero et al. (2007) studied the Sr and Nd isotope geology of the Ollo de Sapo Formation in Villadepera and Hiendelaencina, and found important differences in Sr isotopes: whereas at Villadepera $(^{87}\text{Sr}/^{86}\text{Sr})_{\text{init}}$ is always lower than 0.7075, in Hiendelaencina this

ratio is always higher than 0.7098. Nd isotopes do not show meaningful regional variations, but there are differences between the Nd isotopes of the metavolcanic rocks and the metagranites, especially in Villadepera. In Hiendelaencina, the $(^{87}\text{Sr}/^{86}\text{Sr})_{\text{init}}$ are 0.71039 and 0.70985 for the metavolcanic rocks, and 0.71054, 0.71105 and 0.71168 for the metagranite. In Villadepera, the $(^{87}\text{Sr}/^{86}\text{Sr})_{\text{init}}$ are 0.70700 and 0.70755 for the metavolcanic rocks, and 0.70442 and 0.70521 for the metagranite. The $\text{Nd}_{(\text{CHUR})}$ model age for the Miranda do Douro metagranite, which has the most primitive isotope and chemical composition, is 0.69 Ga, slightly older than the zircon age of the source, whereas for the Villadepera metavolcanic rocks the $\text{Nd}_{(\text{CHUR})}$ model age is 1.0 Ga. In Hiendelaencina, both rock types yielded the same value of 0.94 Ga.

1.3.2. The pre-Middle Ordovician metagranites of the Schist-Graywacke Complex Domain

1.3.2.1. General geology and composition

The metagranites of the Schist-Graywacke Complex Domain were extensively studied in the 1970s, when their analogy with the materials of the Ollo de Sapo was established by Bard et al. (1970). These authors suggested that some augen-gneisses, for example the Santa María de la Alameda outcrop, might be the source of the Ollo de Sapo Formation. Fernández Casals and Capote (1971) studied the petrology of the augen-gneisses of Eastern Guadarrama and proposed they are part of a thick formation made of augen-gneisses with intercalated banded gneisses and granoblastic rocks at the base and micro-augen gneisses at the top, similar therefore to the distribution of coarse-grained and fine-grained facies in the Ollo de Sapo. For this reason, they concluded that this formation could be the Ollo de Sapo intensely migmatized. Additionally, they placed this formation below the Cambrian metamorphic series and assigned it an Upper pre-Cambrian age. Fernández Casals (1974) investigated the gneissic formation of La Morcuera and differentiated: (i) a basal complex of augen-gneisses which represented orthogneisses derived from porphyritic granites with enclaves and aplitic dikes, (ii) a thinner succession of several types of gneisses derived from volcanic and sedimentary rocks. Finally, this author concluded that the gneissic formation of La Morcuera was beneath the Cambrian sedimentary formation and was pre-Cambrian. García de Figuerola and Franco González (1975) mapped and studied the petrology of the metamorphic terranes cropping out to the east of Guijuelo (Salamanca province). In this area, the Bercimuelle gneiss, the origin of which was unclear, crops out. They suggested that it either represents an older Variscan granodiorite intruded in the Mo-

rille series before the second phase of deformation or was derived from a volcanic formation interlayered within the metasediments. Capote et al. (1977) studied the pre-Cambrian series of the Montes de Toledo and the Central System. In the Montes de Toledo, they distinguished a lower formation of shales and graywackes and an upper turbidite formation composed of metapelites. In the Central System, they also differentiated two main formations: (i) the first dominantly composed of augen-gneisses, and (ii) the other mostly composed of metasediments. Navidad (1979) carried out a study of the augen series of western and central Guadarrama. This author differentiated two different facies of augen-gneisses: (i) Homogenous augen-gneisses and (ii) Heterogeneous augen-gneisses, and concluded that their petrological and compositional features suggested that they were derived from a vulcano-sedimentary series.

Navidad and López Ramos (1981) studied the augen-gneisses of El Vellón and Pedrezuela massifs located in the Sierra de Guadarrama. The Vellón massif was composed of three facies: (i) feldspathic gneisses with megacrysts, which represent the most internal facies, (ii) biotitic orthogneisses, which represent the intermediate facies, and (iii) feldspathic micro-augen-gneisses, which represent the most external facies. The Pedrezuela massif consists of two-mica augen-gneisses and different types of leucogneisses. They inferred that both massifs were ancient metagranitic massifs which had different ages, the Pedrezuela massif being younger than El Vellón massif or, at least, younger than the intermediate facies of El Vellón massif. In the same year, Peinado and Álvaro (1981) studied the gneisses of the metamorphic area of “El Escorial” (Central System) and they differentiated two types of gneisses: the biotitic orthogneisses of the Santa Maria de la Alameda and the Abantos orthogneisses. The former were interpreted as I-type porphyritic biotite granites with a homogenous composition and quartz-dioritic enclaves. These granites formed as low volume stratiform bodies intruded in a relatively cold host rock producing an aureole of contact metamorphism that was subsequently affected by the second phase of Variscan deformation. The Abantos orthogneisses, by contrast, represent S-type two-mica granites characterized by the presence of aluminous silicates and abundant associated leucogranites, aplites and pegmatites. These were more voluminous and intruded into the Upper pre-Cambrian-Lower Cambrian regional metasedimentary series at different levels. Saavedra et al. (1984) studied the petrography and the geochemistry of the granitoids and metamorphic rocks of the Morille-Martinamor area (Salamanca province). One of these metamorphic

rocks was the San Pelayo gneiss whose composition is intermediate between the composition of granitoids and the non-calcic metamorphic rocks.

Finally, Bea et al. (1990) described the migmatization processes that occurred in the Almohalla Formation (Anatectic Complex of Pena Negra, Ávila Batholith). This formation is composed of three lithotypes: (i) a pre-Variscan, orthoderived, coarse-grained augen-gneiss, (ii) ophthalmitic and diktynonitic migmatites developed from the orthogneiss, (iii) a garnet-bearing leucogranite of Variscan age, which crops out in the orthogneiss as a small, irregular, subautochthonous body with transitional contacts with the host. Bea et al. (1990) concluded that the two processes of partial fusion that occurred in this formation were very different although they had been generated from the same material.

1.3.2.2. Stratigraphic position

The stratigraphy of the pre-Middle Ordovician metagranites of the Schist-Graywacke Complex Domain was studied for the first time by Capote and Fernández Casals (1975). They identified several pre-Ordovician gneissic formations which have different age and lithology and concluded that : (i) the lower pre-Ordovician Formation was an augen-gneiss with xenoliths (La Morcuera) and could represent a metamorphosed and deformed porphyritic granite or could have derive from a granitic pluton which intruded the surrounding meta-sediments, (ii) a series which overlays the Morcuera 1 and from bottom to top consisted of: (a) pelitic and psammitic metasediments with amphibolitic and calcareous intercalations, in whose base there were rocks with volcanic participation, (b) a lithologic group which came from acid lavas and sediments related to these lavas, (the Hiendelaencina gneiss could be included in this group), and (c) pelitic-psammitic group with salic-silicate rocks and without limestones. Furthermore, the lower pre-Ordovician Formation in Salamanca are below Lower Cambrian formations and for that reason, these authors considered them to be pre-Cambrian formations, although pre-Ordovician series in the Ávila province could be considered to be Lower Cambrian formations.

1.3.2.3. Structural geology

Concerning the structural geology of the pre-Middle Ordovician metagranites of the Schist-Graywacke Complex Domain, one of the most important papers is that of Macaya et al. (1991). These authors described the deformational structures generated in the cover and the basement of the Sierra de Guadarrama through the Variscan orogeny. The gneisses of

the Schist-Graywacke Complex Domain reveal structures attributable to the first and third Variscan deformation phases but they do not display pre-Variscan structures, perhaps because they were obliterated by the intense deformation and high grade metamorphism developed during the Variscan orogeny. Escuder Viruete et al. (1994) studied the tectono-thermic evolution of the Tormes Dome in which two units were distinguished: (i) The Lower Unit which consisted of glandular orthogneisses, ortho-derived leucogneisses and aluminous paragneisses with lenses of marble and calc-silicate rocks, and (ii) the Upper Unit which was a monotonous sequence of slates and quartzose schists. According to these authors, two events could be distinguished in the tectono-thermic evolution of the Tormes Dome. The first event is well represented by a clockwise segment of the P-T path, characteristic of the thickening and stacking of crustal units during the continental collision. The second event was related to the gravitational collapse of the crust, giving way to a short stage of prograde decompression and, then, to a nearly isothermal decompression. The uplift of deep and hot rocks was accompanied by the development of anatexis processes and controlled by extensional ductile deformation which gave rise to their juxtaposition with an Upper Unit that underwent nearly isobaric heating. Following this, both units, as a whole, already at a high crustal level, experienced cooling and a weak decompression.

1.3.2.4. Radiometric dating and isotope geology

The first geochronological research about the pre-Middle Ordovician metagranites of the Schist-Graywacke Complex Domain were by Priem et al. (1970) who dated the Variscan and pre-Variscan plutonic rocks of the western Iberian Peninsula using the Rb-Sr method. They obtained Rb-Sr ages between 460 and 430 Ma for pre-Variscan rocks and concluded that they were Upper Ordovician. Bischoff et al. (1978) undertook a geochemical and geochronological study of the metarhyolites and orthogneisses of eastern Guadarrama. The results confirmed the existence of a thermal event in the transition Carboniferous-Permian period but did not identify a crystallization age for the gneisses due to the homogenization of the Rb-Sr system.

In the 1980s, several important investigations were carried out on the pre-Middle Ordovician metagranites of the Schist-Graywacke Complex Domain. Linares et al. (1987) dated the San Pelayo orthogneiss using the Rb-Sr method and obtained an age of 430 ± 30 Ma that they considered a minimum age of metamorphism, probably produced by the intrusion of the neighboring Variscan granites. Vialette et al. (1987) dated some gneisses from the

Sierra de Guadarrama using the Rb-Sr method. The results were as follows: (i) a crystallization age of 471 ± 12 Ma and a $(^{87}\text{Sr}/^{86}\text{Sr})_{\text{init}}$ of 0.7093 ± 0.010 for the El Villar de Pradena gneiss, (ii) a crystallization age of 474 ± 7 Ma and a $(^{87}\text{Sr}/^{86}\text{Sr})_{\text{init}}$ of 0.7070 ± 0.0017 for the Abantos gneiss, and (iii) a crystallization age of 494 ± 10 Ma and a $(^{87}\text{Sr}/^{86}\text{Sr})_{\text{init}}$ of 0.7087 ± 0.0011 for the Otero de Herreros gneiss. And finally, Wildberg et al. (1989) tried to reconstruct the pre-Variscan and Variscan histories of the Sierra de Guadarrama using U-Pb in zircon concentrates from the El Cardoso and Madarquillos gneisses. These authors obtained a discordia with an upper intersection at approximately 2435 Ma and a lower intersection at $540 +31/-35$ Ma for the El Cardoso gneiss, and a discordia with an upper intersection at 1970 Ma and a lower intersection at 370 ± 8 Ma for the Madarquillos gneiss.

Following this, Valverde-Vaquero and Dunning (2000) dated the El Cardoso, Riaza, Buitrago and La Morcuera gneisses of the Sierra de Guadarrama using U-Pb in zircon concentrates. The results were as follows: (i) a discordia with an upper intersection at 2.64 Ga and a lower intersection at 480 ± 2 Ma for El Cardoso gneiss, (ii) a discordia with an upper intersection at $468 +16/-8$ Ma for the Riaza gneiss, (iii) a discordia with an upper intersection at $488 +10/-8$ Ma for the Buitrago gneiss, and finally, (iv) a discordia with an upper intersection at 477 ± 4 Ma for La Morcuera gneiss.

The most recently published geochronological data for this area are for the “La Almo-halla” gneiss located in the Sierra de Gredos (Central System). The first study was by Bea et al. (2003) which obtained a stepwise evaporation Pb-Pb age of 543 ± 6 Ma, the second study was carried out by Zeck et al. (2004) and obtained an ion-microprobe U-Pb age of 546 ± 3 Ma.

2. SAMPLES AND METHODS

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2.1. Samples

For this work we collected a total of seventy four samples of pre-Middle Ordovician magmatic rocks from the Central Iberian Zone. The number of samples suitable for chemical and isotopic studies was severely limited by the scarcity of fresh-rock outcrops, despite the fact that we used a portable drilling machine wherever the alteration was

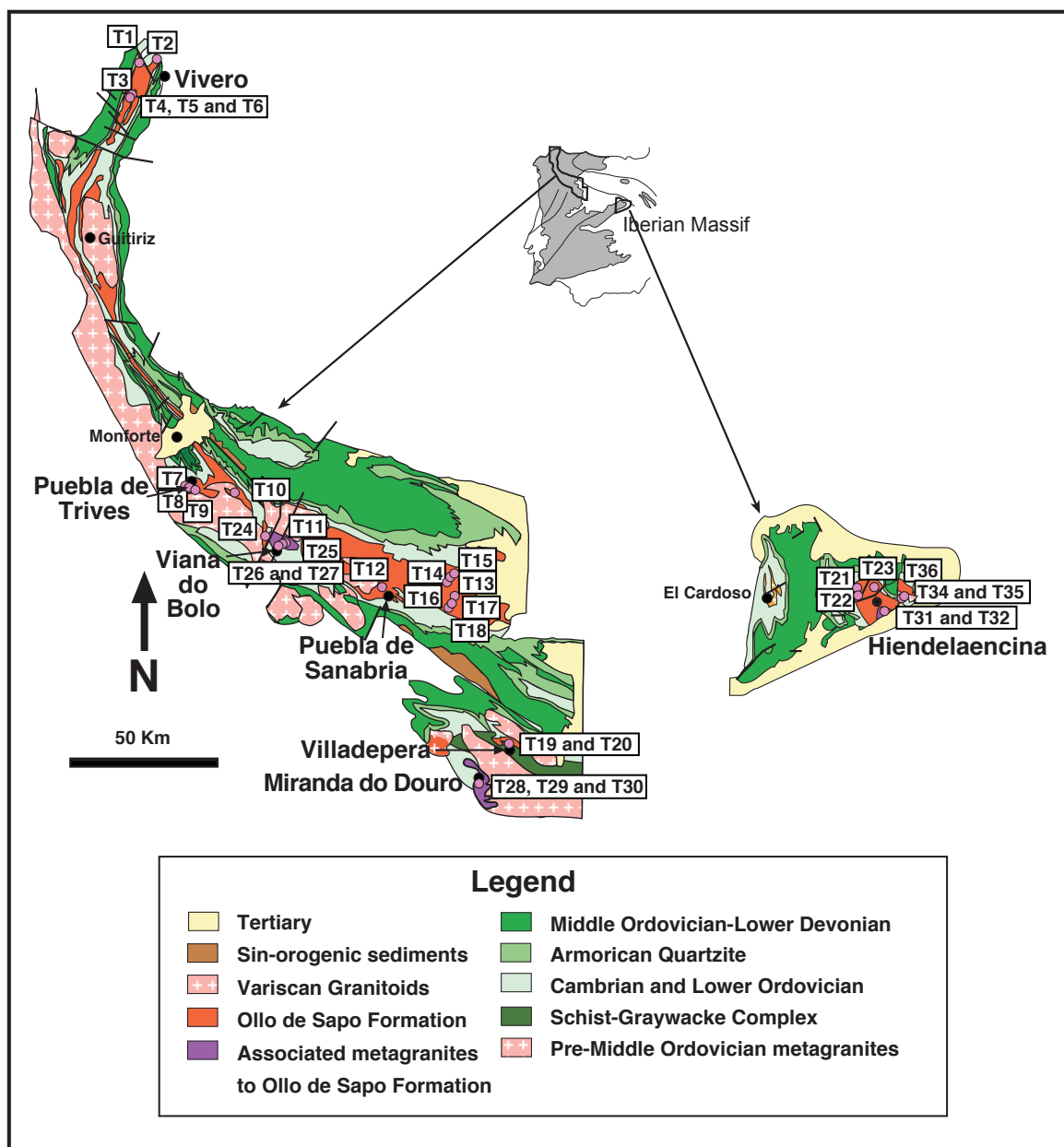


Figure 2.1: Geological scheme of the Ollo de Sapo Domain (Martínez Catalán et al., 2004) in which sampled metavolcanic and metagranitic rocks are located and represented by letter T.

particularly intense. The samples were collected from two main areas: the Ollo de Sapo Domain and the northern part of the Schist-Graywacke Complex Domain. Sixty three samples were analyzed for major and trace elements, including 14 REE, Th and U. A subset of forty two samples was also analyzed for Sr and Nd isotopes. Zircon was separated from thirty five samples, and analyzed using single-grain stepwise Pb-Pb evaporation, and U-Pb ion microprobe and LA-ICPMS methods.

We sampled the metavolcanic rocks of the Ollo de Sapo Formation in five selected areas that represent the whole domain: Vivero-Mañón, Trives, Sanabria, Villadepera and Hiendelaencina (Fig. 2.1). We also sampled four spatially associated metagranites: the Antoñita metagranite in Hiendelaencina, the Miranda do Douro metagranite in Villadepera, and the Viana do Bolo and San Sebastián metagranites in Sanabria. The location and relevant field information about these samples is shown in Table 2.1.

Table 2.1: Samples of the Ollo de Sapo Domain. The letter between parentheses in the column “Rock type” is the sample reference. The coordinates are geographic coordinates.

Locality	Rock type	Coordinates	Description
Vivero	Fine grained Ollo de Sapo (T1)	N 43°43'53.2" W 7°41'56.9"	Deformed gneiss with blue quartz and feldspar phenocrysts. Phenocrysts with pressure shadows Fine-grained matrix. It is located in the upper part of the Ollo de Sapo Formation.
	Coarse-grained Ollo de Sapo (T2).	N 43°44'45.7" W 7°38'55.6"	Deformed gneiss with K-feldspar megacrysts and blue quartz phenocrysts. It is located in the lower part of the Ollo de Sapo Formation.
Mañón	Fine-grained Ollo de Sapo (T3)	N 43°41'24.5" W 7°43'06.7"	Gneiss without feldspar megacrysts. It is located in the upper part of the Ollo de Sapo Formation.
	Coarse-grained Ollo de Sapo (T4, T5, T6)	N 43°41'12.5" W 7°43'21.5"	Intensely deformed gneiss with feldspar megacrysts and blue quartz phenocrysts. It is located in the lower part of the Ollo de Sapo Formation.
Puebla de Trives	Fine grained Ollo de Sapo (T7)	N 42°23'04.0" W 7°27'41.6"	Very deformed gneiss with enclaves and deformed dikes. Blue quartz phenocrysts. Very micaceous matrix. Without feldspar megacrysts. It is located in the upper part of the Ollo de Sapo Formation.
	Coarse-grained Ollo de Sapo (T8, T9, T10)	N 42°23'06.9" W 7°29'30.6" N 42°22'40.0" W 7°26'21.6" N 42°19'50.0" W 7°13'05.7"	Variably deformed gneiss. Feldspar megacrysts. It is located in the lower part of the Ollo de Sapo Formation.

Locality	Rock type	Coordinates	Description
Sanabria	Coarse-grained Ollo de Sapo (T11)	N 42°09'25.9" W 6°54'38.4"	Very deformed gneiss with enclaves and deformed dikes. Feldspar megacrysts. It is located in the lower part of the Ollo de Sapo Formation.
	Fine-grained Ollo de Sapo (T12)	N 42°07'50.3" W 6°38'45.3"	Very deformed gneiss with feldspar and quartz phenocrysts. It is located in the middle part of the Ollo de Sapo Formation.
	Mixed Facies (T13, T14, T15, T16, T17, T18)	N 42°06'04.9" W 6°24'33.4"	Coarse-grained facies. Gneiss with feldspar megacrysts, quartz phenocrysts and black fragments.
		N 42°06'42.1" W 6°24'03.3" N 42°06'59.8" W 6°22'30.6" N 42°03'40.7" W 6°19'25.7" N 42°01'25.7" W 6°20'39.3"	Fine-grained facies. Gneiss with quartz phenocrysts and without feldspar megacrysts. It is located in the upper part of the Ollo de Sapo Formation.
Villadepera	Fine-grained Ollo de Sapo (T19, T20)	N 41°33'37.3" W 6°07'51.6"	Migmatized, extremely deformed gneiss. Cut by dikes and pegmatoid stocks.
Hiendelaencina	Coarse-grained Ollo de Sapo (T21)	N 41°08'20.6" W 3°03'27.3"	Gneiss with feldspar megacrysts and blue quartz phenocrysts. It is located in the lower part of the Ollo de Sapo Formation.
	Fine grained Ollo de Sapo (T22)	N 41°06'28.1" W 3°04'15.4"	Gneiss with blue quartz phenocrysts and without feldspar megacrysts. More micaceous than the coarse grained Ollo de Sapo. It is located in the middle part of the Ollo de Sapo Formation.
	Coarse-grained Ollo de Sapo (T23)	N 41°05'06.0" W 2°59'53.1"	Gneiss with feldspar megacrysts and blue quartz phenocrysts. It is located in the upper part of the Ollo de Sapo Formation.
Viana do Bolo	Viana do Bolo metagranite (T24, T25)	N 42°08'34.6" W 6°56'58.8" N 42°13'52.1" W 7°06'28.8"	Very deformed and migmatized medium-coarse-grained gneiss without feldspar megacrysts.
Porto	San Sebastián metagranite (T26, T27)	N 42°07'54.9" W 6°59'06.9"	Deformed fine-grained gneiss without feldspar megacrysts.
Miranda do Douro	Miranda do Douro metagranite (T28, T29, T30)	N 41°29'29.1" W 6°16'24.6"	Biotitic mesocratic medium-fine grained gneiss with feldspar phenocrysts.
Hiendelaencina	Antoñita metagranite (T31, T32)	N 41°03'24.07" W 2°59'10.0"	Deformed coarse-grained gneiss. Feldspar phenocrysts with SC structures. Associated aplitic dikes.
La Bodera	Antoñita metagranite (T33, T34, T35)	N 41°07'13.9"/ W 41°07'13.9" N 41°07'18.9" W 2°54'23.4"	Deformed gneiss with feldspar phenocrysts. Phenocrysts with pressure shadows.

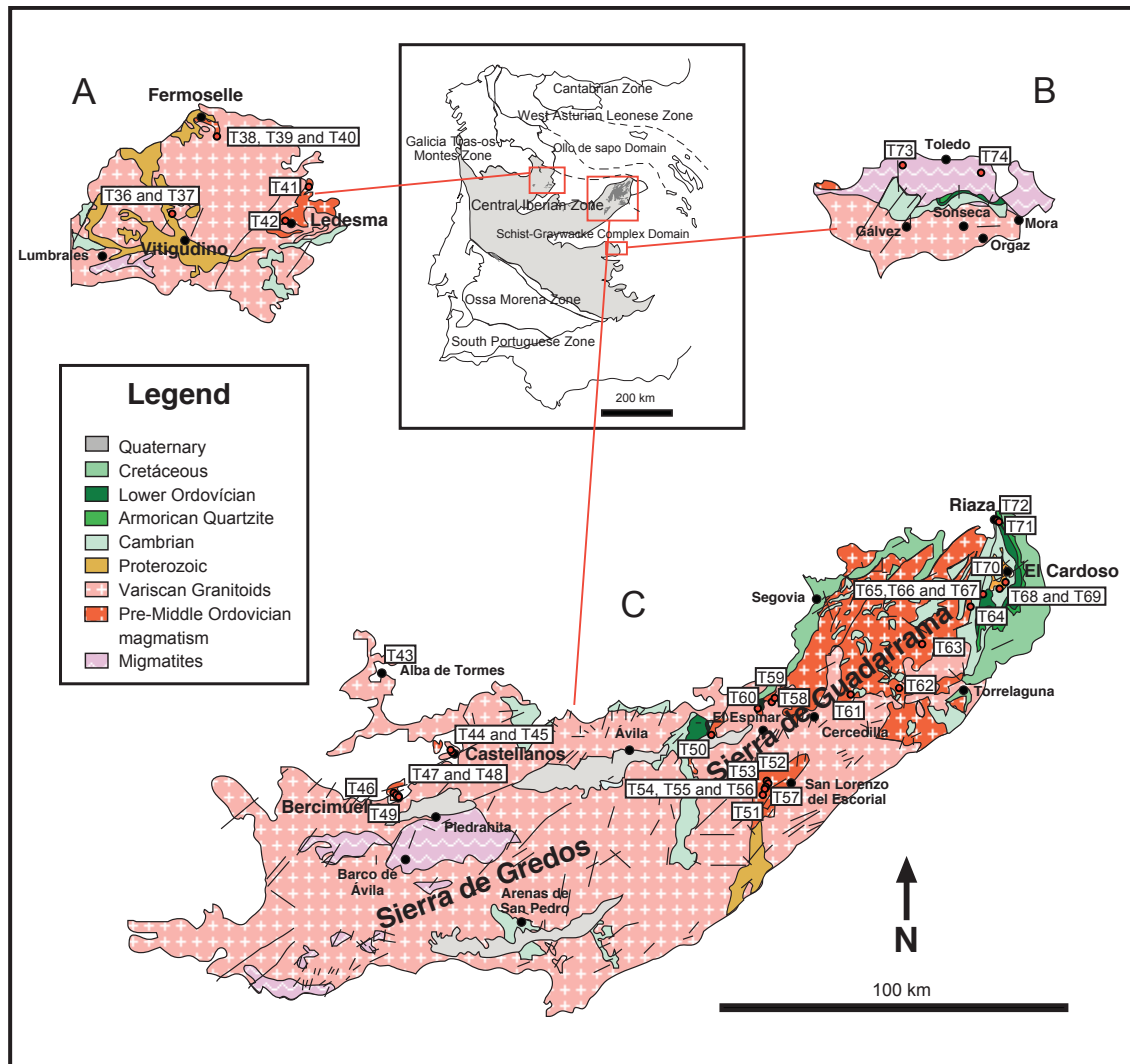


Figure 2.2: Geological scheme of the outcrops of the pre-Middle Ordovician metagranites and metavolcanic rocks of the north of the Schist-Graywacke Complex Domain. Sampled metagranites are represented by letter T. The geological scheme represents sections of the geological map of Spain (Álvaro et al., 1994).

We also sampled the pre-Middle Ordovician metagranites and metavolcanic rocks in four areas of the northern part of the Schist-Graywacke Complex Domain (Fig. 2.2):

- In the Tormes Dome: the Fermoselle, Vitigudino and Ledesma metagranites.
- In the northwest part of the Sierra de Gredos: the San Pelayo, Bercimuelle and Castellanos metagranites.

- In the Sierra de Guadarrama: La Cañada, La Hoya, La Estación I, La Estación II, El Caloco, Vegas de Matute, La Morcuera and Buitrago de Lozoya metagranites, and La Berzosa, El Cardoso and Rianza metavolcanic rocks.
- In the Anatectic Complex of Toledo: the Polán and Mohares metagranites.

The location and relevant field information for the pre-Middle Ordovician metagranites is shown in the Table 2.2.

Table 2.2: Samples of the metagranites and metavolcanic rocks from the north of the Schist-Graywacke Complex Domain. The letter between parentheses in the column “Rock type” is the sample reference. The coordinates are geographic coordinates.

Locality	Rock type	Coordinates	Description
Barceño	Vitigudino metagranite (T36, T37)	N 41°05'01.9" W 6°27'14.5"	Slightly migmatized medium-grained gneiss with few feldspar megacrysts.
Sardón de los Frailes	Fermoselle metagranite (T38, T39, T40)	N 41°14'44.7" W 6°16'22.1" N 41°14'23" W 6°16'32"	Slightly migmatized and deformed medium-coarse grained gneiss with few feldspar megacrysts.
Almeida	Ledesma metagranite (T41)	N 41°19'16.6" W 5°53'14.9"	Deformed and migmatized biotitic gneiss with feldspar megacrysts.
Ledesma	Ledesma metagranite (T42)	N 41°06'15.6" W 6°01'28.0"	Biotitic gneiss with feldspar megacrysts.
Martinamor	San Pelayo metagranite (T43)	N 40°47'36.6" W 5°34'32.1"	Deformed coarse-grained gneiss. Cut by late pegmatoid dikes.
Diegoálvaro	Castellanos metagranite (T44, T45)	N 40°38'48.9" W 5°18'43.0" N 40°39'02.4" W 5°18'46.7"	Deformed biotitic fine-grained gneiss with enclaves.
Bercimuelle	Bercimuelle metagranite (T46, T47, T48, T49)	N 40°31'34.3" W 5°30'30.5" N 40°31'26.6" W 5°30'05.3" N 40°31'25.9" W 5°29'47.8"	Deformed biotitic augen-gneiss.
La Cañada	La Cañada metagranite (T50)	N 40°42'06.4" W 4°26'50.4"	Biotitic augen-gneiss with meta-sedimentary enclaves. Cut by aplitic and pegmatoid dikes.
Navas del Marqués	Navas del Marqués metagranite (T51)	N 40°33'14.4" W 4°16'14.4"	Slightly migmatized augen-gneiss.
Santa María de la Alameda	Santa María de la Alameda metagranite (T52)	N 40°35'24.0" W 4°15'34.2"	Deformed and slightly migmatized biotitic augen-gneiss with enclaves.
Santa María de la Alameda	La Hoya metagranite (T53)	N 40°34'34.7" W 4°15'30.0"	Very migmatized biotitic augen-gneiss.

Samples and methods

Locality	Rock type	Coordinates	Description
La Estación	La Estación I metagranite (T54, T55, T56)	N 40°34'30.6" W 4°15'53.3"	Very deformed biotitic augen-gneiss.
La Estación	La Estación II metagranite (T57)	N 40°34'20.1" W 4°15'55.8"	Deformed biotitic augen-gneiss. Cut by late aplitic dikes.
Vegas de Matute	Vegas de Matute metagranite (T58, T59)	N 40°47'56.1" W 4°14'29.1" N 40°47'34.5" W 4°15'01.3"	Biotitic mesocratic augen-gneiss with calco-silicate enclaves and garnet.
Vegas de Matute	El Caloco metagranite (T60)	N 40°47'21.0" W 4°17'24.9"	Fine-grained leucogneiss.
La Morcuera	La Morcuera metagranite (T61, T62)	N 40°48'21.5" W 3°58'11.9" N 40°49'41.3" W 3°49'49.1"	Deformed biotitic augen-gneiss. Exhibits rapakivi texture and pressure shadows.
Buitrago de Lozoya	Buitrago metagranite (T63)	N 40°56'37.4" W 3°43'03.1"	Slightly migmatized biotitic medium-grained gneiss with few enclaves. Cut by a few leucocratic stocks.
Prádena del Rincón	Prádena del Rincón metagranite (T64)	N 41°02'03.9" W 3°33'07.1"	Slightly migmatized medium-grained leucogneiss with biotite and garnet. Cut by pegmatoid stocks and pygmatic veins
Montejo de la Sierra	La Berzosa metavolcanic rock (T65, T66, T67)	N 41°03'41.7" W 3°30'25.0" N 41°03'41.7" W 3°30'22.0"	Deformed micaceous augen-gneiss
El Cardoso de la Sierra	El Cardoso metavolcanic rock (T68, T69, T70)	N 41°04'28.2" W 3°27'29.3" N 41°05'04.9" W 3°25'45.6"	Deformed biotitic fine-grained gneiss
Riaza	Fine-grained metavolcanic rocks (T71) Coarse-grained metavolcanic rocks (T72)	N 41°16'49.8" W 3°28'13.8" N 41°16'05.6" W 3°27'45.2"	Deformed two-mica gneiss. Deformed augen gneiss with a few leucocratic dikes parallel to the foliation.
Guajaraz	Polán metagranite (T73)	N 39°51'38.9" W 4°09'07.7"	Deformed micaceous augen-gneiss.
Toledo	Mohares metagranite (T74)	N 39°50'05.0" W 3°53'12.2"	Deformed micaceous augen-gneiss.

2.2. Analytical methods

2.2.1. Whole-rock major and trace element analyses

Major and trace elements of sixty three samples were analyzed at the Centro de Instrumentación Científica (CIC) of the University of Granada. Major-element and Zr determinations were performed by X-ray fluorescence after fusion with lithium tetraborate.

Typical precision was better than $\pm 1.5\%$ for an analyte concentration of 10 wt.%, and $\pm 5\%$ for 100 ppm Zr. Trace elements, except Zr, were determined by ICP-mass spectrometry (ICP-MS) after $\text{HNO}_3 + \text{HF}$ digestion in a Teflon-lined vessel, evaporation to dryness, and subsequent dissolution in 100 ml of 4 vol.% HNO_3 . Instrument measurements were carried out in triplicate with a PE SCIEX ELAN-5000A spectrometer using Rh as internal standard. Precision was better than $\pm 5\%$ for analytical concentrations of 10 ppm. As standards we used the natural rocks UBN, WSE, PME, BCR, BRN, DRN, GSP, GA and GH with the values recommended by Govindaraju (1994) and Govindaraju et al. (1994).

Table 2.3: Values of the standards used in a routine run from the Granada laboratory. The subscript M and C represent the measured and the certificated values respectively.

Elements	PMS _M	PMS _C	WSE _M	WSE _C	UBN _M	UBN _C	BEN _M	BEN _C	BR _M	BR _C	AGV _M	AGV _C
Li	7.5	7.3	13.5	13.6	27	27	12.4	13	14.2	13	11	12
Rb	0.9	1	25	25	3.3	4	47.4	47	46.9	47	68.2	67.3
Cs	0.5	0.4	0.6	0.5	10.1	10	0.8	0.8	0.7	0.8	1.3	1.3
Be	0.2	0.5	0.9	1.1	0.1	0.2	1.9	1.9	1.6	1.5	2.2	2.1
Sr	271	280	405	410	7	9	1370	1370	1323	1320	671	662
Ba	142	148	330	338	25	27	1001	1025	1054	1050	1213	1226
Sc	34.3	34	26.9	28	13.3	13	22.3	22	22.1	25	12	12.2
V	192	192	335	340	65	75	235	235	239	235	125	121
Cr	320	314	108	99	2267	2300	365	360	369	380	27	10
Co	49	49	44	44	100	100	61	60	58	52	15	15
Ni	122	115	56	55	1983	2000	264	267	260	260	17	16
Cu	61	59	69	65	25	28	73	72	72	72	58	60
Zn	69	60	129	117	81	85	133	133	171	160	95	88
Ga	16	16	22	23	2	3	18	17	18	19	21	20
Y	11	11	31	30	3	3	28	30	28	30	19	20
Nb	2.1	2.6	15.9	18	0.1	0.1	103.1	105	102.8	98	13	15
Ta	0.3	0.2	1.3	1.2	0.1	0	5.7	6.2	6.1	6.2	0.9	0.9
Zr	37	39	194	195	4	4	259	260	261	260	222	227
Hf	1.07	1.12	5.19	5.3	0.22	0.1	5.87	5.6	5.91	5.6	5.3	5.1
Mo	1.86	1.9	3.74	3.7	0.25	0.55	2.78	2.8	2.45	2.4	2.29	2.7
Sn	2.97	3	17.93	18	0.3	0	1.89	2	2.27	2	4.31	4.2
Tl	0.04	0.04	0.13	0.16	0.03	0.06	0.02	0.04	0.05	0.05	0.29	0.34
Pb	2.06	2.5	13.52	13.8	12.53	13	4	4	4.67	5	36.67	36
U	0.014	0.03	0.631	0.65	0.071	0.07	2.371	2.4	2.438	2.5	1.963	1.92
Th	0.043	0.05	3.039	3	0.103	0.07	10.946	10.4	10.958	11	6.792	6.5
La	2.737	2.8	26.204	27	0.37	0.35	81.147	82	81.265	82	38.823	38
Ce	6.704	6.8	59.384	61	0.818	0.8	148.724	152	148.5	151	68.292	67
Pr	1.03	1.08	7.668	7.8	0.123	0.12	16.844	17.5	17.109	17	8.307	7.6
Nd	5.342	5.5	32.082	33	0.608	0.6	63.92	67	66.363	65	33.05	33
Sm	1.749	1.75	8.461	8.8	0.192	0.2	12.029	12.2	12.106	12.2	5.995	5.9
Eu	1.019	1.07	2.106	2.25	0.069	0.08	3.657	3.6	3.668	3.7	1.647	1.64
Gd	2.053	2	7.059	7.2	0.333	0.3	9.752	9.7	9.831	9.5	4.757	5
Tb	0.342	0.36	1.1	1.1	0.065	0.06	1.396	1.3	1.399	1.25	0.736	0.7
Dy	1.93	2	5.915	6	0.386	0.38	6.141	6.4	6.029	6.4	3.543	3.6
Ho	0.409	0.42	1.186	1.2	0.085	0.09	1.075	1.1	1.079	1.1	0.71	0.67
Er	1.085	1.1	2.865	3	0.269	0.28	2.277	2.5	2.415	2.5	1.724	1.7
Tm	0.165	0.17	0.405	0.43	0.037	0.045	0.295	0.34	0.315	0.34	0.239	0.34
Yb	0.926	1	2.441	2.5	0.275	0.28	1.722	1.8	1.716	1.8	1.572	1.72
Lu	0.145	0.15	0.356	0.37	0.043	0.045	0.233	0.24	0.24	0.25	0.251	0.27

The accuracy of the analytical results can be estimated from the comparison of the results of analyzing these standards as samples in a routine run, and their certified values (Table 2.3)

2.2.2. Sr and Nd isotope analyses

Sr and Nd isotopes of forty seven samples were analyzed at the CIC of the University of Granada. These samples (0.1000 g) were digested with HNO₃ + HF in a Teflon-lined vessel and analyzed by thermal ionization mass spectrometry (TIMS) in a Finnigan Mat 262 RPQ spectrometer after separation with ion-exchange resins using conventional procedures. All reagents were ultra clean. Normalization values were $^{86}\text{Sr}/^{88}\text{Sr} = 0.1194$ and $^{146}\text{Nd}/^{144}\text{Nd} = 0.7219$. Blanks were 0.6 and 0.09 nanograms for Sr and Nd. The external precision (2σ), estimated by analyzing 10 replicates of the standard WS-E (Govindaraju et al., 1994), was better than $\pm 0.003\%$ for $^{87}\text{Sr}/^{86}\text{Sr}$ and $\pm 0.0015\%$ for $^{143}\text{Nd}/^{144}\text{Nd}$. $^{87}\text{Rb}/^{86}\text{Sr}$ and $^{147}\text{Sm}/^{144}\text{Nd}$ were directly determined by ICP-MS at Granada following the method developed by Montero and Bea (1998), with a precision better than $\pm 1.2\%$ and $\pm 0.9\%$ (2σ) respectively.

2.2.3. Zircon separation and dating

Zircons of thirty five samples were separated using conventional magnetic and heavy-liquid techniques. Once mounted and polished, zircon grains were studied by cathodoluminescence imaging using a ZEISS DSM 950 and LEO 1430-VP scanning electronic microscopes (SEM) at the CIC of the University of Granada. Subsequently they were analyzed:

- For U-Th-Pb, using the Cameca IMS1270 ion microprobe of the NORDSIM facility in Stockholm and the SHRIMP II of the Australian Geoscience of Canberra.
- For U-Th-Pb and for Ti and Zr/Hf using the LA-ICPMS of the CIC of the University of Granada.
- For Pb-Pb using the stepwise evaporation method at the TIMS facility of the CIC of the University of Granada.

2.2.3.1. Zircon U-Pb ion microprobe analyses

Ion microprobe analyses were done in a CAMECA IM-1270 instrument at the NORDSIM facility (Stockholm) and in a SHRIMP II at the Australian National University (Canberra). Analytical methods broadly follow those described by Whitehouse et al. (1999) and references mentioned therein. U/Pb and Th/Pb ratios were calibrated using the Geostandards 91500 reference zircon (1065 Ma; Wiedenbeck et al., 1995) and include a propagated error component from replicate analyses of 91500 during the analytical session. Errors on $^{207}\text{Pb}/^{206}\text{Pb}$ ratios are either the observed analytical uncertainty or the counting statistics error, whichever is highest. Common Pb corrections assume that most contaminant Pb is present on the surface of the analyzed grains, introduced from the sample preparation process, and has a composition that can be approximated using the Stacey and Kramers (1975) model for the present day. For grains with ages <1000 Ma, common lead was corrected using the "207-correction" which is calculated by projecting the uncorrected analyses onto concordia from the assumed common $^{207}\text{Pb}/^{206}\text{Pb}$ present day composition. In most cases, however, the amount of common Pb, revealed by monitoring ^{204}Pb , is relatively small and has little influence on the interpreted age. For grains with ages > 1000 Ma, factors for common lead correction were calculated by iteration from the $^{204}\text{Pb}/^{206}\text{Pb}$ and $^{204}\text{Pb}/^{207}\text{Pb}$ ratios provided by the Stacey and Kramers (1975) model at the calculated age, until convergence to a constant value. All ages are calculated using the decay constant recommendations of Steiger and Jäger (1977).

2.2.3.2. Zircon U-Th-Pb, Zr/Hf and Ti laser ablation ICP-MS analyses

LA-ICPMS analyses of Ti, Zr, Hf, Th, U and Pb isotopes were carried out with a Nd-YAG 213 nm Mercantek laser and a torch-shielded quadrupole Agilent 7500 ICP-MS spectrometer. To avoid the isobaric interference caused by $^{96}\text{Zr}^{+2}$, Ti was determined on the isotope ^{49}Ti (5.5%) instead of the most abundant ^{48}Ti (73.8%). The laser beam was set at a diameter of 60 μm , with a repetition rate of 10 Hz and an output energy of 75%. The ablation time was 60 s and the spot was pre-ablated during 45 s with a laser output energy of 50%. The ablation was done in a He atmosphere. ^{91}Zr was used as an internal standard. The external standard was the NIST-610 glass, which contains 434 ppm Ti, 439.9 ppm Zr, 417.7 ppm Hf, 409 ppm Pb, 457.1 ppm U and 450.6 ppm Th (Pearce et al., 1997). The following isotope ratios, determined by TIMS at the University of Granada, were also used: $^{204}\text{Pb}/^{206}\text{Pb} = 0.06$, $^{207}\text{Pb}/^{206}\text{Pb} = 0.9127$, $^{208}\text{Pb}/^{206}\text{Pb} = 2.1898$, $^{206}\text{Pb}/^{238}\text{U} = 0.2501$, $^{208}\text{Pb}/^{232}\text{Th} = 0.5402$. U-Pb LA-ICPMS ages are in good agreement with ion-microprobe

data but show more dispersion and tend to be more discordant. The precision (1σ) estimated on ten replicates of the NIST-610 analyzed in the same run was better than 2.5% for element ratios and ca. 0.3% for isotope ratios. Common lead interferences are significantly higher than in ion microprobe analyses, owing to the larger spot diameter. Data with a discordance factor $(^{206}\text{Pb}/^{238}\text{U})_{\text{age}}/(^{207}\text{Pb}/^{235}\text{U})_{\text{age}} < 0.9$ were always rejected, except they plot in a well-defined discordia line. In the rest, common lead was corrected using the same methods as described in the previous section. The detection limit for Ti was about 0.4 ppm, and the precision (1σ) estimated on the standards NIST-610, NIST-612 and a homemade glass, was 4%, 8% and 20% for Ti concentrations of 434 ppm, 48 ppm, and 12 ppm. Ti-in-zircon temperatures were estimated with the formula of Watson and Harrison (2005): $T (^{\circ}\text{K}) = 5080 / (6.01 - \log_{10}(\text{ppm Ti}))$ assuming a TiO_2 activity = 1.

2.2.3.3. Zircon Pb-Pb stepwise evaporation analyses

Single-zircon evaporation analyses (Kober, 1987) were performed using a SEM-RPQ multicollector Finnigan Mat 262 thermal ionization mass spectrometer with a double filament ion-source arrangement, operated with the RunIt262 Spectromat™ software. The zircon grain was mounted on a canoe-shaped Re evaporation filament and heated until the lead beam was intense enough (~200-400 ^{206}Pb ions per second). The lead was collected on the ionization filament for 20-30 min, and was then analyzed in 5 blocks with 7 scans per block. Once the analysis was finished, a new step was started by heating the zircon on the evaporation filament at a higher temperature than in the previous step (usually increasing the current by 50-100 mA) and analyzing as before the Pb deposited anew on the ionization filament. The procedure was repeated until all the lead was exhausted from the zircon. The number of steps depended on the size and lead content of each zircon. Data acquisition was performed in dynamic mode (peak hopping), using a secondary electron multiplier as detector with the 206-204-206-207-208 mass sequence. The mass-ratio 204/206 was monitored to detect and, if necessary, correct for common lead. Factors for common lead correction were calculated by iteration from the $^{204}\text{Pb}/^{206}\text{Pb}$ and $^{204}\text{Pb}/^{207}\text{Pb}$ ratios provided by the Stacey and Kramers (1975) model at the calculated age, until convergence to a constant value. Mass fractionation in the detector was corrected by multiplying by $\sqrt{(207/206)}$. Standard errors for each step were calculated according to the formula: $\text{SE} = 2 * \sigma / \sqrt{n}$. The confidence interval for the final age is calculated for the 95% confidence level of the mean of all steps.

3. PETROGRAPHY AND MINERALOGY

3. PETROGRAPHY AND MINERALOGY

Detailed description of the petrography and mineralogy of the Cambro-Ordovician rocks of Central Iberia studied by many authors (see section 1.3). The fact that all these materials were variably metamorphosed during the Variscan has resulted in a large variety of textures and mineral assemblages which, when the metamorphic grade is high, obliterates the primary pre-metamorphic features of the rocks. Therefore, since the primary purpose of this work has been to study the geochronology and isotope geology of the magmatic protoliths from which these rocks were derived, our sampling has been markedly biased in favor of low-grade samples. It should also be noted that the ranges of mineral compositions mentioned in the next section refer exclusively to the samples studied in this work. Our intention has been to provide a description of the rocks from which we studied zircons, rather than discussing the compositional spectrum or the metamorphic evolution of the studied bodies. The reader interested in these aspects can find more information in the literature mentioned in section 1.3

3.1 Ollo de Sapo Domain

3.1.1. The Ollo de Sapo Formation gneisses

As noted previously, the Ollo de Sapo Formation is composed of coarse-grained and fine-grained augen gneisses. Whereas the former tend to be located in the lower part of the formation, the later tend to be placed in the upper part. Locally the two types appear as alternating thin (< 1m) layers forming what is customarily known as the mixed facies.

3.1.1.1. The Ollo de Sapo coarse-grained gneisses

These are mesocratic gneisses with large augen within a medium- to fine-grained groundmass (Fig. 3.1). The augen often consist of isolated crystals of K-feldspar, plagioclase or quartz. Less frequently they are made of polycrystalline (though usually monomineralic) aggregates of these minerals.

The augen of K-feldspar are subhedral, or rarely euhedral, stubby prism to ovoidal crystals, locally rapakivi, which may reach up to 15 cm in diameter (Fig. 3.1). When the metamorphic grade is very low, they show Carlsbad twinning and are markedly perthitic indicating that they originally contained an elevated fraction of albite component. They often have inclusions of other minerals, especially plagioclase, quartz, biotite and,



Figure 3.1: Field appearance of the coarse-grained facies of the Ollo de Sapo metavolcanic rocks at Hiendelaencina region. Note the rapakivi texture of some megacrysts.

remarkably, zircon. For this reason, in an area of the Hiendelaencina region where the rock's weathering makes it possible to hand-pick K-feldspar megacrysts, we have been able to extract zircon from them (see section 4.1.1.5.2.).

Subhedral or anhedral, often rounded, quartz phenocrysts are also very abundant. Their size is about 1-3 cm, being, therefore, notably smaller than the K-feldspar megacrysts. If the metamorphic grade is low they often have a characteristic bluish color (Fig. 3.1, see also Fig. 3.5) attributed to the presence of sagenitic needles of rutile. Under the microscope they frequently appear as hexagonal prismatic sections with embayed borders, which strongly suggest a volcanic origin (Fig. 3.2A). As the metamorphic grade increases, the quartz phenocrysts become polygonized, recrystallizing into polycrystalline aggregates that in most cases still keep the original shape of the crystal (Fig. 3.2B)

The phenocrysts of plagioclase consist of subhedral stubby prisms up to 1-2 cm long, of oligoclase or oligoclase-albite, scarcely zoned and often with a marked albite twinning.

Besides these three types of phenocrysts, the low-grade coarse-grained gneisses of the Ollo de Sapo can also locally contain small (\approx 5-10 mm) lithic fragments composed of quartz, plagioclase, biotite, zircon and Fe-Ti oxides (Fig 3.3).

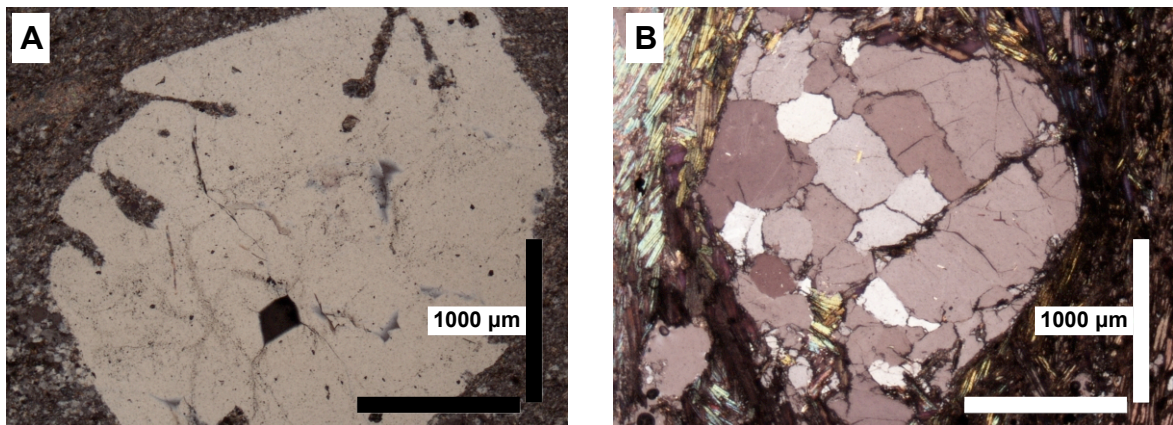


Figure 3.2: A) Quartz phenocrystal with embayed borders of the coarse-grained facies, in parallel nichols); B) Quartz phenocrystal which has been recrystallized forming fine-grained aggregates due to the metamorphism, in crossed nichols.



Figure 3.3: Lithic fragments (black angular specks) in coarse-grained low-grade metavolcanic rocks of the Ollo de Sapo in Sanabria.

The groundmass is fine-grained and variably deformed, the grain size increases with the metamorphic grade. In the samples with the lowest grade of metamorphism, it consists of quartz, oligoclase-albite, K-feldspar, retrograded crystal of biotite, chlorite, illite and, locally, prehnite-pumpellite, epidote-zoisite, carbonates and pseudomorphs of Fe-Mg minerals, the primary nature of which cannot currently be ascertained (Fig. 3.4). As the metamorphic grade increases, the groundmass becomes increasingly recrystallized and strongly foliated, with conspicuous SC structures defined by alternating muscovite + biotite-rich and quartz + feldspar-rich layers. The phenocrysts develop pressure shadows. The most deformed samples have abundant quartz and plagioclase ribbons, the internal structure of which is granoblastic.

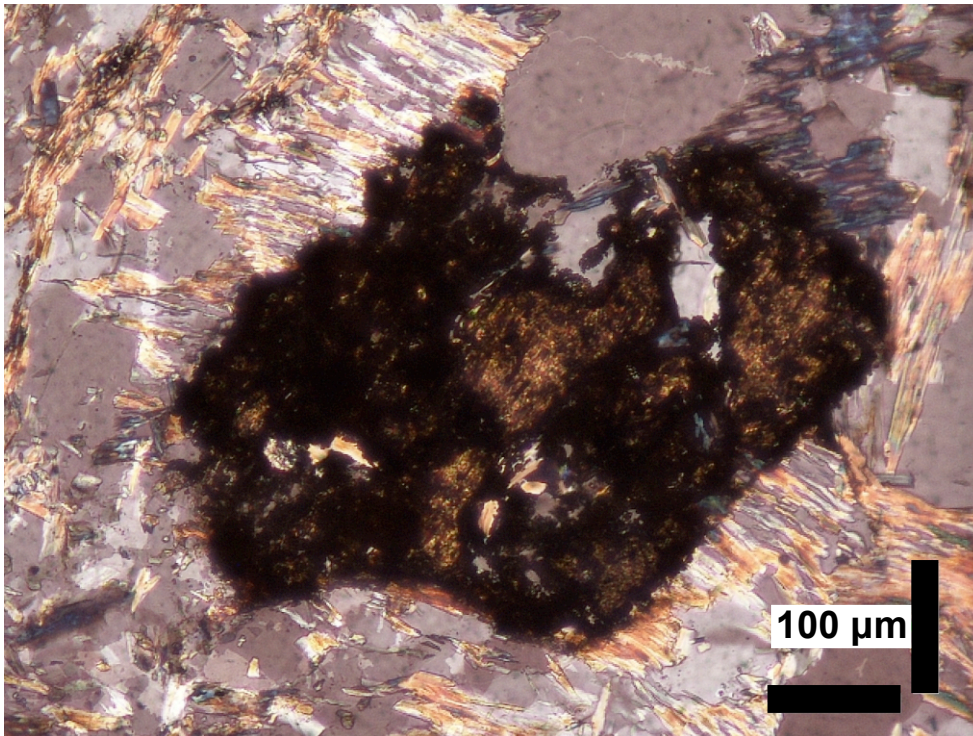


Figure 3.4: Pseudomorph of a Fe-Mg mineral, probably amphibole in the Ollo de Sapo Formation at Sanabria. Crossed nichols.

The accessory mineral assemblage consists of apatite, zircon, monazite, xenotime, titanite, ilmenite, rutile, Fe-oxides, occasional tourmaline, and rare uraninite. The most abundant are apatite and zircon. Apatite comprises stubby subhedral, up to 1 mm, prismatic crystals. Zircon consists of euhedral to subhedral, rarely rounded anhedral, crystals up to 350 μm , of short bipyramidal prism, seldom long prisms or needle-like. Titanite consists of subhedral to anhedral yellow-brown crystals up to 500 μm often closely associated with biotite, thus suggesting it might be secondary. Monazite and xenotime are subhedral and are present as isolated crystals or aggregates up to 300 μm in diameter. Tourmaline is subhedral and has a pleochroism from pale yellow to blue and size up to 250 μm .

The Ollo de Sapo gneisses are often altered, so that the primary mineralogy is, in most cases partially, transformed to secondary, low-temperature. The alteration is especially intense in the lower grade metamorphic samples.

3.1.1.2. The Ollo de Sapo fine-grained gneisses

These are formed of mesocratic to leucocratic augen gneisses (Fig. 3.5) that differ from these described in the previous section by the small size of the augen. The rocks contain

small phenocrysts of K-feldspar, quartz or plagioclase embedded in a fine-grained matrix (see Fig. 1.5C and 1.5D in Chapter 1). The gneisses are commonly deformed showing SC structures (Fig. 3.5).

The phenocrysts of K-feldspar are subhedral short prisms up to 1 cm in length. They often appear twinned and some have perthitic intergrowths. The phenocrysts of quartz are subhedral or rounded anhedral, up to 1-1.5 cm long, often with a characteristic bluish color (Fig. 3.4). The phenocrysts of plagioclase (oligoclase) are subhedral and are up to 6 mm long.

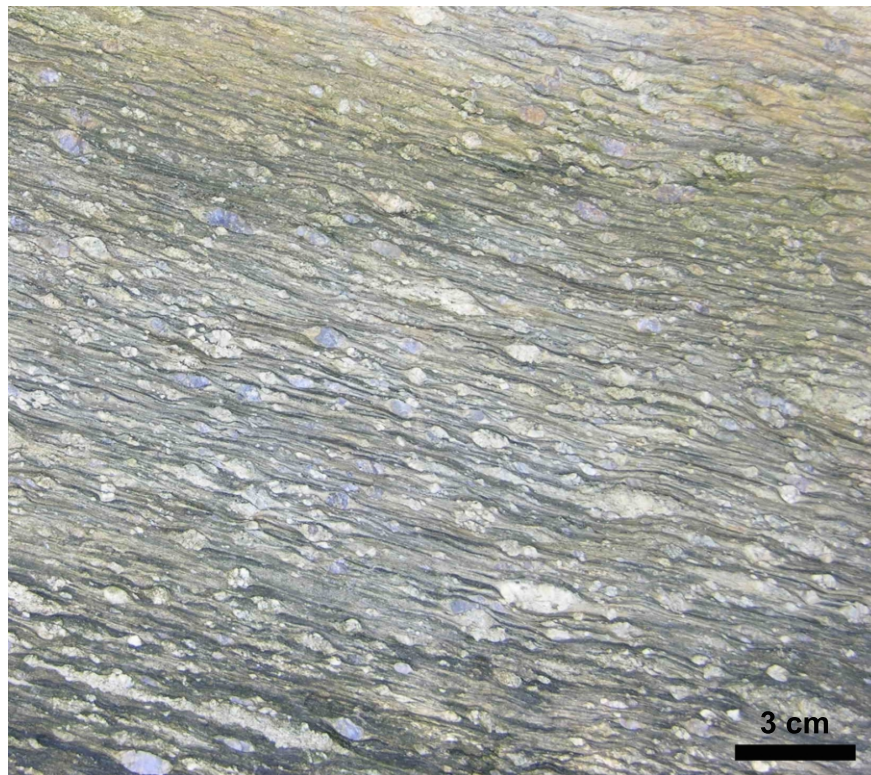


Figure 3.5: Field appearance of the fine-grained facies of the Ollo de Sapo metavolcanic rocks at Vivero. Note the bluish color of quartz phenocrysts.

The groundmass of the Ollo de Sapo fine-grained gneisses is identical to that of the coarse-grained gneisses both in texture and mineralogy. To avoid tedious repetitions, therefore, the reader is referred to section 3.1.1.1.

3.1.2. The Ollo de Sapo Formation metagranites

As mentioned in previous sections, we have studied the four most representative metagranites associated with the Ollo de Sapo Formation: Viana do Bolo, San Sebastián,

Miranda do Douro and Antofñita. Currently, these rocks are mesocratic to leucocratic augen-gneisses characterized by large subhedral K-feldspar megacrysts, up to 15 cm, although often between 3 and 8 cm, with conspicuous aplopegmatitic dikes (Fig. 3.6) and enclaves (Fig. 3.7).

Some petrographic characteristics are common to all the metagranites. The K-feldspar megacrysts often show Carlsbad twinning and locally have Frals inclusions. They are embedded in a medium to coarse-grained (up to 1 cm) foliated groundmass which mainly consists of quartz, plagioclase, K-feldspar, biotite and rare muscovite. Quartz is present as anhedral or subhedral, slightly rounded, crystals. K-feldspar appears as anhedral crystals often with microcline twinning and perthitic intergrowths. Plagioclase is often oligoclase, rarely andesine or albite, which forms subhedral tabular crystals that may reach up to 4 cm and commonly display albite twinning. They are always zoned. The varietal mineralogy is biotite and muscovite, rarely biotite alone, which are present as subhedral, often deformed crystals. When the metamorphic grade reaches the anatexis isograd, these rocks often contain cordierite + sillimanite \pm garnet.

The accessory mineral assemblage is composed of apatite, zircon, monazite, xenotime, tourmaline, ilmenite and Fe-sulfides. The most abundant is apatite, which forms short euhedral prisms up to 600 μm long. Zircon consists of short, occasionally long, bipyramidal, pale-pink to brownish-orange prisms with a maximum dimension up to 300 μm , although sometimes pale-yellow. Monazite and xenotime are subhedral isolated crystals with a maximum dimension up to 30 and 10 μm respectively. Tourmaline appears as euhedral or subhedral crystals with a maximum dimension up to 1.5 mm and has a pleochroism from colorless to yellow-pale blue.

The Viana do Bolo metagranite (Fig. 3.7) is leucocratic and coarse-grained with quartz + K-feldspar + plagioclase + biotite + muscovite as major minerals. Apatite, zircon, monazite and Fe-oxides are accessory minerals.

The San Sebastián metagranite appears forming strongly deformed narrow (< 2 m thick) dikes. It is a fine-grained leucocratic gneiss with quartz + K-feldspar + plagioclase + biotite as major minerals and apatite, zircon, monazite and Fe-Ti-oxides as accessories.

The Miranda do Douro metagranite (Fig. 3.8) is mesocratic and coarse-grained with quartz + plagioclase + biotite + K-feldspar as major minerals and apatite, zircon, monazite,

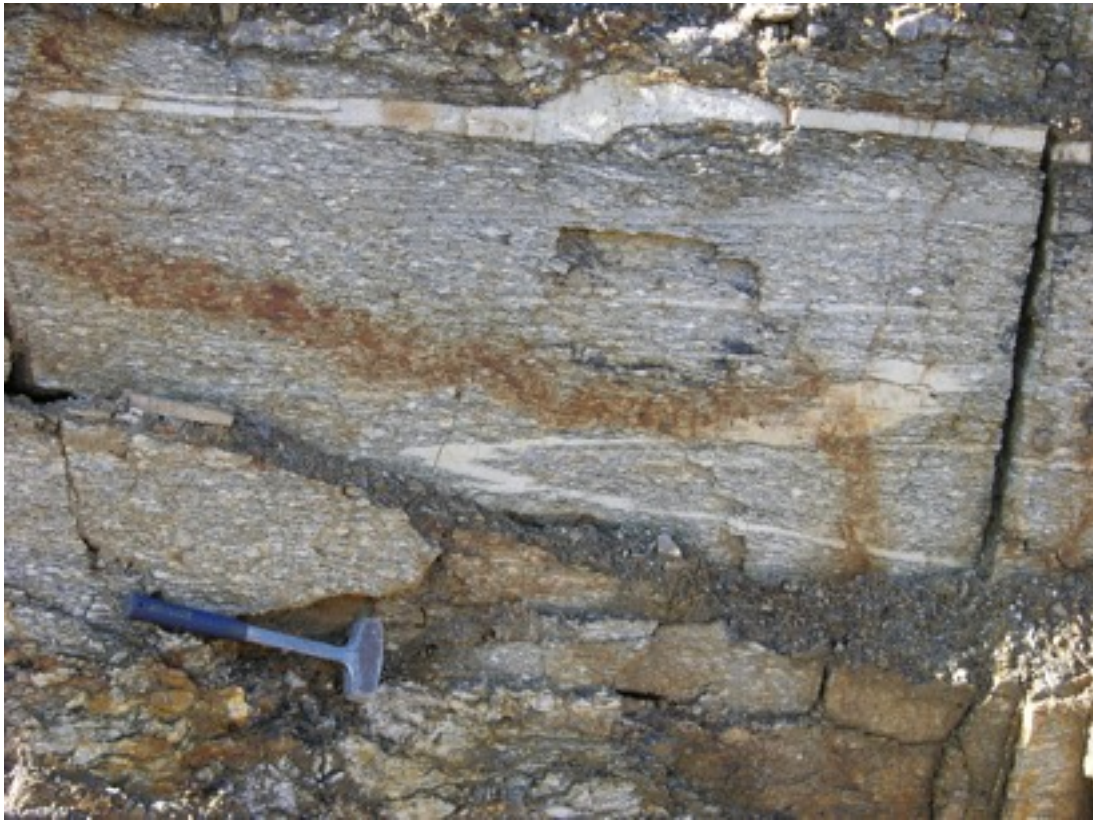


Figure 3.6: Field appearance of the metagranites associated to the Ollo de Sapo Formation. Note the folded aplites in the Antoñita metagranite of Hiendelaencina



Figure 3.7: Field appearance of the metagranites associated to the Ollo de Sapo Formation. Note the microgranular enclave in the Viana do Bolo metagranite

xenotime, Fe-Ti-oxides and Fe-sulfides as accessory minerals.



Figure 3.8: Field appearance of the Miranda do Douro metagranite.

The Antoñita metagranite (Fig. 3.6) is leucocratic and coarse-grained with quartz + plagioclase + K-feldspar + plagioclase + biotite + muscovite as major minerals and tourmaline, apatite, zircon, monazite and xenotime as accessory minerals.

3.2. The Schist-Graywacke Complex Domain

The pre-Middle Ordovician metagranites and metavolcanic rocks of the Schist-Graywacke Complex Domain are mostly found in four areas located in the northern part of the domain. These are the Tormes Dome, Sierra de Gredos, Sierra de Guadarrama and the Anatectic Complex of Toledo. The most relevant petrographic features of the pre-Middle Ordovician rocks of these four areas are as follows:

3.2.1. The Tormes Dome

The pre-Middle Ordovician metagranites of the Tormes Dome studied here, Feroselle, Vitigudino and Ledesma are leucocratic augen gneisses. The augen comprise subhedral to euhedral prisms of K-feldspar megacrysts with a maximum dimension up to 6 cm. They show Carlsbad twinning and perthitic intergrowths and contain numerous minute inclusions of other minerals such as biotite and quartz. In areas where the deformation has been more intense, they also show fractures perpendicular to major axis and asymmetric pressure shadows.



Figure 3.9: Field appearance of the Feroselle metagranite.



Figure 3.10: Field appearance of the Vitigudino metagranite.



Figure 3.11: Field appearance of the Ledesma metagranite. Note the asymmetric pressure shadows

The megacrysts are included in a medium-grained and variably metamorphosed groundmass. When the metamorphic grade is low, the groundmass mainly consists of quartz + K-feldspar + plagioclase (oligoclase-albite) + biotite ± muscovite defining a hypidiomorphic texture. As the metamorphic grade increases, biotite is partially replaced by sillimanite ± cordierite. Locally, the metagranites reached anatexis and produced small leucocratic stocks of Variscan anatectic leucogranites.

The accessory mineral assemblage consists of apatite, zircon, monazite, xenotime, tourmaline, Fe-sulphide and Fe-oxides. Apatite is colorless euhedral crystals with a maximum dimension up to 300 µm. Zircon comprises short, occasionally long, bipyramidal, pale-pink to brownish crystals with a maximum dimension of 450 µm. It can also sometimes be colorless or yellow. Monazite and xenotime are isolated subhedral crystals with a maximum dimension of 200 µm and 170 µm respectively. Tourmaline consists of up to 1 mm long subhedral crystals, and shows a slight pleochroism from yellow-greenish to pale blue.

The Feroselle metagranite (Fig. 3.9) is leucocratic and coarse-grained with quartz + K-feldspar + plagioclase + biotite ± muscovite ± sillimanite ± cordierite as major minerals and apatite, zircon, monazite, xenotime and Fe-oxides as accessory minerals.

The Vitigudino metagranite (Fig. 3.10) is leucocratic and fine-grained with quartz + K-feldspar + plagioclase + biotite ± muscovite as major minerals and apatite, zircon, monazite, xenotime, Fe-oxides and Fe-sulfides as accessory minerals.

The Ledesma metagranite (Fig. 3.11) is leucocratic and coarse-grained with quartz + K-feldspar + plagioclase + biotite ± muscovite ± cordierite ± sillimanite as major minerals and apatite, zircon, monazite, ilmenite and occasional sulfides as accessory minerals.

3.2.2. The Sierra de Gredos

The pre-Middle Ordovician metagranites of the Sierra de Gredos studied in this work are Bercimuelle, Castellanos and San Pelayo, all of them located in the northwest of the Sierra. These metagranites are strongly metamorphosed, even partially migmatized. They are typically mesocratic and coarse grained, though fine-grained leucocratic facies are also common, especially in Castellanos. The augen consist of subhedral prismatic K-feldspar phenocrysts with dimensions between 3-5 cm, that often have tiny inclusions of quartz and biotite. They usually show Carlsbad twinning and perthitic intergrowths.



Figure 3.12: Field appearance of the Bercimuelle metagranite.

The K-phenocrysts are embedded in a medium to coarse-grained groundmass mainly composed of quartz + plagioclase + K-feldspar + biotite \pm muscovite \pm sillimanite \pm cordierite \pm rare garnet.

The accessory mineral assemblage comprises apatite, zircon, tourmaline, monazite, xenotime, rutile, sulfides and Fe-Ti-oxides. Zircon consist of short, sometimes long or slightly rounded, bipiramidal, usually pale-pink to brownish-orange prisms with a maximum dimension of 450 μm , although sometimes is colorless or pale-yellow. Apatite comprises colorless subhedral crystals up to 200 μm . Tourmaline consists of subhedral crystals with a maximum dimension of 4 mm and pleochroism from yellow-blue to pale-yellow-pale-blue. Monazite and xenotime comprise either isolated subhedral crystals or fine-grained aggregates with a maximum dimension of 80 and 120 μm respectively.

The San Pelayo metagranite is mesocratic and coarse-grained composed of K-feldspar + quartz + oligoclase + biotite + muscovite + abundant tourmaline. Apatite, zircon, monazite, xenotime, Fe-Ti-oxides and sulfides are accessory minerals.

The Castellanos metagranite is mesocratic to leucocratic and coarse to fine-grained with quartz + K-feldspar + plagioclase + biotite + muscovite ± cordierite as major minerals and tourmaline, apatite, zircon, monazite and xenotime as accessories.

The Bercimuelle metagranite (Fig. 3.12) is also mesocratic to leucocratic, coarse-grained, and always appears variably migmatized. The major minerals are quartz + K-feldspar + albite-oligoclase + biotite ± cordierite. The accessory assemblage comprises apatite, zircon, monazite, xenotime, Fe-Ti-oxides and sulfides.

3.2.3. The Sierra de Guadarrama

As mentioned in section 1.2.2.2., the Sierra de Guadarrama contains the most voluminous accumulation of pre-Middle Ordovician metagranites and metavolcanic rocks in the Central Iberian Zone. The most relevant petrographic features are as follows:

3.2.3.1. Metagranites

The metagranites of the Sierra de Guadarrama sampled for this work are La Cañada, Santa Maria de la Alameda, La Hoya, La Estación I, La Estación II, Vegas de Matute (melanocratic and leucocratic metagranites), La Morcuera, Buitrago de Lozoya and Prádena del Rincón (see Table 2.2, chapter 2). All these rocks are mesocratic to leucocratic granitic gneisses, either augen or fine-grained, that have been variably deformed and metamorphosed, often to anatexis.

Some petrographic characteristics are common to all the metagranites. The augen consist either of K-feldspar megacrysts, plagioclase phenocrysts or monomineralic aggregates of quartz or plagioclase. K-feldspar megacrysts are subhedral to euhedral prisms, rarely ovoidal, with a maximum dimension up to 8-12 cm. They show Carlsbad twinning and perthitic intergrowths, and may have rapakivi textures. The plagioclase phenocrysts are subhedral crystals of oligoclase-andesine with a maximum dimension up to 6 mm.

The major mineral assemblage of the groundmass consists of K-feldspar + plagioclase + quartz + biotite ± muscovite forming a hypidiomorphic aggregate. At low metamorphic grade, biotite is altered to relicts of biotite and chlorite and an aggregate of illite and other phyllosilicates. As the metamorphic grade increases, garnet, cordierite and sillimanite appear as a consequence of the breakdown of biotite.



Figure 3.13: Mesoscopic appearance of the La Cañada metagranite.



Figure 3.14: Mesoscopic appearance of the La Hoya metagranite.



Figure 3.15: Field appearance of the La Estación I metagranite.



Figure 3.16: Field appearance of the La Estación II metagranite.



Figure 3.17: Field appearance of the melanocratic metagranite of Vegas de Matute.



Figure 3.18: Field appearance of the leucocratic metagranite of Vegas de Matute.

The accessory mineral assemblage consists of apatite, zircon, monazite, xenotime, rutile, Fe-Ti-oxides and Fe-Cu sulfides. Apatite occurs as subhedral to euhedral colorless crystals with a maximum dimension of 1 mm. Zircon consists of short, rarely long or acicular, bipyramidal prisms with a maximum dimension up to 450 μm . It is usually pale-pink to brownish and sometimes yellowish or orange. Monazite and xenotime occurs as subhedral to euhedral isolated crystals or aggregates with a maximum dimensions up to 500 and 200 μm respectively.

The La Cañada metagranite (Fig. 3.13) is mesocratic and coarse-grained with quartz + plagioclase + K-feldspar + biotite \pm cordierite \pm garnet \pm muscovite as major minerals. Apatite, zircon, monazite and Fe-oxides are the accessory minerals.

The La Hoya metagranite (Fig. 3.14) is mesocratic and coarse-grained with K-feldspar + plagioclase + quartz + biotite \pm muscovite. Apatite, zircon, monazite, xenotime and Fe-Ti oxides are the accessory minerals.

The La Estación I metagranite (Fig. 3.15) is mesocratic, coarse-grained, with K-feldspar + oligoclase + quartz + biotite \pm sillimanite \pm chlorite \pm cordierite. Apatite, zircon, tourmaline, monazite, xenotime, Fe-Ti-oxides and Fe-sulfides are the accessory minerals.

The La Estación II metagranite (Fig. 3.16) is mesocratic and coarse-grained with K-feldspar + quartz + oligoclase + biotite \pm chlorite \pm muscovite. Apatite, zircon, monazite, xenotime and Fe-Ti-oxides are the accessory minerals.

The melanocratic metagranite of Vegas de Matute (Fig. 3.17) is coarse-grained with quartz + plagioclase + K-feldspar + biotite \pm muscovite \pm chlorite \pm sillimanite \pm garnet. Apatite, zircon, monazite, xenotime, Fe-Ti oxides and Fe-Cu sulfides are the accessory minerals.

The leucocratic metagranite of Vegas de Matute (Fig. 3.18) is fine-grained with quartz + plagioclase + K-feldspar + biotite + muscovite as major minerals and apatite, zircon, monazite, xenotime and Fe-Ti oxides as accessories.

The La Morcuera metagranite (Fig. 3.19) is mesocratic and coarse-grained with K-feldspar + plagioclase + quartz + biotite \pm muscovite as major minerals and apatite, zircon, monazite, xenotime, rutile and Fe-Ti oxides as accessories.



Figure 3.19: Field appearance of the La Morcuera metagranite. Note the rapakivi textures of some megacrysts.

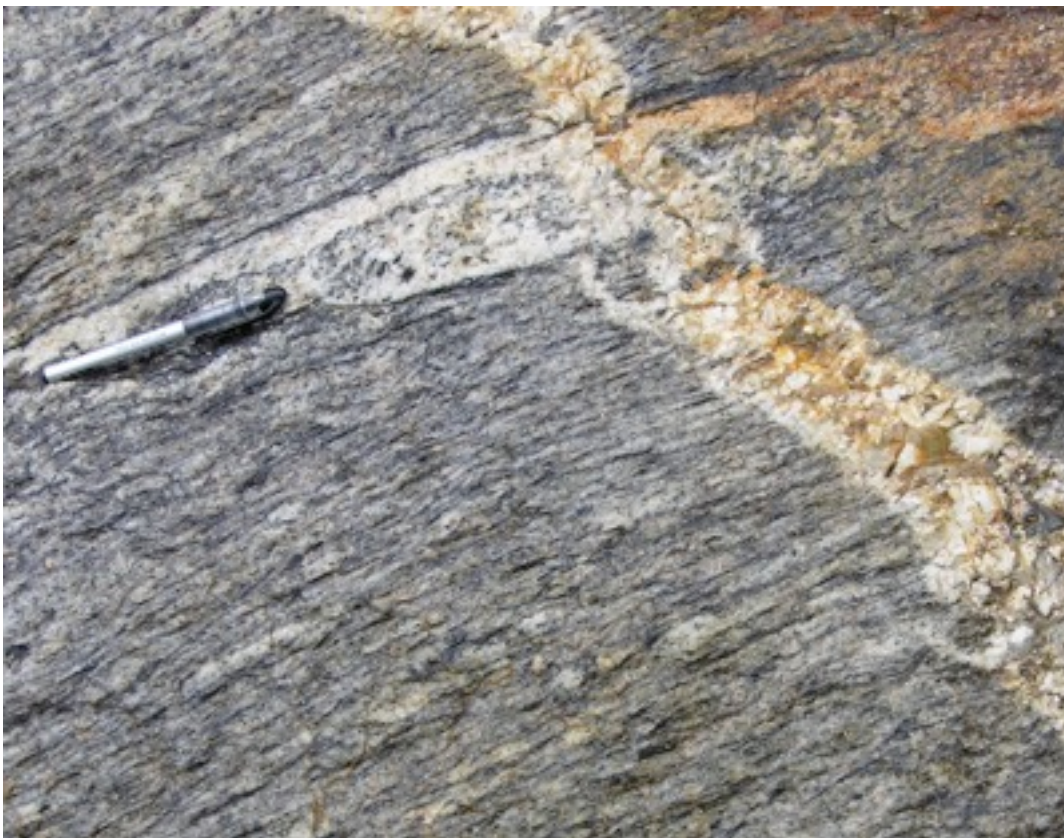


Figure 3.20: Mesoscopic appearance of the Buitrago de Lozoya metagranite cut by pegmatite and aplite veins.



Figure 3.21: Field appearance of the Prádena del Rincón metagranite.

The Buitrago de Lozoya metagranite (Fig. 3.20) is mesocratic and coarse-grained with K-feldspar + plagioclase + quartz + biotite \pm sillimanite \pm muscovite. Apatite, zircon, monazite, xenotime and Fe-sulfides are the accessory minerals.

The Prádena del Rincón metagranite (Fig. 3.21) is leucocratic and fine-grained with quartz + K-feldspar + plagioclase + biotite \pm muscovite \pm garnet as major minerals and apatite, zircon, monazite, Fe-Ti oxides as accessories.

3.2.3.2. *Metavolcanic rocks*

In the Sierra de Guadarrama there are three gneissic bodies, La Berzosa, El Cardoso and Riaza, with features similar to the metavolcanic rocks of the Ollo de Sapo Formation (García Cacho, 1973; Capote and Fernández Casals, 1975). As in the latter formation, a coarse-grained and a fine-grained facies are identified depending on the size of the augen.

The **coarse-grained facies** (Fig. 3.22) is formed of gneisses with abundant augen. These are either K-feldspar megacrysts, phenocrysts of plagioclase and quartz, or



Figure 3.22: Field aspect of the coarse-grained metavolcanites of the Sierra de Guadarrama.



Figure 3.23: Field aspect of the fine-grained metavolcanites of the Sierra de Guadarrama.

polycrystalline aggregates of plagioclase and quartz. The K-feldspar megacrysts are subhedral prisms with a maximum dimension up to 6-8 cm; they show Carlsbad twinning, perthitic intergrowths and numerous small inclusions of biotite and quartz. The phenocrysts of plagioclase are subhedral platy crystals of oligoclase (An_{21-18}) with conspicuous albite twinning. The phenocrysts of quartz are subhedral, often rounded, crystals with a maximum dimension of 2 mm.

The major mineral assemblage of the groundmass depends on the metamorphic grade. In general it consists of a fine grained hypidiomorphic aggregate of quartz, plagioclase, K-feldspar, biotite and muscovite, which are locally partially replaced by subhedral to anhedral crystals of cordierite.

The accessory mineral assemblage consists of apatite, zircon, monazite, xenotime, rutile, Fe-Ti-oxides and Fe-Cu sulfides. Apatite is present as subhedral to euhedral colorless crystals with a maximum dimension up to 500 μm . Zircon consists of short, bipyramidal, pale pink to brownish prisms with a maximum dimension up to 450 μm . Monazite and xenotime occur as subhedral to euhedral isolated crystals with maximum dimensions up to 20 and 15 μm respectively, or as aggregates.

The **fine-grained facies** (Fig. 3.23) do not have K-feldspar megacrysts and the phenocrysts are always less than 1 cm. In general, the gneisses of the fine-grained facies are less metamorphosed than the coarse-grained gneisses. For that reason, the original shapes of the phenocrysts of quartz and embayed borders in quartz can be distinguished. The rest of the features of the fine-grained facies are similar to the features of the coarse-grained facies described above.

3.2.4. The Anatectic Complex of Toledo

As mentioned in the previous section, in the Anatectic Complex of Toledo samples from two metagranitic orthogneisses, Polán and Mohares, located in the north part of the complex (see Table 2.2 in chapter 2) have been studied. In both cases they are mesocratic to leucocratic fine-grained gneisses, locally augen, which were strongly metamorphosed during the Variscan orogeny (Fig. 3.24). The augen consist of K-feldspar megacrysts and quartz phenocrysts. The megacrysts of K-feldspar are subhedral with a maximum dimension of 3-5 cm. They show Carlsbad twinning and small inclusions of plagioclase.



Figure 3.24: Field aspect of the Mohares metagranite of the Anatectic Complex of Toledo. Note the lack of augen and the elevated migmatization.

The phenocrysts of quartz are anhedral or slightly rounded, and are rarely larger than 1 cm.

The augen are embedded in a medium-grained groundmass with a major mineral assemblage mainly composed of quartz, plagioclase, K-feldspar and biotite. Plagioclase appears as subhedral to anhedral albite-twinning crystals. K-feldspar is present as anhedral crystals. Biotite appears as subhedral, often deformed, crystals. In the most metamorphosed samples, biotite is partially replaced by subhedral to euhedral crystals of garnet and cordierite.

The accessory mineral assemblage is composed of apatite, zircon, monazite and Fe-oxides. Apatite appears as euhedral crystals with a maximum dimension up to 300 μm . Zircon consists of short, although occasionally long, bipiramidal, yellowish to brownish prisms with a maximum dimension up to 450 μm . Monazite appears as pale-brown euhedral crystals with a maximum dimension of 160 μm .

4. GEOCHRONOLOGY

4. GEOCHRONOLOGY

For this study more than two thousand zircon grains were hand-picked under the binocular microscope, mounted and studied by cathodoluminescence imaging. A selection of them were dated using spot U-Th-Pb and stepwise evaporation Pb-Pb methods as described in Chapter 2. The cathodoluminescence studies have also been used for calculating the proportion of inherited cores contained in the zircons. The crystallization ages of all studied bodies are summarized in Table 4.1, at the end of this chapter.

4.1. The Ollo de Sapo Domain

4.1.1. The Ollo de Sapo Formation

We separated zircon from ten different samples of the metavolcanic rocks of the Ollo de Sapo Formation. The exact location and coordinates of the studied samples are given in Table 2.1. The zircons from all samples show many similarities. Most of them are euhedral or subhedral short prisms, rarely long prisms or needle-like, often with pale-pink-brownish, rarely grayish, or colorless. The average size is about 200 x 90 μm ; occasionally they may reach sizes up to 350 x 135 μm . The cathodoluminescence imaging reveal that these zircons stand out because of the ubiquitous presence of inherited cores (Figs. 4.1z1, 4.1z2, 4.1z3, 4.1z6, 4.1z7, 4.1z8, 4.1z10, 4.1z11, 4.1z12, 4.1z13 and 4.1z14), present in no less than 70-80%, locally up to 100% of the grains (Fig. 4.2). Most cores show magmatic oscillatory zoning. Most rims also display typical magmatic oscillatory zoning, though it is

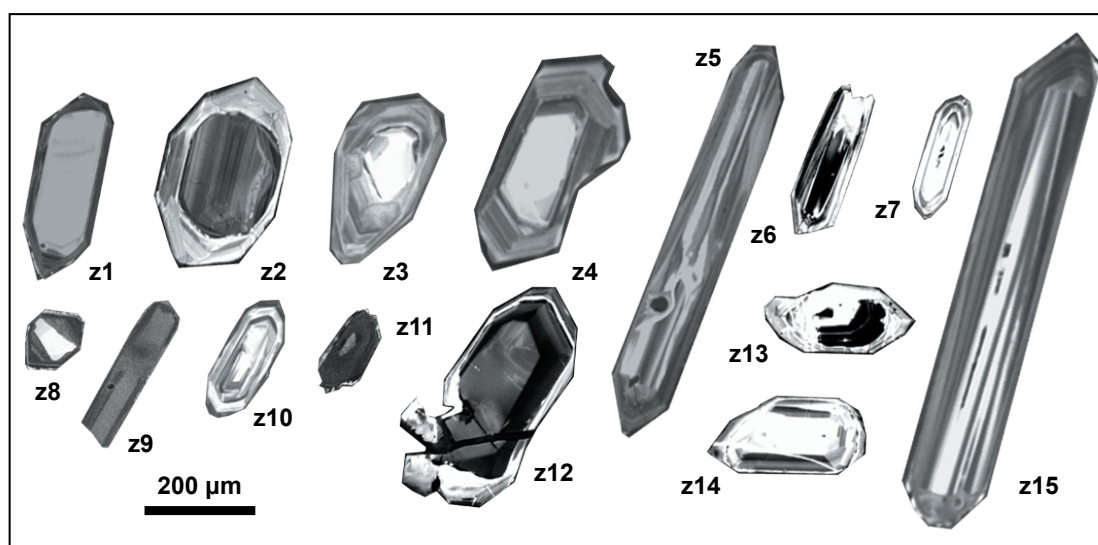


Figure 4.1: Cathodoluminescence images of the most representative internal textures of the metavolcanic rocks of the Ollo de Sapo Formation.

often sharply discordant with the cores (Fig. 4.1z4, 4.1z5, 4.1z9 and 4.1z15); convolute zoning is also found locally in some grains.

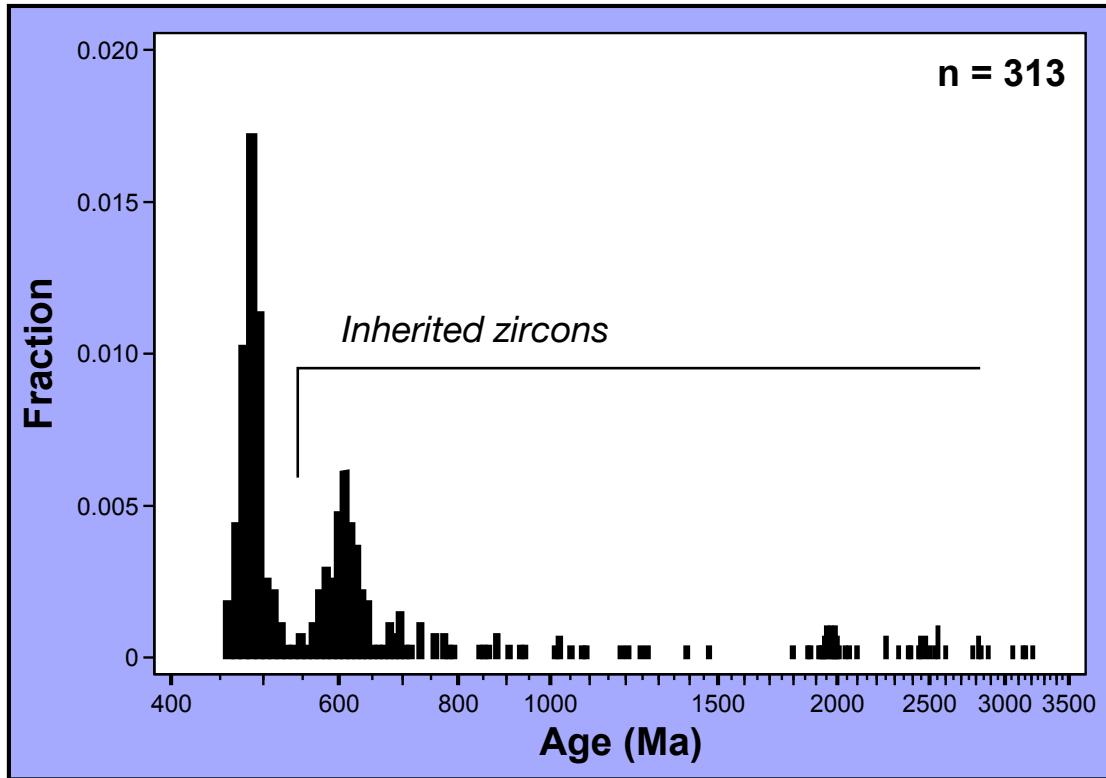


Figure 4.2: Age distribution pattern found in zircons of the metavolcanic and metagranitic rocks of the Ollo de Sapo Domain. Built from the U-Pb SIMS, SHRIMP and LA-ICPMS data.

The results of zircon dating in the Ollo de Sapo Formation can be summarized as follows:

4.1.1.1. *Vivero*

In the Vivero region we studied three samples of metavolcanic rocks, two (one coarse-grained, another fine-grained) from the Ria del Barquero in the Cantabric Coast and one fine-grained facies from the Mañón area. The three samples were studied using U-Pb LA-ICPMS (Table 4.2) and Pb-Pb stepwise evaporation (Table 4.3). U-Pb dating of the three samples yielded the same results (Figs. 4.3A and 4.3C): a very abundant concordant to subconcordant Cambro-Ordovician population with a mean age of 486 ± 3 Ma (Fig. 4.3B), another also abundant concordant to subconcordant population with a mean age of 608 ± 16 Ma, two concordant points at ca. 1000 Ma, several concordant points at ca. 2000 Ma, and a discordia line with an upper intercept at ca. 2600 Ma.

Twelve zircon grains were analyzed by the Pb-Pb evaporation method. Seventeen out of 48 evaporation steps yielded a plateau-age at 486 ± 3 Ma, identical to the U-Pb age of the youngest concordant population (Fig. 4.3D), which we therefore consider the crystallization age of the Ollo de Sapo in Vivero. It is remarkable that most crystals contain inherited cores. The most common ages for these cores are ca. 540 and ca. 600 Ma and minimum ages of ca. 800, 2000 and 2800 Ma are also present.

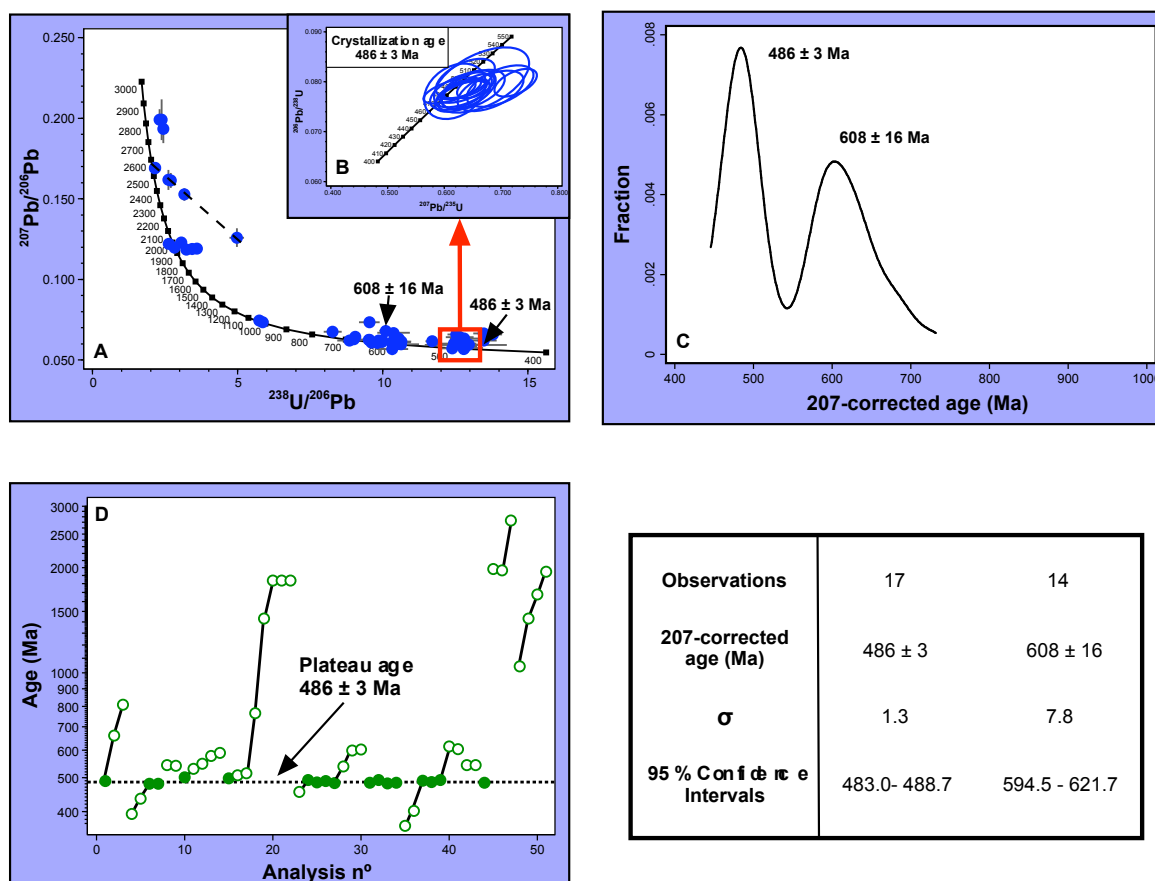


Figure 4.3: Geochronological data of the Vivero metavolcanic rocks: A) Tera-Wasserburg concordia; B) Wetherill concordia; C) Distribution of concordant 207-corrected ages and the statistical data of the two most abundant populations; and D) Pb-Pb stepwise evaporation age.

According to the cathodoluminescence images, the Cambro-Ordovician population (486 ± 3 Ma) was present in rims of grains containing older cores (Fig. 4.4LAz17) and in grains with uniform ages from core to rim (Fig. 4.4LAz3). The Ediacaran population (608 ± 16 Ma), however, was present in the cores of zircons (Fig. 4.4LABz1), most of which were overgrown by Cambro-Ordovician rims (Fig. 4.4LAz17). The older-than-Ediacaran ages always correspond to relictic cores (Fig. 4.4LAz9).

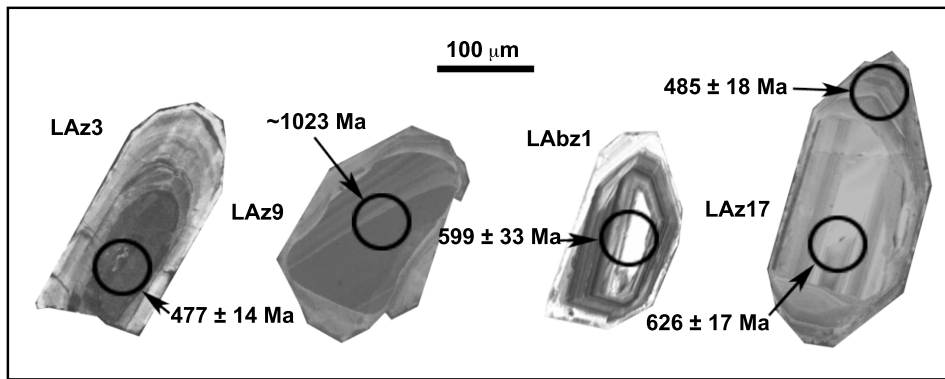
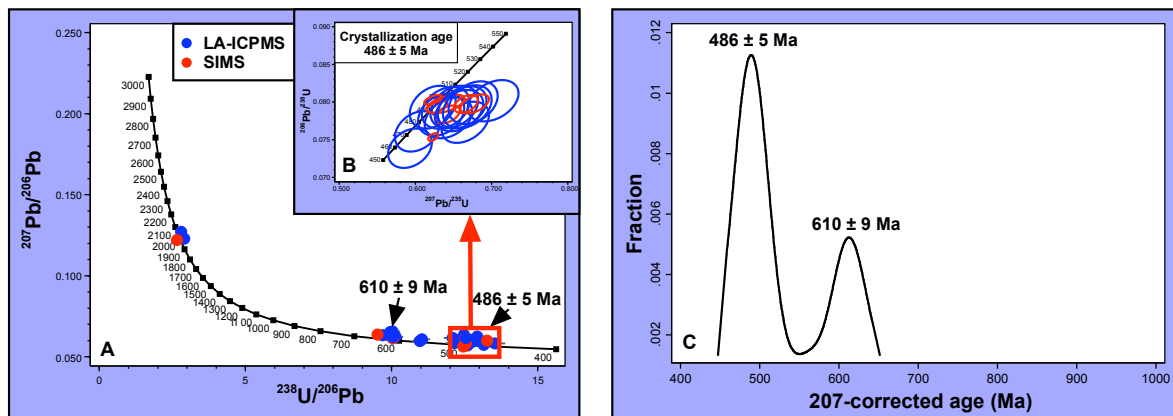


Figure 4.4: Cathodoluminescence images and ages of selected zircons from the Vivero metavolcanic rocks.

4.1.1.2. Trives

In the Trives region, a coarse-grained metavolcanic rock was dated using U-Pb ion microprobe and LA-ICPMS (Fig. 4.5 and Table 4.4). These data reveal the presence of three concordant population (Fig. 4.5A and 4.5C). The first one, very abundant, is at 458-499 Ma, yields a mean age of 486 ± 5 Ma and is considered the age of crystallization (Fig. 4.5B). The second population, also abundant, at 605-616 Ma yields a mean age of 610 ± 9 Ma. Between these two populations are two points at 558 and 562 Ma which probably represent a mixing age between them. The third population, less numerous and



Observations	207-corrected age (Ma)	σ	95 % Confidence Intervals
20	486 ± 5	2.3	481.2 - 490.8
8	610 ± 9	4.8	607.5 - 613.0

Figure 4.5: Geochronological data of the Trives metavolcanic rocks: A) Tera-Wasserburg concordia; B) Wetherill concordia; and C) Distribution of concordant 207-corrected ages and the statistical data of the two most abundant populations.

also concordant, yielded ages between 1988 Ma and 2059 Ma, with a mean of 2015 ± 40 Ma.

As revealed by the cathodoluminescence images, the Cambro-Ordovician population (486 ± 5 Ma) is present in rims of grains containing older cores and in grains with uniform ages from core to rim (Fig. 4.6z10 and 4.6z12). The Ediacaran (610 ± 9 Ma) (Fig. 4.6z4) and the Orosirian (1988 Ma and 2059 Ma populations) (Fig. 4.6z1) were always obtained in cores of zircons with Cambro-Ordovician rims.

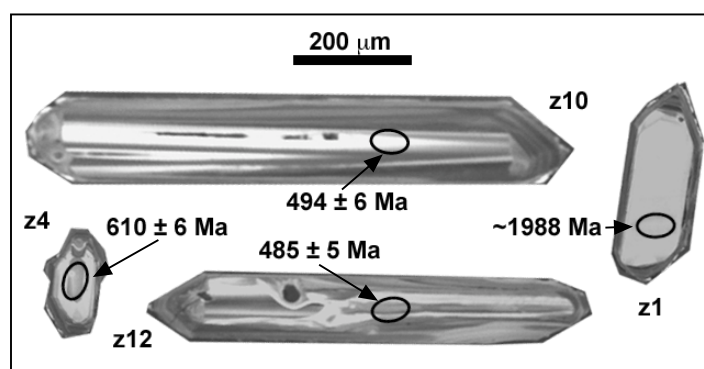


Figure 4.6: Cathodoluminescence images and ages of selected zircons from the Trives metavolcanic rocks.

4.1.1.3. Sanabria

In the Sanabria region, the metavolcanic rocks form a sequence with a thickness of about 1500 m (Díez Montes et al., 2006). To determine the time span recorded in the sequence, we analyzed samples located at different stratigraphic positions: one from the lowermost coarse-grained facies, another from the uppermost fine-grained facies and another from the upper mixed facies.

4.1.1.3.1. Lower metavolcanic rocks

The sample from the lower metavolcanic rocks was analyzed by U-Pb ion-microprobe and LA-ICPMS (Table 4.5). The results revealed a complex age distribution with two concordant populations and several grains scattered at older ages (Figs. 4.7A and 4.7C). The most abundant concordant population is Cambro-Ordovician, with a precise mean age of 492 ± 4 Ma which is therefore interpreted as the crystallization age of the lower metavolcanic rocks of the Ollo de Sapo Formation in the Sanabria region (Fig. 4.7B). The second most abundant concordant population is Ediacaran, with a mean age of 605 ± 4 Ma.

There are also two points at 680 Ma, another two at 1.0 Ga, one at ca. 2.0 Ga and some others defining a discordia line with an upper interception at ca. 2.6 Ga.

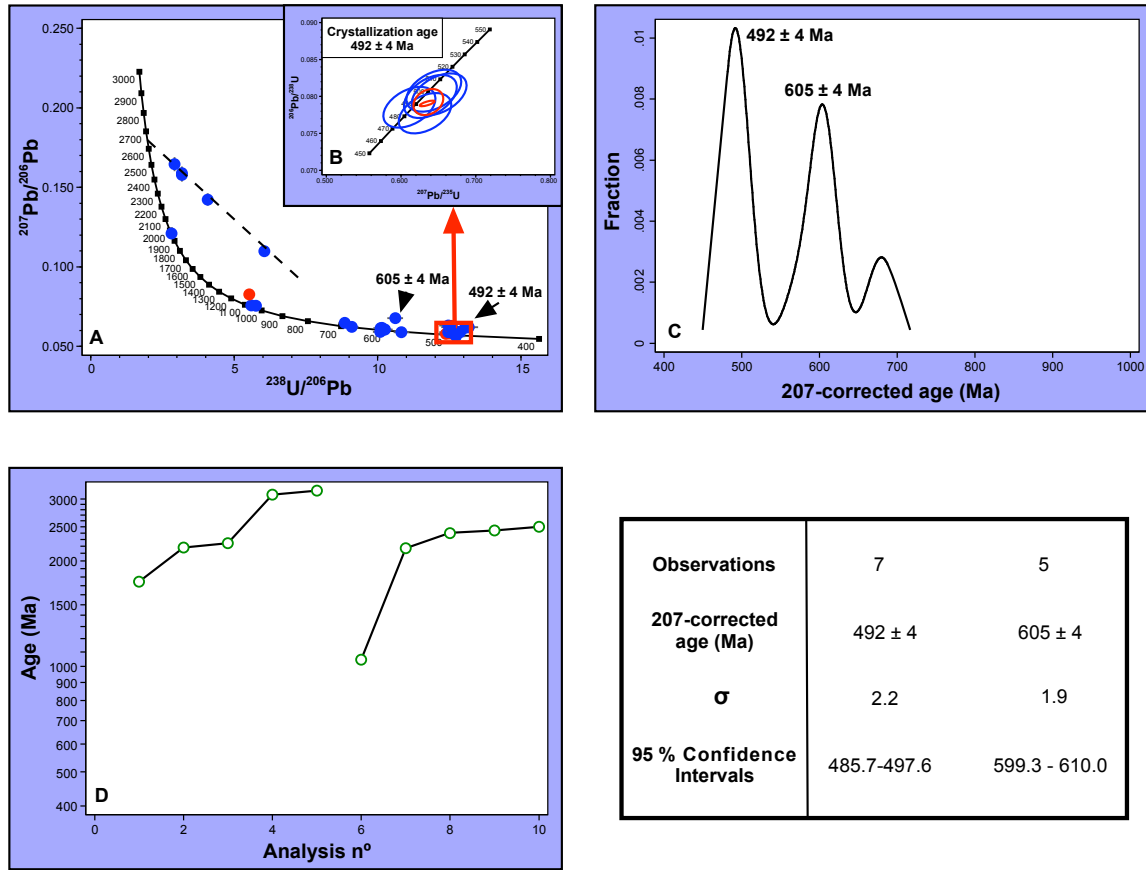


Figure 4.7: Geochronological data of the lower metavolcanic rocks of the Sanabria region: A) Tera-Wasserburg concordia; B) Wetherill concordia; C) Distribution of concordant 207-corrected ages and the statistical data of the two most abundant populations; and D) Pb-Pb stepwise evaporation age.

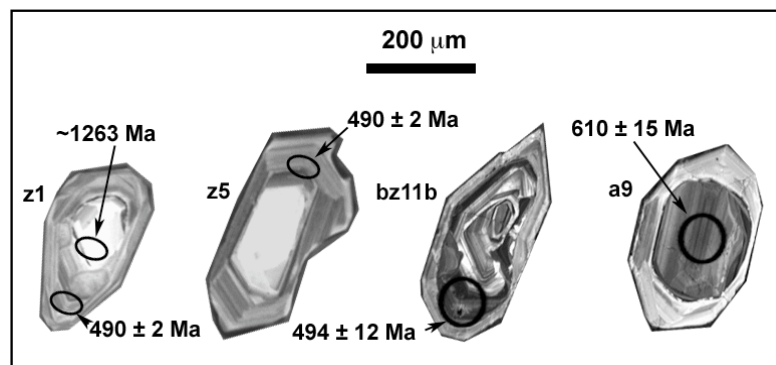


Figure 4.8: Cathodoluminescence images and ages of selected zircons from the lower meta-volcanic rocks of Sanabria region.

Some zircon grains analyzed with the Pb-Pb evaporation method did not yield a suitable estimation of the crystallization age, but revealed the presence of Archean cores with minimum ages of 3-3.2 Ga (Fig. 4.7D and Table 4.6).

The Cambro-Ordovician ages, according to the cathodoluminescence imaging, are represented as rims of grains containing older relictic cores (Fig. 4.8z1) and in uniform grains (Figs. 4.8z5 and 4.8bz11b). The Ediacaran ages were obtained in cores of grains rimmed by Cambro-Ordovician overgrowths (Fig. 4.8a9) or in scarce totally relictic zircons. All older-than-Ediacaran ages were obtained in relictic zircons and, more often, in cores rimmed by Cambro-Ordovician overgrowths (Fig. 4.8z1).

4.1.1.3.2. Upper metavolcanic rocks

The two samples belonging to the upper metavolcanic rocks and mixed facies were analyzed with the U-Pb (ion-microprobe and LA-ICPMS) and Pb-Pb stepwise evaporation methods (Tables 4.7 and 4.8). The two samples have exactly the same U-Pb age

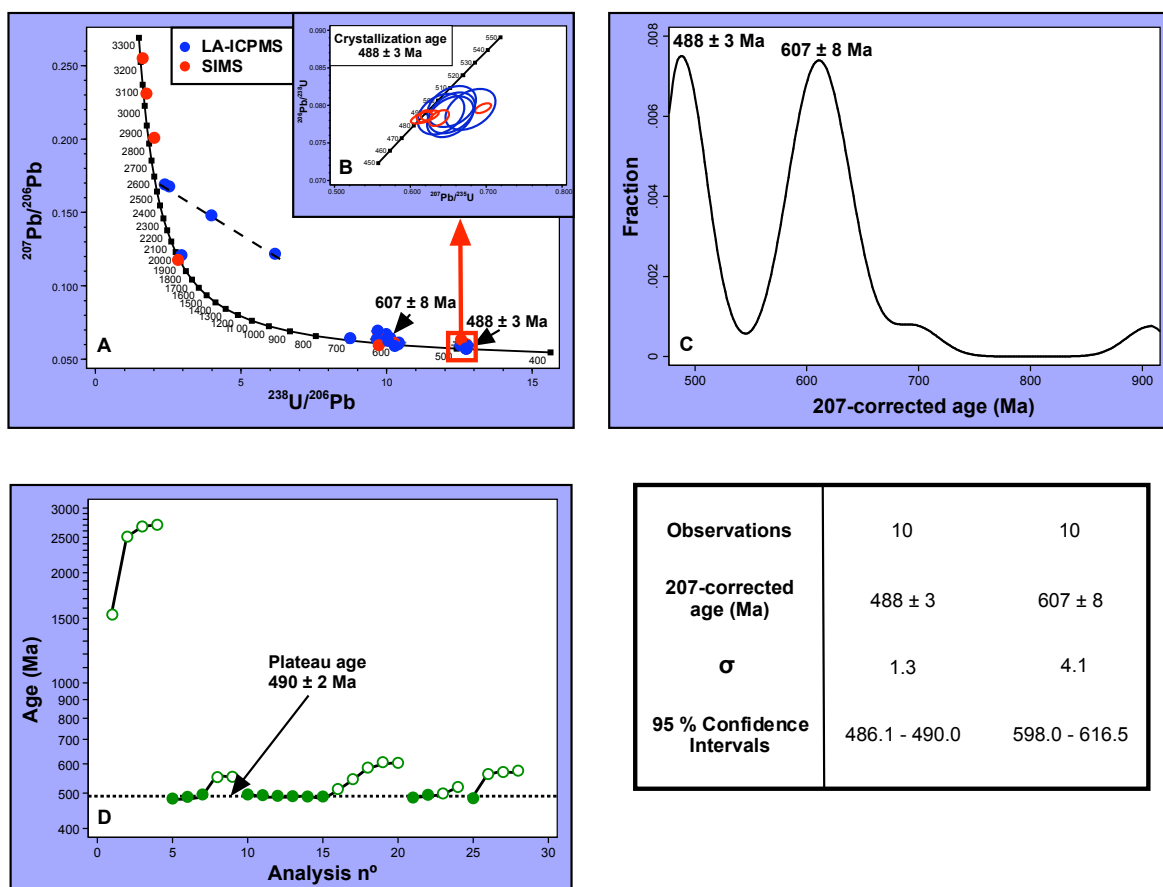


Figure 4.9: Geochronological data of the upper metavolcanic rocks of the Sanabria region: A) Tera-Wasserburg concordia; B) Wetherill concordia; C) Distribution of concordant 207-corrected ages and the statistical data of the two most abundant populations; and D) Pb-Pb stepwise evaporation age.

distribution, with abundant concordant Cambro-Ordovician and Ediacaran populations, a few concordant Orosirian ages, and several points yielding a discordia line with an upper interception at ca. 2.6 Ga (Figs. 4.9A and 4.9C).

The Cambro-Ordovician population is the most abundant. It yielded a precise mean age of 488 ± 3 Ma which represents, therefore, the crystallization age of the upper metavolcanic rock in the Sanabria region (Fig. 4.9B). The Ediacaran zircons yielded a mean age of 607 ± 8 Ma and, as in previous cases, made up the main inherited population.

Six zircon grains were analyzed by the Pb-Pb stepwise evaporation method. Twelve of 28 evaporation steps yielded a precise plateau-age of 490 ± 2 Ma (Fig. 4.9D) which represents, therefore, the best estimate of the crystallization age of these rocks.

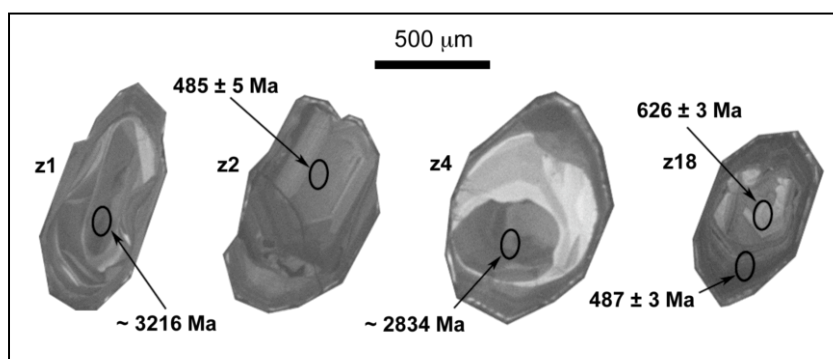
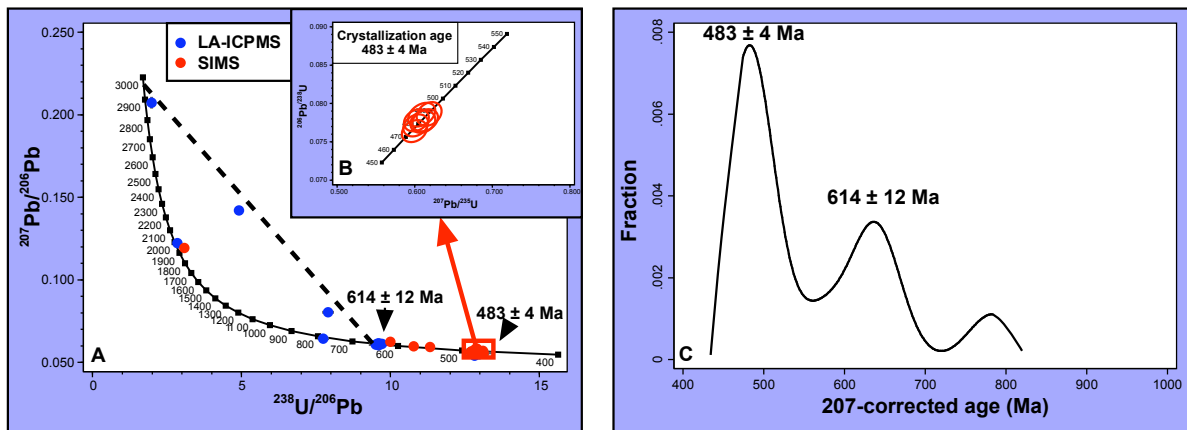


Figure 4.10: Cathodoluminescence images and ages of selected zircons from the upper metavolcanic rocks of the Sanabria region.

According to the cathodoluminescence images, the Cambro-Ordovician population is present in rims that overgrow over older cores (Fig. 4.10z18) and, less frequently, from grains with uniform ages (Fig. 4.10z2). The Ediacaran and older populations were obtained in cores of grains rimmed by Cambro-Ordovician overgrowths (Figs. 4.10z4 and 4.10z18).

4.1.1.4. Villadepera

The metavolcanic rocks of the Villadepera region were dated using U-Pb ion-microprobe and LA-ICPMS (Table 4.9). Both techniques yielded similar results (Fig. 4.11A and 4.11C) and reveal the presence of two concordant populations and few older ages. As before, the most abundant is Cambro-Ordovician and the second Ediacaran. There are also two concordant points about 2000 Ma, and some discordant points that define a discordia line with an upper interception at ca. 3000 Ma.



Observations	207-corrected age (Ma)	σ	95 % Confidence Intervals
7	483 ± 4	1.8	486.5 - 496.6
4	614 ± 12	6.2	586.8 - 620.4

Figure 4.11: Geochronological data of the Villadepera metavolcanic rocks: A) Tera-Wasserburg concordia; B) Wetherill concordia; and C) Distribution of concordant 207-corrected ages and the statistical data of the two most abundant populations.

The Cambro-Ordovician population yielded a precise mean age of 483 ± 4 Ma which represents, therefore, the crystallization age of the Ollo de Sapo Formation in the this region (Fig. 4.11B). The Ediacaran, has an mean age of 614 ± 12 Ma. The former, according to cathodoluminescence imaging, resulted either from rims of grains with inherited cores or from grains with uniform ages (Fig. 4.12z6). The later (Fig. 4.12z5), as well as the older ages (Fig. 4.12z7), were obtained from the restitic cores of zircons rimmed by Cambro-Ordovician overgrowths.

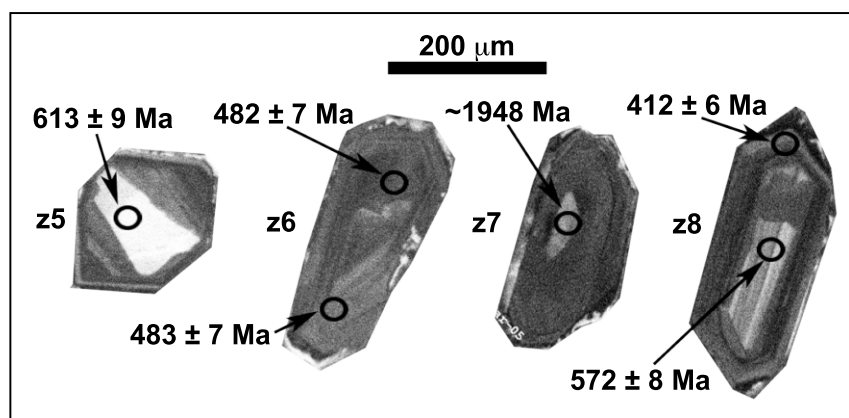


Figure 4.12: Cathodoluminescence images and ages of selected zircons from the Villadepera metavolcanic rocks.

4.1.1.5. Hiendelaencina

In the Hiendelaencina region, as in Sanabria, the metavolcanic rocks form a thick sequence of about 2000 m (González Lodeiro, 1981b). In order to determine the time span recorded in this sequence, three samples in different stratigraphic positions were analyzed: two samples from the lower metavolcanic rocks, one augen and one fine-grained, and one sample from the upper metavolcanic rocks.

4.1.1.5.1. Lower metavolcanic rocks

U-Pb ion microprobe and LA-ICPMS data from the two samples from the lower metavolcanic rocks are indistinguishable, so that they will be discussed together (Table 4.10). These data reveal the presence of several concordant or nearly concordant populations and few older ages (Figs. 4.13A and 4.13C). The first concordant population is Cambro-Ordovician, with a mean age of 495 ± 5 Ma which is interpreted as the crystallization age of the lower metavolcanic rocks of the Ollo de Sapo in the

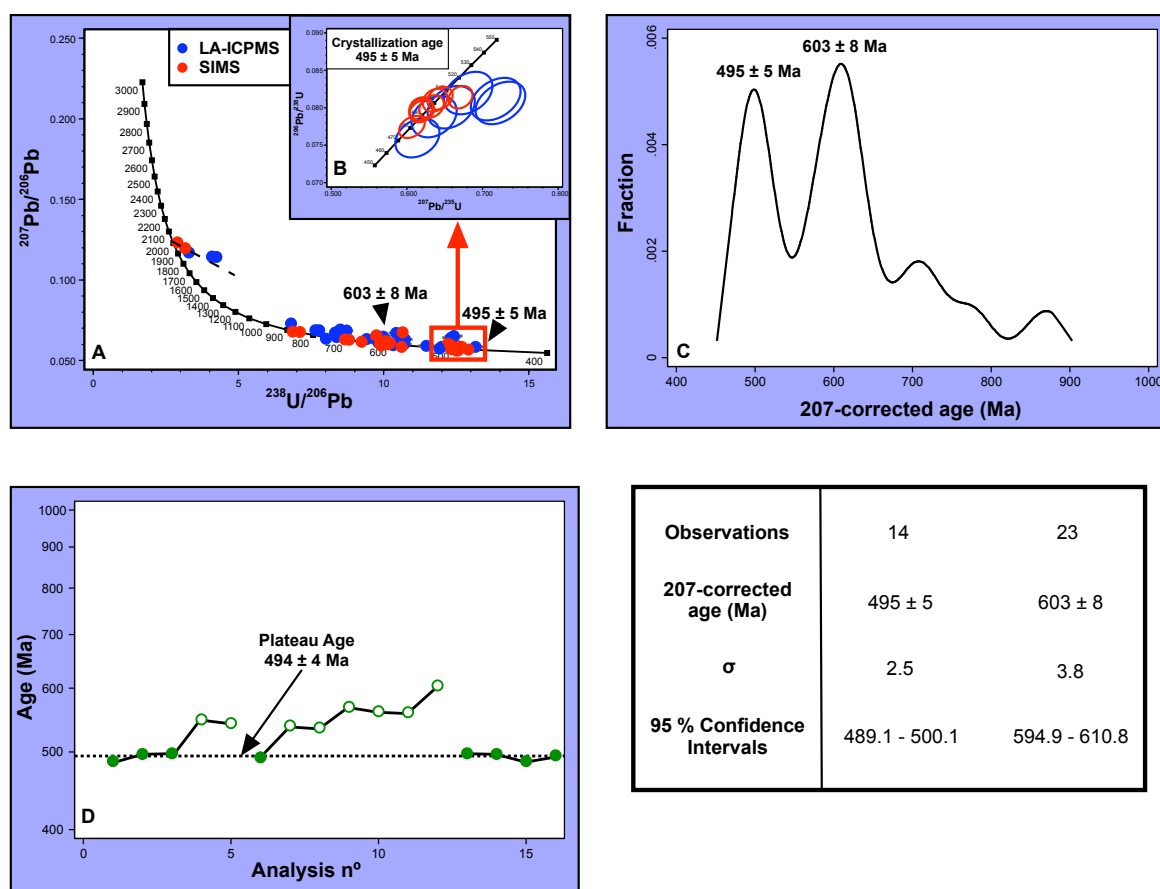


Figure 4.13: Geochronological data of the lower metavolcanic rocks of the Hiendelaencina region: A) Tera-Wasserburg concordia; B) Wetherill concordia; C) Distribution of concordant 207-corrected ages and the statistical data of the two most abundant populations; and D) Pb-Pb stepwise evaporation age.

Hiendelaencina region (Fig. 4.13B). The second concordant population has a mean age of 603 ± 8 Ma and represents, therefore, the main inherited population. The third concordant population spreads from 662 Ma to 695 Ma and the last one from 864 Ma to 878 Ma. The older ages are discordant and yield a discordia line with an upper interception at 2016 ± 49 Ma. It must be considered that the 662 Ma to 695 Ma population may be an artifact, owing to mixed Ediacaran and Tonian ages.

Three zircons were analyzed by the Pb-Pb stepwise evaporation method (Table 4.11 and Fig. 4.13D). Eight of the 16 performed evaporation steps yielded a plateau-age of 494 ± 4 Ma, which matches the age obtained by U-Pb analysis. Two of these grains had a restitic core with a minimum age of 540 and about 600 Ma.

The crystallization ages were obtained in uniform grains and, more often, in rims over older cores (Fig. 4.14z32). The 572-627 Ma (Fig. 4.14z32), the 662-695 Ma and the 864-878 Ma populations (Fig. 4.14z28) were mainly obtained in cores of rimmed grains or, less frequently, in totally restitic zircons, and the ≈ 2000 Ma ages were obtained in totally restitic grains (Fig. 4.14z26).

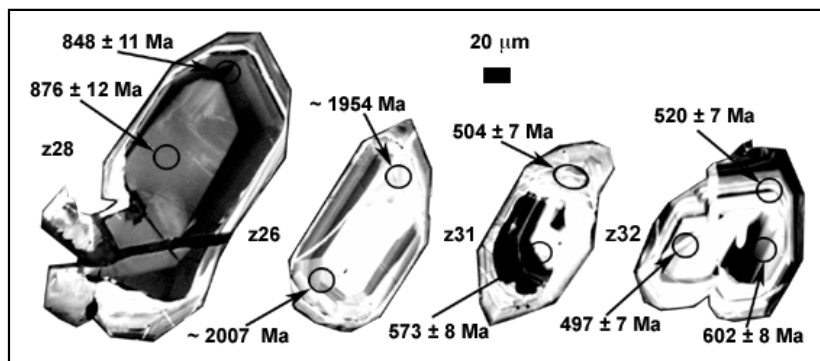


Figure 4.14: Cathodoluminescence images and ages of selected zircons from the lower metavolcanic rocks of the Hiendelaencina region.

4.1.1.5.2. Upper metavolcanic rocks

To date the upper part of the Hiendelaencina sequence we extracted zircons from a concentrate of fresh K-feldspar megacrysts from the coarse-grained facies. These zircons were analyzed by the U-Pb ion microprobe and LA-ICPMS and Pb-Pb stepwise evaporation method (Table 4.12 and Table 4.13)

U-Pb data revealed the presence of two main concordant or nearly concordant populations (Figs. 4.15A and 4.15C), the most abundant at 463-504 Ma, yielded a mean

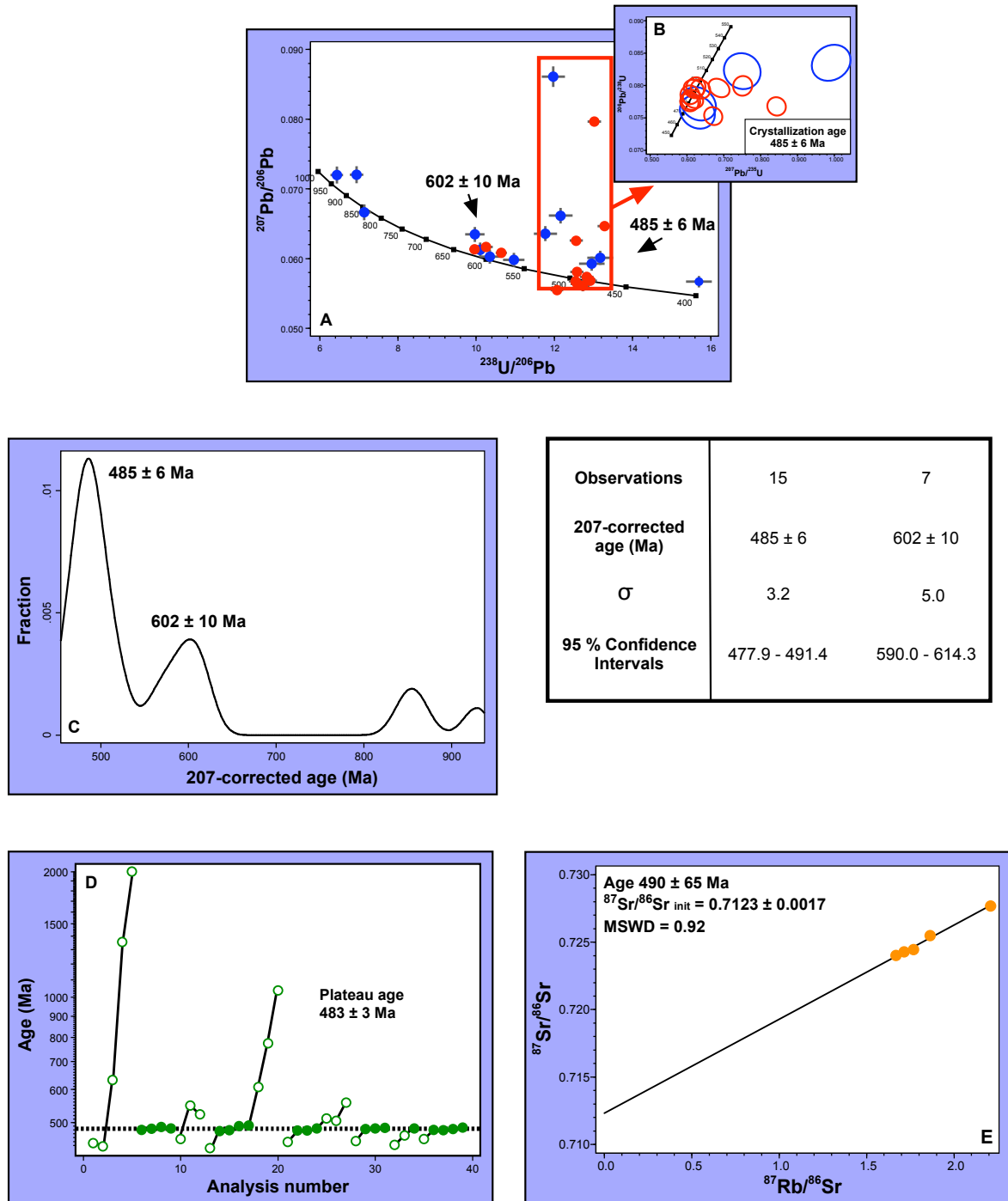


Figure 4.15: Geochronological data of the upper metavolcanic rocks of the Hiendelaencina region: A) Tera-Wasserburg concordia; B) Wetherill concordia; C) Distribution of concordant 207-corrected ages and the statistical data of the two most abundant populations; D) Pb-Pb stepwise evaporation age; and E) Rb-Sr internal isochron of a sliced K-feldspar megacryst.

age of 485 ± 6 Ma and represents the time of crystallization (Fig. 4.15B). The second population at 578-616 Ma, has a mean age of 602 ± 10 Ma and represents the main inherited population. There are also a few concordant or nearly-concordant grains yielding ages between 850 and 900 Ma and finally there is a younger-than-crystallization age which corresponds to texturally discontinuous rims, the significance of which will be discussed in section 4.3.3.

Eight zircon grains were also studied with the Pb-Pb stepwise evaporation method. Nineteen out of 39 evaporation steps yielded a very precise plateau age of 483 ± 3 Ma (Fig. 4.15D). Since this value totally agrees with the U-Pb data, it is considered the best estimate of the crystallization age. Seven grains yielded younger overgrowths with a minimum age of ca. 445 Ma and four grains also had older cores, in one case with a minimum age of ca. 2000 Ma.

Also, five slabs sliced from one of these megacrysts have sufficient Rb/Sr variation to obtain a Rb/Sr internal isochron (Table 4.14). This yielded an age of 490 ± 65 Ma with $(^{87}\text{Sr}/^{86}\text{Sr})_{\text{init}} = 0.7123 \pm 0.0017$ and $\text{MSWD} = 0.92$ (Fig. 4.15E). Although the confidence interval is large, because the $^{87}\text{Sr}/^{86}\text{Sr}$ coordinate only ranges from 1.67 to 2.21, the excellent fit and the coincidence with U-Pb and Pb-Pb zircon data permits consideration of isochron as a good estimate of the crystallization age of the K-feldspar megacryst.

The 463-504 Ma ages obtained in rims of grains overgrowing over older restitic cores and in uniform grains (Figs. 4.16z27 and 4.16z32). The 578-616 Ma ages were mainly obtained in the core of 463-504 Ma rimmed zircons (Figs. 4.16z23 and 4.16z24).

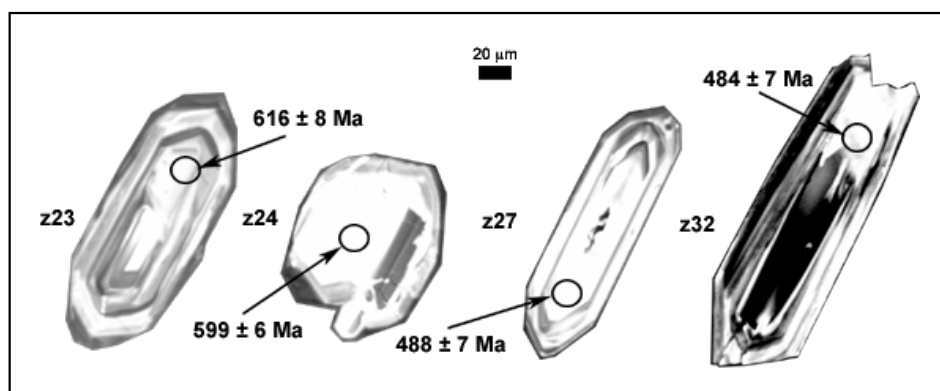


Figure 4.16: Cathodoluminescence images and ages of selected zircons from the upper metavolcanic rocks of the Hiendelaencina region.

4.1.2. The metagranites of the Ollo de Sapo Formation

Zircon grains from the four metagranites associated with the Ollo de Sapo Formation, Antofñita, Miranda do Douro, San Sebastián and Viana do Bolo are often pale-pink to yellow or, less frequent, orange to brownish or colorless. Most are short prismatic, euhedral or subhedral and only a few are long prismatic or acicular. Their average size is about $180 \times 75 \mu\text{m}$, and the maximum up to $450 \times 180 \mu\text{m}$. As in the metavolcanic rocks, cathodoluminescence imaging reveals that a significant fraction of them, no less than 70%, except in the San Sebastián metagranite with less than 20%, contain relict cores (Fig. 4.2) which in most cases display a oscillatory zoning (Figs. 4.17z5, 4.17z8, 4.17z12, 4.17z13 and 4.17z14).

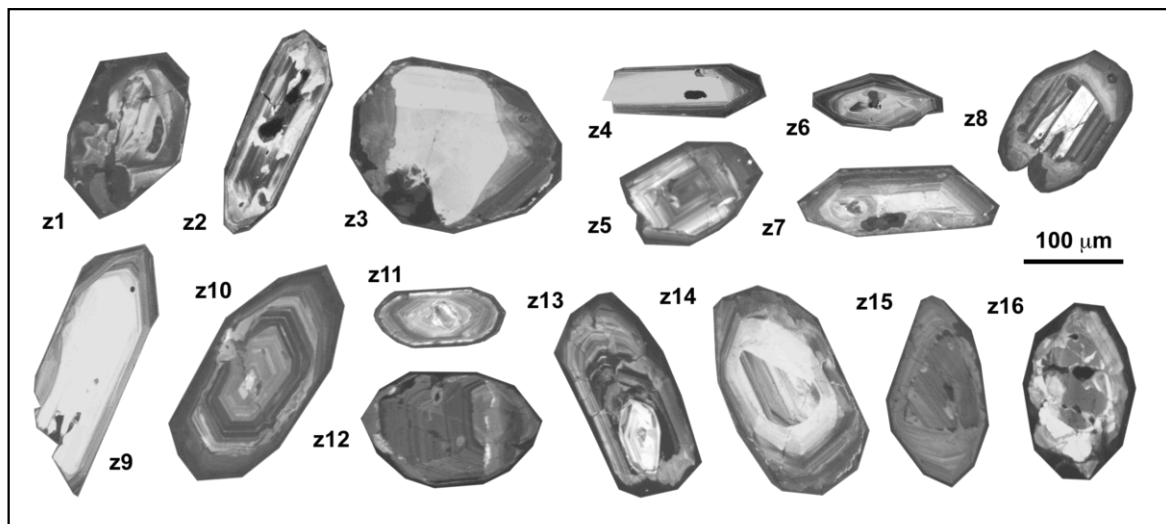


Figure 4.17: Cathodoluminescence images of the most representative internal textures of the metagranites of the Ollo de Sapo Domain.

4.1.2.1. Viana do Bolo

The Viana do Bolo metagranite outcrops in the Sanabria region and is spatially associated to the Ollo de Sapo metavolcanic rocks (see Fig. 3 in Lancelot et al., 1985). We have dated the common facies at Covelo using the U-Pb LA-ICPMS (Table 4.15) and Pb-Pb stepwise evaporation methods (Table 4.16). The U-Pb data revealed the existence of two concordant populations, one Cambro-Ordovician and the other Ediacaran, and few older ages defining a discordia line with an upper interception at ca. 2.6 Ga (Figs. 4.18A and 4.18C). The Cambro-Ordovician population is the most abundant and concordant. It is present either in rims or in totally uniform grains (Figs. 4.19LA050a and 4.19LA024a). Its mean age is $487 \pm 4 \text{ Ma}$ (Fig. 4.18B) which we consider, therefore, the crystallization age

of the Viana do Bolo metagranite. The Ediacaran population is less abundant and also less concordant. It comes entirely from zircon cores (Figs. 4.19LA018a and 4.19LA048a) and yielded a mean age of 609 ± 10 Ma.

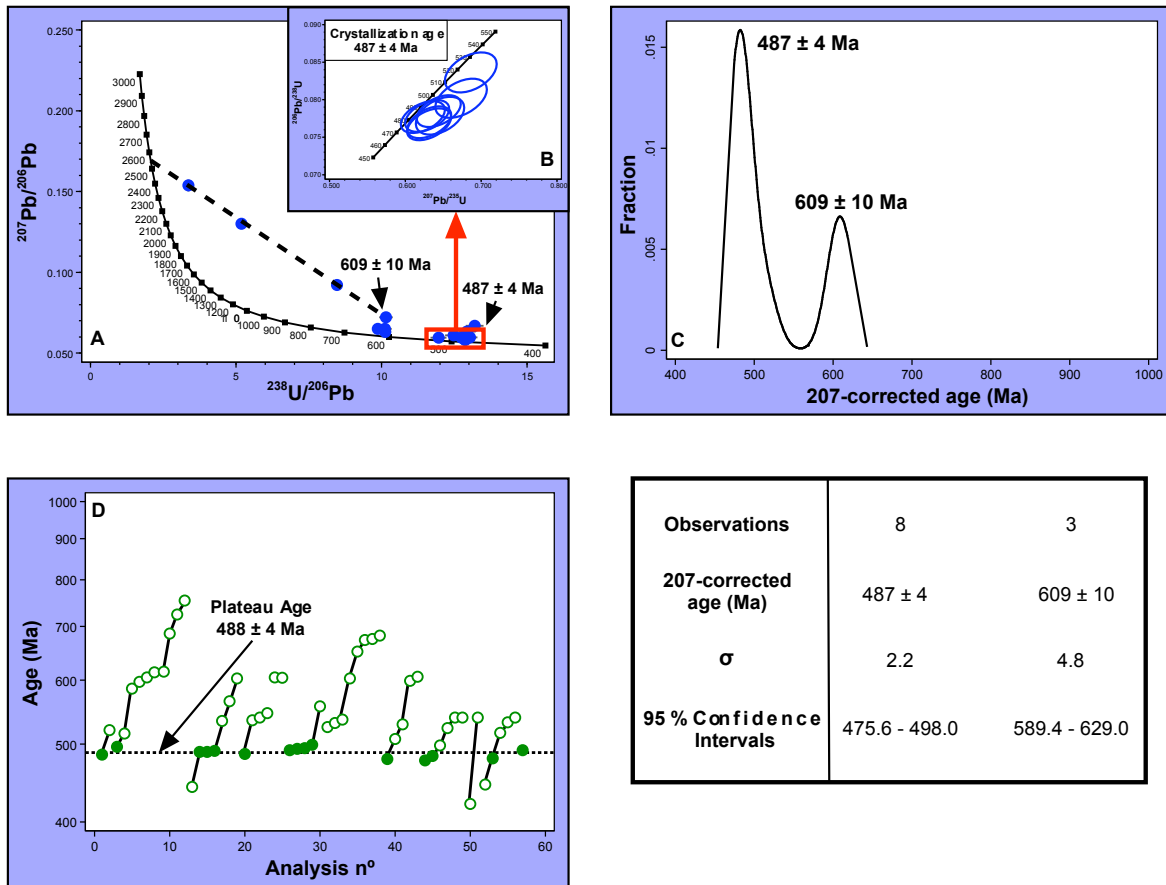


Figure 4.18: Geochronological data of the Viana do Bolo metagranite: A) Tera-Wasserburg concordia; B) Wetherill concordia; C) Distribution of concordant 207-corrected ages and the statistical data of the two most abundant populations; D) Pb-Pb stepwise evaporation age.

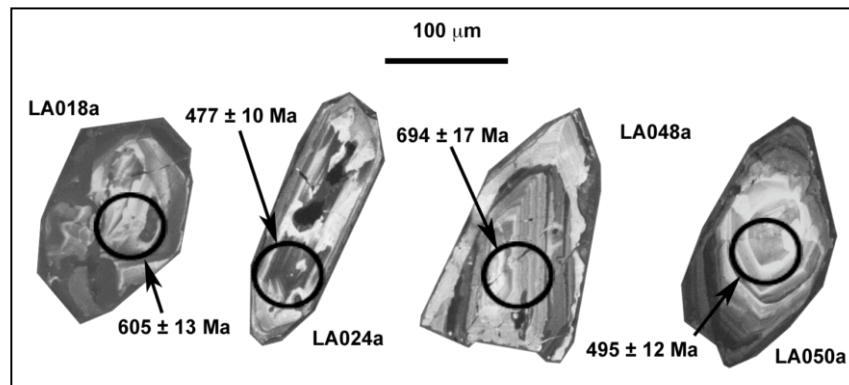


Figure 4.19: Cathodoluminescence images and ages of selected zircons from the Viana do Bolo metagranite.

Thirteen zircon grains were analyzed by the Pb-Pb stepwise evaporation method. Fifteen out of 58 evaporation steps provided a plateau age of 488 ± 4 Ma (Fig. 4.18D), which is practically identical to the youngest U-Pb concordant population. We consider this value, therefore, as the best estimate of the crystallization age of the Viana do Bolo metagranite. Again, the fraction of grains with a pre-magmatic core (minimum age ≈ 540 -600 Ma) is remarkably elevated.

4.1.2.2. San Sebastián

The San Sebastián metagranite is spatially associated with the Viana do Bolo metagranite and the volcanic rocks of Sanabria. In contrast with them, the San Sebastián metagranite has little zircon inheritance, being to date, so far, the only known case of Central Iberian Cambro-Ordovician rocks with a percentage of zircon grains with pre-magmatic core less than 20%. Not surprisingly, therefore, Lancelot et al. (1985) using zircon concentrate analyses were able to determine an upper intercept at 465 ± 10 Ma. We studied this metagranite by U-Pb ion microprobe (Table 4.17), and found that most grains

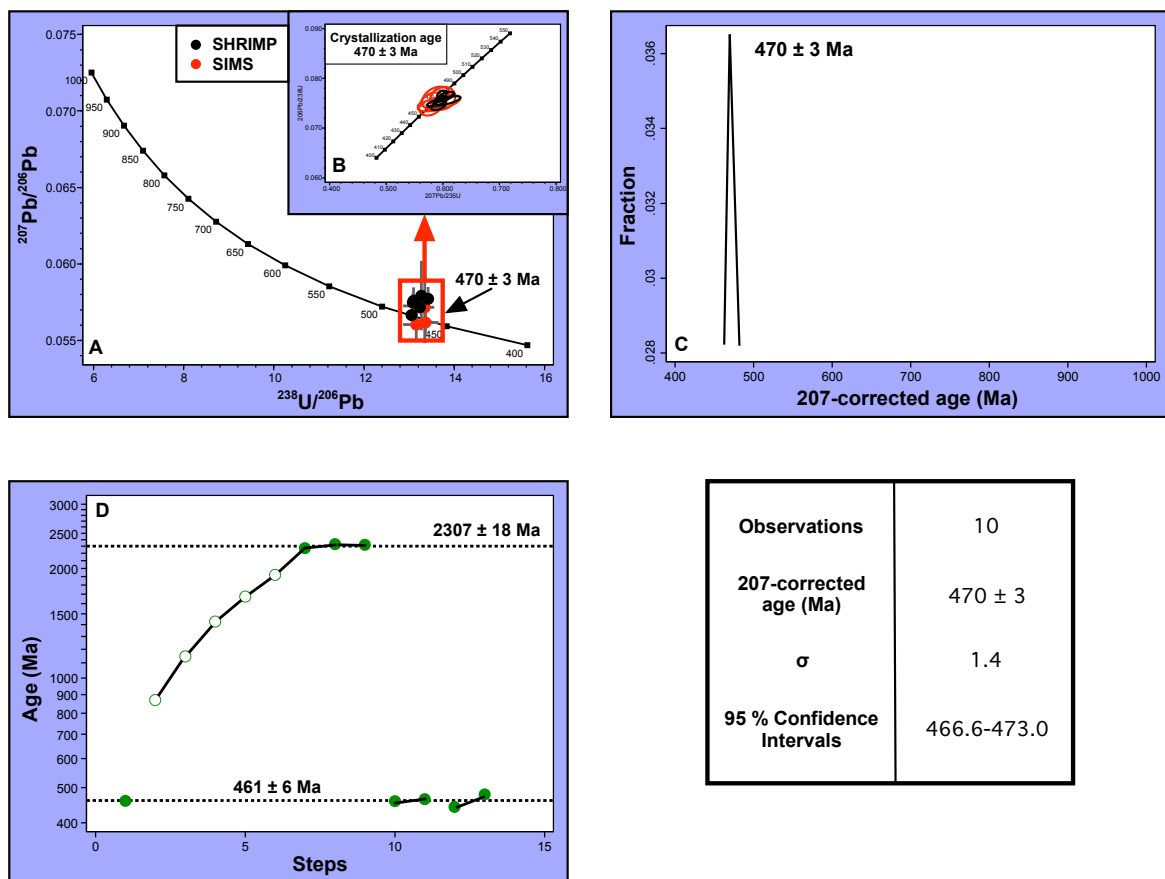


Figure 4.20: Geochronological data of the San Sebastián metagranite: A) Tera-Wasserburg concordia; B) Wetherill concordia; C) Distribution of concordant 207-corrected ages and the statistical data of the crystallization population; D) Pb-Pb stepwise evaporation age.

are concordant with an average 207-corrected age of 470 ± 3 Ma (Figs. 4.20A, 4.20B and 4.20C). The San Sebastián metagranite, therefore, is the youngest dated Cambro-Ordovician magmatic body of the Ollo de Sapo formation, about 15 Ma younger than any other from the Sanabria region.

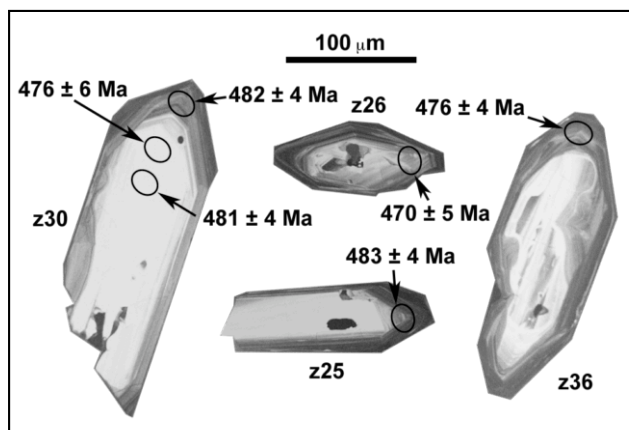


Figure 4.21: Cathodoluminescence images and ages of selected zircons from the San Sebastián metagranite.

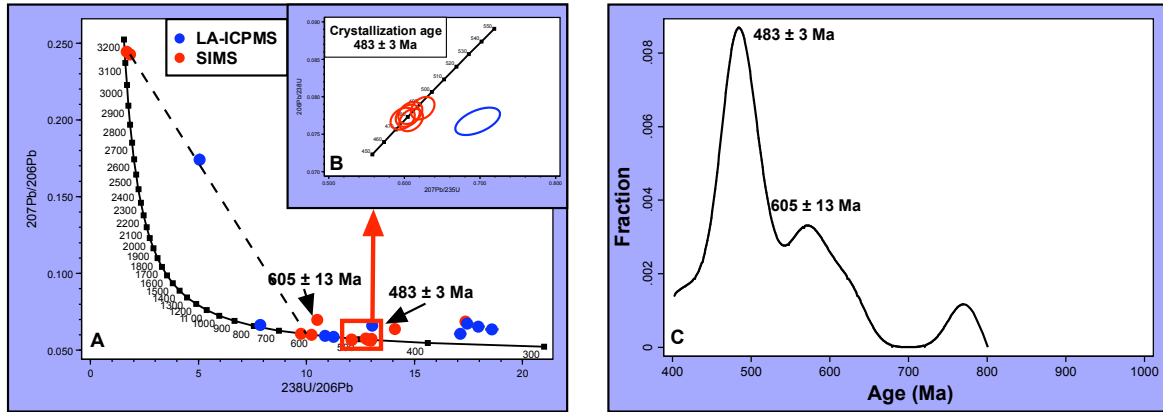
Pb-Pb data for four zircons (Table 4.18) yielded a badly defined plateau at 461 ± 6 Ma and a plateau corresponding to the age of the restitic core of one of the zircons at 2307 ± 18 Ma (Fig. 4.20D).

4.1.2.3. *Miranda do Douro*

The Miranda do Douro metagranite is spatially associated with the metavolcanic rocks of the Villadepera region. It has been dated by U-Pb ion microprobe and LA-ICPMS obtaining two main concordant or nearly concordant populations (Figs. 4.22A and 4.22C and Table 4.19). The first and most abundant population is Cambro-Ordovician with a mean age of 483 ± 3 Ma which represents, therefore, the age of crystallization (Fig. 4.22B). The second most abundant population is Ediacaran, with a mean age of 605 ± 13 Ma. The latter represents the main inherited population. There are also other younger and highly discordant points at about 350 Ma, and a few older discordant ages which yielded a discordia line with an upper interception at $3.167 \text{ Ma} +57/-48 \text{ Ma}$.

The crystallization ages were acquired in uniform grains (Fig. 4.23z10) and in rims of crystals that contain premagmatic cores. The Ediacaran population comprises a few grains with uniform ages and, more often, cores of crystals rimmed by large Cambro-Ordovician

overgrowths (Fig. 4.23z18). As in previous cases, younger-than-crystallization ages are mostly found in narrow discontinuous rims. Regarding the older ages, one corresponds to a core rimmed by an Ordovician overgrowth and the other two are the rim and the core of a single Archean grain (Fig. 4.23z19).



Observations	207-corrected age (Ma)	σ	95 % Confidence Intervals
6	483 ± 3	1.7	477.4 - 488.9
3	603 ± 13	6.7	539.8 - 666.4

Figure 4.22: Geochronological data of the Miranda do Douro metagranite: A) Tera-Wasserburg concordia; B) Wetherill concordia; C) Distribution of concordant 207-corrected ages and the statistical data of the two most abundant populations.

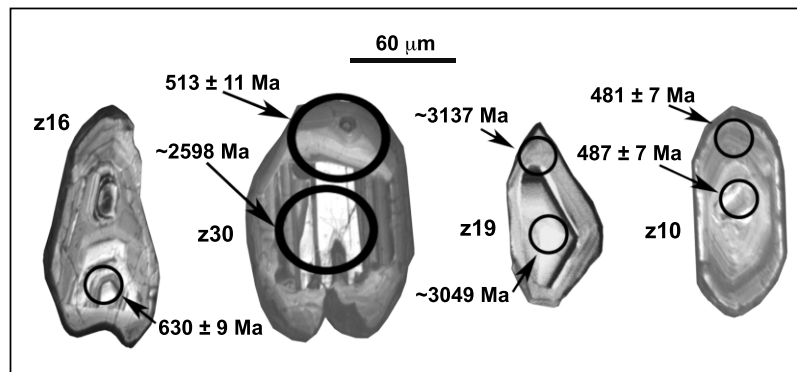


Figure 4.23: Cathodoluminescence images and ages of selected zircons from the Miranda do Douro metagranite.

4.1.2.4. Antoñita

The Antoñita metagranite is spatially associated with the metavolcanic rocks of the Hiendelaencina region. Our attempts to date the Antoñita metagranite using U-Pb either by

LA-ICPMS or ion microprobe were not successful due to high discordance. However, we succeed using the Pb-Pb stepwise evaporation method. Ten zircon grains yielded 46 evaporation steps, 19 of which define a plateau with a mean age of 474 ± 4 Ma (Table 4.20 and Fig. 4.24A). Most zircon grains also have older cores, the precise age of which cannot be estimated because of the lack of plateaus, except in one case that yielded a Neo-Archaean age of 2.68 Ga. Six grains also have younger overgrowths with ages down to 400 Ma, a value that roughly agrees with the ca. 380 Ma U-Pb lower intercept previously found by Wildberg et al. (1989). These anomalous younger ages are frequent in Hiendelaencina despite not being migmatized during the Variscan. Their significance is discussed in section 4.3.3.

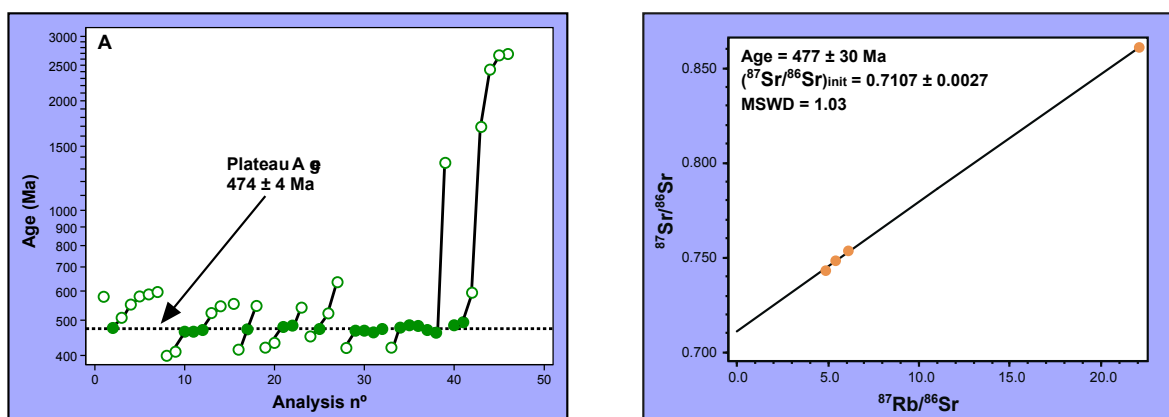


Figure 4.24: Geochronological data of the Antoñita metagranite: A) Pb-Pb stepwise evaporation age; and B) Rb-Sr isochron.

Four whole-rock samples from the Antoñita metagranite were analyzed for Sr isotopes (Table 4.21). Surprisingly, despite the intense Variscan deformation that affected them, the five samples fit well into a Rb-Sr isochron (Fig. 4.24B) that yielded an age of 477 ± 30 Ma, with $(^{87}\text{Sr}/^{86}\text{Sr})_{\text{init}} = 0.7107 \pm 0.0027$ and a MSWD = 1.03. This value is very close to the 474 Ma Pb-Pb age, and we considered it, therefore, as a good estimate of the crystallization age of the Antoñita gneiss.

4.2. The Schist-Graywacke Complex Domain

More than five hundred zircons from the metagranites and metavolcanic rocks of the northern part of the Schist-Graywacke Complex Domain have been studied using optical and cathodoluminescence imaging. These zircons show a great variety of colors: pale-pink, yellow, orange, brownish and colorless. Most are short prismatic, euhedral or subhedral and only a few are long prismatic or acicular. Their average size is about 180 x 75 μm , and the maximum up to 450 x 180 μm .

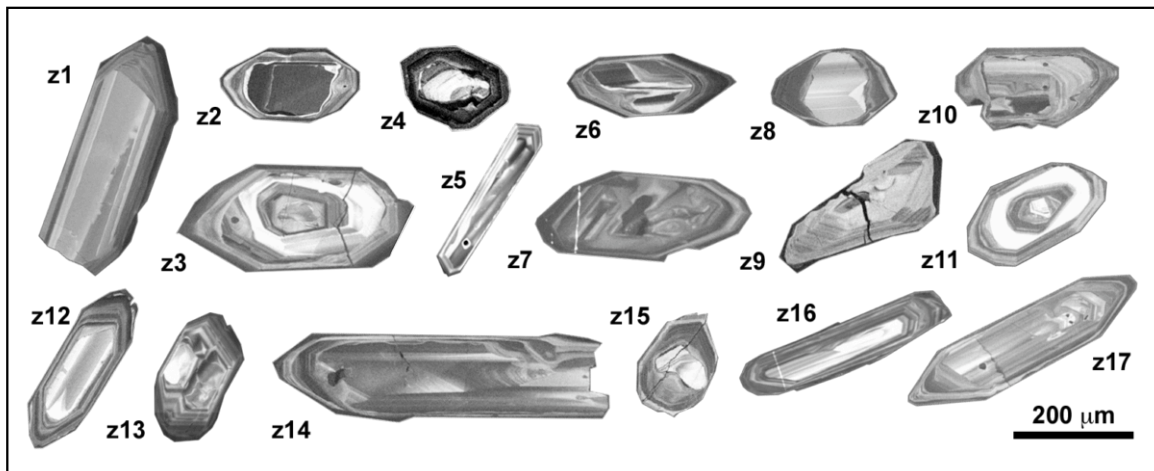


Figure 4.25: Cathodoluminescence images of the most representative internal textures of the metagranites and metavolcanic rocks of the north of the Schist-Graywacke Complex Domain.

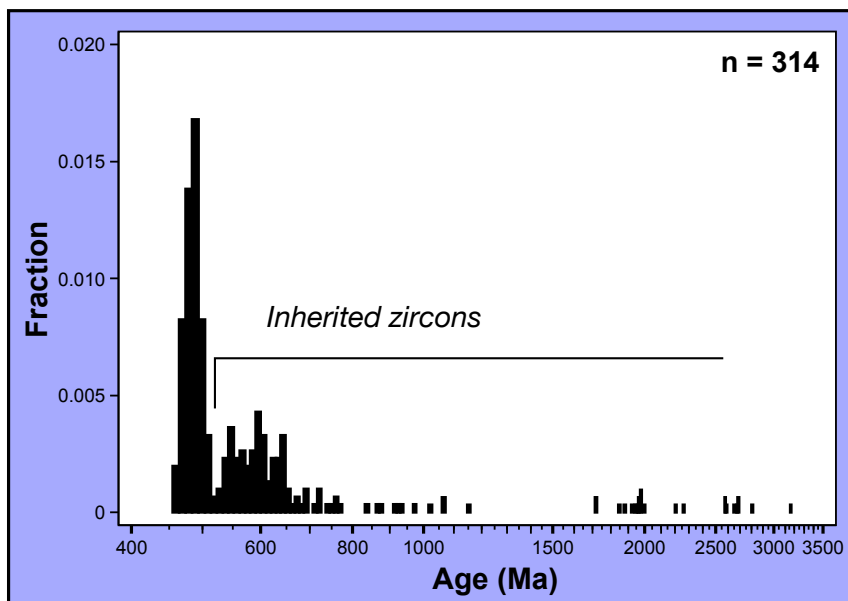


Figure 4.26: Age distribution pattern found in zircons of the metagranites and metavolcanic rocks of the north of the Schist-Graywacke Complex Domain. Built from U-Pb SIMS, SHRIMP and LA-ICPMS data.

Cathodoluminescence images reveal that most of the zircons have well-defined oscillatory zoning (Figs. 4.25z1, 4.25z5, 4.25z12, 4.25z13, 4.25z14, 4.25z16 and 4.25z17) and few also have convolute zoning (Figs. 4.25z6 and 4.25z7). Between 70 and 90 % of these zircons contain inherited cores (Fig. 4.26), except in the Sierra de Guadarrama where the percentage of inherited cores decreases up to 50-70%. These inherited cores often show a marked zoning discordant with that of the rims (Figs. 4.25z3, 4.25z4, 4.25z8, 4.25z9 and 4.25z11).

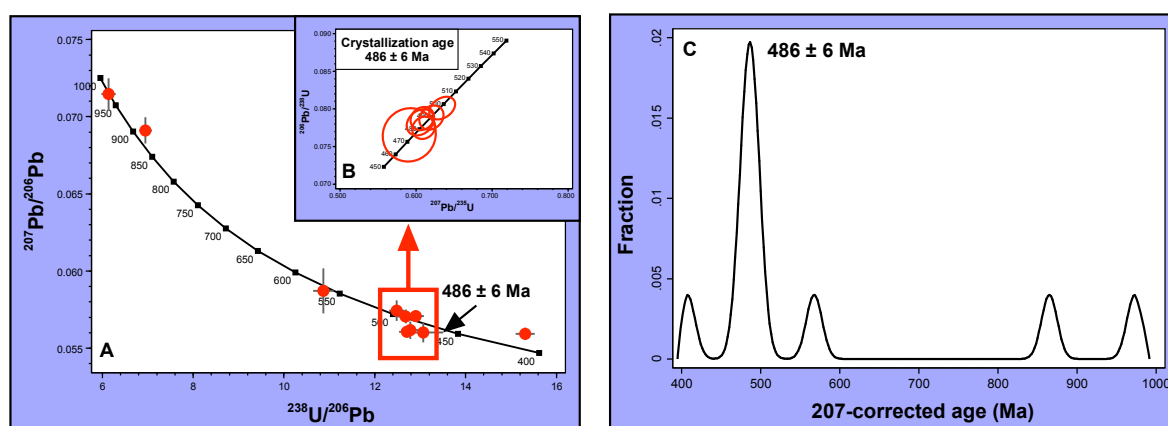
In the northern of the Schist-Graywacke Complex Domain, there are numerous small bodies of Cambro-Ordovician metagranites that define a broad band roughly parallel to the boundary with the Ollo de Sapo Formation. These rocks have been sampled in four areas: (1) the Tormes Dome, (2) the Gredos sector and (3) the Guadarrama sector of the Ávila batholith, and (4) Toledo.

4.2.1. The Tormes Dome

In the Tormes Dome, we sampled three metagranites: Fermoselle, Vitigudino and Ledesma. These metagranites have been dated using U-Pb ion microprobe.

4.2.1.1. Fermoselle

Here we obtained ten concordant to subconcordant ages (Figs. 4.27A and 4.27C and Table 4.22). Six of them are Cambro-Ordovician, with a mean age of 486 ± 6 Ma which is



Observations	207-corrected age (Ma)	σ	95 % Confidence Intervals
6	486 ± 6	3.0	478.6 - 493.8

Figure 4.27: Geochronological data of the Fermoselle metagranite: A) Tera-Wasserburg concordia; B) Wetherill concordia; and C) Distribution of concordant 207-corrected ages and the statistical data of the crystallization population.

therefore considered as the age of crystallization (Fig. 4.27B). The other three are subconcordant older points at 568 Ma, 865 Ma and 973 Ma and finally there is a younger point at approximately 407 Ma. According to the cathodoluminescence images, the crystallization ages were obtained either in uniform grains (Fig. 4.28z33) or in rims over older cores. The three older points were all obtained in the core of grains rimmed by 486 ± 6 overgrowth (Fig. 4.28z23) or in totally restitic grains (Fig. 4.28z22).

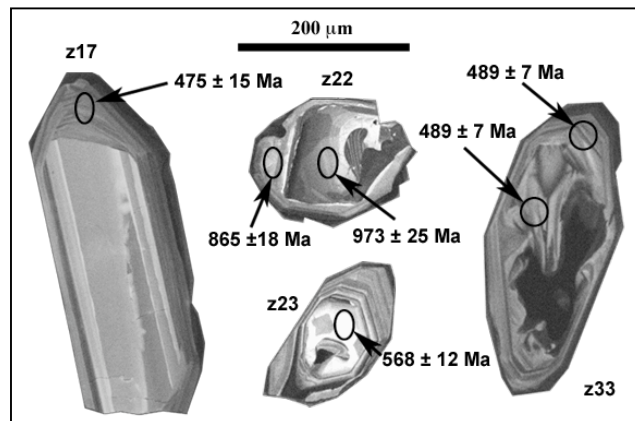
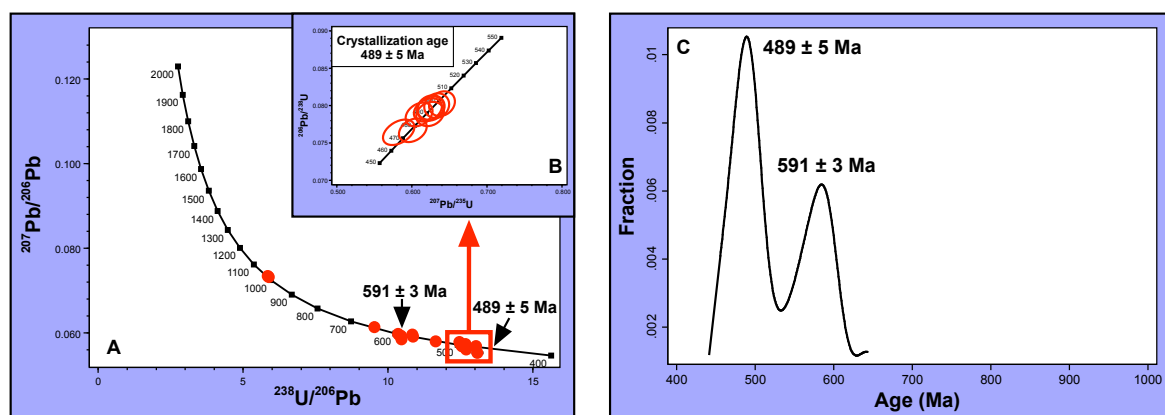


Figure 4.28: Cathodoluminescence images and ages of selected zircons from the Feroselle metagranite.

4.2.1.2. Vitigudino

The U-Pb data of the Vitigudino metagranite reveal the existence of two nearly concordant populations, one Cambro-Ordovician and the other Ediacaran, and two



Observations	207-corrected age (Ma)	σ	95 % Confidence Intervals
9	489 ± 5	2.7	483.2 - 495.7
4	591 ± 3	1.3	587.1 - 595.5

Figure 4.29: Geochronological data of the Vitigudino metagranite: A) Tera-Wasserburg concordia; B) Wetherill concordia; and C) Distribution of concordant 207-corrected ages and the statistical data of the two populations.

concordant points at ca. 1.0 Ga (Figs. 4.29A and 4.29C and Table 4.23). The Cambro-Ordovician population is the most abundant, with a mean age of 489 ± 5 Ma, which we consider the age of crystallization (Fig. 4.29B). The Ediacaran population, less numerous, has a mean age of 591 ± 3 Ma. There are also two points at 531 Ma and 643 Ma that probably resulted from analyzing the transition between the main components. The Ordovician ages were obtained in the rim of grains with inherited cores (Fig. 4.30z10) or grains with uniform ages (Fig. 4.30z1) and older ages were obtained in the cores of rimmed zircons (Figs. 4.30z7, 4.30z10 and 4.30z15).

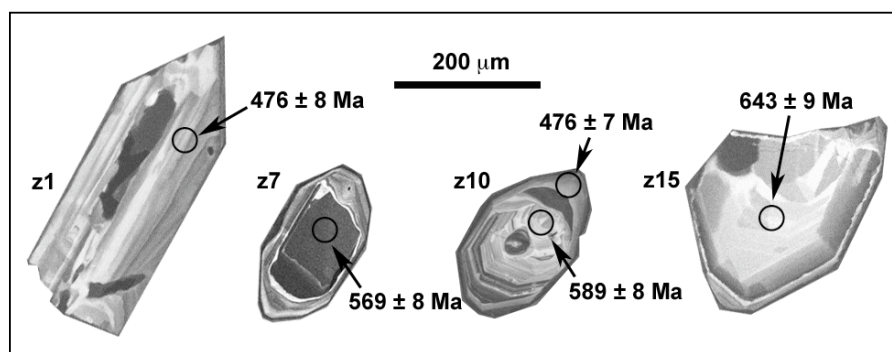
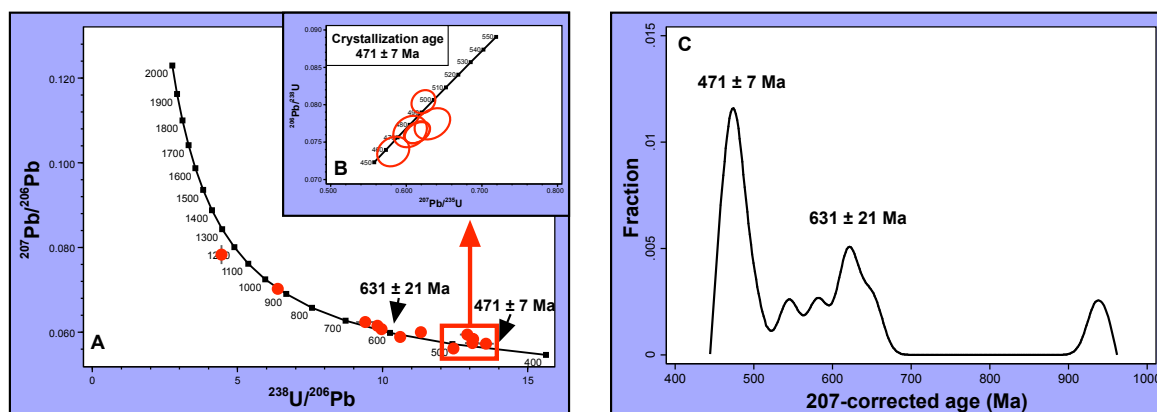


Figure 4.30: Cathodoluminescence images and ages of selected zircons from the Vitigudino metagranite.

4.2.1.3. Ledesma

The U-Pb data of the Ledesma metagranite also reveal the occurrence of two nearly



Observations	207-corrected age (Ma)	σ	95 % Confidence Intervals
5	471 ± 7	3.6	461.5 - 481.3
3	631 ± 21	10.4	585.9 - 675.7

Figure 4.31: Geochronological data of the Ledesma metagranite: A) Tera-Wasserburg concordia; B) Wetherill concordia; and C) Distribution of concordant 207-corrected ages and the statistical data of the two populations.

concordant populations, one Cambro-Ordovician and the other Ediacaran, and older ages at 938 Ma and 1156 Ma (Figs. 4.31A and 4.31C and Table 4.24). The Cambro-Ordovician population is the most abundant and the most concordant, with a mean age of 471 ± 7 Ma (Fig. 4.31B), which is therefore regarded as the crystallization age of the Ledesma metagranite. The Ediacaran population has a mean age of 631 ± 21 Ma. The two nearly concordant points at 545 and 582 Ma probably represent mixed ages resulting from analyzing an internal boundary. According to the cathodoluminescence images, whereas the Cambro-Ordovician population was obtained in rims of zircons containing inherited cores (Figs. 4.32z22, 4.32z29 and 4.32z30) and in grains with uniform ages (Fig. 4.32z17), the Ediacaran population and older grains were always obtained in the core of rimmed grains (Figs. 4.32z29 and 4.32z30).

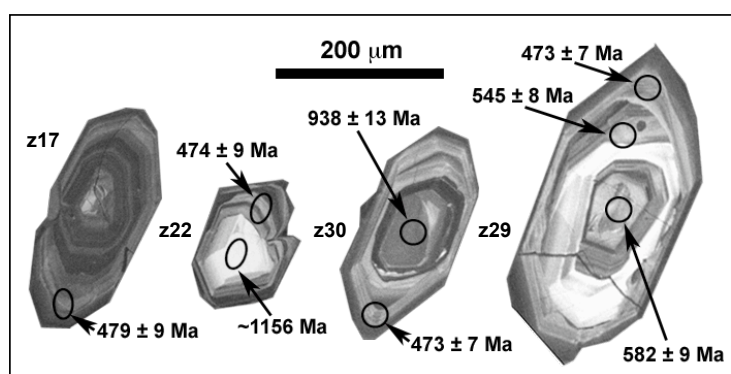


Figure 4.32: Cathodoluminescence images and ages of selected zircons from the Ledesma metagranite.

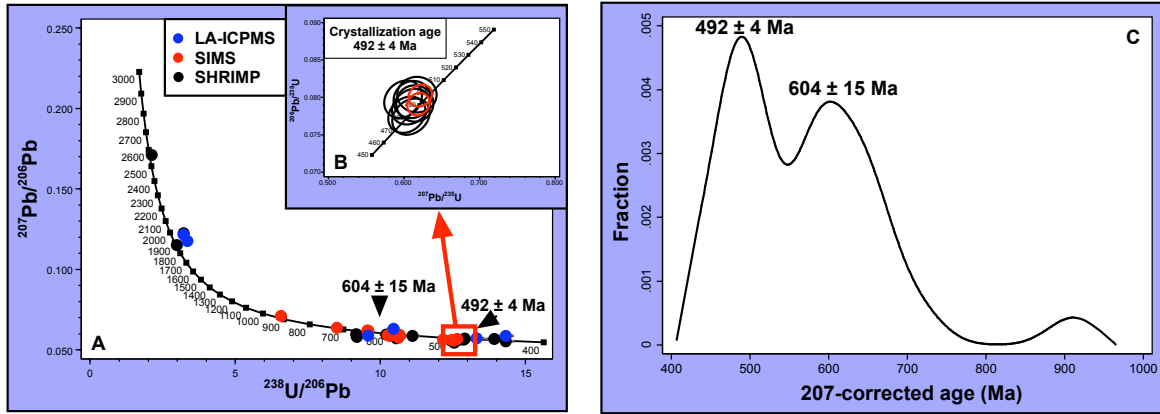
4.2.2. The Gredos Sector

In the northwest of the Sierra de Gredos, we sampled three metagranites: Bercimuelle, Castellanos and San Pelayo. The exact location and coordinates of the samples are given in Table 2.2. They have been dated using U-Pb ion microprobe and LA-ICPMS. The results are detailed as follows:

4.2.2.1. San Pelayo

As in all other cases, the San Pelayo metagranite contains two concordant populations, one Cambro-Ordovician and the other Ediacaran, and a few isolated older grains (Figs. 4.33A and 4.33C and Table 4.25). The Cambro-Ordovician population yields a mean age of 492 ± 4 Ma, which represents the age of crystallization (Fig. 4.33B). The Ediacaran population has a mean age of 604 ± 15 Ma. Besides this, there are several older concordant

ages: one point at about 717 Ma, one point at about 912 Ma, four points at about 2000 Ma and one point at about 2.6 Ga. Finally, as in previous cases, there are younger-than-crystallization ages (see discussion in section 4.3.3).



Observations	207-corrected age (Ma)	σ	95 % Confidence Intervals
9	492 ± 4	2.2	486.5 - 496.6
11	604 ± 15	7.6	586.8 - 620.4

Figure 4.33: Geochronological data of the San Pelayo metagranite: A) Tera-Wasserburg concordia; B) Wetherill concordia; and C) Distribution of concordant 207-corrected ages and the statistical data of the two most abundant populations.

The Cambro-Ordovician population comes from rims of grains with inherited cores (Fig. 4.34z28) and, less frequently, from uniform zircons. The Ediacaran population comes either from cores of rimmed grains or from totally restitic zircons (Fig. 4.34z27 and 4.34z30). The oldest ages were obtained, in all cases, from cores of zircons rimmed by Cambro-Ordovician overgrowths (Fig. 4.34z29).

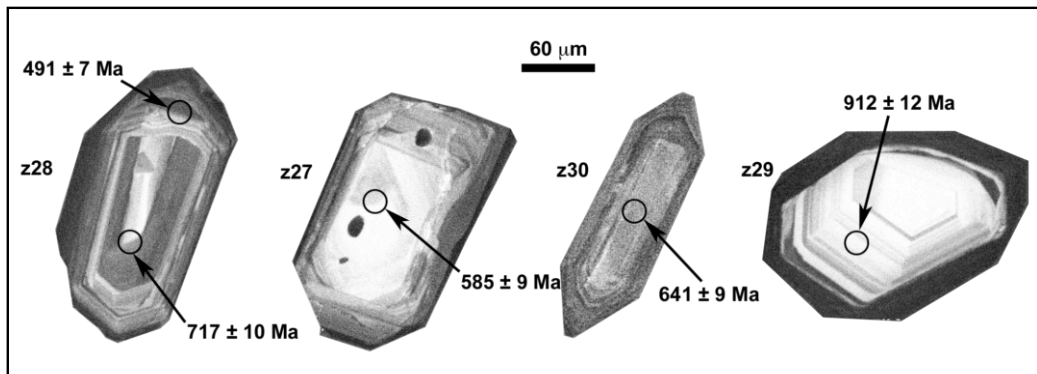


Figure 4.34: Cathodoluminescence images and ages of selected zircons from the San Pelayo metagranite.

4.2.2.2. *Castellanos*

The U-Pb data of the Castellanos metagranite demonstrate the existence of three main concordant populations (Figs. 4.35A and 4.35C and Table 4.26). The most abundant is also the youngest; it is Cambro-Ordovician, with a mean age of 498 ± 4 Ma, and represents the age of crystallization of this body (Fig. 4.35B). The second most abundant population, in contrast to all previous bodies, is not Ediacaran but Early Cambrian, with a mean age of 548 ± 10 Ma. The third is Ediacaran, with a mean age of 612 ± 10 Ma. Besides this, we found some other concordant or nearly concordant points: one at 683 Ma, another at 1.8 Ga and another about 3.2 Ga.

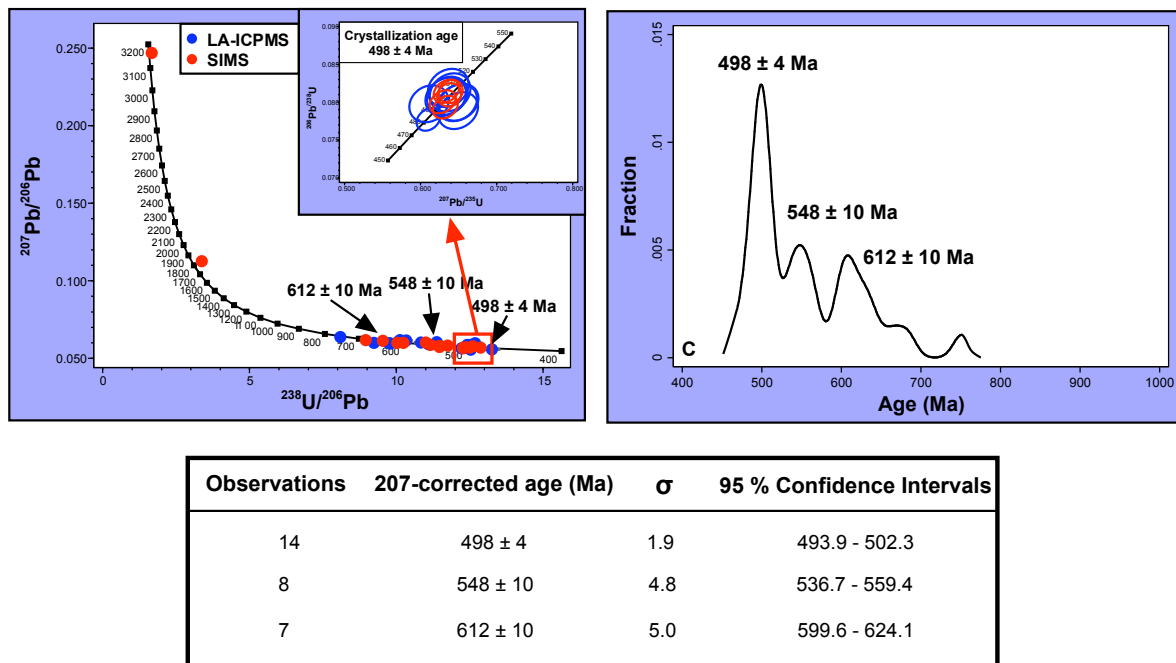


Figure 4.35: Geochronological data of the Castellanos metagranite: A) Tera-Wasserburg concordia; B) Wetherill concordia; and C) Distribution of concordant 207-corrected ages and the statistical data of the three most abundant populations.

According to the cathodoluminescence images, the 498 ± 4 Ma population comes either from the rims of zircons that have inherited cores, or from uniform grains (Fig. 4.36z20). The 548 ± 10 Ma population also comes either from cores or rims of older grains (Fig. 4.36z11). Remarkably this is the crystallization age of the neighboring Almohalla orthogneiss (543 ± 6 Ma, Bea et al., 2003). The older ages were acquired in inherited cores and totally restitic zircons (Figs. 4.36z11, 4.36 z14 and 4.36 15).

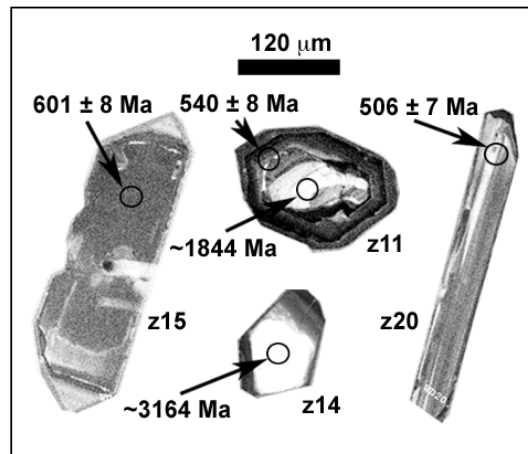
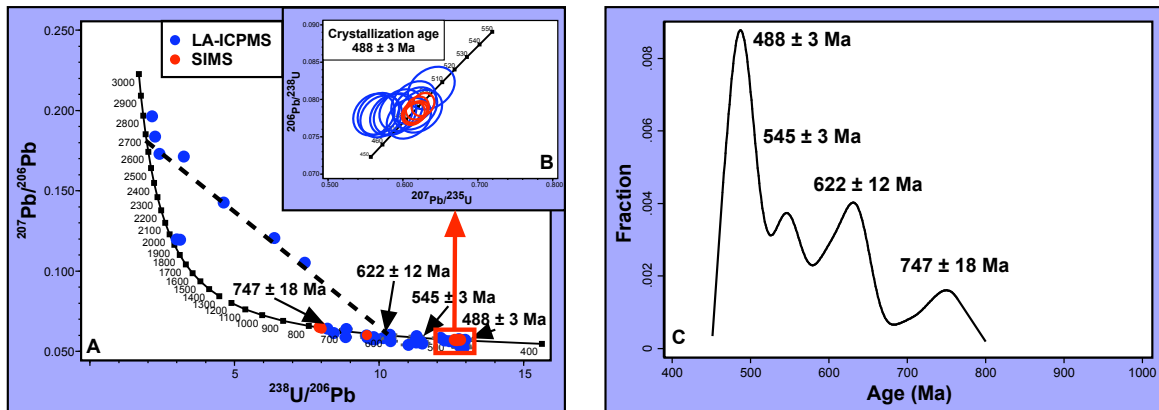


Figure 4.36: Cathodoluminescence images and ages of selected zircons from the Castellanos metagranite.

4.2.2.3. Bercimuelle

The U-Pb ion microprobe data of the Bercimuelle metagranite reveals a highly complex age distribution, with four concordant populations (Figs. 4.37A and 4.37C and Table 4.27). The youngest is Cambro-Ordovician, with a mean age of 488 ± 3 Ma (Fig. 4.37B). It is preceded by another population at 545 ± 3 Ma, which as in the Castellanos metagranite, is coincident with the age of the neighboring Almohalla orthogneiss. The third is Ediacaran,



Observations	207-corrected age (Ma)	σ	95 % Confidence Intervals
16	488 ± 3	1.4	485.1 - 491.1
6	545 ± 3	1.5	540.9 - 548.7
11	622 ± 12	6.0	608.3 - 634.9
4	747 ± 18	8.9	718.8 - 775.3

Figure 4.37: Geochronological data of the Bercimuelle metagranite: A) Tera-Wasserburg concordia; B) Wetherill concordia; and C) Distribution of concordant 207-corrected ages and the statistical data of the four most abundant populations.

with a mean age of 622 ± 12 Ma. And the fourth is Cryogenian, clustering around 747 ± 18 Ma. Besides this, there are some data clustering around 2.0 Ga, and a discordia line with upper interception close to 2.7 Ga, all very similar to what was found in the metavolcanic rocks of the Ollo de Sapo Formation.

According to cathodoluminescence imaging, crystallization ages were obtained in uniform grains (Fig. 4.38z7) and in rims of grains with inherited cores (Fig. 4.38z3). The other Cambrian, Ediacaran and Cryogenian populations were obtained in inherited cores and, rarely, in rims of grains with inherited cores (Fig. 4.38z2).

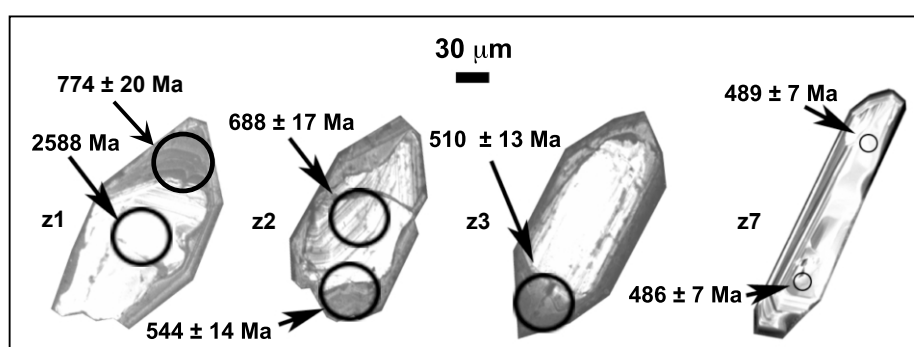


Figure 4.38: Cathodoluminescence images and ages of selected zircons from the Bercimuelle metagranite.

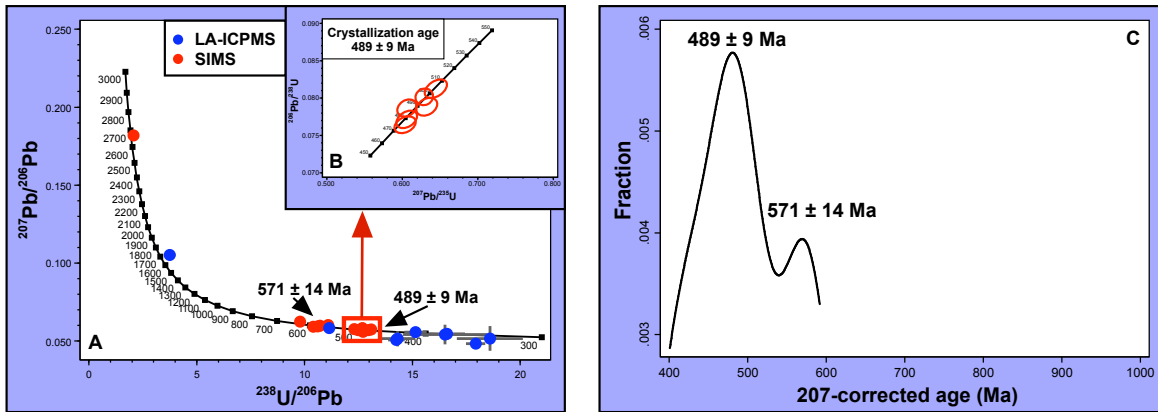
4.2.3. The Guadarrama Sector

In the Sierra de Guadarrama, where the largest accumulation of Cambro-Ordovician granitoids of Iberia occur, we sampled the following orthogneisses of probable metagranitic origin: La Cañada, La Hoya, La Estación I, La Estación II, Vegas de Matute, El Caloco, La Morcuera and Buitrago de Lozoya. The exact location and coordinates of the samples are given in Table 2.2. We also sampled three orthogneisses that probably derived from metavolcanic rocks: La Berzosa, El Cardoso and Riaza. All of them were dated using U-Pb ion microprobe and LA-ICPMS. The results are described below:

4.2.3.1. La Cañada

U-Pb ion microprobe and LA-ICPMS data from La Cañada metagranite display the presence of two concordant or nearly concordant populations and some isolated ages (Figs. 4.39A and 4.39C and Table 4.28). The first population is Cambro-Ordovician and yields a mean age of 489 ± 9 Ma that represents the age of crystallization of this body (Fig. 4.39B). The second population is Ediacaran, with a mean age of 571 ± 14 Ma. There are two older

concordant ages, at 1718 and 2670 Ma and, again, several younger-than-crystallization ages between about 338 and 440 Ma.



Observations	207-corrected age (Ma)	σ	95 % Confidence Intervals
6	489 ± 9	4.4	477.1 - 500.0
5	571 ± 14	7.2	551.3 - 591.4

Figure 4.39: Geochronological data of La Cañada metagranite: A) Tera-Wasserburg concordia; B) Wetherill concordia; and C) Distribution of concordant 207-corrected ages and the statistical data of the two most abundant populations.

The crystallization ages were usually obtained in the rims of grains containing inherited cores (Figs. 4.40z45 and 4.40z49) and, rarely, in uniform grains (Fig. 4.40z41). The inherited population was always obtained in cores of zircons rimmed by Cambro-Ordovician overgrowths (Figs. 4.40z45 and 4.40z50), as well as the two older ages (Fig. 4.40z49). Younger-than-crystallization ages were, in all cases, obtained in the outermost part of some grains.

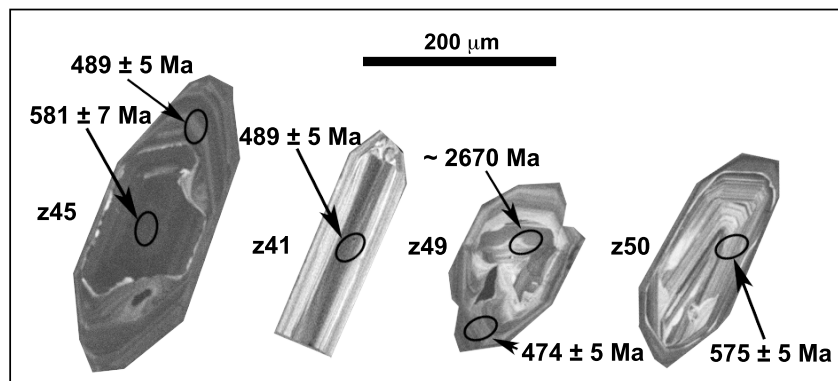
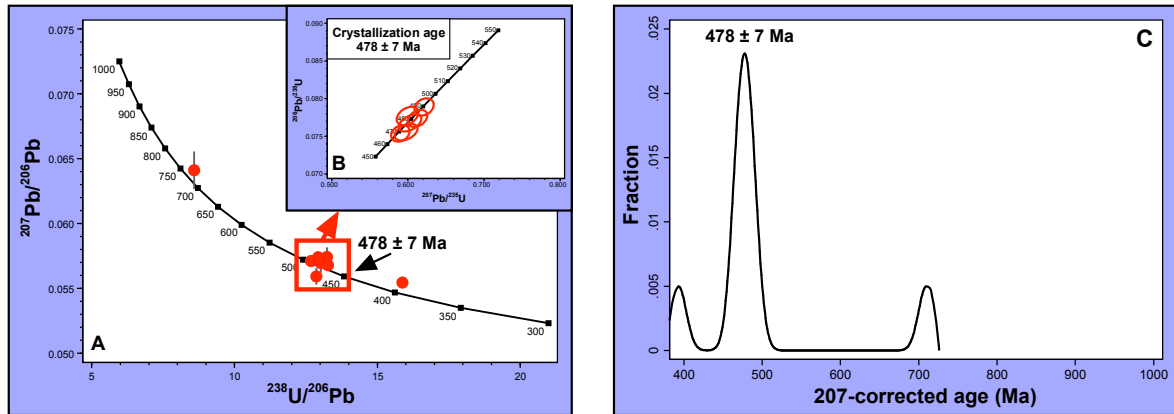


Figure 4.40: Cathodoluminescence images and ages of selected zircons from La Cañada metagranite.

4.2.3.2. La Hoya

The U-Pb ion microprobe data of La Hoya metagranite reveal the existence of a Cambro-Ordovician nearly concordant population with a mean age of 478 ± 7 Ma that we interpret as representing the age of crystallization (Fig. 4.41 and Table 4.29). There is also a subconcordant point at 710 Ma and a younger-than-crystallization age at 393 Ma.



Observations	207-corrected age (Ma)	σ	95 % Confidence Intervals
6	478 ± 7	3.4	468.9 - 486.4

Figure 4.41: Geochronological data of La Hoya metagranite: A) Tera-Wasserburg concordia; B) Wetherill concordia; and C) Distribution of concordant 207-corrected ages and the statistical data of the crystallization population.

According to the cathodoluminescence images, the crystallization population was obtained in uniform grains (Fig. 4.42z2) and, less frequently, in rims of zircons with older cores (Fig. 4.42z16). The 710 Ma age was obtained in a core of a grain with a Cambro-Ordovician rim (Fig. 4.42z16).

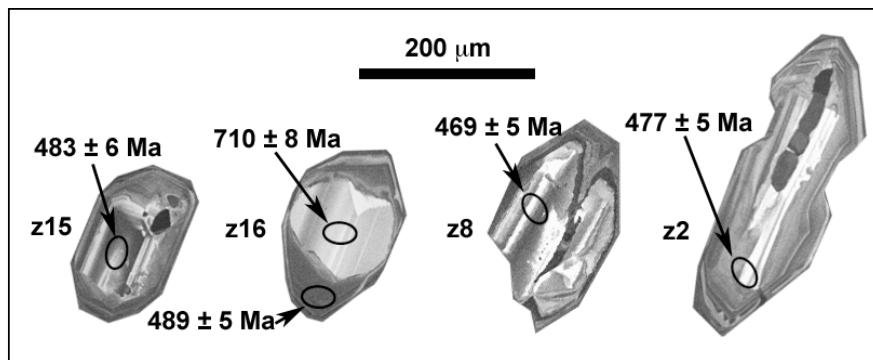
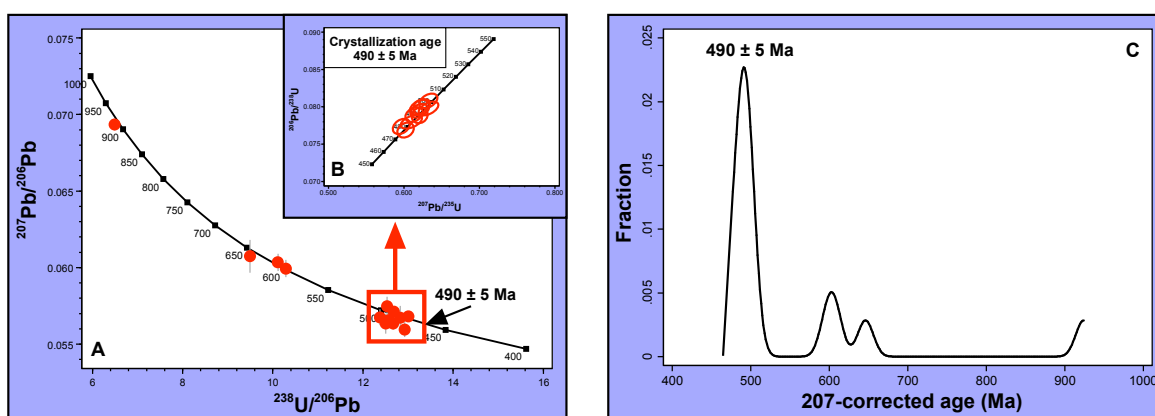


Figure 4.42: Cathodoluminescence images and ages of selected zircons from La Hoya metagranite.

4.2.3.3. La Estación

In La Estación I metagranite, the U-Pb data reveal the existence of a concordant and abundant population and few older ages (Fig. 4.43A and 4.43C and Table 4.30). The concordant population is Cambro-Ordovician, with a mean age of 490 ± 5 Ma, which represents the age of crystallization of this metagranite (Fig. 4.43B). There are also two concordant points at about 600 Ma, another point at 646 and another, also concordant, at about 924 Ma.



Observations	207-corrected age (Ma)	σ	95 % Confidence Intervals
10	490 ± 5	2.4	484.9 - 496.8

Figure 4.43: Geochronological data of La Estación I metagranite: A) Tera-Wasserburg concordia; B) Wetherill concordia; and C) Distribution of concordant 207-corrected ages and the statistical data of the crystallization population.

The Cambro-Ordovician population was obtained in uniform grains (Fig. 4.44z30) and in rims of grains containing older cores (Figs. 4.44z29, 4.44z33 and 4.44z36). All the older ages were obtained in cores of zircons rimmed by Cambro-Ordovician overgrowths (Figs. 4.44z29, 4.44z33 and 4.44z36).

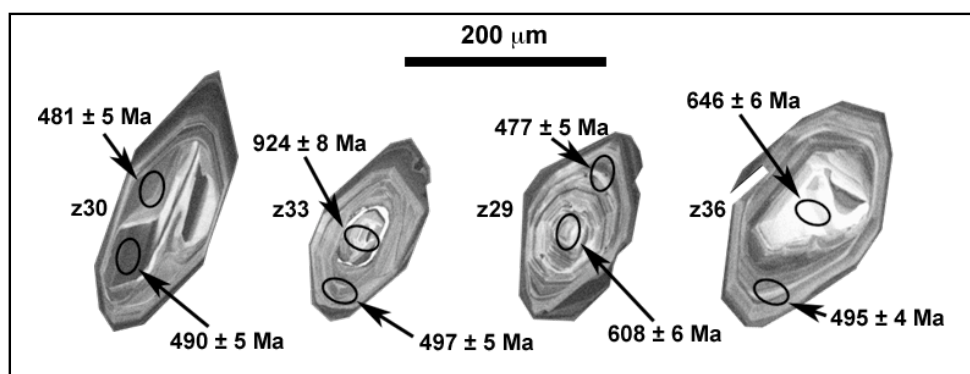
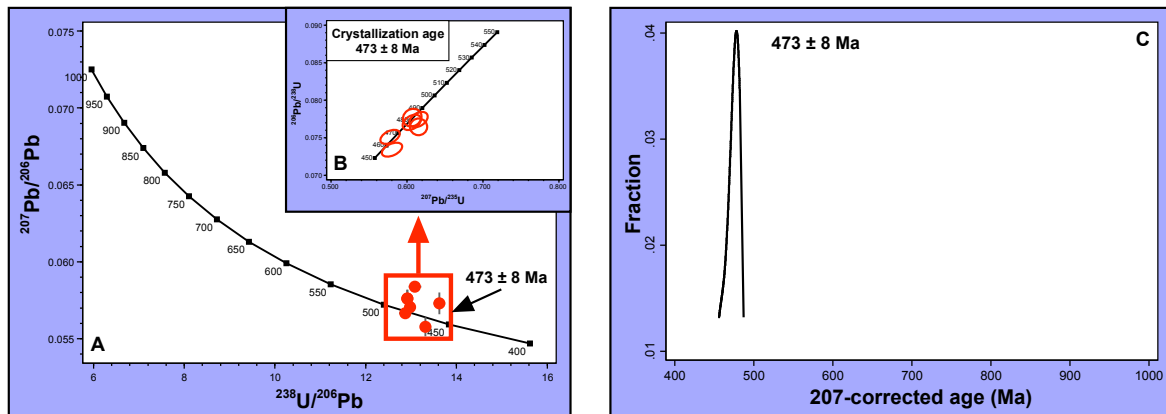


Figure 4.44: Cathodoluminescence images and ages of selected zircons from La Estación I metagranite.

4.2.3.4. La Estación II

In La Estación II metagranite, the U-Pb data from five zircons reveal the presence of a nearly concordant population, Cambro-Ordovician (Figs. 4.45A and 4.45C and Table 4.31), with a mean age of 473 ± 8 Ma, which seems to be the best estimate of the age of crystallization of this body (Fig. 4.45B). No older inherited cores were found, nevertheless, according to the cathodoluminescence images, there are inherited cores (Fig. 4.46).



Observations	207-corrected age (Ma)	σ	95 % Confidence Intervals
6	473 ± 8	4.0	465.2-483.3

Figure 4.45: Geochronological data of La Estación II metagranite: A) Tera-Wasserburg concordia; B) Wetherill concordia; and C) Distribution of concordant 207-corrected ages and the statistical data of the crystallization population.

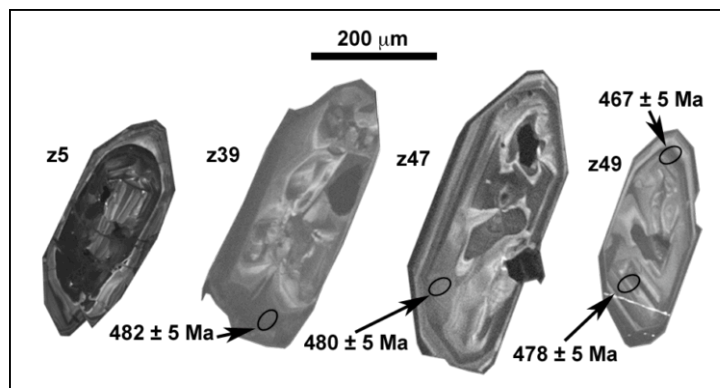
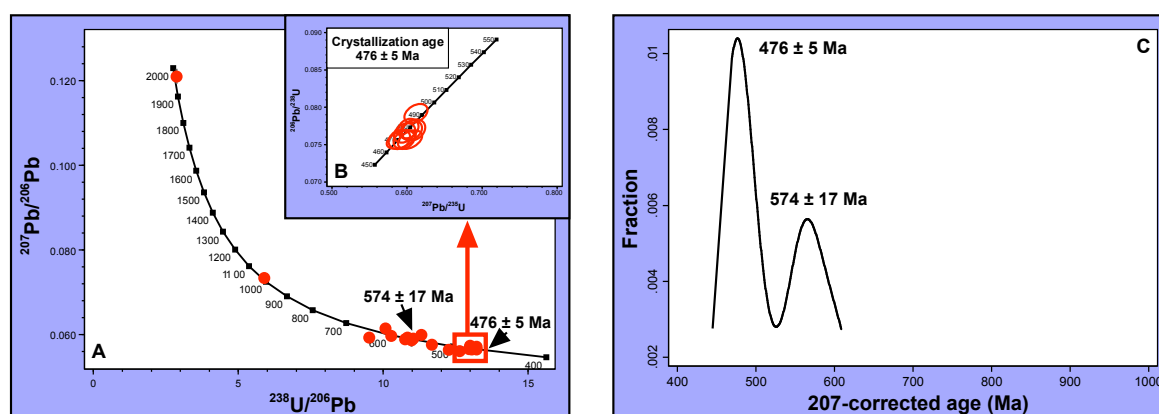


Figure 4.46: Cathodoluminescence images and ages of selected zircons from La Estación II metagranite.

According to the cathodoluminescence images, all crystallization ages were obtained in uniform grains (Figs. 4.46z39, 4.46z47 and 4.46z49), but some images also reveal the existence of inherited cores although they have not appeared in the analyzed zircons (Fig. 4.46z5).

4.2.3.5. Vegas de Matute (Melanocratic metagranites)

The U-Pb ion microprobe data from the Vegas de Matute metagranite reveal the presence of one concordant population, another nearly concordant one and two isolated points (Figs. 4.47A and 4.47C and Table 4.32). The first population is Cambro-Ordovician, with a mean age of 476 ± 5 Ma which represents the age of crystallization (Fig. 4.47B). The second population, which is less abundant than the previous, is Ediacaran ranging from 545 to 609 Ma with a mean age of 574 ± 17 Ma. Between these two population there is a point with an age of 530 Ma which, although concordant, can result from the mixture of the last two concordant components. Apart from that, there are two older ages, one at 1024 Ma and the other at 1971 Ma.



Observations	207-corrected age (Ma)	σ	95 % Confidence Intervals
9	476 ± 5	2.4	470.2 - 481.2
7	574 ± 17	8.5	552.7 - 594.4

Figure 4.47: Geochronological data of the Vegas de Matute (Melanocratic metagranites): A) Tera-Wasserburg concordia; B) Wetherill concordia; and C) Distribution of concordant 207-corrected ages and the statistical data of the two populations.

As usual, the crystallization ages were obtained in rims of crystals containing older cores (Figs. 4.48z48 and 4.48z50) and also in uniform crystals (Fig. 4.48z57). The inherited population (545-609 Ma) was always obtained in cores of crystals rimmed by

Cambro-Ordovician overgrowths (Fig. 4.48z48). The two older ages were also obtained in restitic cores (Figs. 4.48z50 and 4.48z54) rimmed by Cambro-Ordovician overgrowths.

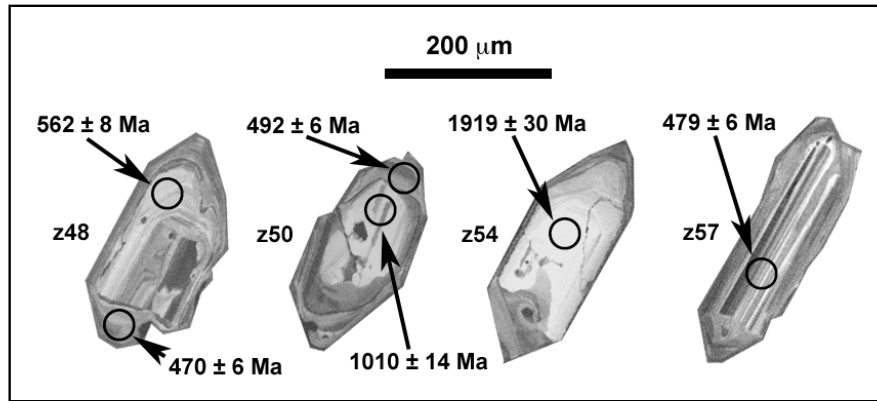
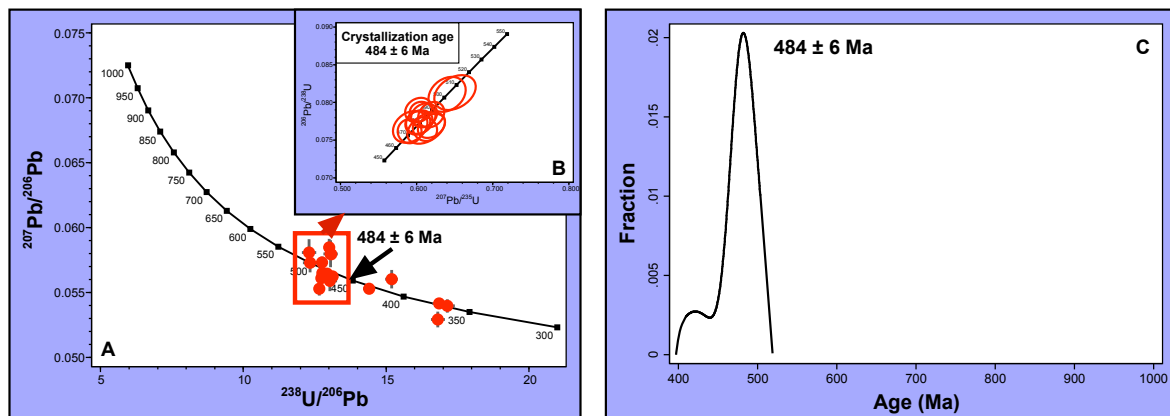


Figure 5.48: Cathodoluminescence images and ages of selected zircons from the Vegas de Matute (Melanocratic metagranites).

4.2.3.6. Vegas de Matute (Leucocratic metagranites: El Caloco)

In the leucocratic El Caloco metagranite we found a single population, Cambro-Ordovician (Figs. 4.49A and 4.49C and Table 4.33), with a mean age of 484 ± 6 Ma (Fig. 4.49B) which represents the crystallization age. Most grains have narrow Variscan rims caused by the intense migmatization that affects this body. Owing to the thinness of these rims, the ages close to 360 Ma shown in Fig. 4.49A must be considered maximum estimates.



Observations	207-corrected age (Ma)	σ	95 % Confidence Intervals
13	484 ± 6	2.9	477.9 - 490.1

Figure 4.49: Geochronological data of the Vegas de Matute (Leucocratic metagranites): A) Tera-Wasserburg concordia; B) Wetherill concordia; and C) Distribution of concordant 207-corrected ages and the statistical data of the crystallization population.

No inherited cores were found and all data were acquired in uniform zircons (Figs. 4.50z33, 4.50z43 and 4.50z45). Despite this, the cathodoluminescence images reveal the presence of inherited cores (Fig. 4.50z4).

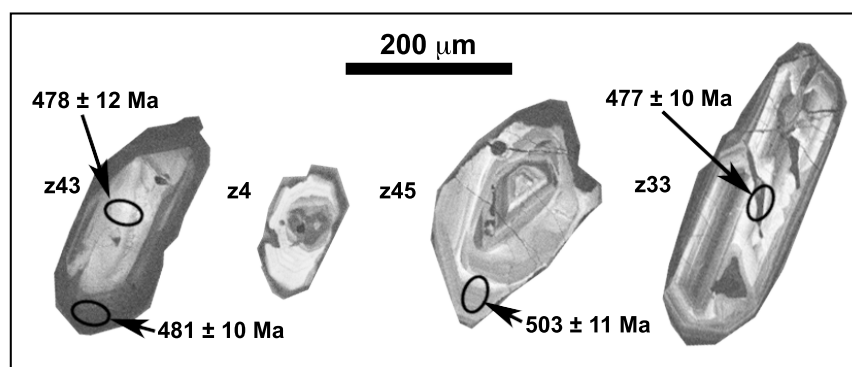
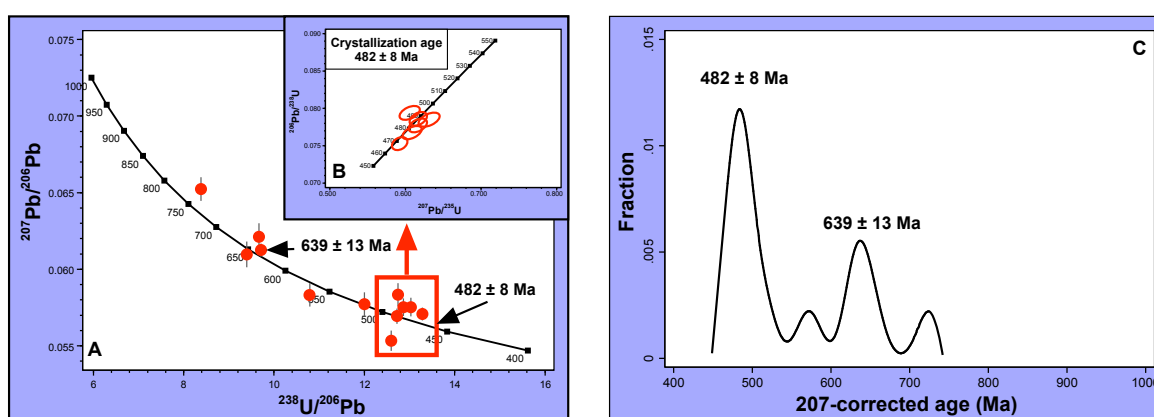


Figure 4.50: Cathodoluminescence images and ages of selected zircons from the Vegas de Matute (Leucocratic metagranites).

4.2.3.7. La Morcuera

The U-Pb ion microprobe data from La Morcuera metagranite reveal the presence of two nearly concordant populations (Figs. 4.51A and 4.51C and Table 4.34). The first one, Cambro-Ordovician, yields a mean age of 482 ± 8 Ma that represents the best estimate of the crystallization age (Fig. 4.51B). The second population, Ediacaran, yields a mean age of 639 ± 13 Ma. Between these two populations are two nearly concordant point at 516 and 572 Ma, which probably represent mixed ages between the two populations mentioned before. There is also another nearly concordant point at about 725 Ma.



Observations	207-corrected age (Ma)	σ	95 % Confidence Intervals
6	482 ± 8	3.8	472.5 - 491.9
3	639 ± 13	6.7	610.3 - 667.8

Figure 5.51: Geochronological data of La Morcuera metagranite: A) Tera-Wasserburg concordia; B) Wetherill concordia; and C) Distribution of concordant 207-corrected ages and the statistical data of the two populations.

According to the cathodoluminescence images, the Cambro-Ordovician population is present in the rims of grains having inherited cores (Figs. 4.52z48 and 4.52z54) and in uniform grains (Fig. 4.52z57). The inherited population was, in all cases, acquired in the core of crystals rimmed by Cambro-Ordovician overgrowths (Figs. 4.52z48, 4.52z50 and 4.52z54).

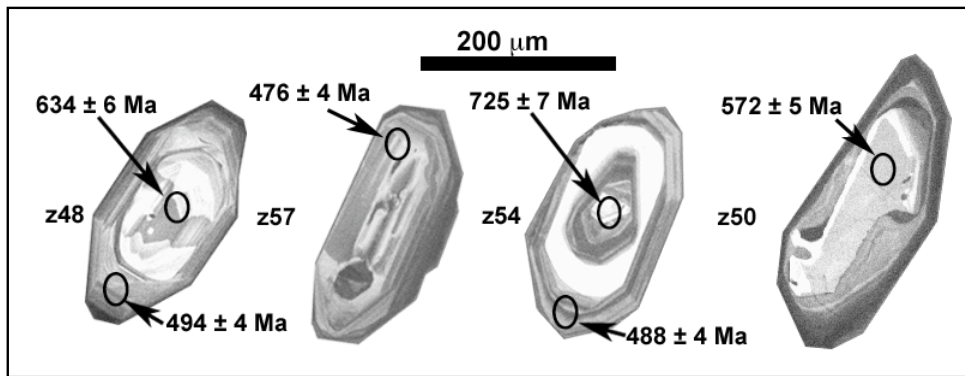
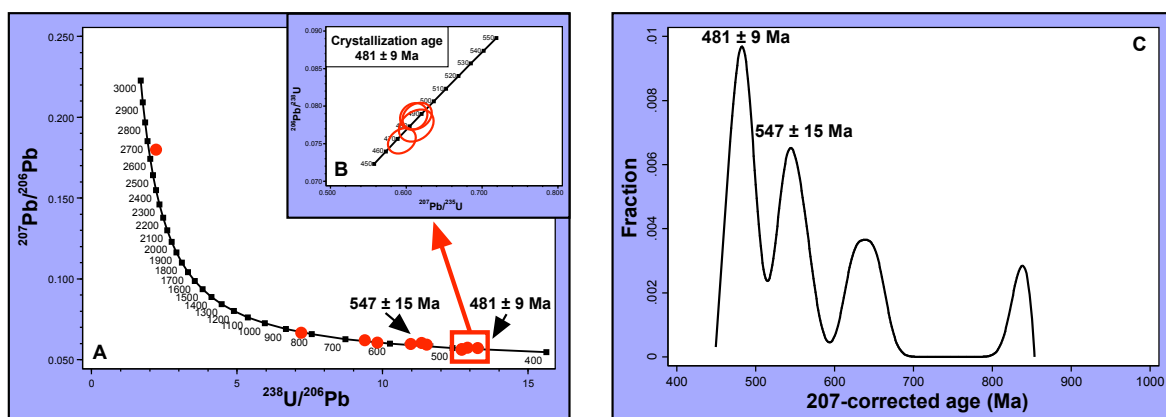


Figure 4.52: Cathodoluminescence images and ages of selected zircons from La Morcuera metagranite.

The 482 ± 8 Ma age matches the crystallization age obtained by Valverde-Vaquero and Dunning (2000) using conventional U-Pb dating on zircon concentrates (477 ± 4 Ma).

4.2.3.8. Buitrago de Lozoya

In the Buitrago de Lozoya metagranite, the U-Pb ion microprobe data reveal the occurrence of two concordant populations and few isolated ages (Figs. 4.53A and 4.53C



Observations	207-corrected age (Ma)	σ	95 % Confidence Intervals
4	481 ± 9	4.8	465.8 - 496.6
3	547 ± 15	7.7	514.0 - 580.4

Figure 4.53: Geochronological data of the Buitrago de Lozoya metagranite: A) Tera-Wasserburg concordia; B) Wetherill concordia; and C) Distribution of concordant 207-corrected ages and the statistical data of the two most abundant populations.

and Table 4.35). The first population is Cambro-Ordovician, with a mean age of 481 ± 9 Ma that represents the age of crystallization (Fig. 4.53B). The second population is Early Cambrian, with a mean age of 547 ± 15 Ma and represents the main inherited population. There are also two more concordant points at 625 and 652 Ma, a third point at 839 Ma and an another one at 2653 Ma .

The crystallization ages were acquired in rims of crystals with restitic cores (Figs. 4.54z17 and 4.54z23). The main inherited population was obtained in uniform zircons (Fig. 4.54z11). All other older ages were acquired in cores of rimmed crystals (Fig. 4.54z16, 4.54z17 and 4.54z23).

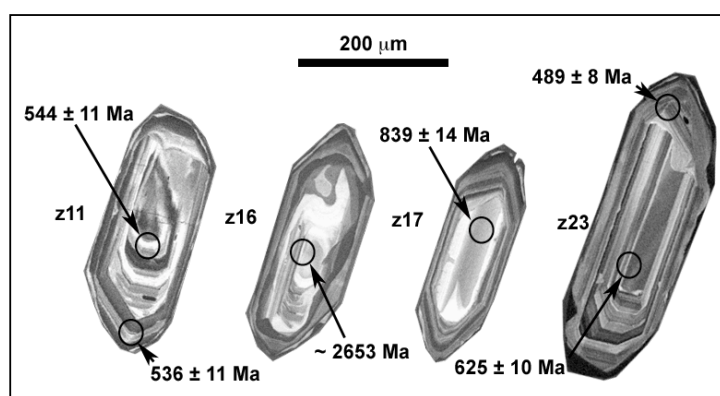


Figure 4.54: Cathodoluminescence images and ages of selected zircons from the Buitrago de Lozoya metagranite.

The 481 ± 9 Ma age matches, within error range, the crystallization age obtained by Valverde-Vaquero and Dunning (2000) using conventional U-Pb dating on zircon concentrates ($488 \pm 10 / -8$ Ma).

4.2.3.9. *La Berzosa*

The U-Pb ion microprobe data of La Berzosa metavolcanic rocks reveal the existence of two nearly concordant populations (Fig. 4.55A and 4.55C and Table 4.36). The first population is Cambro-Ordovician, with a mean age of 462 ± 11 Ma (Fig. 4.55B). It should be emphasized that only four of these determinations were close to concordia ($^{206}\text{Pb}/^{238}\text{U}_{\text{age}} / ^{206}\text{Pb}/^{238}\text{U}_{\text{age}} > 0.95$) so that this value must be simply considered as a provisional estimate of the crystallization age of La Berzosa rocks. The second, less abundant, population has between 593 and 635 Ma and has a mean age of 613 ± 15 Ma. There are also three older concordant ages at 721, 1965 and 2682 Ma.

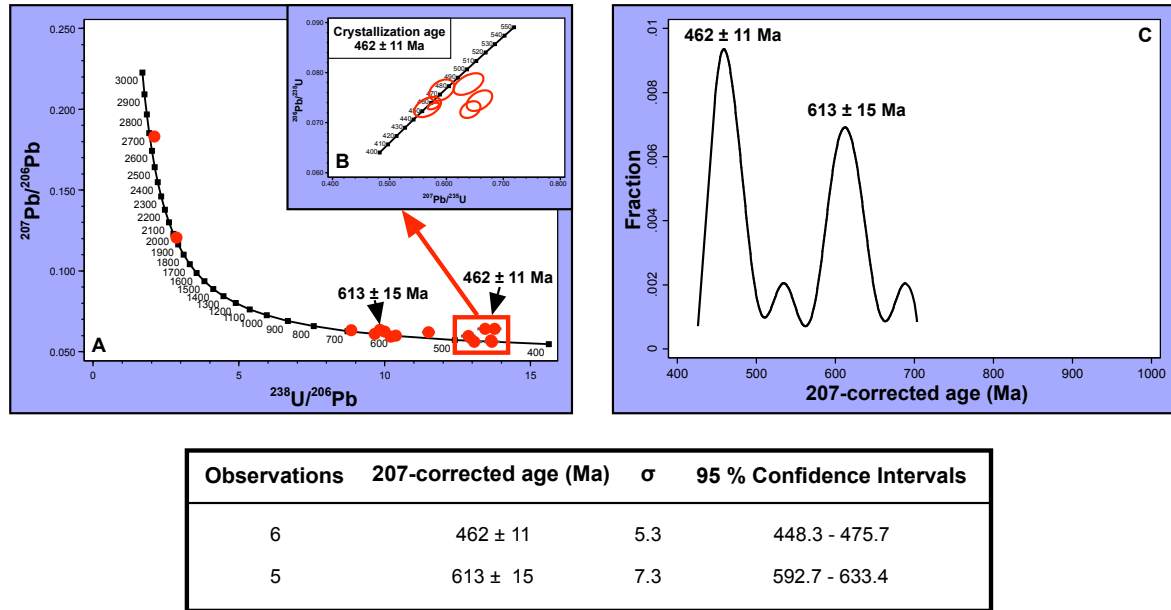


Figure 4.55: Geochronological data of La Berzosa metagranite: A) Tera-Wasserburg concordia; B) Wetherill concordia; and C) Distribution of concordant 207-corrected ages and the statistical data of the two populations.

The crystallization ages were obtained in uniform zircons (Figs. 4.56z1 and 4.56z8) and in rims of grains containing restitic cores. The inherited population was obtained in uniform zircons (Fig. 4.56z16) and in cores of rimmed zircons.

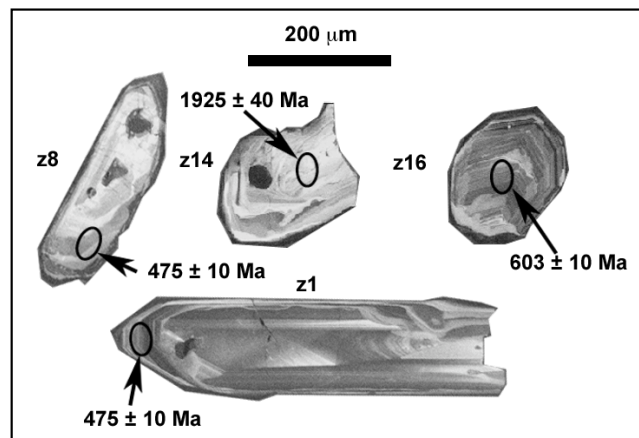
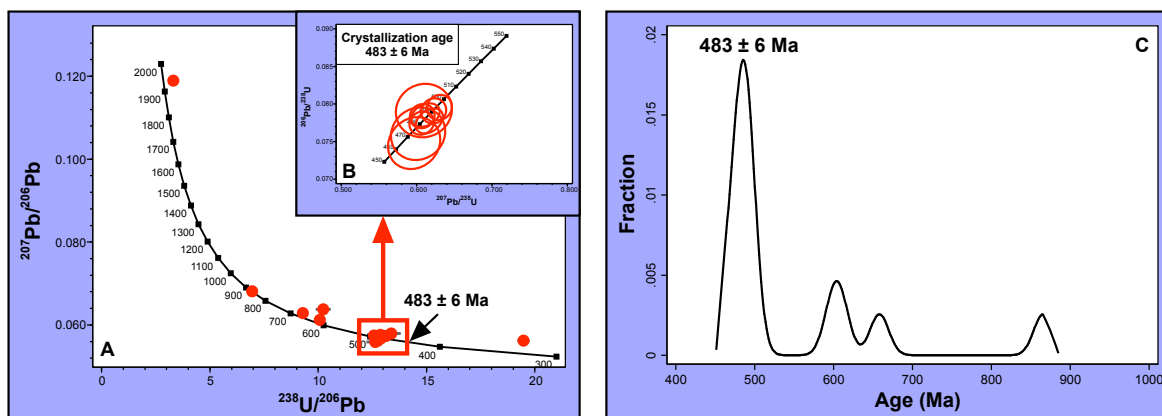


Figure 4.56: Cathodoluminescence images and ages of selected zircons from La Berzosa metagranite.

4.2.3.10. El Cardoso

The U-Pb dating of El Cardoso metavolcanic rocks (Figs. 4.57A and 4.57C and Table 4.37) yielded a Cambro-Ordovician concordant population with a mean age of 483 ± 6 Ma

which represents the crystallization age (Fig. 4.57B). There are also some subconcordant points at 599, 609, 658 and 866 Ma and a discordant point near 1.95 Ga.



Observations	207-corrected age (Ma)	σ	95 % Confidence Intervals
9	483 ± 6	3.2	475.7-490.7

Figure 4.57: Geochronological data of El Cardoso metagranite: A) Tera-Wasserburg concordia; B) Wetherill concordia; and C) Distribution of concordant 207-corrected ages and the statistical data of the crystallization population.

According to the cathodoluminescence images, the crystallization ages were acquired in uniform grains (Fig. 4.58z11) and rims of grains with inherited cores (Fig. 4.58z2). The inherited populations were obtained in cores of rimmed zircons (Figs. 4.58z2 and 4.58z16) and in uniform zircons (Fig. 4.58z1).

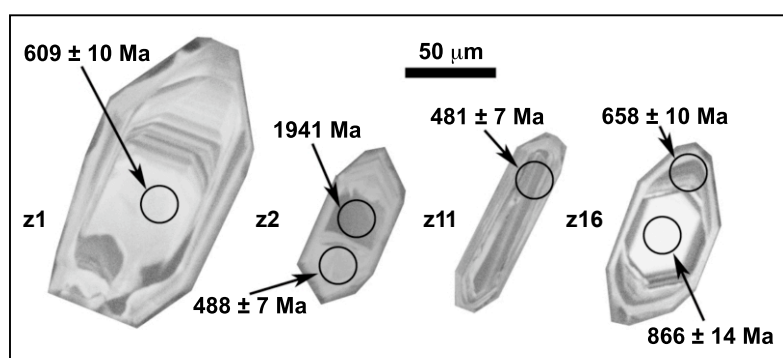
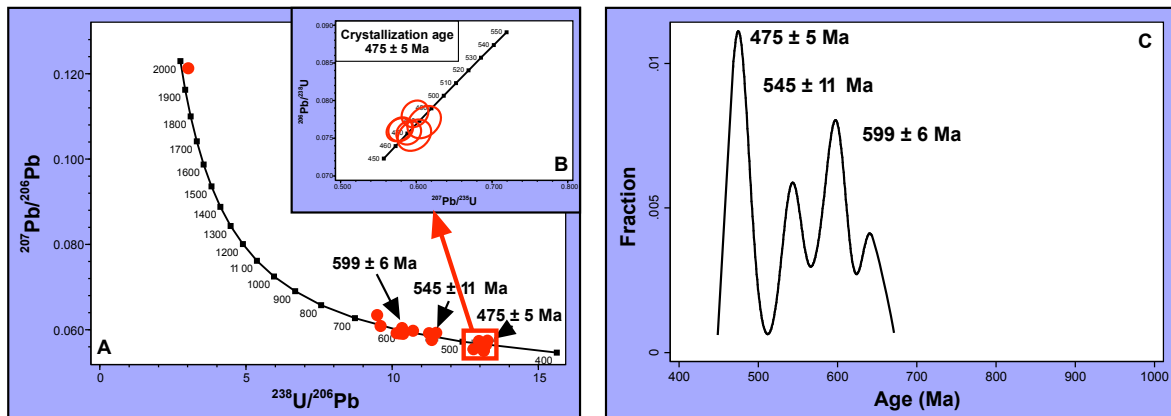


Figure 4.58: Cathodoluminescence images and ages of selected zircons from El Cardoso metagranite.

The 483 ± 6 Ma age matches, within error range, the crystallization age obtained by Valverde-Vaquero and Dunning (2000) using conventional U-Pb dating on zircon concentrates (480 ± 2 Ma).

4.2.3.11. Rianza

The U-Pb ion microprobe data of the Rianza metavolcanic rocks reveal three concordant populations (Fig. 4.59A and 4.59C and Table 4.38). The youngest is Cambro-Ordovician, with a mean age of 475 ± 5 Ma, which represents the crystallization age (Fig. 4.59B). The second is Early Cambrian, with a mean age of 545 ± 11 Ma. The oldest is Ediacaran, clustering at around 599 ± 6 Ma.



Observations	207-corrected age (Ma)	σ	95 % Confidence Intervals
6	475 ± 5	2.7	468.3 - 482.3
4	545 ± 11	5.5	530.4 - 565.2
4	599 ± 6	2.8	590.0 - 607.7

Figure 4.59: Geochronological data of La Rianza metagranite: A) Tera-Wasserburg concordia; B) Wetherill concordia; and C) Distribution of concordant 207-corrected ages and the statistical data of the three populations.

The crystallization ages, according to the cathodoluminescence images, were obtained in uniform grains and in rims of grains with inherited cores (Fig. 4.60z5). The inherited populations were acquired in cores of rimmed grains (Figs. 4.60z5 and 4.60z4) and in uniform grains (Fig. 4.60z13).

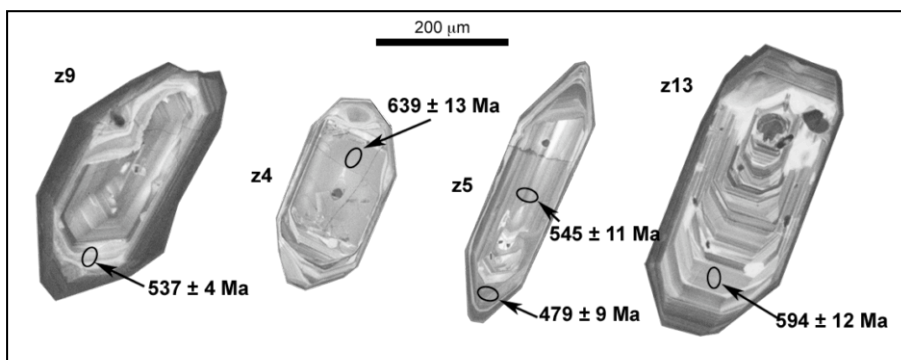


Figure 4.60: Cathodoluminescence images and ages of selected zircons from the Rianza metagranite.

The 475 ± 5 Ma age is different from, but still within error range, of the value obtained by Valverde-Vaquero and Dunning (2000) using conventional U-Pb dating on zircon concentrates ($468 +19/-8$ Ma).

4.2.4. The Anatectic Complex of Toledo

In the Anatectic Complex of Toledo we sampled two metagranites: Polán and Mohares. They were dated using U-Pb ion microprobe with the following results:

4.2.4.1. Polán

The Polán metagranite has a concordant Cambro-Ordovician population, two older concordant points at 521 and 587 Ma and one younger-than-crystallization ages at 429 Ma

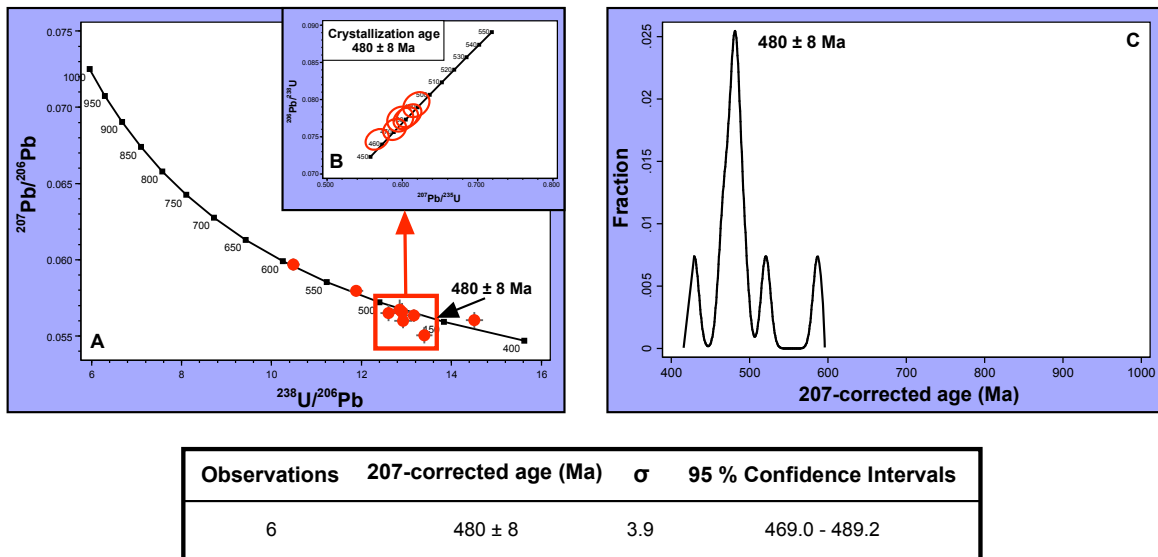


Figure 4.61: Geochronological data of the Polán metagranite: A) Tera-Wasserburg concordia; B) Wetherill concordia; and C) Distribution of concordant 207-corrected ages and the statistical data of the crystallization population.

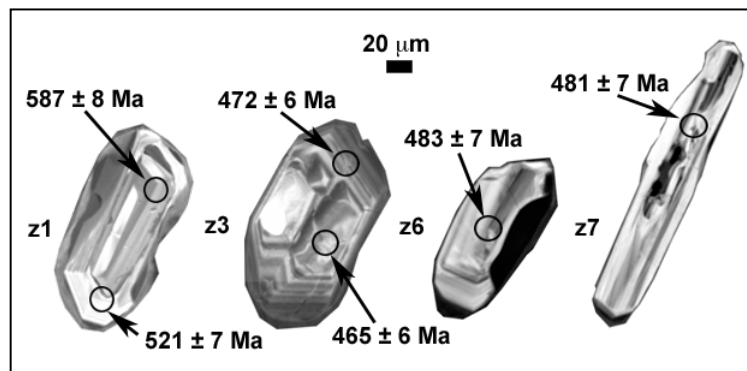


Figure 4.62: Cathodoluminescence images and ages of selected zircons from the Polán metagranite.

(Fig. 4.61A and 4.61C and Table 4.39). The Cambro-Ordovician population has a mean age of 480 ± 8 Ma, which represents the crystallization age (Fig. 4.61B). All ages from the crystallization population were acquired in uniform zircons (Figs. 4.62z3, 4.62z6 and 4.62z7) and the same occurred with the rest of the ages, except the 587 Ma age, which corresponds to an inherited core of a rimmed crystal (Fig. 4.62z1).

4.2.4.2. Mohares

The U-Pb dating of the Mohares metagranite reveal the presence of an unique, small and concordant Cambro-Ordovician population with a mean age of 489 ± 7 Ma (Figs. 4.63A and 4.63C and Table 4.40), which we regard as the crystallization age (Fig. 4.63B). No inherited cores were found and all data were obtained in uniform crystals (Fig. 4.64).

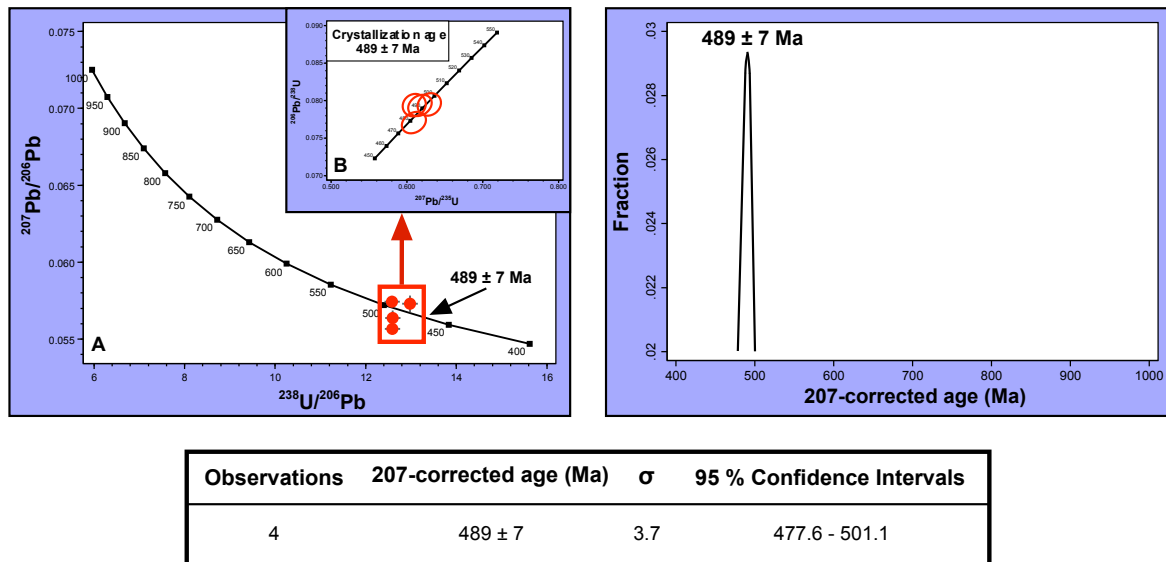


Figure 4.63: Geochronological data of the Mohares metagranite: A) Tera-Wasserburg concordia; B) Wetherill concordia; and C) Distribution of concordant 207-corrected ages and the statistical data of the crystallization population.

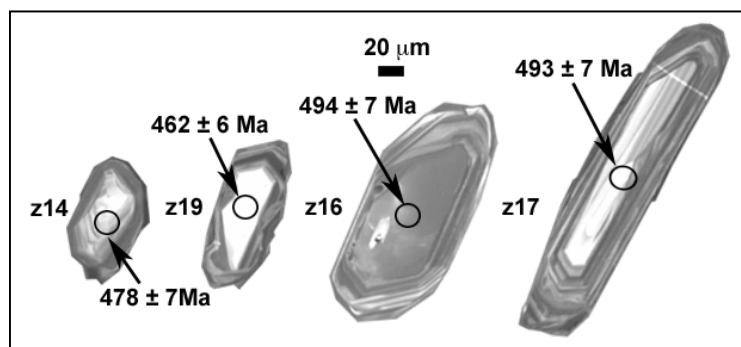


Figure 4.64: Cathodoluminescence images and ages of selected zircons from the Mohares metagranite.

4.3. Summary

4.3.1. The Ollo de Sapo Domain

The geochronological data of the Ollo de Sapo Formation and the associated metagranites, reveal that here the magmatism was synchronous and short lived, except in the Hiendelaencina area, where it was exceptionally long lived. In the northwestern area of the Ollo de Sapo Domain all the magmatic activity occurred between 492 Ma and 486 Ma, an age span comparable to the Urra Formation (495 Ma to 488 Ma, Sola et al., 2008) located in the western extreme of the Central Iberian Zone. The crystallization age is more variable in the southeastern area of the formation. In the Hiendelaencina region there are at least two volcanic episodes, one at 495 Ma and the other at 483 Ma, followed by a plutonic episode at 474 Ma accounting, in total, for about 20 million years of episodic magmatic activity. In the Villadepera region, the volcanic and intrusive rocks have the same age, 483 Ma, coincidental with the youngest volcanism in Hiendelaencina.

As mentioned before, one of the most remarkable features of the Cambro Ordovician rocks of Central Iberia is the elevated percentage of zircons with premagmatic cores, usually about 70-80%, in occasions up to near 100%. In the Ollo de Sapo metavolcanic rocks and metagranites about 80% of the inherited zircons components are Ediacaran, with a mean age of 609 ± 3 Ma (Fig. 4.2). This population was found in all rocks types, irrespective of their locality or whether they are metavolcanic or metagranitic. The Cryogenian (650-700 Ma) and Tonian (850-900 Ma) populations are only well-represented in Hiendelaencina, which is, therefore, another unique feature of this region. The second most important inherited population is Orosirian, with concordant ages clustering around 2.0 Ga. It is far less abundant than the Ediacaran and is much better represented in metavolcanic rocks than in metagranites. Orosirian ages appear mostly in the cores of Cambro-Ordovician zircons, though crystals with a Cambro-Ordovician rim enclosing a large Ediacaran core that contains an internal Orosirian core are also found. There are also two small and poorly defined Archean populations (Fig. 4.2), one Neoproterozoic (approximately 2.6 Ga), discordant, present in Sanabria and Vivero, and the other Mesoarchean, with concordant ages around 3.0-3.2 Ga, which appears in all studied regions except in Trives and Hiendelaencina.

Finally, apart from these inherited populations, there are sometimes younger, and always discordant, ages fluctuating between 450 Ma and 350 Ma which, at least in part,

correspond to mixed Cambro-Ordovician and Variscan ages (see section 4.3.3), the later caused by incipient migmatization.

Table 4. 1. Summary of the crystallization ages of the metavolcanic rocks and metagranites of the Ollo de Sapo Domain and the north of the Schist-Graywacke Complex Domain. Note that the $(^{206}\text{Pb}/^{238}\text{U})_{\text{cor}}$ ages have been obtained using LA-ICPMS, SIMS and SHRIMP methods, $(^{207}\text{Pb}/^{206}\text{Pb})$ ages using stepwise evaporation method and (Rb/Sr)ages using TIMS method.

Ollo de Sapo Domain			
	Ages (Ma)		
	$(^{206}\text{Pb}/^{238}\text{U})_{\text{cor}}$	$(^{207}\text{Pb}/^{206}\text{Pb})$	(Rb/Sr)
<i>Ollo de Sapo Formation</i>			
Vivero	486 ± 3	486 ± 3	-
Trives	486 ± 5	-	-
Sanabria (Lower)	492 ± 4	-	-
Sanabria (Upper)	488 ± 3	490 ± 2	-
Villadepera	483 ± 4	-	-
Hiendelaencina (Lower)	495 ± 5	494 ± 4	-
Hiendelaencina (Upper)	485 ± 6	483 ± 3	490 ± 65
<i>Associated metagranites</i>			
San Sebastián	470 ± 3	461 ± 3	-
Viana do Bolo	487 ± 4	488 ± 4	-
Miranda do Douro	483 ± 3	-	-
Antoñita	-	474 ± 4	477 ± 30
Schist-Graywacke Complex Domain			
<i>The Tormes Dome</i>			
Fermoselle	486 ± 6	-	-
Vitigudino	489 ± 5	-	-
Ledesma	471 ± 7	-	-
<i>The Sierra de Gredos</i>			
San Pelayo	492 ± 4	-	-
Castellanos	498 ± 4	-	-
Bercimuelle	488 ± 3	-	-
<i>The Sierra de Guadarrama</i>			
La Cañada	489 ± 9	-	-
La Hoya	478 ± 7	-	-
La Estación I	490 ± 5	-	-
La Estación II	473 ± 8	-	-
Vegas de Matute (Melanocratic)	476 ± 5	-	-
Vegas de Matute (Leucocratic)	484 ± 6	-	-
La Morcuera	482 ± 8	-	-
Buitrago de Lozoya	481 ± 9	-	-
La Berzosa	462 ± 11	-	-
El Cardoso	483 ± 6	-	-
Riaza	475 ± 5	-	-
<i>The Anatectic Complex of Toledo</i>			
Polán	480 ± 8	-	-
Mohares	489 ± 7	-	-

4.3.2. The Schist-Graywacke Complex Domain

Though the high metamorphic grade affecting most Cambro-Ordovician rocks of the Schist-Graywacke Complex Domain makes it difficult to recognize their protolith, it seems that in most cases it was plutonic. Only in the Sierra del Guadarrama there are some gneisses: El Cardoso, Riaza, and La Berzosa, which were probably derived from volcanic rocks. The duration of the magmatic activity was not the same in the four studied areas. The magmatic activity occurred between: (i) 489 Ma and 471 Ma in the Tormes Dome; (ii) 498 Ma and 488 Ma in the Sierra de Gredos; (iii) 489 Ma and 480 Ma, in the Anatectic Complex of Toledo; (iv) 490 Ma and 462 Ma in the Sierra de Guadarrama, though this lower limit should be considered cautiously because of the imprecision in the determination of the crystallization age of the La Berzosa orthogneiss (see section 4.2.3.9). It seems, therefore, that whereas the magmatism in the Tormes Dome and Toledo was fully coeval with the Ollo de Sapo, in Sierra de Gredos started a little earlier and in the Sierra de Guadarrama started later and lasted longer.

A high fraction of zircon grains from the Cambro-Ordovician rocks of the northern Schist-Graywacke Complex Domain also have inherited cores. In the small bodies of the Sierra de Gredos, Toledo and the Tormes Dome, the inheritance is similar to that of the metavolcanic rocks and metagranites of the Ollo de Sapo Formation, but in the Sierra del Guadarrama the fraction of the zircons with pre-magmatic cores is smaller, oscillating between 50 and 70 %. The distribution of ages of these restitic cores is slightly different from what was found in the Ollo de Sapo Formation (Fig. 4.26). There are three main Neoproterozoic populations (521-646 Ma) that represent about 73 % of the inherited cores. The first two have mean ages of 540 ± 3 Ma and 571 ± 4 Ma, and have been found in the four sampled areas. The third has a mean age of 617 ± 5 Ma and has been found in the Tormes Dome, Sierra de Gredos and Sierra de Guadarrama. There are also other less abundant populations with ages Cryogenian (650-700 Ma), Tonian (850-950 Ma) and Orosirian (1900-2000). Cryogenian and Tonian populations have been observed in the Tormes Dome, Sierra de Gredos and Sierra de Guadarrama, and Orosirian population only has been found in Sierra de Gredos and Sierra de Guadarrama. Finally, there are some nearly concordant Neoproterozoic ages (2.5-2.6 Ga), which have been obtained in the Sierra de Gredos and Sierra de Guadarrama, and a nearly concordant Mesoarchean age which has only been found in the Castellanos metagranite (Sierra de Gredos).

As in the Ollo de Sapo Formation, there is a small population of younger and almost always discordant ages fluctuating between 338 Ma and 450 Ma which may correspond to mixed ages between Cambro-Ordovician and Variscan ages (see next section).

4.3.3. The significance of the younger ages

Many studied zircons have small overgrowths, rarely thicker than 20 μm , with younger-than-Ordovician ages. They are often highly discordant, with 207-corrected values between 450 Ma and 338 Ma. We have found that these younger ages are the consequence of two different processes: (i) solid-state overgrowths on zircons included in Zr-rich feldspars probably caused by the exsolution of Zr (Bea et al., 2006b); and (ii) Variscan overgrowths probably caused by incipient migmatization. In the first case (Fig. 4.65A) the overgrowths are highly cathodoluminescent due to the low concentrations of U, Th and HREE. In the second case, they show a low cathodoluminescence (Fig. 4.65B) due to elevated concentrations of U. To discuss the origin of these overgrowths in detail is beyond the scope of this work. Here it is sufficient to consider that most often the overgrowths yielded highly discordant ages from which no useful information can be derived. For that reason, they will not be considered further.

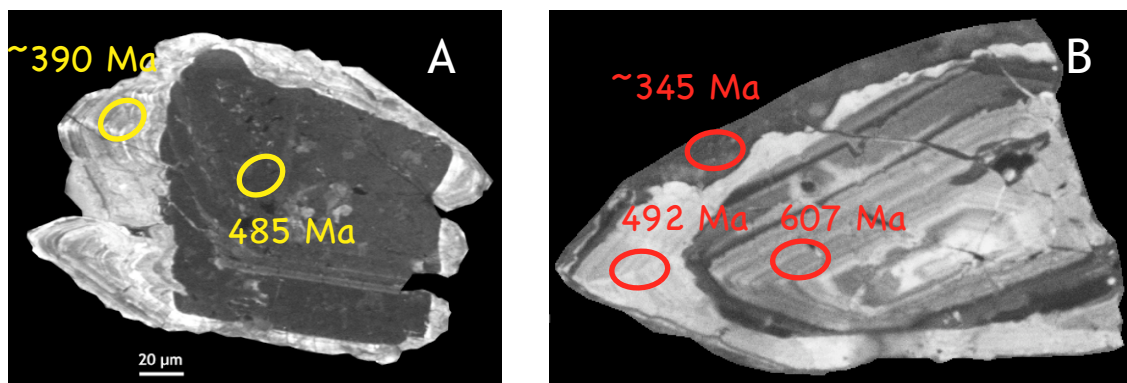


Figure 5.65.- The reason for younger ages. A) Overgrowth in a zircon included in a K-feldspar megacryst from the upper metavolcanic rocks of Hiendelaencina. The elevated cathodoluminescence is caused by low contents of U, Th and HREE. Bea et al. (2006b) attributed them to the exsolution of Zr from Zr-rich feldspars. B) U-rich overgrowth caused by incipient migmatization in the Viana do Bolo metagranite. Ellipses represent zones analyzed with the ion-microprobe. The results for either overgrowth are highly discordant ages.

5. CHEMICAL COMPOSITION AND Sr-Nd ISOTOPE GEOLOGY

5. CHEMICAL COMPOSITION AND Sr-Nd ISOTOPE GEOLOGY

As mentioned before, the Cambro-Ordovician felsic rocks of Iberia form two contrasting geochemical associations (Fig. 1.1A). The most abundant comprises intermediate to silicic peraluminous rocks ($\text{mol. Al}_2\text{O}_3/(\text{CaO} + \text{Na}_2\text{O} + \text{K}_2\text{O}) \approx 1.34 \pm 0.23$) which according to the classification of Frost et al. (2001) are magnesian and alkaline-calcic to calc-alkaline (Fig. 1.1B). The less abundant association is composed of ferroan, alkaline to calc-alkaline, and predominantly high-silica rocks (Fig. 1.1B) which often contain Na pyriboles. Accordingly, Montero et al., (in press and in rev.) named these two different rock associations as peraluminous and peralkaline, with it being well understood that a few rocks of the first group have $\text{mol. Al}_2\text{O}_3/(\text{CaO} + \text{Na}_2\text{O} + \text{K}_2\text{O}) < 1.05$ and about 50-70% of the rocks of the second group, depending on the massif, have $\text{mol. (Na}_2\text{O} + \text{K}_2\text{O})/\text{Al}_2\text{O}_3 \leq 1$, that is, they are not strictly peralkaline.

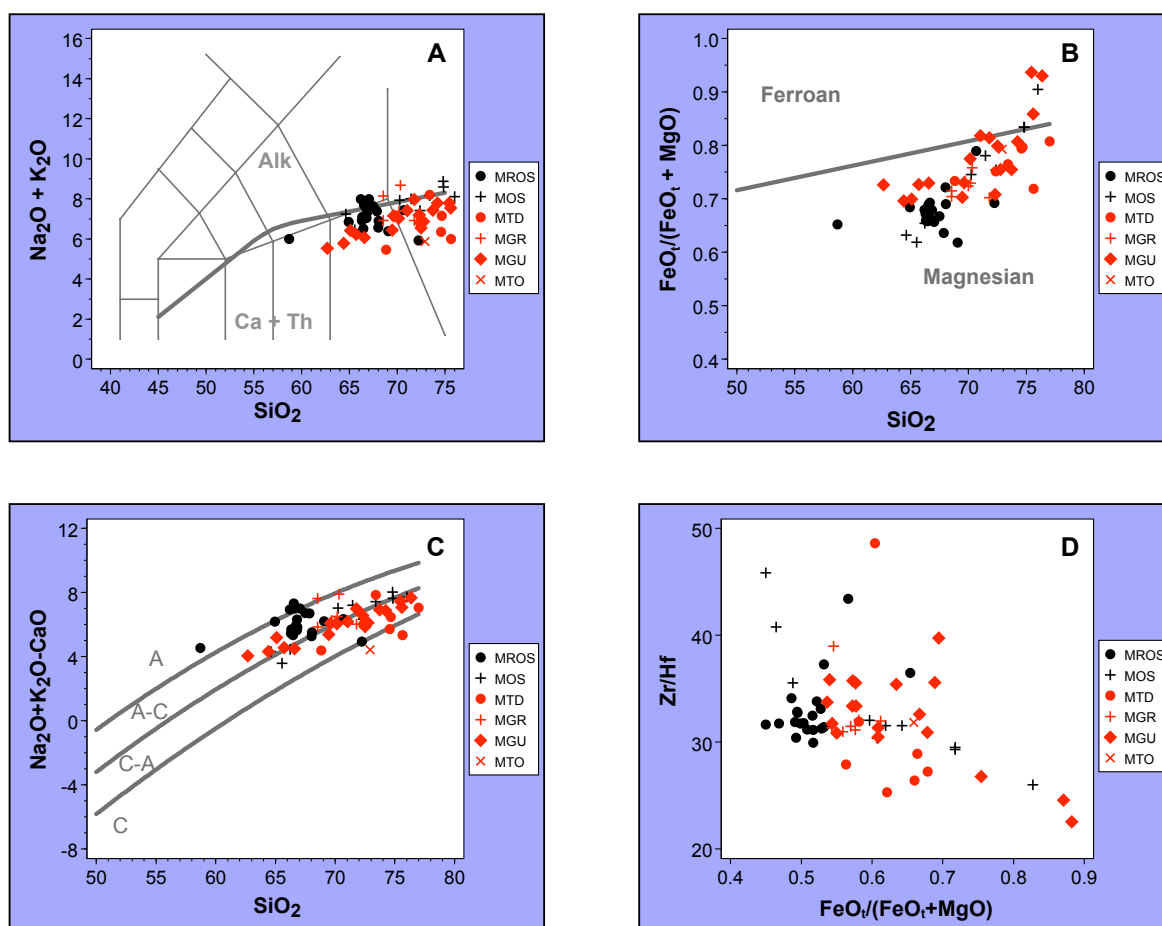


Figure 5.1: A) TAS plot; B and C) Frost et al. (2001) plots; D) Zr/Hf-FeO/(FeO_T+MgO) plot. The meaning of the abbreviations of the legend is: MROS: Metavolcanic rocks of the Ollo de Sapo Domain; MOS: Metagranites of the Ollo de Sapo Domain; MGR: Metagranites of the Sierra de Gredos; MGU: Metagranites and metavolcanic rocks of the Sierra de Guadarrama; MTD: Metagranites of the Tormes Dome; MTO: Metagranites of the Anatectic Complex of Toledo.

All metavolcanic rocks and metagranites of the Central Iberian Zone belong to the peraluminous association. Those of the Ollo de Sapo Domain have ASI (mol. $Al_2O_3/(CaO + Na_2O + K_2O)$) clustering around 1.15 in the metagranites and around 1.50 in the metavolcanic rocks and all of them are magnesian and alkaline-calcic (Fig. 5.1). The metavolcanic rocks have a relatively homogeneous major element composition which corresponds to K-rich peraluminous dacites (Fig. 5.1). The metagranites are more variable, either less silicic (granodioritic) or more silicic (leucogranitic) than the metavolcanic rocks. The trace element composition of both metavolcanic rocks and metagranites is characterized by marked enrichments in alkali elements, especially Li, and precipitous negative anomalies in Sr (Fig. 5.2). The chondrite-normalized REE patterns decrease smoothly from La to Sm, have a marked negative Eu anomaly, and decrease again from Gd to Lu, the slope being steeper in the metagranites than in the metavolcanic rocks (Fig. 5.3). The latter showing a large variation in HREE. When compared to averages of Palaeozoic European S-type collision-related granites and I-type subduction-related granites (Fig. 5.2C), the trace element compositions of both the metagranites and the metavolcanic rocks of the Ollo de Sapo Formation bear a stronger resemblance with the former.

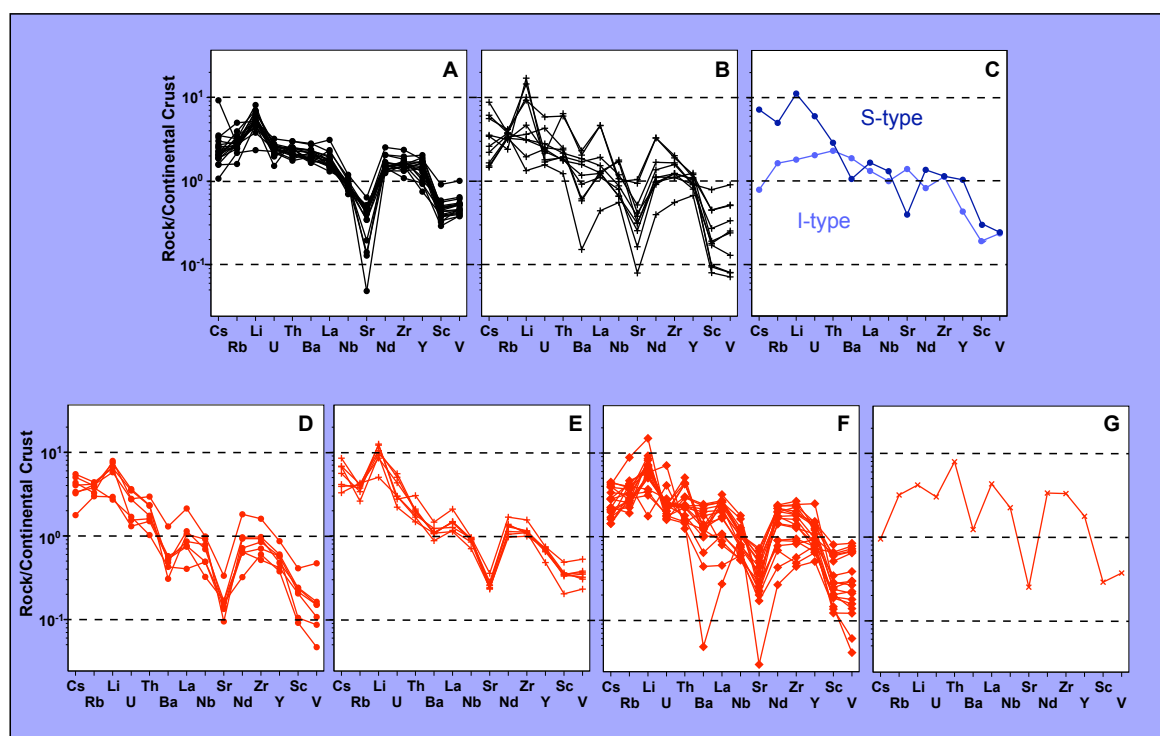


Figure 5.2: Continental-crust normalized trace-element plots of: A) Metavolcanic rocks of the Ollo de Sapo Domain; B) Metagranites of the Ollo de Sapo Domain; C) References averages of the Paleozoic European I and S types granites; D) Metagranites of the Tormes Dome; E) Metagranites of the Sierra de Gredos; F) Metagranites and metavolcanic rocks of the Sierra de Guadarrama; G) Metagranites of the Anatectic Complex of Toledo.

The metagranites and the metavolcanic rocks of the northern Schist-Graywacke Complex Domain do not show any significant difference from the rocks of the Ollo de Sapo Formation. They are equally peraluminous ($ASI \approx 1.4$, Table 5.1), and equally alkaline-calcic to calc-alkaline and magnesian, so that in the Frost et al. (2001) diagrams the rocks from the two domains plot in the same area (Fig. 5.1). The evolution from magnesian to ferroan as SiO_2 increases (Fig. 5.1B) is due to the effects of magmatic differentiation, as evidenced by the sympathetic decrease of the differentiation index e.g., Zr/Hf (Fig. 5.1D). The trace element composition of the metagranites and metavolcanic rocks of the northern Schist-Graywacke Complex Domain is also nearly identical to the metagranites of the Ollo de Sapo Formation, characterized by the enrichment in Li and a strong depletion in Sr (Fig. 5.2). In Guadarrama, the metagranites show the largest compositional spectrum, from granodioritic to leucogranitic, the later showing depleted chondrite-normalized REE-patterns with precipitous negative Eu anomalies and secondary Nd negative anomalies (Fig. 5.3).

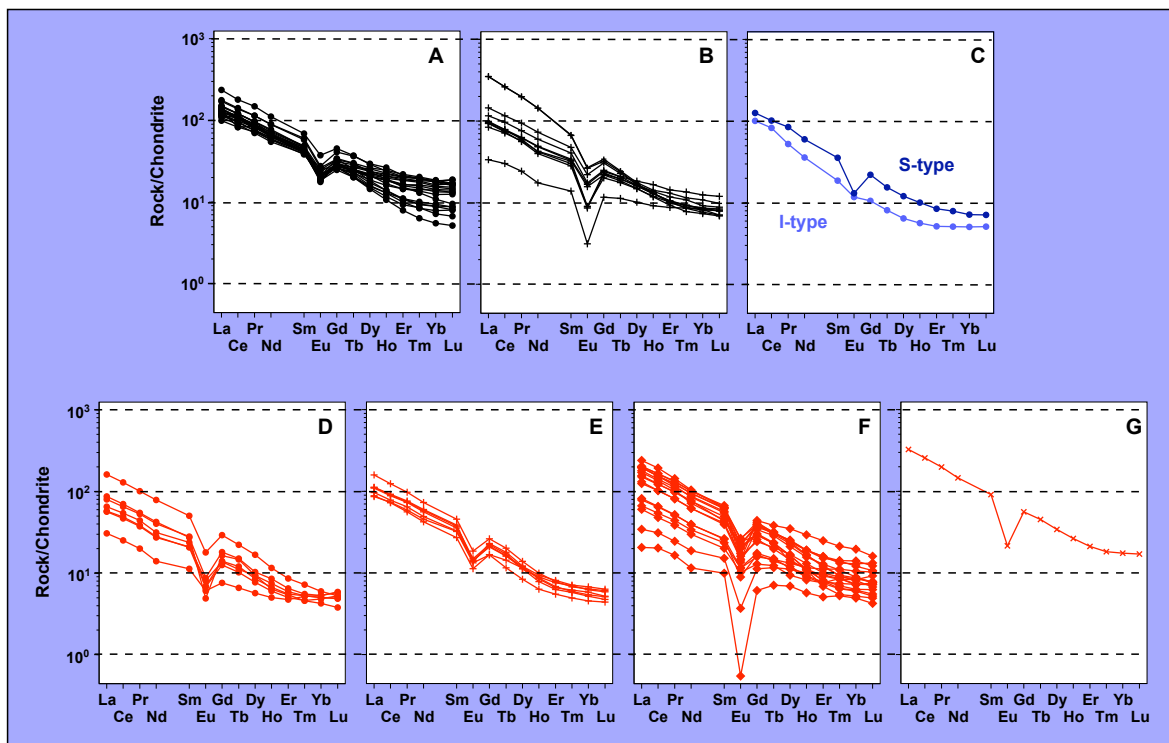


Figure 5.3: Chondrite-normalized REE plots of: A) Metavolcanic rocks of the Ollo de Sapo Domain; B) Metagranites of the Ollo de Sapo Domain; C) References averages of the I and S types granites; D) Metagranites of the Tormes Dome; E) Metagranites of the Sierra de Gredos; F) Metagranites and metavolcanic rocks of the Sierra de Guadarrama; G) Metagranites of the Anatectic Complex of Toledo.

Table 5.1: Alumina saturation index (ASI) and average of the Alumina saturation index (ASI*) of: OSD, Ollo de Sapo Domain; MROS, metavolcanic rocks of the Ollo de Sapo Domain; metagranites of the Ollo de Sapo Domain; SGCD, Schist-Graywacke Complex Domain; MTD, metagranites of the Tormes Dome; MGR, metagranites of the Sierra de Gredos, MGU, metagranites and metavolcanic rocks of the Sierra de Guadarrama; MTO, metagranites of the Anatectic Complex of Toledo.

Domain	Group	ASI	ASI*
OSD	MROS	1.50	1.37
	MOS	1.15	
SGCD	MTD	1.45	1.40
	MGR	1.35	
	MGU	1.40	
	MTO	1.22	

The Sr and Nd isotope composition of the metavolcanic rocks of the Ollo de Sapo Formation (Fig. 5.4 and Table 5.2) clusters around $(^{87}\text{Sr}/^{86}\text{Sr})_{490\text{Ma}} \approx 0.710$ (ranging from 0.707 to 0.713), $\epsilon\text{Nd}_{490\text{Ma}} \approx -4.6$ (ranging from -5.8 to -3.3), and $T_{(\text{DM})} \approx 1.8$ Ga (ranging

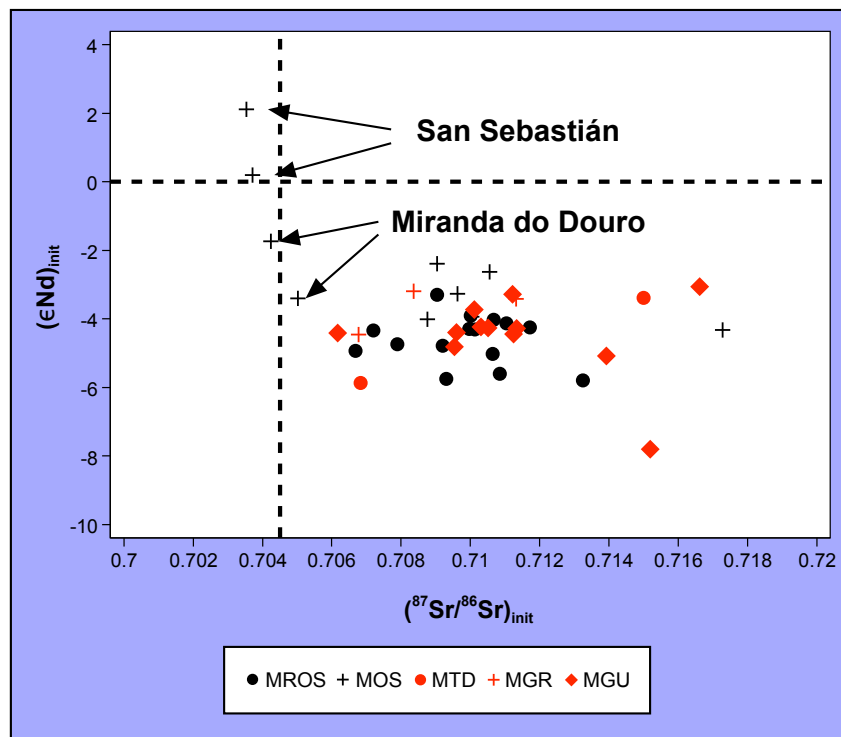


Figure 5.4: $(\epsilon\text{Nd})_{\text{init}}$ vs. $(^{87}\text{Sr}/^{86}\text{Sr})_{\text{init}}$ of the: A) MROS, metavolcanic rocks of Ollo de Sapo Domain; B) MOS, metagranites of the Ollo de Sapo Domain; C) MGR, metagranites of the Sierra de Gredos; D) MGU, metagranites and metavolcanic rocks of the Sierra de Guadarrama, E) MTD, metagranites of the Tormes Dome.

from 1.6 to 2.0), values that suggest they were derived from old continental materials. The metagranites of the Ollo de Sapo Formation have, in general, slightly higher $\epsilon\text{Nd}_{490\text{Ma}}$ than the metavolcanic rocks and show a wider range of $(^{87}\text{Sr}/^{86}\text{Sr})_{490\text{Ma}}$. The isotope composition of two metagranites, Antofñita and Viana do Bolo, overlaps with the metavolcanic rocks, but the other two, Miranda do Douro and San Sebastián, are more primitive, with $(^{87}\text{Sr}/^{86}\text{Sr})_{490\text{Ma}} \approx 0.704$, $T_{(\text{DM})} \approx 1.3$ Ga and variable $\epsilon\text{Nd}_{490\text{Ma}}$ which ranges from -3.4 to -1.7 in Miranda do Douro and from 0.2 to 2.1 in San Sebastián (Fig. 5.4). The San Sebastián metagranite, therefore, besides being the youngest and having the lowest zircon inheritance, also has the most primitive isotope composition of the whole Ollo de Sapo Formation.

The Sr and Nd isotope composition of the metagranites and metavolcanic rocks of the northern Schist-Graywacke Complex Domain are very monotonous and similar to the metavolcanic rocks of the Ollo de Sapo Formation. Neither in Guadarrama, Gredos or Tormes Dome, we have found metagranites with such a primitive isotope composition such as Miranda do Douro or San Sebastián. By contrast, the metagranites of the northern Schist-Graywacke Complex Domain have slightly higher average values of $(^{87}\text{Sr}/^{86}\text{Sr})_{490\text{Ma}} \approx 0.711$ and $T_{(\text{DM})} \approx 1.9$ (Table 5.2). In detail, the metagranites of the Sierra de Gredos show the most primitive isotope composition having $(^{87}\text{Sr}/^{86}\text{Sr})_{490\text{Ma}} \approx 0.709$, $\epsilon\text{Nd}_{490\text{Ma}} \approx -3.7$ and $T_{(\text{DM})} \approx 1.8$ Ga. The metagranites and the metavolcanic rocks of the Sierra de Guadarrama and the metagranites of the Tormes Dome have $(^{87}\text{Sr}/^{86}\text{Sr})_{490\text{Ma}} \approx 0.711$ and $T_{(\text{DM})} \approx 1.9$ Ga and slightly variable $\epsilon\text{Nd}_{490\text{Ma}}$, -4.5 and -4.6 respectively.

Table 5.2: Average Sr and Nd isotope data of: MROS, metavolcanic rocks of Ollo de Sapo Domain; MOS, metagranites of the Ollo de Sapo Domain; MTD, metagranites of the Tormes Dome; MGR, metagranites of the Sierra de Gredos; MGU, metagranites and metavolcanic rocks of Sierra de Guadarrama

Group	$(^{87}\text{Sr}/^{86}\text{Sr})_{490\text{Ma}}$	$(\epsilon\text{Sr})_{490\text{Ma}}$	$(^{143}\text{Nd}/^{144}\text{Nd})_{490\text{Ma}}$	$(\epsilon\text{Nd})_{490\text{Ma}}$	Nd(CHUR) (Ga)	$T_{(\text{DM})}$ (Ga)
MROS	0.7098	84	0.511771	-4.6	1.0	1.8
MOS	0.7082	61	0.511888	-2.3	0.8	1.6
MTD	0.7109	99	0.511771	-4.6	1.1	1.9
MGR	0.7088	70	0.511818	-3.7	1.0	1.8
MGU	0.7113	105	0.511777	-4.5	1.2	1.9

6. DISCUSSION

6. DISCUSSION

6.1. The nature of the magmatic sources

The distribution of 627 concordant or nearly concordant ion-microprobe and LA-ICPMS ages for the Cambro-Ordovician rocks of Central Iberia (Figs. 4.2 and 4.26) shows that a significant fraction of zircons contain inherited cores (Fig. 6.1; see section 6.3) and that a significant fraction of these are Ediacaran. Whereas in the Ollo de Sapo Domain and the Tormes Dome the age of the Ediacaran inherited component is unimodal, dominated by 605 Ma to 620 Ma ages, in Guadarrama and the two eastern most metagranites of Gredos, Castellanos and Bercimuelle, besides a widespread 605-620 Ma component, there is another well-represented inherited component with ages at 540-550 Ma and another, which is less abundant, at 570-580 Ma. It is worth mentioning that Gredos is the only area of the Central Iberian zone where orthogneisses with a crystallization age of 540 Ma to 550 Ma have been identified so far (Bea et al., 2003).

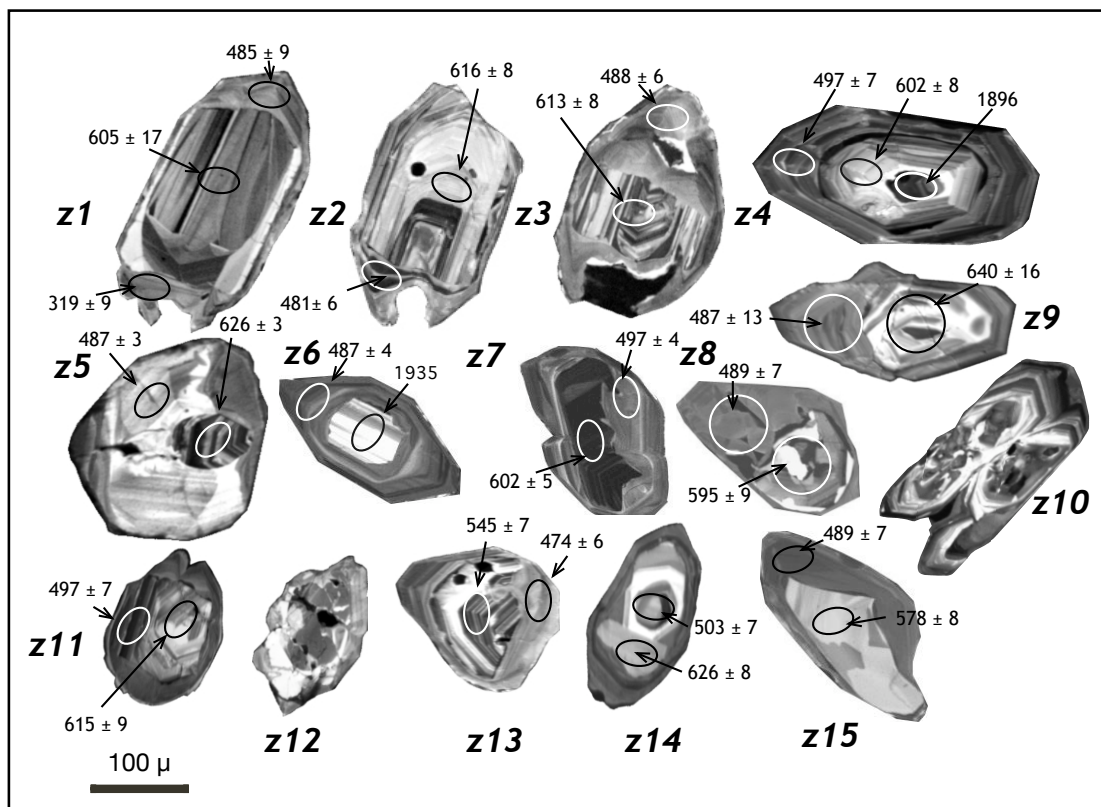


Figure 6.1: Cathodoluminescence images of the representative zircons of the Cambro-Ordovician rocks of Central Iberia. Numbers represent concordant U-Pb ages determined either with ion-microprobe (ellipses) or LA-ICP-MS (circles).

The marked dominance among the inherited cores of the Ediacaran components suggests that the protolith of the Cambro-Ordovician magmas of Iberia mostly consisted of igneous rocks of late Pan-African age, or young (at the time of melting) immature sediments derived from them. In principle, the highly radiogenic Sr and Nd isotope composition of the Cambro-Ordovician igneous rocks might be compatible with derivation from the Cambro-Ediacaran terrigenous series underlying them, either the Angon Formation, the Viana Series or the Schist-Graywacke Complex. Putting aside the fact that the chemical composition of these terrigenous sediments is hardly capable of producing melts with a composition similar to the Cambro-Ordovician igneous rocks, there is compelling evidence derived from zircon studies that negates this possibility. First, the studies of Fernández-Suárez et al. (2000) on detrital zircons of the Cambro-Ediacaran series in the north of the Central Iberian and the Western-Asturian-Leonese Zone indicate

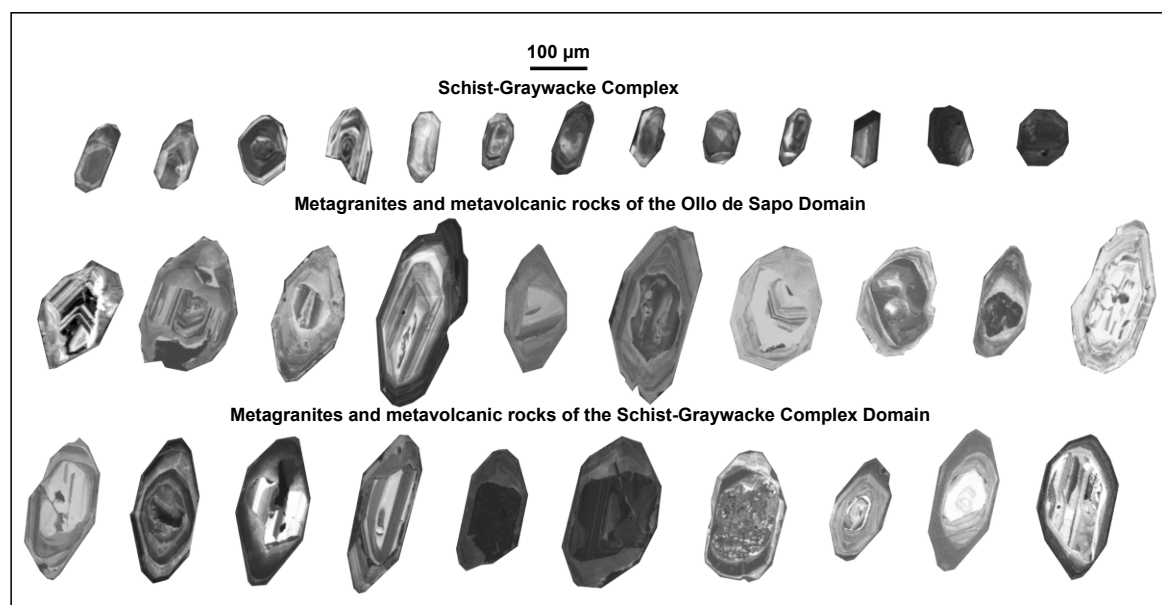


Figure 6.2: Cathodoluminescence images of selected zircons of Schist-Graywacke Complex, metagranites and metavolcanic rocks of the Ollo de Sapo Domain and the north of the Schist-Graywacke Complex Domain. Note the differences of zircon size between the Cambro-Ordovician rocks and the Schist-Graywacke Complex, the latter being smaller than most inherited cores.

that ages between 650 Ma and 950 Ma are among the most abundant, and these ages are not represented in the inherited components of the Cambro-Ordovician metagranites and metavolcanic rocks. Second, our ongoing studies on the materials of the Schist-Graywacke Complex reveal that these rocks contain very few zircons, and that these are very small, smaller even than the inherited cores of the Cambro-Ordovician gneisses (Fig. 6.2), which

must have necessarily been partially dissolved during the Cambro-Ordovician magmatism (see section 6.3.4.). Neither the ages nor the abundance and size of the zircons of the terrigenous series, therefore, are compatible with them being the source of the Cambro-Ordovician magmatism.

Due to vigorous Variscan crustal reworking, intermediate to felsic calc-alkaline igneous rocks of late Pan-African rocks are poorly exposed in Iberia, except for some small dioritic to granitic massifs in the Mérida region to the south of the Central Iberian Zone (Bandrés et al., 2004), and a few granodioritic gneisses in the west of the Ávila batholith (Bea et al., 2003). Calc-alkaline plutonism of Ediacaran age, however, was one of the most important magmatic events of the neighboring Anti-Atlas region of Morocco (Gasquet et al., 2005), a region attached to Iberia during the Ediacaran (Ennih and Liegeois, 2001; Ennih and Liegeois, 2003) but which was subsequently little affected by the Variscan orogeny.

Sr and Nd isotopes are consistent with the idea that late Pan-African rocks similar to those of the Anti-Atlas could have been the source of the Cambro-Ordovician magmatism

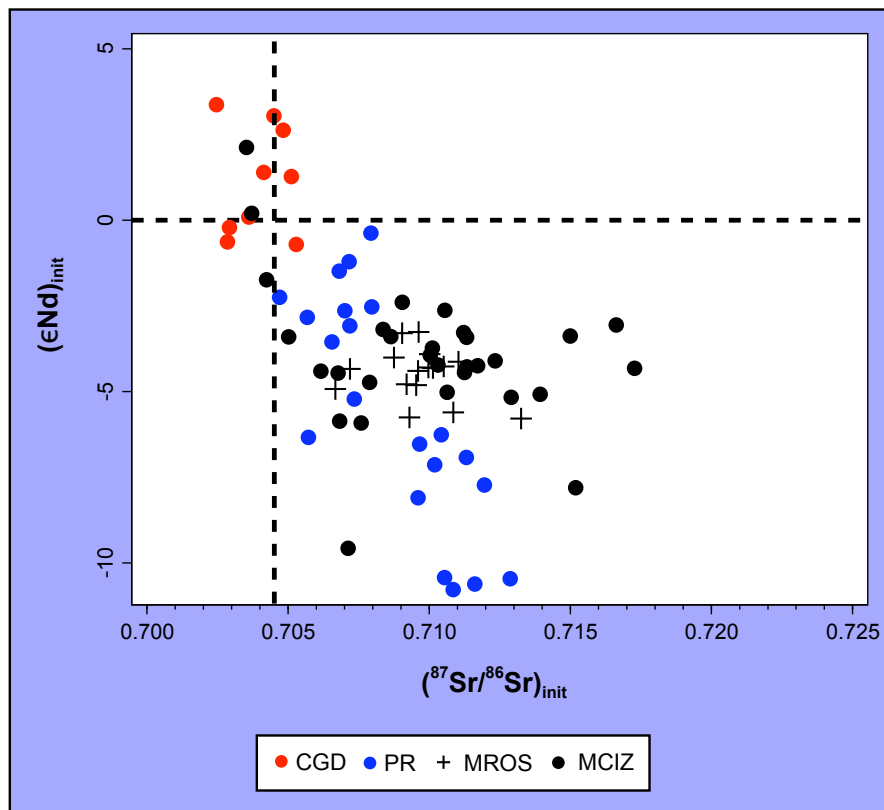


Figure 6.3: $(\epsilon\text{Nd})_{\text{init}}$ vs. $(^{87}\text{Sr}/^{86}\text{Sr})_{\text{init}}$ of the: CGD, Carrascal gabbrodiorites; PR, Pan-African rocks which correspond to Anti-Atlas late-Pan-African granites and volcanics; MROS, metavolcanic rocks of the Ollo de Sapo Domain; MCIZ, metagranites of the Central Iberian Zone.

of Iberia. Figure 6.3 represents the data of the Cambro-Ordovician igneous rocks of the Central Iberian Zone with data from Anti-Atlas late-Pan-African granites and volcanics (Gasquet et al., 2005; unpublished data of P. Montero) recalculated at 490 Ma and data from the Carrascal gabbro-diorites, the only mafic Cambro-Ordovician body of the Central Iberian Zone (Sola et al., 2005; Sola, 2007). Despite the noise caused by the Variscan metamorphism, it is evident in this plot that the Cambro-Ordovician magmas of Iberia can have been generated by partial melting of rocks similar to the Anti-Atlas with little material contribution of mafic magmas except in the most primitive metagranites, Miranda do Douro and, especially, San Sebastián.

6.2. The reasons for the abnormally elevated zircon inheritance

The investigations described in Chapter 3 have revealed an unusually large proportion of zircon crystals with pre-magmatic, inherited zircon cores. If we accept that discordant cores such as those shown in Fig. 6.2 are restitic (which was confirmed wherever U-Pb

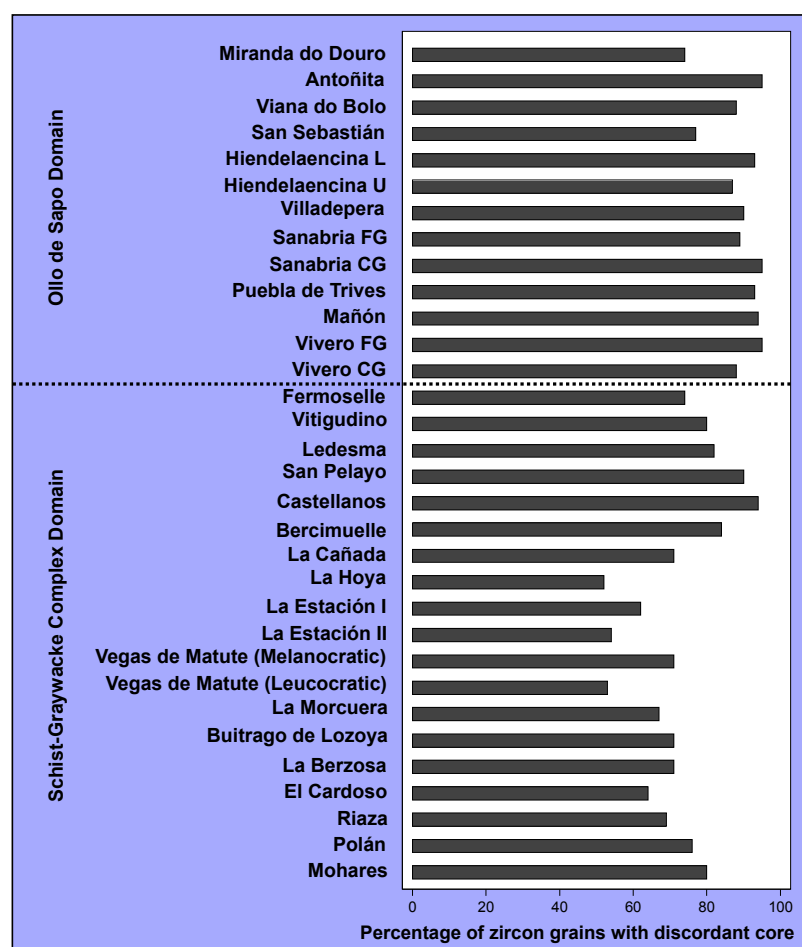


Figure 6.4: Percentage of zircon grains with pre-magmatic cores in 32 Cambro-Ordovician massifs of the Central Iberia. L, lower; U, upper, CG, coarse-grained; and FG, fine-grained.

spot analyses are available), the examination of some 2000 zircons by cathodoluminescence imaging has shown that the proportion of grains with inherited cores varies from 75% to 95% in the metagranites, and from 70% to 80%, locally up to 100%, in the metavolcanic rocks, except in Guadarrama where this number decreases to 50% -70 % (Fig. 6.4). To the authors' knowledge, such a high degree of inheritance is common in low melt-fraction migmatites (e.g. Montero et al., 2004b) or low-T pegmatoid mobilisates (e.g. Gilotti and McClelland, 2005) but is extremely rare in high-level granites or rhyodacitic volcanic rocks.

The solubility of zircon in felsic magmas has been thoroughly studied experimentally by Watson and Harrison (1983) and Harrison and Watson (1983), who have demonstrated that in peraluminous or metaluminous melts solubility only depends of the temperature and composition (basically the silica content) of the magma. This experimental model has been checked on many occasions (e.g. Bea and Montero, 1999) so that it seems to be accurate beyond any reasonable doubt. Since the expected variations in the composition of crustal melts have far less influence than the variations in temperature (for example, melts with 68 % and 72 % of SiO₂ dissolve 148 and 131 ppm Zr at 800°C, but dissolve 253 and 223 ppm Zr at 850°C), the first hypothesis that might explain the abnormally elevated zircon inheritance of the Cambro-Ordovician magmas is that they never reached a high enough temperature to dissolve all the zircon present in the source rock, which for some unknown reason was unusually elevated.

To check this hypothesis we consider the zircon saturation temperature, calculated according to the Watson and Harrison (1983) model, with the Ti-in-zircon temperature (Watson and Harrison, 2005; and Watson et al., 2006). Whereas the former represents an estimate of the temperature required to avoid zircon inheritance (Hancher and Watson, 2003), the later represents a minimum estimate of the temperature attained by the magmas at their locus of melting.

6.2.1. Zircon saturation temperatures

The zircon saturation temperature is calculated with the expression:

$$T_{Zr} = 12900 / (\log(497600/Zr) + 3.8 + (0.85 * (M-1))) - 273; \quad (1)$$

where:

T_{Zr} is the zircon saturation temperature,

Zr is in ppm,

$M = (\text{Ca} + \text{Na} + \text{K}) / (\text{Al} * \text{Si} / (\text{Si} + \text{Ti} + \text{Al} + \text{Fe} + \text{Mg} + \text{Mn} + \text{Ca} + \text{Na} + \text{K}))$, and all major elements are expressed as atomic equivalents.

The application of this expression yielded a slightly skewed unimodal distribution, tailed at lower values due to loss of Zr during crystallization of the granitoids, with the mode value at 815°C (Fig. 6.5). If we exclude the most silicic samples, and hence the most prone to have undergone magmatic differentiation processes, the average T_{Zr} is $822 \pm 5^\circ\text{C}$ which would then represent the temperature required to dissolve all zircon present in the rock.

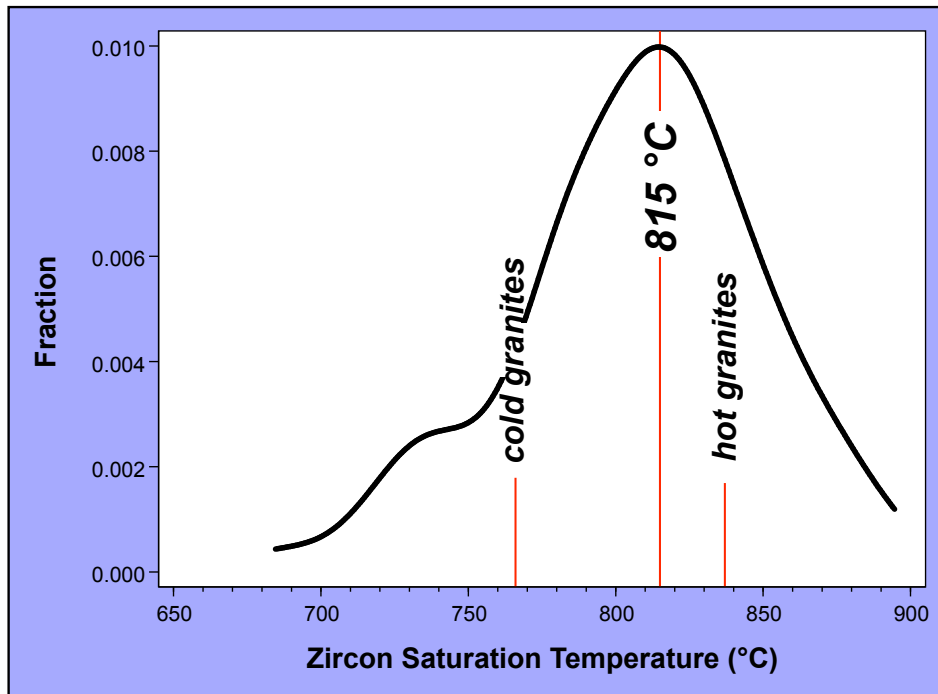


Figure 6.5: Zircon saturation temperature for representative samples of the metavolcanic rocks and metagranites of the Ollo de Sapo Domain and the north of the Schist-Graywacke Complex Domain calculated from data of Table 6.1. The zircon saturation temperature of these samples is compared to the temperature of cold and hot granites.

Miller et al. (2003) have demonstrated that zircon inheritance and zircon saturation temperatures are antithetic. These authors divided North-American granites into "hot" (average $T_{\text{Zr}} = 837^\circ\text{C}$, little or no inheritance) and "cold" (average $T_{\text{Zr}} = 766^\circ\text{C}$; high zircon inheritance, most often clustering around 50%) types. With minor modifications, this two-fold categorization seems applicable worldwide (e.g., Asrat et al., 2004; Rowe et al., 2007; Bryan et al. 2008), irrespective of granite age and typology, implying that the

inverse relation between inheritance and T_{Zr} is probably a reflection of the conditions that most typically occur during granite petrogenesis. Remarkably, the Central Iberian Cambro-Ordovician igneous rocks do not follow this rule, because they have simultaneously an average high T_{Zr} of 822 °C, characteristic of "hot" granites, and a zircon inheritance that not only matches, but in most cases nearly duplicates that of "cold" granites (Fig. 6.6). As discussed below, this unusual combination suggests disequilibrium processes that do not

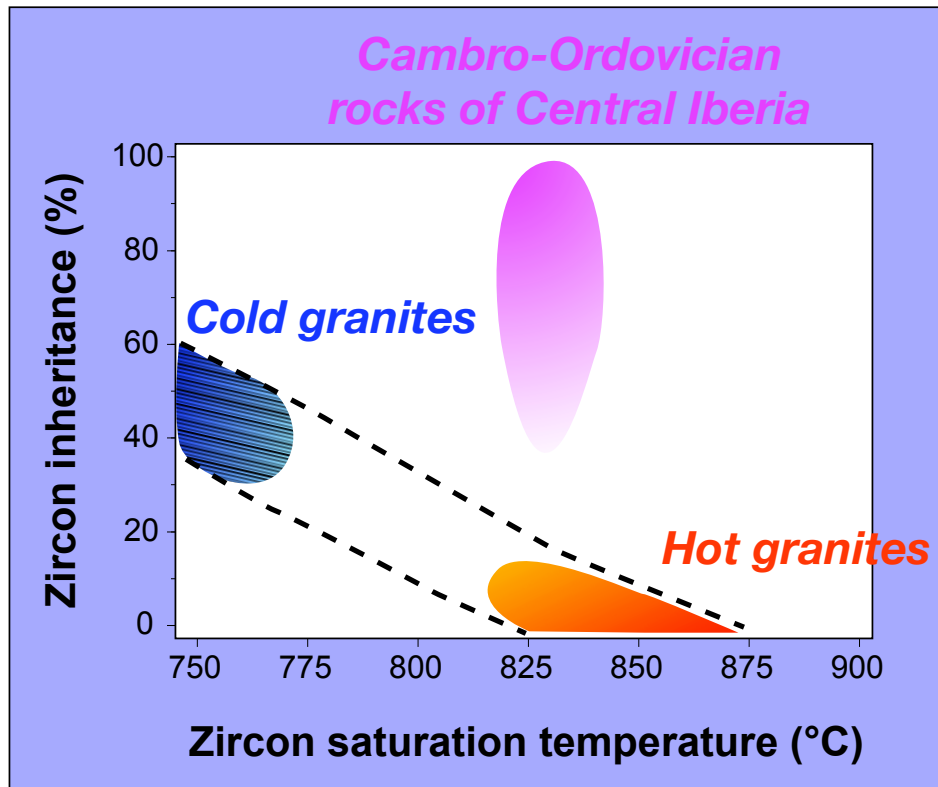


Figure 6.6: Zircon inheritance vs zircon saturation temperature of Cambro-Ordovician rocks of the Central Iberia. These rocks are compared to cold and hot granites of Miller et al. (2003).

normally occur during granite magmas.

To understand the atypical processes involved in the generation of the Iberian Cambro-Ordovician magmas, we must consider the following: Zircon solubility in common crustal melts depends on melt composition and temperature, but it is almost independent of pressure and water content (Watson and Harrison, 1983). Since the variations to be expected in the melt bulk-composition have less influence than the variations to be expected in temperature, the first hypothesis that might explain the abnormally elevated zircon inheritance of the Cambro-Ordovician magmas is that they were never sufficiently

hot to dissolve all the zircon grains entrained from the source, these, for some unspecified reason, being exceptionally abundant.

6.2.2. Ti-in-zircon temperatures

The minimum temperature attained by the magma can be estimated by applying the Ti-in-zircon thermometer (Watson and Harrison, 2005; Watson et al., 2006) to the Cambro-Ordovician rims of the zircon grains. To this end, we analyzed the low-abundance (5.5%) but interference-free ^{49}Ti isotope, plus ^{238}U , ^{207}Pb and ^{206}Pb (to ascertain the age), ^{178}Hf (to calculate Zr/Hf) and ^{92}Zr and ^{29}Si (as internal standards) with a LA-ICPMS system, ablating 60 μm diameter spots to get enough sensitivity. One of the main problems with this geothermometer is the sensitivity to inclusions of Ti-bearing minerals in the ablated volume, something that cannot be totally avoided by careful microscopic examination prior to analyses. For this reason, we filtered out our results excluding analyses that yielded $\text{Ti} > 100$ ppm ($\approx 994^\circ\text{C}$) because higher values are indicative of inclusions and $\text{Zr/Hf} > 45$ to avoid residual crystals formed from lower T melts (Bea et al., 2006).

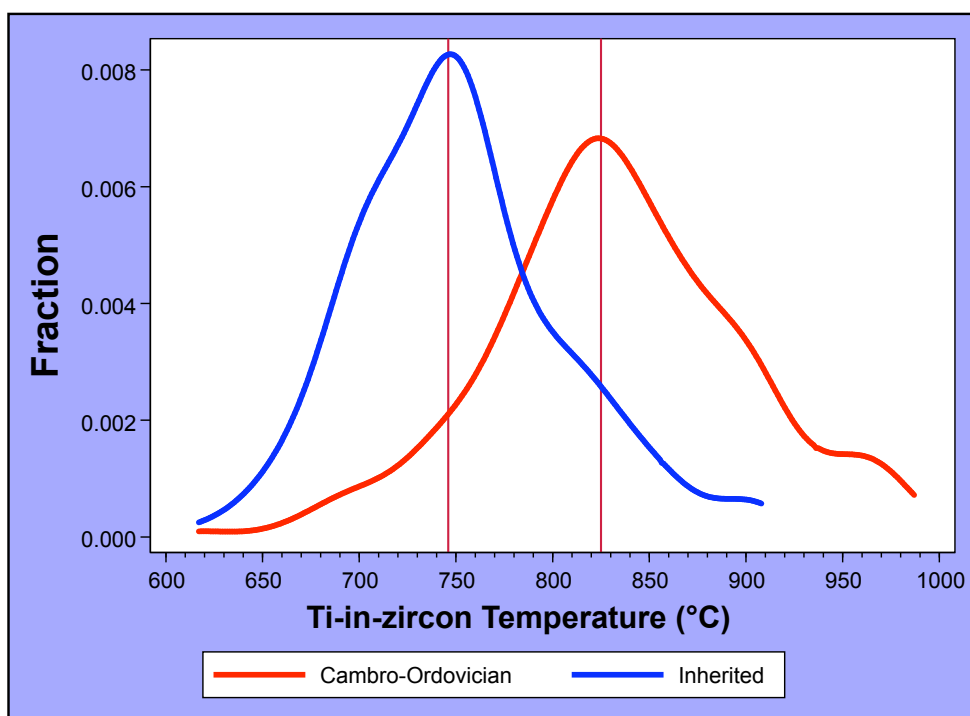


Figure 6.7: Gaussian distribution of the Ti-in-zircon temperature of the Cambro-Ordovician zircons and inherited cores.

With this method we have studied 11 massifs: 2 metavolcanic and 9 metagranitic (Table 6.1). Among these, only one from Guadarrama (La Morcuera metagranite) was included,

the rest of metagranites have zircons with Zr/Hf decreasing from 60 to 20, which strongly suggest that they have undergone magmatic fractionation processes. The results reveal that in all cases the crystallization temperatures of the inherited zircons, which peak at 745°C (Fig. 6.7), are considerably lower than the crystallization temperatures of the neoformed Cambro-Ordovician grains, which peak at 826°C and have some values exceeding 950°C. Remarkably, the average Ti-in-zircon temperature is always higher than the average zircon saturation temperature, except in La Morcuera, where magmatic fractionation almost certainly occurred.

Therefore, if we accept that the Ti-in-zircon temperature represents a minimum estimate of the thermal peak of the magmas (see discussion in section 6.3.3.4), we must conclude that the temperature attained by the Cambro-Ordovician magmas was not lower than 900-950°C, that is, high enough to dissolve all entrained zircons. The reasons why so many of them survived should be, therefore, related to the zircon dissolution kinetics.

Table 6.1: Summary of the Ti-in-zircon and zircon saturation temperatures. T_{cao} , Ti-in-zircon temperatures of the Cambro-Ordovician zircons; $T_{\text{pre-CAO}}$, Ti-in-zircon temperatures of pre-Cambro-Ordovician cores; and T_{Zr} , zircon saturation temperature.

Massif	T_{CAO}	$T_{\text{pre-CAO}}$	T_{Zr}
Vivero	840	767	843
Sanabria	831	770	843
Viana do Bolo	828	786	803
San Sebastian	844	765	778
Miranda do Douro	829	754	817
Fermoselle	820	744	767
Bercimuelle	827	713	796
Ledesma	865	795	794
Vitigudino	821	732	780
Castellanos	823	738	802
Morcuera	814	746	869

6.2.3. Zircon dissolution kinetics in the Cambro-Ordovician magmas of Iberia

During partial melting of a rock, the main factors that may hinder or may delay the dissolution of zircon into the melt are: (i) shielding by major minerals, (ii) the small volume of the magma pores surrounding zircon grains, (iii) exceedingly fast heating/cooling rates.

The first factor, shielding by major minerals, can be easily excluded because the mineral that contains most inclusions of zircon is biotite (Bea, 1996), and this mineral has, necessarily, been involved in anatectic reactions. In the same way, the high mobility of magmas capable of eruption on the Earth's surface implies a total connectivity of melt, excluding therefore the second factor, small volume of melt reservoirs. Accordingly, the most probable hypothesis is that the elevated zircon inheritance was the result of extremely fast heating/cooling rates. To check this hypothesis, we can use the experimental work of Watson (1996).

6.2.3.1. Watson's equation

Watson's equation (1996, eq. 17) for calculating the instant dissolution rate of spherical zircon crystals is:

$$dr/dt \times 10^{17} = -U \left(\left[(1.25 \times 10^{10}/r) \exp(-28380/T) + 7.24 \times 10^8 \exp(-23280/T) \right] \right); \quad (2)$$

where:

dr/dt is the instant dissolution rate (cm s^{-1}),

r is the radius of a spherical zircon crystal (cm),

T is the absolute temperature ($^{\circ}\text{K}$),

U is the difference between the current Zr concentration of the melt and the concentration required for zircon saturation according to the experimental model of Watson and Harrison (1983). In partially molten systems, calculating U requires knowledge of the Zr concentration and bulk-rock composition of the protolith, and the volume of the melt reservoirs.

This equation can be used stepwise to calculate the variations of the zircon radius as a function of time for a given heating (or cooling) gradient. This requires inputting: (i) the new T , (ii) the new zircon radius that resulted from the previous step, (iii) the new U value calculated considering the amount of zircon dissolved in the previous step and the change in zircon solubility caused by the variation in T and melt composition.

6.2.3.2. Choice of parameters

The data of Bea et al. (2003) indicate that the 543 Ma Almohalla metagranite contains zircons with maximum dimensions of $300 \times 100 \times 100 \mu\text{m}$ and a volume equivalent to a $89 \mu\text{m}$ radius sphere and an average Zr concentration of 283 ppm. The data of Bandrés et al. (2004) indicate that the Pan-African granitoids of Mérida contain zircon grains with

maximum dimensions of 270 x 100 x 100 μm (a volume equivalent to a 86 μm radius sphere) and have an average concentration of Zr of 125 ppm. This value, however, is probably an underestimate because the samples were analyzed after acid digestion; our XRF data for the same rocks reveal concentrations between 150 and 270 ppm Zr. Additionally, the ca. 600 Ma granites of Morocco, with average XRF Zr concentrations of 227 ppm, contain zircon grains with dimensions most frequently around 150 x 70 x 50 μm (volume equivalent to a 56 μm radius sphere) and only the largest ones reach 300 x 120 x 90 μm (volume equivalent to a 92 μm radius sphere). Neither the zircon grain size, nor the Zr concentration of these rocks are exceptional, but instead are close to what one would expect for common granodiorites and granites.

Mérida, Morocco, and to a lesser extent Almohalla, granitoids can give an idea, at least approximately, about the grain size distribution of zircon, the Zr concentration and the bulk-rock composition of a potential protolith, needed for the calculations on zircon solution kinetics made in the next section. Therefore, we assumed the following conservative initial conditions:

1. Beginning of melting: 700°C
2. Maximum temperature reached by the magma: 900°C
3. Volume of melt reservoirs: infinite
4. Major elements composition of the melt (to calculate zircon solubility): calculated using the equations of (Winther, 1995) for a granodioritic protolith at 8 kbar and 2% H₂O: SiO₂ = 70.73; TiO₂ = 0.42; Al₂O₃ = 14.97; FeO_{tot} = 1.89; MgO = 0.75; MnO = 0.03; CaO = 1.68; Na₂O = 2.91; K₂O = 4.74.
5. Zr concentration in the protolith = 225 ppm.

6.2.3.3. Estimation of heating rates

Figure 6.8 shows the calculated time-temperature coordinates at which spherical zircons with radii of 25 μm , 50 μm , 75 μm , 100 μm , 150 μm and 200 μm will dissolve totally in the melt as a function of the heating rate. The following stands out: Zircons with a spherical radius of 50 μm , i.e. with a volume similar to the population most commonly found in the protolith, would only survive if heating occurred at a rate of 0.1 °C/year or faster. Zircons with a radius of 100 μm , that is with a volume larger than the largest zircons of the protolith, would only survive if the heating rate was 0.025 °C/year or faster. In

summary, the survival of protolith zircons requires less than 10,000 years, probably around 2,000 years, from the beginning of melting (700°C) up to the thermal peak (900°C).

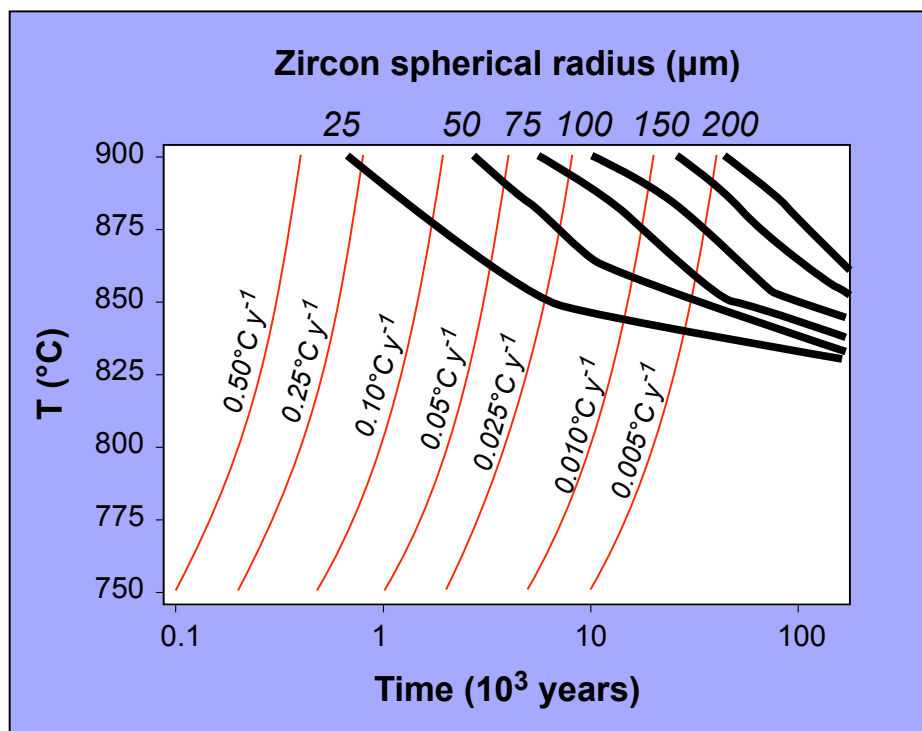


Figure 6.8: Disappearance curves (Black lines) of spherical zircon grains as a function of the heating rate (Red lines) calculated with Watson's equation (1996, eq. 17).

6.2.3.4. Estimation of cooling rates

Zircon grains that were partially dissolved during heating will continue dissolving during cooling as long as the melt does not become zircon saturated. Therefore, we calculated again the time-temperature coordinates at which spherical zircon survivors with radii of 20 µm, 40 µm, 60 µm and 80 µm will totally dissolve in the melt as a function of the cooling rate. To estimate the Zr concentration in the melt at 900°C we considered that all protolith zircon grains had a spherical radius of 100 µm and that, during heating, the radius of the dissolving crystal was coupled to the Zr concentration of the melt so that the total concentration of Zr in the system (melt plus crystals) was always equal to that of the protolith: 225 ppm. Results are shown in Fig. 6.9, from which it follows that the survival of 20 µm, 40 µm, 60 µm and 80 µm zircons require linear cooling rates from 900°C to 830°C faster than 0.4°C/year, 0.1 °C year⁻¹, 0.05°C/year, and 0.025°C/year, respectively. Since the cooling rates of erupted felsic magmas are often much faster than these (e. g. Harris et al.,

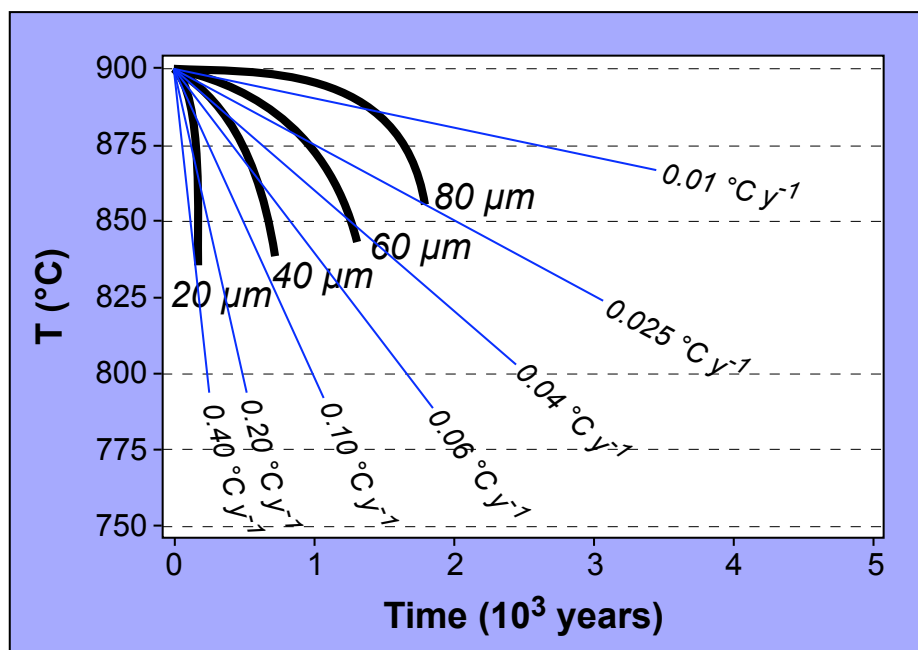


Figure 6.9: Disappearance curves (Black lines) of spherical zircon survivors as a function of the cooling rate (Blue lines) calculated with Watson's equation.

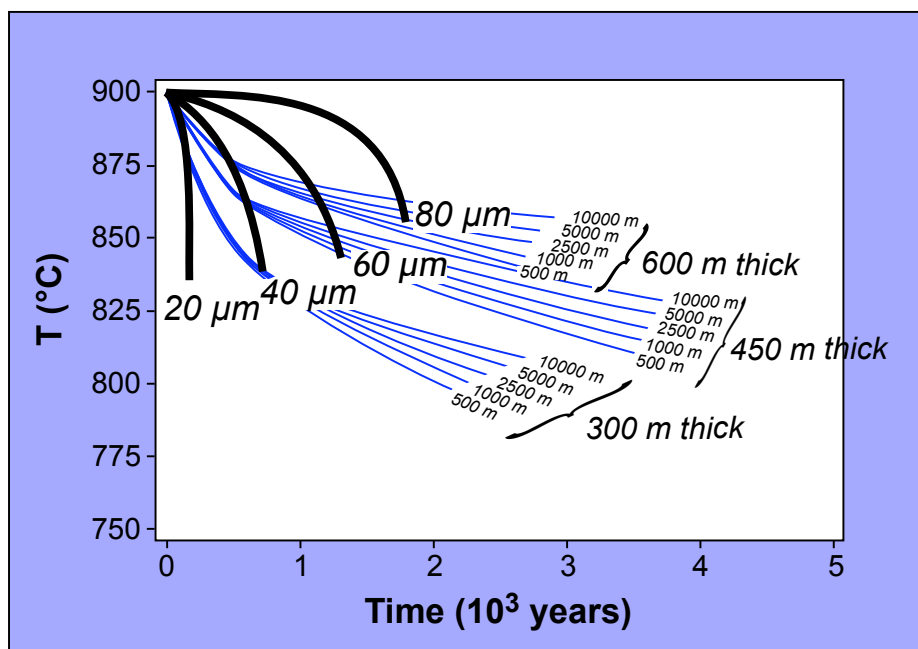


Figure 6.10: Disappearance curves (Black lines) of spherical zircon survivors as a function of the cooling rate calculated as in Fig. 6.9, compared with the 1D cooling curves (Blue lines) of granitic sills of different thickness emplaced at depths from 500 m to 10.000 m.

2002), we can conclude that most survivors would have a chance of cooling with little size reduction if the magma crystallized quickly by being rapidly transported from the melting region to Earth's surface.

The situation, however, is different if the magmas crystallized at depth, such as in the case of the metagranites. According to Iglesias Ponce de Leon and Ribeiro, (1981), González Lodeiro, (1981) and Lancelot et al. (1985), the metagranites of the Ollo de Sapo Formation occur as high-level sills or laccoliths about 300-400 m thick. A similar geometry can be presumed for the metagranites of the north of the Schist-Graywacke Complex Domain except, perhaps, in Guadarrama. To assess the behavior of zircon under these conditions we calculated the 1D cooling paths of granitic sills with the same initial T of 900°C, different thickness (300 m, 450 m and 600 m) which were emplaced at different depths (500 m, 1000 m, 2500 m, 5000 m and 10000 m), and compared them with the curves of zircon disappearance during cooling as calculated with Watson's equation (1996, eq. 17).

The results of the above are shown in Figure 6.10, from which the following stands out: During cooling of a 300 m thick body, independent of the depth of intrusion (up to 10 km), all zircons with a radius larger than 40 μm will survive. If the sill thickness increases up to 450 m, only zircons with a radius larger than 70 μm would survive. If the sill thickness increases up to 600 m, even the zircons with a radius of 80 μm will disappear unless the sill is emplaced at a depth less than 2 km. It seems, therefore, that the critical parameter governing zircon survival in granite magmas emplaced in the upper crust is the thickness of the magmatic body, with the depth of intrusion seemingly playing a secondary role.

These calculations indicate that the elevated zircon inheritance of the Cambro-Ordovician igneous rocks of Central Iberia resulted from the swiftness of the magmatic pulses, which apparently lasted just a few thousands of years from the beginning of melting to final emplacement. The fact that the fraction of zircons with inherited cores decreases from > 70-80% in all rocks of the Ollo de Sapo Domain and most metagranites of the Schist-Graywacke Complex Domain to about 50% in Guadarrama, can be easily understood because in this area the accumulation of Cambro-Ordovician granitoids occurred at batholithic scale, which undoubtedly resulted in slower cooling rates and, therefore, enhanced zircon dissolution.

6.2.3.5. Discussion of results

The applicability of the above calculations to geological systems depends on the validity of the numerical model of zircon dissolution, the proper choice of the initial conditions and model parameters, and the deviations caused by (i) the residence of Zr in minerals other than zircon and (ii) the variable grain size and non-spherical shape of zircon crystals. These circumstances can be evaluated as follows:

Watson's (1996) equation relies on two factors: First, on zircon solubility in silicic melts, which for non-peralkaline liquids depends primarily on the temperature and the melt major-element composition (Watson and Harrison, 1983). Second, on Zr diffusion in the melt, which depends additionally on the H₂O content of the melt (Harrison and Watson, 1983). The equation treats temperature and melt composition as independent variables (the latter for calculating zircon solubility), but it assumes a constant 3% H₂O for the melt. In principle, this assumption might be a serious limitation. In practice, however, it does not critically affect the model. First, because crustal magmas rarely have less than 2-3 % H₂O (e.g., Clemens, 1984; Carrington and Harley, 1996; etc.). Second, because the effects on Zr diffusion mostly occur in the first 2-3% dissolved H₂O (Harrison and Watson, 1983).

It should also be considered that Watson's equation is a simplification. According to this author, however, it deviates less than 10% from the results of more rigorous, moving boundary, finite-difference methods (Watson, 1996). Therefore, since this deviation is perfectly tolerable for our purposes and nothing indicates that the Cambro-Ordovician magmas of Central Iberia were exceptionally H₂O-poor, we can accept the numerical foundation for calculating the curves of zircon disappearance plotted in Figs. 6.8 and 6.9.

The choice of a Pan-African protolith is justified by the dominance of inherited 600-620 Ma ages (Figs. 4.2 and 4.26). Moreover, this choice is not critical because the zircon grain-size distribution and the Zr concentration estimated for this protolith are typical for common granitoids. Only if the protolith had zircons with a spherical radius larger than 150 μm , or a bulk rock concentration of Zr higher than 400-500 ppm, would we expect significant departures from the model. Since such features are mostly limited to peralkaline rocks, and these are unknown among the 600-620 Ma magmatism of north-west Gondwana (e.g. Gasquet et al., 2005), we can safely exclude this possibility. Neither is the major-element composition of the melt critical, because it was necessarily silicic and variations of less than 5% SiO₂ have little effect.

In contrast, the determination of the peak temperature reached by the melts is crucial, especially if it were overestimated. The Ti-in-zircon thermometer requires TiO_2 activity equal to one (Watson et al., 2006). In the present case, the presence of primary ilmenite and rutile inclusions indicates that such a condition was satisfied. If not, it would have caused underestimation, which would not invalidate our conclusions but, instead, indicate even faster heat-transfer rates. More important, perhaps, is that Ti-in-zircon temperatures can be easily overestimated if there were minute inclusions of Ti-bearing minerals or glass within the analyzed volume. The problem is especially serious when Ti is determined using a LA-ICPMS system such as the one we used here, which to get reasonable ^{49}Ti sensitivity requires ablation of craters with a diameter of 60 μm and a depth of about 40 μm . Nevertheless, the careful selection, under the microscope, of the areas to be analyzed and, especially, the nearly Gaussian distribution of the results, notably exempt of outliers (see Fig. 6.7), indicates that inclusions have caused little trouble in the present case. Consequently, we can assume that the maximum temperatures recorded by the Ti-in-zircon thermometer (900-950°C) represent a minimum estimate of the magma's thermal peak, an assumption totally consistent with a large body of experimental data indicating that temperatures of this order are required for generating large volumes of silicic crustal magma in vapor-absent conditions (e.g. Clemens, 2003; and references therein mentioned).

A final consideration is that Watson's equation assumes that all zircon grains are spheres of the same size, and that all Zr resides in zircon. Real rocks, however, have non-spherical zircon crystals, and these are of different shapes and sizes. Real rocks also have a variable fraction of Zr residing in minerals other than zircon such as titanite, amphibole or garnet (e.g., Bea et al., 2006). Certainly, all these differences may affect the zircon dissolution rate. Shapes other than a sphere would increase it, since a sphere represents the smallest surface/volume ratio. Nevertheless, the existence of a large variety of zircon sizes could delay the dissolution of the largest grains because the Zr concentration in the melt would increase rapidly owing to the fast dissolution of the smallest grains. Zircon dissolution would also be delayed if phases other than zircon release Zr to the melt but, on the other hand, it would be accelerated if a Zr-bearing mineral such as garnet appears as a product of melting reactions and extracts Zr from the melt. However, on balance, none of these effects are likely to be of great importance and, what is more, they tend to mutually cancel, so we can accept that the results of Watson's equation acceptably match the real situation, i.e. that the magmatic pulses that generated the Central Iberian Cambro-Ordovician magmas were

very fast, probably taking around two thousand years from the beginning of melting until their eruption or emplacement as thin sills or laccoliths at upper crustal levels.

6.3. Origin and geodynamic significance of the Cambro-Ordovician magmas of the Central Iberian Zone.

The origin and significance of the Cambro-Ordovician magmatism of Central Iberia has been controversial. Whereas most authors agree that the coeval peralkaline association found in the Ossa Morena and Galicia-Tras-Os-Montes zones was generated in a rifting environment (e.g., Ribeiro 1987; Ribeiro and Floor 1987; Montero et al. 1998; Santos Zalduegui et al. 1995; Montero and Floor 2004a), the peraluminous association, Ollo de Sapo and Guadarrama orthogneisses included, has been interpreted by other authors as evidence of an active margin setting, simply because of the supposed calc-alkaline chemical signature (e.g., Gebauer et al. 1993; Valverde-Vaquero and Dunning 2000; von Raumer et al. 2003; Bea and Navidad 2004).

However, the interpretation of the peraluminous Cambro-Ordovician rocks of Iberia as a product of arc-related magmatism does not find support from chemical, isotopic, and geochronologic or geologic evidence. First, because the link between the geochemical signature and geodynamic setting is not definitive and may be equally explained as a legacy from their protoliths, as proposed for the Cambro-Ordovician rocks of the northern Bohemian Massif by Klimas-August (1990) and Floyd et al. (2000). Second, and more importantly, because the chemical composition of these rocks is not calc-alkaline but alkaline-calcic (Fig. 5.1), and bears no resemblance with the composition of subduction-related igneous rocks with similar SiO₂ contents (Figs. 5.2 and 5.3). On the contrary, it shows a striking similarity with the composition of collision-related granites derived from partial melting of old continental materials, which is totally consistent with the high abundance of zircon crystals with old, pre-magmatic, cores. The fact that the most abundant inherited age population, found in all rock types irrespective of their locality is Ediacaran, with a mean age of 605-625 Ma (Figs 4.2 and 4.26) suggests that the source of the Cambro-Ordovician magmas must have been mainly composed of igneous rocks of Ediacaran age, or sediments derived from them. Sr and Nd isotopes point towards the same conclusion. The isotope composition of the metavolcanic rocks and metagranites, with $(^{87}\text{Sr}/^{86}\text{Sr})_{\text{init}} \approx 0.709\text{-}0.711$, $\epsilon\text{Nd}_{\text{init}} \approx -5.8$ to -3.3 , and $T_{(\text{DM})}$ around 1.8-1.9 Ga, does not reflect the input of juvenile arc-material but indicates derivation from materials with a long

crustal residence. The only exception are two metagranites, Miranda do Douro and San Sebastián, which are isotopically more primitive probably as the result of some hybridization with juvenile mafic magmas similar to the Cambro-Ordovician Carrascal gabbrodiorites at the southern boundary of the Central Iberian Zone (Fig. 6.3). But, as discussed below, this is not an acceptable argument in favor of arc-related magmatism.

The geodynamic environment proposed for the Cambro-Ordovician magmas must be compatible with the following facts:

1. Very fast generation of crustal melts, as revealed by zircon inheritance.
2. Short duration and synchronicity of the magmatism all over the Central Iberian Zone.
3. The lack of coeval regional metamorphism and deformation.

With respect to the first point, it should be considered that the only heating mechanism capable of melting crustal materials at the required rate is the advection of heat by mafic magmas, as revealed by the numerical analysis of Huppert and Sparks (1989). Other crustal-heating mechanisms have much large time-constants, from 10^5 - 10^6 years for the displacement of isotherms caused by tectonics, burial or erosion (e.g., Chapman and Furlong, 1992; Zen, 1995; Huerta et al., 1999) to 1 - 3×10^7 years for radiogenic heating (e.g. Vanderhaeghe and Teyssier, 2001; Bea et al., 2003). The analysis of Huppert and Sparks (1989) also predicts that felsic magmas were generated following the intrusion of mafic magmas would have peak-temperatures of 900°C , abundant pre-magmatic crystals and a highly porphyritic character, a set of features found in the Central Iberia Cambro-Ordovician rocks which are difficult to explain by any other mechanism. This gives additional support to the idea that heat for crustal melting was supplied by mantle-derived mafic magmas. Additionally, the imperceptible hybridization between these and the felsic magmas, except perhaps in San Sebastián and Miranda do Douro metagranites suggests that the meltable crustal material directly overlay the mafic intrusions, where the density difference and the quick solidification of the mafic magma at the contact would make mixing unlikely (Huppert and Sparks, 1989). This places the mafic intrusions at the crust-mantle interface. Lastly, the fast melt transport to upper crustal levels points to extensional rather than compressional forces. All these reasons, therefore, strongly suggests that the Central Iberian Cambro-Ordovician magmas were generated during the rifting of a continental crust caused by an upwelling mantle-plume, which likely occurred during the

early Paleozoic fragmentation and dispersal of terranes from the northern margin of west Gondwana (Crowley et al., 2000; Matte, 2001) and ultimately led to the formation of the Iberian microplate. This mechanism would, additionally, explain the synchronicity of the Cambro-Ordovician magmatic event in Central Iberia, the lack of coeval regional metamorphism and deformation, and the short duration of the magmatism throughout this zone.

In this scenario crustal deformation depends to a significant extent on the mechanical coupling between mantle and crust (Burov and Guillou-Frottier, 2005); if the coupling is weak, the concentration of plume-related extension in the mantle lithosphere has little effect on the crust. This might explain the scarcity of contemporaneous deformation. The lack of any Cambro-Ordovician metamorphic imprint in all exposed mid-crustal sections of Central Iberia may also be explained by the swiftness of the generation and emplacement of the crustal magmas. Metamorphism involves conductive heat transfer, an inherently slow process that requires much more than a few thousands of years to be perceptible at a crustal scale. Here, the heat advected to the lower crust by mantle magmas was first consumed by melting reactions and then quickly transported to the uppermost crustal sections by the so-produced magmas causing, therefore, a negligible thermal impact on most of the crustal section above the melting zone except, perhaps, the first lowermost 1000-2000 m.

6.4. The age of the Cambro-Ordovician rifting in Iberia

When the geochronological data obtained in this work is combined with all currently available crystallization ages for the Iberian Cambro-Ordovician magmatic rocks (Fig. 6.11, and references therein), the following stands out: In the Central Iberian Zone the oldest crystallization ages of the metavolcanic rocks cluster around 495 Ma, and the youngest around 483 Ma, with the maximum production of magma at 490-485 Ma. After 483 Ma, the magmatic activity was only plutonic and was mostly concentrated in the Guadarrama and Hiendelaencina regions (Fig. 1.8). The exceptions are the 470 Ma San Sebastián metagranite, and the 471 Ma Carrascal massif, spatially associated to the Urra Formation, which is the only Cambro-Ordovician mafic body of Central Iberia (Solá et al. 2005). The last magmatic event in the Central Iberian Zone prior to ~495 Ma produced the 543 Ma Almohalla metagranite (Bea et al., 2003) and the 570-580 Ma Merida-Aljucén massif (Bandrés et al. 2004; Talavera et al. 2008a). After ~470 Ma, the magmatic activity

ceased totally until the intrusion of the first Variscan granitoids at ~350 Ma (Montero et al. 2004b).

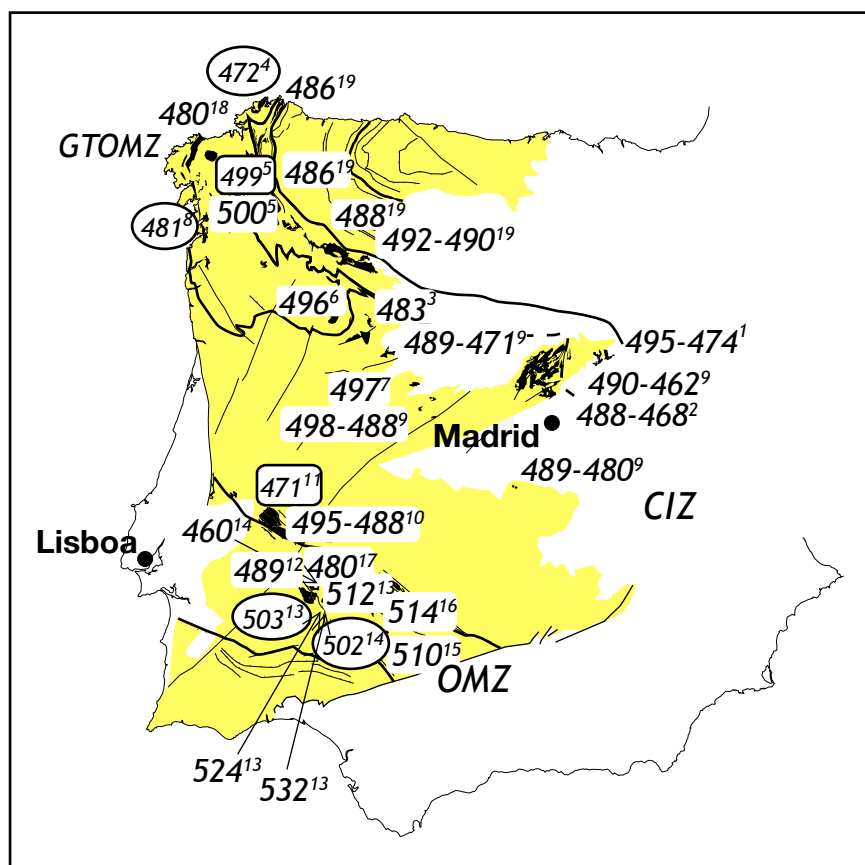


Figure 6.11: Zircon U-Pb crystallization ages of the Iberian Cambro-Ordovician magmas. Numbers within ellipses represent peralkaline rocks. Numbers with squares represent mafic bodies. The rest are felsic peraluminous rocks. Superscripts represent data sources: 1, Montero et al., 2007*; 2, Valverde-Vaquero and Dunning, 2000; 3, Bea et al., 2006*; 4, Valverde-Vaquero et al., 2005; 5, Abati et al., 1999; 6, Dallmeyer and Quesada, 1992; 7, Bea et al., 2009*; 8, Montero et al., 2008; 9, unpublished data of this thesis; 10, Solá et al., 2008; 11, Solá et al., 2005; 12, Schäfer, 1990; 13, Ochsner, 1993; 14, Montero et al., 2000; 15, Montero et al., 1999; 16, Ordóñez Casado, 1998; 17, Expósito Ramos et al., 2003; 18, Talavera et al., 2008b; 19, **Montero et al., in press***. Note the references with * superscript contain data of this thesis.

In the Ossa Morena Zone, to the south (Fig. 6.11), the Cambro-Ordovician magmatism spanned a longer period than in Central Iberia, from 515 Ma to 460 Ma without a clear maximum of magma production, and with an episode of peralkaline magmatism at

500-505 Ma (Ochsner 1993; Salman 2002). In the Galicia-Tras-os-Montes Zone, to the north (Fig. 6.11), despite the lithologies of the Cambro-Ordovician rocks being similar to Ossa Morena, the crystallization ages are akin to Central Iberia, with the oldest clustering around 500-495 Ma (Dallmeyer and Quesada, 1992 and Abati et al., 1999) and the youngest around 470 Ma (Valverde-Vaquero et al., 2005). Remarkably, the Galician peralkaline rocks have ages from 482 Ma (the Galiñeiro massif, Montero et al., 2008) to 475 Ma (alkaline rhyolites of the Schistose Domain, Valverde-Vaquero et al. 2005). Therefore, they are 20-25 Ma younger than their counterparts in Ossa Morena to the south and, in contrast to that zone, apparently closed the Cambro-Ordovician magmatic cycle.

It follows that in the Central Iberian and Galicia-Tras-os-Montes Zones, the Cambro-Ordovician rifting started at ~500-495 Ma, reached a maximum activity between 495-483 Ma, depending on the particular location, and ceased at ~470 Ma. In Ossa Morena, on the other hand, the situation is not so clear, because there was episodic magmatism throughout the Cambrian. The beginning of the rifting would have occurred after the emplacement of the the 532 Ma Culebrín tonalite, the composition of which is hardly compatible with a rifting environment (Salman, 2002), and rifting must have continued to 500-505 Ma, when the peralkaline magmas crystallized (Fig. 6.11 and references therein). In addition, both the 515-510 Ma and the 460 Ma peraluminous granites were probably generated during the Cambro-Ordovician rifting (Exposito Ramos et al., 2003). Therefore, these dates, 515 Ma to 460 Ma, would mark the minimum duration of the rifting in this region. If so, the extension began before, and was longer, than in the rest of Iberia, which suggests that the Ossa Morena crust overlay a zone of the lithospheric mantle that was particularly active during the Lower Paleozoic extension.

7. SUMMARY AND CONCLUSIONS

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7.1. Summary and conclusions

The Cambro-Ordovician magmatism of the Central Iberian Zone is concentrated in three NW-SE roughly parallel curved lineaments (Fig.1.1A), two of them in the northern half of the Central Iberian Zone, and the other to the south, at the boundary with the Ossa Morena Zone. From north to south, these three lineaments are:

1. The metavolcanic rocks and metagranites of the Ollo de Sapo Domain.
2. The metagranites and metavolcanic rocks of the northern Schist-Graywacke Complex Domain, which are concentrated in four areas: the Tormes Dome, the northwestern part of the Sierra de Gredos, the Sierra de Guadarrama and the Anatectic Complex of Toledo.
3. The Urra Formation, located in Portugal.

This work has concentrated on the two first lineaments. Its main objectives have been to determine accurately the U-Pb (ion microprobe and LA-ICPMS) and/or Pb-Pb (stepwise sequential evaporation) crystallization ages and Sr and Nd isotope geology of these rocks. From these data, the genesis and geodynamic significance of the Cambro-Ordovician magmatism of Central Iberia is discussed.

<i>Metavolcanic rocks of the Ollo de Sapo Formation</i>	
Vivero	486 ± 3 Ma
Trives	486 ± 5 Ma
Sanabria	492 ± 4 Ma (lower metavolcanic rocks) 490 ± 2 Ma (upper metavolcanic rocks)
Villadepera	483 ± 4 Ma
Hiendelaencina	495 ± 5 Ma (lower metavolcanic rocks) 483 ± 3 Ma (upper metavolcanic rocks)

From north to south, the average crystallization ages of the metavolcanic rocks of the Ollo de Sapo Formation are:

<i>Metagranites associated with the Ollo de Sapo Formation</i>	
Viana do Bolo (Covelo) (Sanabria region)	487 ± 4 Ma
San Sebastián (Sanabria region)	470 ± 3 Ma
Miranda do Douro (Villadepera region)	483 ± 3 Ma
Antoñita (Hiendelaencina region)	474 ± 4 Ma

<i>Tormes Dome</i>	
Fermoselle	486 ± 6 Ma
Vitigudino	489 ± 5 Ma
Ledesma	471 ± 7 Ma
<i>NW Gredos</i>	
San Pelayo	492 ± 4 Ma
Castellanos	498 ± 4 Ma
Bercimuelle	488 ± 3 Ma
<i>Guadarrama</i>	
La Cañada	489 ± 9 Ma
La Hoya	478 ± 7 Ma
Santa María de la Alameda	490 ± 5 Ma (<i>Estación I</i>) 473 ± 8 Ma (<i>Estación II</i>)
Vegas de Mature	476 ± 5 Ma (<i>melanocratic</i>) 484 ± 6 Ma (<i>leucocratic</i>)
La Morcuera	482 ± 8 Ma
Buitrago de Lozoya	481 ± 9 Ma
La Berzosa (<i>metavolcanic</i>)	462 ± 11 Ma
El Cardoso (<i>metavolcanic</i>)	483 ± 6 Ma
Riaza (<i>metavolcanic</i>)	475 ± 5 Ma
<i>Toledo</i>	
Polán	480 ± 8 Ma
Mochares	489 ± 7 Ma

The average crystallization age of the metagranites spatially associated with the metavolcanic rocks of the Ollo de Sapo Formation, again from north to south, are:

The average crystallization age of the metagranites and metavolcanic rocks of the northern area of the Schist-Graywacke Complex are:

These data reveal that in the Ollo de Sapo Domain, the magmatic activity was synchronous and short lived, except in the Hiendelaencina region where was exceptionally prolonged. In the northwest of the Ollo de Sapo Domain and Sanabria and Villadepera regions, the whole magmatic activity occurred at the same time, between 492 and 483 Ma. After this time, there is only one small igneous body in the Sanabria region, San Sebastián metagranite (470 Ma). In the Hiendelaencina region, there are, at least, two volcanic episodes at 495 and 483 Ma, which occurred before the plutonic activity, at 474 Ma.

In the north of the Schist-Graywacke Complex Domain, the Cambro-Ordovician magmatism began in the Sierra de Gredos, where it took place from 498 Ma to 488 Ma. Apparently, the magmatism ceased in this region at the same time that started in the Tormes Dome (489 Ma to 471 Ma), Toledo (489 Ma to 480 Ma) and Guadarrama, where it occurred almost uninterruptedly from 490 Ma to 462 Ma producing the only batholithic accumulation of Cambro-Ordovician granitoids of Iberia.

All the Cambro-Ordovician igneous rocks of the Ollo de Sapo Domain and the north of the Schist-Graywacke Complex Domain belong to the peraluminous association, as defined by Montero et al. (2008). In these regions no rocks of the peralkaline association have been found so far. The rocks from both domains have comparable major and trace element compositions: they are markedly peraluminous, magnesian and alkaline-calcic (Fig. 5.1) characterized by enrichments in the most incompatible elements with a positive Li anomaly and a negative Sr anomaly (compared to the average continental crust, Fig. 5.2), Th and U are elevated, and REE patterns decrease smoothly from La (100-200 x chondrite) to the HREE (5-15 x chondrite) with a constant negative Eu anomaly (Fig. 5.3). Such a composition is typical of granites derived from crustal materials with little, if any, input of juvenile mantle materials.

In the same way, the metavolcanic rocks and some of the metagranites of both areas also have similar Sr and Nd isotope compositions. The metavolcanic rocks of the Ollo de Sapo Formation (Fig. 5.4 and Table 5.2) cluster around $(^{87}\text{Sr}/^{86}\text{Sr})_{490\text{Ma}} \approx 0.710$, $\epsilon\text{Nd}_{490\text{Ma}} \approx -4.6$, and $T_{(\text{DM})} \approx 1.8$ Ga. Two Ollo de Sapo Formation metagranites, Antofñita and Viana do Bolo, have the same composition but the other two, Miranda do Douro and San Sebastián, are more primitive, with $(^{87}\text{Sr}/^{86}\text{Sr})_{490\text{Ma}} \approx 0.704$, $T_{(\text{DM})} \approx 1.3$ Ga and variable $\epsilon\text{Nd}_{490\text{Ma}}$ which ranges from -3.40 to -1.74 in Miranda do Douro and from 0.2 to 2.1 in San Sebastián (Fig. 5.4). The Sr and Nd isotope composition of the metagranites and metavolcanic rocks from the north of the Schist-Graywacke Complex Domain is also similar to the metavolcanic rocks of the Ollo de Sapo Formation, with $(^{87}\text{Sr}/^{86}\text{Sr})_{490\text{Ma}} \approx 0.711$, $\epsilon\text{Nd}_{490\text{Ma}} \approx -3.6$ to -4.7 and $T_{(\text{DM})} \approx 1.9$. These values support the inference drawn from the major and trace element compositions, that the source rocks from which the Cambro-Ordovician magmas of Central Iberia derived were continental crust with little material contribution from mafic magmas, except in the most primitive metagranites, Miranda do Douro and San Sebastián.

One of the most remarkable features of all the studied rocks is the anomalously elevated proportion of zircons with inherited cores, between 70 and 80%, on occasions up to nearly 100% in the Ollo de Sapo Domain, and between 50 and 90% in the north of the Schist-Graywacke Complex Domain. The age distribution of the inherited cores reveals a vast preponderance of Ediacaran components, though there are some differences between the two domains. In both there are: (i) a 605-625 Ma population, which is more abundant in the Ollo de Sapo Domain; (ii) a Cryogenian (650-700 Ma) population; (iii) a Tonian (850-900) population; (iv) an Orosirian (1900-2000 Ma) population, which is more abundant in the Ollo de Sapo Domain; (v) a Neoproterozoic population; and (vi) a Mesoarchean population. But in the metagranites of the northern Schist-Graywacke Complex Domain we have additionally found two younger populations at 540 Ma and 571 Ma which have not been observed in the Ollo de Sapo Domain. The occurrence of these populations is related to the existence of 540 Ma late Pan-African granitoids such as the Almohalla orthogneiss in the Sierra de Gredos (Bea et al., 2003). The dominance of Ediacaran cores indicates that the magmatic source was dominated by igneous rocks of this age, or immature sediments derived from them. The isotopic composition of the Cambro-Ordovician rocks of Central Iberia is compatible with them being derived from late-Pan-African rocks similar to the ones currently found in the Moroccan Anti-Atlas.

The high proportion of inherited cores was maintained despite the fact that the peak temperature of the magmas, estimated with the Ti-in-zircon geothermometer (Watson and Harrison, 2005; Watson et al., 2006) at 826°C and some values exceeding the 950°C, surpassed the rock's zircon saturation temperature, estimated with Watson and Harrison (1983) model at 815°C. Therefore, according to the calculations of heating and cooling rates estimated with Watson's equation for the kinetics of zircon dissolution (Watson, 1996), abnormally elevated zircon inheritance of the Central Iberia can be explained by swift magmatic pulses, which lasted just a few thousands of years from the beginning of melting to final emplacement and solidification.

The Cambro-Ordovician magmatism of Central Iberia is not a product of arc-related magmatism as has been repeatedly suggested solely because of its supposedly "calc-alkaline" signature. Our data reveal that neither the chemical nor the isotopic composition of these rocks bears any resemblance to the composition of subduction-related igneous rocks with similar SiO₂ contents but are, instead, akin to collision-related granites derived from partial melting of old continental materials. This is totally consistent with the high abundance of zircon crystals with old pre-magmatic cores.

The very fast rate of generation of crustal melts revealed by zircon inheritance can only be produced by advective heating caused by mafic magmas. Additionally, the imperceptible hybridization between these and the felsic magmas, except perhaps in San Sebastián and Miranda do Douro metagranites, suggests that the meltable crustal material just overlay the mafic intrusions, where the density difference and the quenching of the mafic magma at the contact would make mixing unlikely (Huppert and Sparks, 1989). This places the mafic intrusions at the crust-mantle interface. Lastly, the fast melt transport to upper crustal levels points to extensional rather than compressional forces. All these reasons, therefore, strongly suggest that the Central Iberian Cambro-Ordovician magmas were generated during rifting of a continental crust caused by an upwelling mantle-plume, which likely occurred during the early Paleozoic fragmentation and dispersal of terranes from the northern margin of west Gondwana (Crowley et al., 2000; Matte, 2001) and ultimately led to the formation of the Iberian microplate. This mechanism would, additionally, explain the synchronicity of the Cambro-Ordovician magmatic event in Central Iberia, the lack of coeval regional metamorphism and deformation, and the short duration of the magmatism throughout this zone.

Finally, when our data are combined with all the available geochronological data from the Central Iberian, Galicia Tras-os-Montes and Ossa Morena Zones, it is possible to determine the timing of the Cambro-Ordovician rifting in Iberia. In Central Iberia and Galicia Tras-os-Montes it started at approximately 500-495 Ma, reached a maximum between 495-483 Ma, depending on the particular area, and finished at approximately 470 Ma. In the Ossa Morena Zone, however, the situation is not so clear because there were many different magmatic episodes during the Cambrian. Our estimate is that 515 Ma to 460 Ma would mark the minimum duration of the Ossa Morena rifting so that in this region the extension started before and was more prolonged than in the rest of Iberia.

7.2. Conclusiones

El magmatismo Cambro-Ordovícico de la Zona Centroibérica se concentra en tres bandas curvadas y casi paralelas, con dirección NW-SE (Fig. 1.1A). Dos de ellas se localizan en la mitad norte de la Zona Centroibérica y la tercera se localiza en el sur, en el límite con la Zona de Ossa Morena. De norte a sur, estas tres bandas son:

1. Las rocas metavolcánicas y los metagranitos del Dominio de Olla de Sapo.
2. Las rocas metavolcánicas y los metagranitos del norte del Dominio del Complejo Esquito-Grauváquico, que se localizan en cuatro áreas: en el Domo del Tormes, en el noroeste de la Sierra de Gredos, en la Sierra de Guadarrama y en el Complejo Anatético de Toledo.
3. La Formación Urra, situada en Portugal.

Esta tesis se ha concentrado en las dos primeras bandas y sus principales objetivos han sido la determinación precisa de las edades de cristalización U-Pb por microsonda iónica y ablación láser, y/o Pb-Pb por evaporación secuencial, y el estudio de la geología isotópica de Sr y Nd. A partir de estos datos, se ha discutido la posible génesis y el significado geodinámico del magmatismo Cambro-Ordovícico de la Zona Centroibérica.

De norte a sur, las edades de cristalización de las rocas metavolcánicas de la Formación Olla de Sapo son:

<i>Rocas metavolcánicas de la Formación Olla de Sapo</i>	
Vivero	486 ± 3 Ma
Trives	486 ± 5 Ma
Sanabria	492 ± 4 Ma (Rocas metavolcánicas inferiores) 490 ± 2 Ma (Rocas metavolcánicas superiores)
Villadepera	483 ± 4 Ma
Hiendelaencina	495 ± 5 Ma (Rocas metavolcánicas inferiores) 483 ± 3 Ma (Rocas metavolcánicas superiores)

De norte a sur, las edades de cristalización de los metagranitos asociados espacialmente a las rocas metavolcánicas de la Formación Olló de Sapo son:

<i>Metagranitos asociados a la Formación Olló de Sapo</i>	
Viana do Bolo (Covelo) (Región de Sanabria)	487 ± 4 Ma
San Sebastián (Región de Sanabria)	470 ± 3 Ma
Miranda do Douro (Región de Villadepera)	483 ± 3 Ma
Antoñita (Región de Hiendelaencina)	474 ± 4 Ma

Las edades de cristalización de los metagranitos y las rocas metavolcánicas de la zona norte del Dominio del Complejo Esquisto-Grauváquico son:

<i>Domo del Tormes</i>	
Fermoselle	486 ± 6 Ma
Vitigudino	489 ± 5 Ma
Ledesma	471 ± 7 Ma
<i>Noroeste de Gredos</i>	
San Pelayo	492 ± 4 Ma
Castellanos	498 ± 4 Ma
Bercimuelle	488 ± 3 Ma
<i>Guadarrama</i>	
La Cañada	489 ± 9 Ma
La Hoya	478 ± 7 Ma
Santa María de la Alameda	490 ± 5 Ma (<i>Estación I</i>) 473 ± 8 Ma (<i>Estación II</i>)
Vegas de Mature	476 ± 5 Ma (<i>melanocrático</i>) 484 ± 6 Ma (<i>leucocrático</i>)
La Morcuera	482 ± 8 Ma
Buitrago de Lozoya	481 ± 9 Ma
La Berzosa (<i>Rocas metavolcánicas</i>)	462 ± 11 Ma
El Cardoso (<i>Rocas metavolcánicas</i>)	483 ± 6 Ma
Riaza (<i>Rocas metavolcánicas</i>)	475 ± 5 Ma
<i>Toledo</i>	
Polán	480 ± 8 Ma
Mochares	489 ± 7 Ma

Estos datos revelan que la actividad magmática en el Dominio de Olló de Sapo fue sincrónica y de corta duración, excepto en la región de Hiendelaencina donde fue excepcionalmente larga. En el noroeste del Dominio de Olló de Sapo y en las regiones de Sanabria y Villadepera, la actividad magmática ocurrió al mismo tiempo, entre los 492 y los 483 Ma. Después de este momento, no se observa actividad magmática en estas áreas salvo un pequeño cuerpo ígneo en la región de Sanabria, el metagranito de San Sebastián (470 Ma). En la región de Hiendelaencina, se observan, al menos, dos episodios volcánicos a 495 y a 483 Ma que ocurrieron antes de la actividad plutónica que se produjo a los 474 Ma.

En el norte del Dominio del Complejo Esquito-Grauváquico, el magmatismo Cambro-Ordovícico comenzó en la Sierra de Gredos a los 498 y terminó a los 488 Ma. Aparentemente, el magmatismo cesa en esta región al mismo tiempo que comienza en el Domo del Tormes (entre los 489 y los 471 Ma), Toledo (entre los 489 y los 480 Ma) y Guadarrama, donde tuvo lugar casi ininterrumpidamente desde los 490 a los 462 Ma generando la única acumulación de granitoides Cambro-Ordovícicos en Iberia con dimensiones batolíticas.

Todas las rocas ígneas Cambro-Ordovícicas del Dominio de Olló de Sapo y del norte del Dominio del Complejo Esquito-Grauváquico pertenecen a la asociación peraluminosa definida por Montero et al. (2008). En estas regiones no se han encontrado rocas de la asociación peralcalina. La composición de elementos mayores y trazas rocas de las rocas de ambos dominios es similar: son marcadamente peraluminosas, magnesianas y alcalino-cálcicas (Fig. 5.1) caracterizadas por un enriquecimiento en los elementos más incompatibles con una anomalía positiva de Li y una anomalía negativa de Sr (comparado con la composición media de la corteza continental, Fig. 5.2), un alto contenido de U y Th y un patrón de REE que decrece suavemente desde el La (100-200 x condrito) a las HREE (5-15 x condrito) con una anomalía negativa de Eu (Fig. 5.3). Esta composición es típica de granitos que proceden de materiales corticales antiguos con una pequeña contribución de materiales jóvenes mantélicos.

La composición isotópica de Sr y Nd de las rocas metavolcánicas y los metagranitos de los dos dominios es similar. Las rocas metavolcánicas de la Formación Olló de Sapo (Fig. 5.4 y Tabla 5.2) tienen un relación $(^{87}\text{Sr}/^{86}\text{Sr})_{490\text{Ma}} \approx 0.710$, $\epsilon_{\text{Nd}490\text{Ma}} \approx -4.6$, y $T_{(\text{DM})} \approx 1.8$ Ga. Dos de los cuatro metagranitos asociados a la Formación Olló de Sapo, Antoñita y

Viana do Bolo, tienen la misma composición isotópica que las rocas metavolcánicas mientras que los otros dos metagranitos, Miranda do Douro y San Sebastián son más primitivos, con una relación $^{87}\text{Sr}/^{86}\text{Sr}_{490\text{Ma}} \approx 0.704$, $T_{(\text{DM})} \approx 1.3$ Ga y $\epsilon\text{Nd}_{490\text{Ma}}$ variable que oscila entre -3.40 y -1.74 en Miranda do Douro y entre 0.2 y 2.1 en San Sebastián (Fig. 5.4). La composición isotópica de Sr y Nd de los metagranitos y las rocas metavolcánicas del norte del Dominio del Complejo Esquisto-Grauváquico es similar a la de las rocas metavolcánicas de la Formación Olló de Sapo, con una relación $(^{87}\text{Sr}/^{86}\text{Sr})_{490\text{Ma}} \approx 0.711$, un $\epsilon\text{Nd}_{490\text{Ma}}$ que oscila entre -3.6 y -4.7 y un $T_{(\text{DM})} \approx 1.9$. Estos valores apoyan la idea, inferida a partir de la composición de elementos mayores y trazas de estas rocas, de que la fuente de los magmas Cambro-Ordovícicos de Iberia Central fue una corteza continental con una pequeña contribución de magmas máficos, excepto en las rocas más primitivas, Miranda do Douro y San Sebastián.

Una de las características más destacadas de las rocas estudiadas es la elevada proporción de circones con núcleos heredados, que oscila entre un 70 y un 80%, llegando, en ocasiones, a alcanzar el 100 % en el Dominio del Olló de Sapo, y entre un 50 y un 90 % en el norte del Dominio del Complejo Esquisto-Grauváquico. La distribución de edades de los núcleos heredados pone de manifiesto el predominio de un componente Ediacariense, aunque existen diferencias entre los dos dominios. Ambos tienen: (i) una población de 605-625 Ma, que es más abundante en el Dominio del Olló de Sapo; (ii) una población Criogeniense (650-700 Ma); (iii) una población Toniense (850-900); (iv) una población Orosiriense (1900-2000 Ma), que es más abundante en el Dominio del Olló de Sapo; (v) una población Neoarcaica; y (vi) una población Mesoarcaica. Pero en los metagranitos del norte del Dominio del Complejo Esquisto-Grauváquico además se han observado dos poblaciones más jóvenes a 540 y a 571 Ma, que no se han encontrado en el Dominio de Olló de Sapo. La presencia de estas poblaciones se ha relacionado con la existencia de granitoides Panafricanos como el ortogneis de la Almohalla de la Sierra de Gredos (Bea et al., 2003). La abundancia de núcleos Ediacarienses indica que la fuente magmática estaba compuesta en gran parte por rocas de esta edad o sedimentos inmaduros derivados de ellas. La composición isotópica de las rocas Cambro-Ordovícicas de Iberia Central es compatible con aquellas que han sido derivadas de rocas Panafricanas similares a las que actualmente se encuentran en el Anti-Atlas marroquí.

Una alta proporción de núcleos heredados sobrevivió a pesar de que el pico de temperatura de los magmas, estimada con el geotermómetro de Ti en circón (Watson y

Harrison, 2005; Watson et al., 2006) a 826°C, excediendo algunos valores los 950°C, superó la temperatura de saturación de circón en la roca, estimada con el modelo de Watson y Harrison (1983) a 815°C. Por tanto, de acuerdo con los cálculos de los gradientes de calentamiento y enfriamiento estimados a partir de la ecuación de Watson aplicada a la cinética de disolución de circón (Watson, 1996), la elevada proporción de núcleos heredados de Iberia Central puede explicarse a partir de la existencia de pulsos magmáticos rápidos, que tuvieron una duración de apenas unos pocos miles de años desde el comienzo de la fusión hasta su emplazamiento final y su solidificación.

El magmatismo Cambro-Ordovícico de Iberia Central no es el producto de un magmatismo de arco como se ha sugerido, en repetidas ocasiones, apoyándose en su supuesta signatura “calcoalcalina”. Nuestros datos revelan que la composición química e isotópica de estas rocas no se parece a la composición de las rocas ígneas de ambientes de subducción con contenidos similares de SiO₂, sino a la composición de granitos de colisión derivados por fusión parcial de materiales continentales antiguos. Esto es compatible con la elevada proporción de circones con núcleos pre-magmáticos antiguos.

La herencia de estos circones apunta hacia una rápida generación de fundidos corticales que sólo pueden producirse por la transferencia de calor procedente de magmas máficos. Además, la inapreciable hibridación entre los magmas máficos y los félsicos, excepto quizá en los metagranitos de San Sebastián y Miranda do Douro, sugieren que el material cortical fértil estaba justo a techo de estas intrusiones máficas, siendo la mezcla entre ambos poco probable debido a la diferencia de densidad y a la rápida solidificación de los magmas máficos (Huppert y Sparks, 1989). El lugar dónde se sitúan las intrusiones máficas se localiza en la interfase corteza-manto. Por último, el rápido transporte del fundido hasta niveles de la corteza superior indican la existencia de fuerzas extensionales y no compresivas. Por tanto, todas estas razones sugieren que los magmas Cambro-Ordovícicos de Iberia Central se generaron durante el rifting de una corteza continental producida durante el ascenso de una pluma mantélica, que probablemente ocurrió durante la fragmentación y la dispersión de terrenos en el margen norte del oeste de Gondwana durante Paleozoico Inferior (Crowley et al., 2000; Matte, 2001) y que llevó a la formación de la microplaca Iberia. Este mecanismo además podría explicar el sincronismo del evento magmático Cambro-Ordovícico, la ausencia de metamorfismo regional y deformación contemporáneo y la corta duración del magmatismo a lo largo de esta zona.

Por último, si se combinan los datos de esta tesis con los datos geocronológicos disponibles de la Zonas Centroibérica, Galicia Tras-os-Montes y Ossa Morena, es posible determinar la duración del rifting Cambro-Ordovícico en Iberia. En Iberia Central y en Galicia Tras-os-Montes, comenzó aproximadamente a 500-495 Ma, alcanzando su máximo entre los 495-483 Ma, dependiendo de la zona en particular, y terminó aproximadamente a 470 Ma. En la zona de Ossa-Morena, la situación no es tan clara porque se observan diferentes episodios magmáticos durante el Cámbrico. Nuestras estimaciones señalan que la duración mínima del rifting en Ossa Morena sería de unos 55 Ma, comenzando a 515 Ma y finalizando a 460 Ma. Por lo tanto, en esta región la extensión comenzó antes y duró algo más que en el resto de Iberia.

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APPENDIX

9. APPENDIX

9.1. Tables of the geochronological data

9.1.1. The Ollo de Sapo Domain

Table 4.2: U-Pb laser ablation results of the Vivero metavolcanic rocks.

Grain	Element concentration (ppm)			Isotope ratios				Ages (Ma)			
	U	Th	Pb	(²⁰⁷ Pb/ ²⁰⁶ Pb) _{meas}	σ (%)	(²⁰⁶ Pb/ ²³⁸ U) _{meas}	σ (%)	²⁰⁶ Pb/ ²³⁸ U	²⁰⁷ Pb/ ²³⁵ U	²⁰⁷ Pb/ ²⁰⁶ Pb	(²⁰⁶ Pb/ ²³⁸ U) _{cor}
LA014a	99	78	11	0.06806	5.24	0.09905	2.15	609	667	870	603 ± 14
LA018a	312	69	122	0.16145	2.89	0.37122	1.83	2035	2260	2471	-
LA020a	377	91	126	0.15277	2.54	0.31572	3.37	1769	2066	2377	-
LA022a	717	529	76	0.06096	6.12	0.09677	6.06	596	604	638	595 ± 38
LA024a	306	50	36	0.06759	4.07	0.12088	3.8	736	766	856	732 ± 30
LA026a	277	55	21	0.05865	4.48	0.07832	1.7	486	498	554	485 ± 9
LA036a	319	37	24	0.05907	4.75	0.07775	2.63	483	498	570	481 ± 13
LA038a	221	129	24	0.06196	3.2	0.10014	2.87	615	628	673	614 ± 19
LA044a	353	364	73	0.07441	2.75	0.17422	2.6	1035	1041	1053	-
LA046a	439	572	180	0.11847	2.41	0.30947	4.26	1738	1829	1933	-
LA058a	272	35	28	0.07343	2.46	0.10487	3.7	643	735	1026	633 ± 25
LA062a	309	26	28	0.0668	3.91	0.09643	5.36	594	645	832	588 ± 33
LA070a	239	134	26	0.06127	4.38	0.10031	3.63	616	623	649	616 ± 23
LA072a	271	75	22	0.06595	2.59	0.0798	2.45	495	554	805	489 ± 13
LA074a	513	48	39	0.05729	4.36	0.08067	3.48	500	501	503	500 ± 18
LA080a	211	46	18	0.06163	4.33	0.08546	1.95	529	554	661	526 ± 11
LA082a	365	29	27	0.06009	2.7	0.07914	2.19	491	512	607	489 ± 11
LA092a	791	342	65	0.06406	4.35	0.07908	3.8	491	538	744	486 ± 19
LA094a	926	2724	108	0.06645	3.47	0.07419	2.69	461	527	821	455 ± 13
LA096a	365	53	26	0.06583	4.74	0.07253	2.31	451	514	801	446 ± 11
LA104a	252	57	19	0.06025	4.64	0.0784	3.64	487	509	613	485 ± 18
LA106a	305	96	24	0.06326	4.49	0.07816	3.77	485	528	717	481 ± 19
LA108a	322	36	93	0.11883	3.04	0.29165	3.71	1650	1781	1939	-
LA110a	288	54	60	0.12593	4.63	0.20122	5.09	1182	1526	2042	-
LA116a	452	383	37	0.06214	3.73	0.0742	3.28	1838	1886	1940	458 ± 16
LA120a	272	152	117	0.16182	3.83	0.38335	2.76	461	500	679	-
LA122a	307	135	92	0.11905	2.27	0.27846	2.61	2092	2292	2475	-
LAa1	362	153	29	0.05803	2.54	0.07842	1.72	1584	1744	1942	486 ± 9
LAa3	210	67	16	0.05874	4.47	0.07812	2.16	487	495	531	484 ± 11
LAa4	1606	560	727	0.19353	4.57	0.41111	6.55	485	498	558	-
LAa5b	376	129	37	0.06124	2.55	0.09673	1.89	2220	2521	2773	594 ± 12
LAa5c	407	183	40	0.0633	3.21	0.09484	1.54	595	606	648	581 ± 9
LAa8b	117	36	13	0.0629	4.16	0.11103	2.01	584	612	718	678 ± 14
LAa8c	116	57	14	0.06434	3.22	0.11055	1.65	679	685	705	674 ± 12
LAa6	861	709	390	0.12207	1.6	0.38057	2.68	676	694	753	-
LAa7c	1395	424	704	0.16915	2.03	0.46443	2.56	2079	2033	1987	-
LAa9	201	44	22	0.06208	3.88	0.11294	4.77	2459	2509	2549	690 ± 35
LAbz1	380	206	39	0.05672	6.38	0.09691	5.33	690	687	677	599 ± 33
LAbz3	746	91	75	0.06258	2.72	0.10483	3.54	596	573	481	642 ± 24
LAbz5b	568	208	59	0.06208	3.33	0.10152	2.54	623	635	677	622 ± 17
LAbz5c	488	197	52	0.06088	3.2	0.10409	1.87	638	638	635	638 ± 13
LAbz6	161	60	16	0.05998	4.6	0.09535	1.66	587	590	603	587 ± 10
LAbz7	189	211	113	0.16903	1.84	0.46611	2.22	2467	2512	2548	-

Appendix

LAbz8c	520	43	183	0.11972	3.19	0.35461	2.35	1957	1954	1952	-
LAbz8b	529	50	171	0.1229	1.94	0.32685	4.48	1823	1907	1999	-
LAbz9	543	177	269	0.19904	3.36	0.43269	1.92	2318	2595	2818	-
LAz2	468	136	36	0.06008	6.37	0.08036	5.75	498	518	607	496 ± 30
LAz3	815	74	60	0.05948	4.72	0.07711	2.89	479	498	585	477 ± 14
LAz4	665	116	50	0.05689	5.84	0.07823	2.36	486	486	487	486 ± 12
LAz6	767	251	73	0.06152	5.47	0.09405	3.26	579	596	658	578 ± 20
LAz9	560	250	100	0.07334	4.5	0.17077	2.33	1016	1019	1023	-
LAz10	1288	393	121	0.05979	3.62	0.09415	5.88	580	583	596	580 ± 36
LAz11	255	102	20	0.0575	2.58	0.07809	5.06	485	489	511	484 ± 25
LAz12	204	110	17	0.05843	5.05	0.07812	3.44	485	496	546	484 ± 17
LAz14	3744	342	1692	0.19907	6.37	0.43128	6.42	2312	2592	2819	-
LAz15	891	126	66	0.0593	7.81	0.07779	11	483	500	578	481 ± 55
LAz17	1094	121	81	0.05946	3.73	0.07838	3.49	486	504	584	485 ± 18
LAz17c	622	666	77	0.06057	2.64	0.10195	2.54	626	625	624	626 ± 17
LAz20	219	116	25	0.06247	3.75	0.10508	3.21	644	654	690	643 ± 22

Table 4.3: Single-grain stepwise evaporation $^{207}\text{Pb}/^{206}\text{Pb}$ data of the Vivero metavolcanic rocks.

Grain	Step	$(^{204}\text{Pb}/^{206}\text{Pb})_{\text{meas}}$	$(^{207}\text{Pb}/^{206}\text{Pb})_{\text{meas}}$	$(^{207}\text{Pb}/^{206}\text{Pb})_{\text{cor}}$	2σ	Age (Ma)
z1	1	0.000152	0.059006	0.056932	0.5	489 ± 11
	2	0.000165	0.063833	0.061439	1.2	660 ± 26
	3	0.000046	0.066605	0.066103	0.6	810 ± 12
z2	1	0.000237	0.05788	0.054555	0.5	394 ± 11
	2	0.000187	0.058143	0.055553	0.6	435 ± 13
	3	0.000051	0.067341	0.056737	0.6	481 ± 13
	4	0.000052	0.057351	0.056732	0.3	481 ± 7
z3	1	0.000153	0.060463	0.058383	0.6	544 ± 13
	2	0.000101	0.059621	0.058296	1.2	541 ± 34
z8	1	0.000054	0.057895	0.057248	0.39	501 ± 8
	2	0.000054	0.058647	0.058003	1.1	530 ± 24
	3	0.000048	0.059031	0.058476	0.6	548 ± 13
	4	0.00003	0.059561	0.059269	0.22	577 ± 5
	5	0.00004	0.060027	0.059591	0.7	589 ± 15
z11	1	0.000074	0.058099	0.057161	0.4	498 ± 18
	2	0.000083	0.058506	0.057439	0.6	508 ± 9
	3	0.000099	0.058903	0.057604	0.6	515 ± 9
	4	0.000127	0.066386	0.064713	0.5	765 ± 10
	5	0.000085	0.091242	0.090274	0.3	1431 ± 6
	6	0.000031	0.112577	0.112429	0.4	1839 ± 7
	7	0.000031	0.112644	0.112497	0.27	1840 ± 5
	8	0.000033	0.112651	0.112477	0.18	1840 ± 3
z12	1	0.000097	0.057345	0.056068	0.7	455 ± 16
	2	0.00008	0.058024	0.057006	0.3	492 ± 6
	3	0.000059	0.057554	0.056833	0.3	485 ± 6
	4	0.000047	0.057488	0.056942	0.2	489 ± 4
	5	0.000047	0.057334	0.056788	0.18	483 ± 4
	6	0.000086	0.059345	0.058237	0.2	539 ± 4
	7	0.000053	0.060508	0.059885	0.5	599 ± 11
	8	0.000056	0.060632	0.059966	0.3	602 ± 6
z2	1	0.000145	0.05879	0.056818	0.6	484 ± 13
	2	0.00002	0.057197	0.057044	0.5	493 ± 11
	3	0.000018	0.056876	0.056751	0.25	482 ± 6
	4	0.000025	0.057038	0.056812	0.4	484 ± 9
z3	1	0.000107	0.055268	0.053841	0.7	364 ± 15
	2	0.0001	0.056073	0.054747	0.4	402 ± 9
	3	0.000037	0.05736	0.05696	0.4	490 ± 9
	4	0.000034	0.057206	0.056849	0.5	486 ± 11
	5	0.000039	0.057458	0.057029	0.36	493 ± 8
	6	0.000417	0.066205	0.060312	0.4	615 ± 9
	7	0.00016	0.062191	0.060018	0.35	604 ± 7
z4	1	0.000076	0.059324	0.058362	0.6	543 ± 13
	2	0.000082	0.059437	0.058388	1	544 ± 22
z5	1	0.000236	0.060108	0.056813	0.5	484 ± 11
z6	1	0.000026	0.12187	0.12182006	0.19	1983 ± 4
	2	0.000048	0.1212	0.12085538	0.22	1968 ± 4
	3	0.000056	0.189127	0.188927	0.29	2733 ± 5
z7	1	0.000155	0.076139	0.07411095	0.24	1044 ± 5
	2	0.000111	0.091553	0.09022258	0.29	1430 ± 5
	3	0.00034	0.107858	0.10344833	0.6	1678 ± 7
	4	0.000144	0.120999	0.11935885	0.38	1947 ± 4

Table 4.4: U-Pb laser ablation (LA) and ion microprobe (IM) results of the Trives metavolcanic rocks. Analyses references ending in b mean rims and c mean cores.

Grain	Element concentration (ppm)			Isotope ratios				Ages (Ma)			
	U	Th	Pb	$(^{207}\text{Pb}/^{206}\text{Pb})_{\text{meas}}$	σ (%)	$(^{206}\text{Pb}/^{238}\text{U})_{\text{meas}}$	σ (%)	$^{206}\text{Pb}/^{238}\text{U}$	$^{207}\text{Pb}/^{235}\text{U}$	$^{207}\text{Pb}/^{206}\text{Pb}$	$(^{206}\text{Pb}/^{238}\text{U})_{\text{cor}}$
<i>Laser-Ablation ICPMS data</i>											
LA008	296	161	111	0.123	1.79	0.344	2.53	1908	1952	1998	-
LA014	181	228	20	0.061	1.79	0.091	2.53	559	574	633	558 ± 14
LA018	366	77	28	0.06	1.79	0.079	2.53	492	516	621	490 ± 12
LA020	582	140	45	0.061	1.79	0.079	2.53	492	519	641	490 ± 12
LA022	606	152	47	0.059	1.79	0.078	2.53	486	499	557	485 ± 12
LA024	265	110	25	0.06	1.79	0.091	2.53	563	568	592	562 ± 14
LA032	253	238	29	0.062	1.79	0.099	2.53	611	626	682	609 ± 15
LA042	272	116	28	0.066	1.79	0.1	2.53	613	653	795	608 ± 15
LA044	89	33	9	0.063	1.79	0.103	2.53	633	652	719	631 ± 16
LA046	345	32	26	0.06	1.79	0.079	2.53	492	509	586	491 ± 12
LA050	249	117	95	0.127	1.79	0.356	2.53	1965	2011	2059	-
LA062c	49	70	6	0.064	1.79	0.1	2.53	613	642	745	610 ± 15
LA062b	102	84	9	0.059	1.79	0.079	2.53	492	502	550	491 ± 12
LA074a	635	346	53	0.064	1.79	0.08	2.53	497	540	725	493 ± 12
LA078a	1004	402	101	0.062	1.79	0.099	2.53	606	624	686	605 ± 15
LA080a	631	148	48	0.06	1.79	0.078	2.53	484	504	598	482 ± 12
LA010a2	589	173	61	0.065	1.79	0.1	2.53	617	654	784	613 ± 15
LA018a6	522	542	48	0.058	1.79	0.074	2.53	459	473	538	458 ± 11
LA020a7	276	26	21	0.06	1.79	0.079	2.53	490	513	620	487 ± 12
LA022a8	722	70	55	0.061	1.79	0.08	2.53	495	524	652	493 ± 12
LA032a11	708	139	54	0.062	1.79	0.077	2.53	480	518	688	477 ± 12
LA036a13	807	44	62	0.058	1.79	0.082	2.53	510	517	545	509 ± 13
LA038a14	2337	213	224	0.063	1.79	0.101	2.53	618	637	706	616 ± 15
LA040a15	510	103	37	0.057	1.79	0.076	2.53	473	479	508	472 ± 12
LA044a17	470	55	38	0.062	1.79	0.083	2.53	512	541	662	510 ± 13
LA062a24	341	70	26	0.057	1.79	0.079	2.53	492	492	489	492 ± 12
<i>Ion-microprobe data</i>											
IMZ1c	n.d.	n.d.	n.d.	0.1221	1.43	0.3754	1.4	2055	2022	1988	-
IMz2c	n.d.	n.d.	n.d.	0.0614	1.12	0.0995	1.63	612	620	652	611 ± 10
IMz3b	n.d.	n.d.	n.d.	0.0607	1.07	0.0797	1.07	494	519	627	492 ± 5
IMz3b2	n.d.	n.d.	n.d.	0.06	0.47	0.0754	0.49	468	492	603	466 ± 2
IMz4c	n.d.	n.d.	n.d.	0.0623	1.73	0.0995	1.05	612	627	684	610 ± 6
IMz4b	n.d.	n.d.	n.d.	0.0575	1.23	0.0793	0.79	492	495	509	492 ± 4
IMz9c2	n.d.	n.d.	n.d.	0.0563	1.28	0.0803	0.58	498	492	464	499 ± 3
IMz7c	n.d.	n.d.	n.d.	0.0638	0.56	0.1051	0.61	644	665	734	642 ± 4
IMz12c	n.d.	n.d.	n.d.	0.0594	1.3	0.0783	1.1	486	503	583	485 ± 5
IMz10c	n.d.	n.d.	n.d.	0.0582	2.61	0.0797	1.18	495	502	536	494 ± 6
IMz16c2	n.d.	n.d.	n.d.	0.0616	1.55	0.0797	1.15	495	525	662	492 ± 6

Table 4.5: U-Pb laser ablation (LA) and ion microprobe (IM) results of the lower metavolcanic rocks of Sanabria region. Analyses references ending in b mean rims and c mean cores.

Grain	Element concentration (ppm)			Isotope ratios				Ages (Ma)			
	U	Th	Pb	$(^{207}\text{Pb}/^{206}\text{Pb})_{\text{meas}}$	σ (%)	$(^{206}\text{Pb}/^{238}\text{U})_{\text{meas}}$	σ (%)	$^{206}\text{Pb}/^{238}\text{U}$	$^{207}\text{Pb}/^{235}\text{U}$	$^{207}\text{Pb}/^{206}\text{Pb}$	$(^{206}\text{Pb}/^{238}\text{U})_{\text{cor}}$
<i>Laser-Ablation ICPMS data</i>											
LAbz10b	439	47	32	0.062	2.7	0.076	2.55	472	508	676	469 ± 12
LAbz11b	772	118	59	0.063	2.7	0.08	2.55	498	538	712	494 ± 12
LAbz13	570	123	43	0.058	2.7	0.081	2.55	501	505	522	500 ± 13
LAbz14	605	31	64	0.065	2.7	0.113	2.55	690	708	768	688 ± 17
LAbz15	451	68	34	0.06	2.7	0.079	2.55	491	513	611	489 ± 12
LAbz3	879	486	165	0.075	2.7	0.174	2.55	1034	1049	1080	-
LAbz3b	522	111	51	0.061	2.7	0.097	2.55	600	605	627	599 ± 15
LAbz4	593	92	214	0.165	2.7	0.343	2.55	1903	2207	2504	-
LAbz5	208	39	19	0.059	2.7	0.092	2.55	570	568	562	570 ± 14
LAbz8	454	28	33	0.062	2.7	0.077	2.55	476	510	665	473 ± 12
LAbz8b	522	41	39	0.06	2.7	0.08	2.55	495	512	587	493 ± 12
LAbz9	422	325	174	0.121	2.7	0.356	2.55	1961	1965	1970	-
LAc1	255	26	25	0.061	2.7	0.099	2.55	609	618	654	608 ± 15
LAc10	467	50	35	0.058	2.7	0.08	2.55	496	502	530	496 ± 12
LAc11	340	27	25	0.056	2.7	0.079	2.55	487	484	470	488 ± 12
LAc3	251	91	89	0.158	2.7	0.315	2.55	1766	2093	2433	-
LAc4	148	23	25	0.11	2.7	0.165	2.55	986	1273	1796	939 ± 23
LAc6	242	28	19	0.059	2.7	0.08	2.55	498	509	561	497 ± 12
LAc7	65	25	23	0.159	2.7	0.315	2.55	1766	2099	2443	-
LAA10	929	35	67	0.059	2.7	0.078	2.55	483	497	566	481 ± 12
LAA11	405	217	48	0.062	2.7	0.11	2.55	672	675	684	672 ± 17
LAA2	289	46	28	0.062	2.7	0.099	2.55	606	617	658	605 ± 15
LAA4	274	23	21	0.062	2.7	0.08	2.55	495	529	678	492 ± 12
LAA5	358	270	73	0.076	2.7	0.179	2.55	1060	1070	1089	-
LAA5b	343	42	25	0.059	2.7	0.078	2.55	483	499	573	481 ± 12
LAA6	169	41	16	0.068	2.7	0.094	2.55	581	641	859	575 ± 14
LAA8	442	128	118	0.142	2.7	0.246	2.55	1417	1789	2256	-
LAA9	195	60	19	0.059	2.7	0.099	2.55	609	601	569	610 ± 15
LAA9b	491	123	48	0.061	2.7	0.098	2.55	602	607	626	602 ± 15
<i>Ion-microprobe data</i>											
IMPz1b	n.d	n.d	n.d	0.0583	1.15	0.079	0.35	490	499	540	490 ± 2
IMPz1c	n.d	n.d	n.d	0.0827	1.95	0.1811	1.45	1073	1138	1263	-
IMPz5b	n.d	n.d	n.d	0.063	0.53	0.0796	0.41	494	534	707	490 ± 2
IMPz12b	n.d	n.d	n.d	0.0625	1.43	0.0799	1.4	495	532	691	492 ± 7
IMPz7b	n.d	n.d	n.d	0.0583	1.12	0.0792	1.63	491	500	541	491 ± 8

Table 4.6: Single-grain stepwise evaporation $^{207}\text{Pb}/^{206}\text{Pb}$ data of the lower metavolcanic rocks of the Sanabria region.

Grain	Step	$(^{204}\text{Pb}/^{206}\text{Pb})_{\text{meas}}$	$(^{207}\text{Pb}/^{206}\text{Pb})_{\text{meas}}$	$(^{207}\text{Pb}/^{206}\text{Pb})_{\text{cor}}$	2σ	Age (Ma)
z1	1	0.000025	0.106876	0.10679323	0.8	1745 ± 14
	2	0.000097	0.137302	0.13636582	0.4	2182 ± 6
	3	0.000086	0.142068	0.14129601	0.6	2243 ± 10
	4	0.000006	0.234751	0.23525394	0.5	3088 ± 8
	5	0.00021	0.249008	0.24734885	0.5	3168 ± 8
z2	1	0.000024	0.074321	0.07415859	2.5	1046 ± 49
	2	0.000063	0.136121	0.13562584	0.8	2172 ± 14
	3	0.000071	0.155188	0.15466218	0.7	2398 ± 12
	4	0.000061	0.159199	0.15881523	0.39	2443 ± 12
	5	0.000067	0.164715	0.16427656	0.28	2500 ± 5

Table 4.7: U-Pb laser ablation (LA) and ion microprobe (IM) results of the upper metavolcanic rocks of Sanabria region. Analyses references ending in b mean rims and c mean cores.

Grain	Element concentration (ppm)			Isotope ratios				Ages (Ma)			
	U	Th	Pb	$(^{207}\text{Pb}/^{206}\text{Pb})_{\text{meas}}$	σ (%)	$(^{206}\text{Pb}/^{238}\text{U})_{\text{meas}}$	σ (%)	$^{206}\text{Pb}/^{238}\text{U}$	$^{207}\text{Pb}/^{235}\text{U}$	$^{207}\text{Pb}/^{206}\text{Pb}$	$(^{206}\text{Pb}/^{238}\text{U})_{\text{cor}}$
<i>Laser-Ablation ICPMS data</i>											
LA020a	277	214	54	0.122	1.96	0.162	2.45	970	1336	1984	908 ± 22
LA022a	199	188	96	0.168	1.96	0.395	2.45	2148	2353	2536	-
LA026a	584	89	55	0.067	1.96	0.1	2.45	616	667	846	610 ± 15
LA038a	747	37	55	0.059	1.96	0.079	2.45	489	503	565	488 ± 12
LA042c	153	98	16	0.064	1.96	0.099	2.45	607	640	757	604 ± 14
LA042b	279	113	22	0.06	1.96	0.079	2.45	490	511	606	488 ± 12
LA056a	134	131	17	0.064	1.96	0.114	2.45	698	712	758	696 ± 17
LA058a	605	297	64	0.063	1.96	0.101	2.45	618	636	699	616 ± 15
LA060a	442	166	44	0.06	1.96	0.096	2.45	592	596	609	592 ± 14
LA066a	247	221	128	0.169	1.96	0.42	2.45	2260	2415	2548	-
LA068a	53	115	27	0.121	1.96	0.339	2.45	1881	1924	1971	-
LA082a	1557	244	118	0.06	1.96	0.078	2.45	487	510	616	485 ± 12
LA084a	557	100	51	0.061	1.96	0.096	2.45	591	603	651	590 ± 14
LA088a	309	95	25	0.062	1.96	0.079	2.45	493	526	674	490 ± 12
LA090a	240	180	30	0.069	1.96	0.103	2.45	633	694	897	627 ± 15
LA096a	52	59	7	0.064	1.96	0.104	2.45	635	657	732	633 ± 15
LA108a	293	146	25	0.059	1.96	0.08	2.45	495	511	585	494 ± 12
LA110a	150	254	20	0.061	1.96	0.098	2.45	600	609	644	599 ± 14
LA112a	58	60	7	0.059	1.96	0.097	2.45	599	592	566	599 ± 14
LA124a	68	62	22	0.148	1.96	0.251	2.45	1444	1840	2323	-
<i>Ion-microprobe data</i>											
IMPz1	n.d.	n.d.	n.d.	0.255	0.46	0.6186	0.9	3105	3173	3216	-
IMPz2c	n.d.	n.d.	n.d.	0.0591	0.95	0.0783	0.96	486	501	570	485 ± 5
IMPz3c	n.d.	n.d.	n.d.	0.0634	0.82	0.0796	0.63	494	536	721	490 ± 3
IMPz4c	n.d.	n.d.	n.d.	0.2009	0.28	0.4957	0.94	2595	2732	2834	-
IMPz7c	n.d.	n.d.	n.d.	0.0596	1.22	0.1027	0.51	631	622	590	631 ± 3
IMPz13	n.d.	n.d.	n.d.	0.0572	2.02	0.0785	0.92	487	489	500	487 ± 4
IMP13c	n.d.	n.d.	n.d.	0.1177	0.31	0.3532	0.75	1950	1936	1922	-
IMP18b	n.d.	n.d.	n.d.	0.0572	0.86	0.0785	0.64	487	489	499	487 ± 3
IMP18c	n.d.	n.d.	n.d.	0.0695	1	0.1032	0.46	633	698	913	626 ± 3
IMP19c	n.d.	n.d.	n.d.	0.231	1.51	0.5719	2.03	2916	3001	3059	-
IMP15c	n.d.	n.d.	n.d.	0.0627	0.49	0.0994	0.69	611	630	697	609 ± 4
IMPz6b	n.d.	n.d.	n.d.	0.0572	0.31	0.0787	0.6	488	490	500	488 ± 3

Table 4.8: Single-grain stepwise evaporation $^{207}\text{Pb}/^{206}\text{Pb}$ data of the upper metavolcanic rocks of the Sanabria region.

Grain	Step	$(^{204}\text{Pb}/^{206}\text{Pb})_{\text{meas}}$	$(^{207}\text{Pb}/^{206}\text{Pb})_{\text{meas}}$	$(^{207}\text{Pb}/^{206}\text{Pb})_{\text{cor}}$	2σ	Age (Ma)
z1	1	0.000006	0.095317	0.09546472	1.1	1537 ± 21
	2	0.000002	0.164924	0.16529885	1.1	2511 ± 18
	3	0.000003	0.18207	0.182475	0.3	2675 ± 5
	4	0.000001	0.185691	0.18612911	0.3	2705 ± 5
z3	1	0.000159	0.058964	0.05678859	0.8	483 ± 17
	2	0.000157	0.059065	0.05691939	0.4	488 ± 8
	3	0.000133	0.058891	0.05709528	0.5	495 ± 11
	4	0.000101	0.059957	0.05863314	0.37	553 ± 8
	5	0.000103	0.059991	0.05863814	0.27	554 ± 5
z1	1	0.001563	0.079676	0.05708818	1.5	495 ± 33
	2	0.0002864	0.0610698	0.05704238	0.4	493 ± 9
	3	0.00026789	0.06075226	0.05699445	0.3	491 ± 6
	4	0.0001	0.058269	0.05695249	0.24	490 ± 5
	5	0.000112	0.058436	0.05694494	0.28	489 ± 6
z4	1	0.000143	0.058881	0.05693909	0.6	489 ± 13
z3	1	0.000045	0.058083	0.05756832	0.7	513 ± 15
	2	0.000109	0.059846	0.05840504	0.8	545 ± 17
	3	0.000071	0.060443	0.05955791	0.7	587 ± 15
	4	0.000038	0.060492	0.060008688	0.7	607 ± 15
	5	0.000106	0.061433	0.06004269	0.5	605 ± 11
z4	1	0.000021	0.057034	0.0568661	0.8	486 ± 17
	2	0.000171	0.059417	0.05706862	0.6	494 ± 13
	3	0.000015	0.057281	0.05720126	0.5	499 ± 11
	4	0.00014	0.059611	0.05771662	0.6	519 ± 13
z5	1	0.000041	0.057264	0.05680502	1	484 ± 22
	2	0.000151	0.060953	0.05890497	0.5	563 ± 10
	3	0.00006	0.059835	0.05910778	0.7	571 ± 15
	4	0.000367	0.0604401	0.05922245	0.5	575 ± 11

Table 4.9: U-Pb laser ablation (LA) and ion microprobe (IM) results of the Villadepera metavolcanic rocks. Analyses references ending in b mean rims and c mean cores.

Grain	Element concentration (ppm)			Isotope ratios				Ages (Ma)			
	U	Th	Pb	$(^{207}\text{Pb}/^{206}\text{Pb})_{\text{meas}}$	σ (%)	$(^{206}\text{Pb}/^{238}\text{U})_{\text{meas}}$	σ (%)	$^{206}\text{Pb}/^{238}\text{U}$	$^{207}\text{Pb}/^{235}\text{U}$	$^{207}\text{Pb}/^{206}\text{Pb}$	$(^{206}\text{Pb}/^{238}\text{U})_{\text{cor}}$
<i>Laser-Ablation ICPMS data</i>											
z2	165	78	13	0.054	2.38	0.078	1.97	484	465	373	486 ± 10
z3	60	75	27	0.122	2.38	0.351	1.97	1939	1963	1989	-
z5	703	52	55	0.082	2.38	0.077	1.97	480	639	1249	465 ± 10
z6	127	46	10	0.058	2.38	0.078	1.97	482	489	523	481 ± 10
z7	440	174	58	0.064	2.38	0.129	1.97	783	776	757	784 ± 16
z8	163	81	18	0.061	2.38	0.104	1.97	639	639	638	639 ± 13
z9	93	61	38	0.159	2.38	0.357	1.97	1969	2214	2450	-
z10	195	103	21	0.061	2.38	0.103	1.97	632	636	650	631 ± 13
z11	233	76	30	0.089	2.38	0.115	1.97	705	893	1394	682 ± 15
z12	1279	351	169	0.08	2.38	0.126	1.97	768	889	1204	753 ± 16
z15	164	73	13	0.056	2.38	0.076	1.97	473	468	443	474 ± 10
z16	139	69	33	0.142	2.38	0.203	1.97	1194	1631	2251	-
z17	258	83	20	0.055	2.38	0.077	1.97	479	466	406	480 ± 10
z18	439	43	44	0.061	2.38	0.105	1.97	645	642	634	645 ± 13
z19	124	97	78	0.207	2.38	0.505	1.97	2636	2779	2885	-
<i>Ion-microprobe data</i>											
z1a	239	24	20	0.0562	1.21	0.0786	1.34	488	483	463	488 ± 7
z1b	279	38	24	0.0561	1.01	0.0778	1.34	483	479	456	484 ± 7
z2a	299	16	24	0.0569	0.96	0.0763	1.35	474	477	489	474 ± 7
z2b	123	97	14	0.0594	1.41	0.0883	1.39	545	553	583	545 ± 8
z3	527	375	51	0.1208	1.06	0.0839	1.34	519	888	1968	478 ± 7
z4a	291	29	24	0.0565	1	0.0771	1.34	479	478	472	479 ± 7
z4c	314	28	27	0.0586	0.94	0.0781	1.36	485	497	551	484 ± 7
z4d	349	21	30	0.0571	0.91	0.0786	1.34	488	489	494	488 ± 7
z5	92	23	11	0.0625	2	0.1001	1.34	615	632	692	613 ± 9
z6a	543	78	46	0.057	0.72	0.0776	1.35	482	484	493	482 ± 7
z6b	369	88	32	0.0571	1	0.0778	1.36	483	485	496	483 ± 7
z7	239	161	103	0.1194	0.46	0.3243	1.34	1811	1875	1948	-
z8c	156	65	17	0.0597	1.46	0.0928	1.34	572	576	591	572 ± 8
z8b	674	44	48	0.0558	0.73	0.0661	1.34	413	418	445	412 ± 6

Table 4.10: U-Pb laser ablation (LA) and ion microprobe (IM) results of the lower metavolcanic rocks of the Hiendelaencina region. Analyses references ending in b mean rims and c mean cores.

Sample	Grain	Element concentration (ppm)			Isotope ratios				Ages (Ma)			
		U	Th	Pb	$\frac{^{207}\text{Pb}}{^{206}\text{Pb}}_{\text{meas}}$	σ (%)	$\frac{^{206}\text{Pb}}{^{238}\text{U}}_{\text{meas}}$	σ (%)	$\frac{^{206}\text{Pb}}{^{238}\text{U}}$	$\frac{^{207}\text{Pb}}{^{235}\text{U}}$	$\frac{^{207}\text{Pb}}{^{206}\text{Pb}}$	$\frac{^{206}\text{Pb}}{^{238}\text{U}}_{\text{cor}}$
<i>Laser-Ablation ICPMS data</i>												
T-21	a1	430	258	44	0.063	2.38	0.093	1.97	575	603	712	572 ± 12
	a2	495	148	41	0.058	2.38	0.083	1.97	516	522	544	516 ± 11
	a3b	346	58	29	0.059	2.38	0.087	1.97	539	544.6	569	538 ± 11
	a6b	476	118	54	0.086	2.38	0.106	1.97	651	826	1333	631 ± 13
	a6c	68	72	21	0.114	2.38	0.236	1.97	1366	1576	1870	1317 ± 29
	a9	495	152	50	0.06	2.38	0.101	1.97	618	616	607	618 ± 13
	bz1	274	74	41	0.073	2.38	0.147	1.97	883	922	1015	878 ± 18
	bz10	1291	377	138	0.063	2.38	0.106	1.97	651	667	720	649 ± 13
	bz3	362	125	48	0.069	2.38	0.129	1.97	780	810	895	776 ± 16
	bz4	643	292	55	0.064	2.38	0.081	1.97	503	549	745	499 ± 10
	bz5	495	269	52	0.067	2.38	0.096	1.97	590	645	841	585 ± 12
	bz7b	513	42	37	0.059	2.38	0.076	1.97	472	486	552	471 ± 10
	bz9	63	47	9	0.069	2.38	0.131	1.97	792	819	891	789 ± 16
c1	145	31	18	0.064	2.38	0.125	1.97	759	750.6	727	759 ± 16	
c4	363	309	43	0.065	2.38	0.103	1.97	630	664	780	623 ± 13	
T-22	a11	312	138	39	0.067	2.38	0.12	1.97	731	760	847	728 ± 15
	a12	208	55	16	0.059	2.38	0.079	1.97	489	500	549	488 ± 10
	a15	608	663	74	0.062	2.38	0.102	1.97	625	637	679	624 ± 13
	a16	401	121	39	0.061	2.38	0.098	1.97	604	611	635	604 ± 12
	a17	202	69	20	0.06	2.38	0.097	1.97	596	594	589	596 ± 12
	a20	427	123	134	0.117	2.38	0.304	1.97	1710	1803	1913	1681 ± 38
	a21	700	724	67	0.06	2.38	0.08	1.97	497	514	591	495 ± 10
	a5	530	934	71	0.061	2.38	0.099	1.97	606	610	627	605 ± 12
	a7	96	71	12	0.069	2.38	0.115	1.97	700	746	889	695 ± 14
	a8	174	132	49	0.114	2.38	0.245	1.97	1412	1606	1870	1366 ± 30
	a9	422	207	52	0.069	2.38	0.118	1.97	717	765	905	712 ± 15
	bz10	66	68	10	0.193	2.38	0.097	1.97	594	1294	2772	494 ± 13
	bz11	458	102	37	0.06	2.38	0.082	1.97	508	527	614	506 ± 10
	bz12	425	252	46	0.061	2.38	0.102	1.97	626	630	647	625 ± 13
	bz2	257	78	25	0.065	2.38	0.1	1.97	615	650	771	612 ± 13
bz6	404	50	36	0.06	2.38	0.094	1.97	576	582	607	576 ± 12	
bz8	682	1113	74	0.065	2.38	0.081	1.97	500	553	779	495 ± 10	
bz9	377	241	38	0.061	2.38	0.095	1.97	584	599	654	583 ± 12	
<i>Ion-microprobe data</i>												

T-22	z26a	201	119	82	0.1199	1.59	0.3158	1.38	1769	1856	1954	-
	z26b	238	150	107	0.1234	0.6	0.3455	1.34	1913	1958	2007	-
	z28c	381	277	71	0.068	0.58	0.1456	1.37	876	874	869	876 ± 12
	z28b	469	286	83	0.0676	0.53	0.1407	1.34	849	851	858	848 ± 11
	z31b	410	36	36	0.0575	0.91	0.0813	1.35	504	505	510	504 ± 7
	z31c	418	250	50	0.0675	0.79	0.0939	1.36	579	638	855	573 ± 8
	z32a	277	46	25	0.0573	1.01	0.0801	1.34	497	498	502	497 ± 7
	z32b	295	395	42	0.0607	0.87	0.0979	1.34	602	608	629	602 ± 8
	z32c	635	72	58	0.0575	0.8	0.0839	1.34	519	518	510	520 ± 7
	z33	200	63	26	0.0629	1.35	0.1136	1.34	694	696	704	693 ± 10
	z35	538	41	68	0.063	0.68	0.1154	1.36	704	705	709	704 ± 10

Table 4.11: Single-grain stepwise evaporation $^{207}\text{Pb}/^{206}\text{Pb}$ data of the lower metavolcanic rocks of the Hiendelaencina region.

Grain	Step	$(^{204}\text{Pb}/^{206}\text{Pb})_{\text{meas}}$	$(^{207}\text{Pb}/^{206}\text{Pb})_{\text{meas}}$	$(^{207}\text{Pb}/^{206}\text{Pb})_{\text{cor}}$	2σ	Age (Ma)
z2	1	0.000061	0.057628	0.0568784	1.2	487 ± 26
	2	0.00005	0.0577351	0.0571456	0.5	497 ± 11
	3	0.000048	0.057726	0.071664	0.6	498 ± 13
	4	0.000074	0.059441	0.0585085	0.4	549 ± 8
	5	0.000058	0.05906	0.0583592	0.5	543 ± 11
z3	1	0.00002	0.057176	0.0570231	0.5	492 ± 11
	2	0.000017	0.058337	0.058231	0.18	539 ± 4
	3	0.000016	0.058252	0.0581603	0.24	536 ± 6
	4	0.000027	0.059315	0.0590663	0.3	569 ± 6
	5	0.000032	0.059181	0.0588591	0.29	562 ± 7
	6	0.000031	0.059129	0.0588215	0.35	560 ± 7
	7	0.000033	0.060362	0.0600291	0.2	605 ± 5
z4	1	0.000018	0.057281	0.0571575	0.18	498 ± 4
	2	0.000016	0.057235	0.0571406	0.3	497 ± 7
	3	0.000015	0.056952	0.0568714	0.3	487 ± 7
	4	0.000018	0.057216	0.0570923	0.3	495 ± 7

Table 4.12: U-Pb laser ablation (LA) and ion microprobe (IM) results of the upper metavolcanic rocks of the Hiendelaencina region. Analyses references ending in b mean rims and c mean cores.

Grain	Element concentration (ppm)			Isotope ratios				Ages (Ma)			
	U	Th	Pb	$(^{207}\text{Pb}/^{206}\text{Pb})_{\text{meas}}$	σ (%)	$(^{206}\text{Pb}/^{238}\text{U})_{\text{meas}}$	σ (%)	$^{206}\text{Pb}/^{238}\text{U}$	$^{207}\text{Pb}/^{235}\text{U}$	$^{207}\text{Pb}/^{206}\text{Pb}$	$(^{206}\text{Pb}/^{238}\text{U})_{\text{cor}}$
<i>Laser-Ablation ICPMS data</i>											
a1	49	29	8	0.072	1.7	0.144	2.5	868	902	987	863 ± 22
a2	224	78	17	0.059	1.7	0.077	2.5	479	497	577	478 ± 12
a4	426	100	37	0.086	1.7	0.084	2.5	517	699	1340	499 ± 13
a5	583	52	42	0.06	1.7	0.076	2.5	472	496	608	470 ± 12
bz1	166	126	13	0.072	1.7	0.064	2.5	397	498	991	388.7
bz2	192	80	19	0.061	1.7	0.099	2.5	609	617	649	608 ± 15
bz5	205	117	34	0.072	1.7	0.155	2.5	931	947	986	929 ± 23
bz6	449	153	44	0.06	1.7	0.097	2.5	595	599	614	594 ± 15
bz7	509	233	48	0.06	1.7	0.091	2.5	563	570	597	562 ± 14
c2	130	74	20	0.067	1.7	0.14	2.5	846	841	827	846 ± 21
c3	308	378	39	0.063	1.7	0.1	2.5	616	640	724	614 ± 15
c4	302	35	24	0.066	1.7	0.082	2.5	509	568	812	504 ± 13
b8	268	40	22	0.064	1.7	0.085	2.5	526	565	728	522 ± 13
<i>Ion-microprobe data</i>											
z21	2707	987	220	0.0797	0.67	0.0768	1.34	477	621	1189	463 ± 6
z22	485	40	39	0.0647	0.95	0.0753	1.34	468	522	763	463 ± 6
z23	354	131	41	0.0613	0.8	0.1004	1.34	617	624	651	616 ± 8
z24	201	21	22	0.0617	1.1	0.0975	1.68	600	613	663	599 ± 10
z26b	275	7	23	0.0564	1.13	0.0795	1.35	493	488	467	493 ± 7
z27	196	38	17	0.0561	1.44	0.0785	1.34	488	482	457	488 ± 7
z28	335	30	29	0.0581	1.13	0.0795	1.34	493	500	533	493 ± 7
z29a	258	32	23	0.0567	1.4	0.0797	1.34	495	492	482	495 ± 7
z30b2	1184	43	100	0.0626	0.6	0.0796	1.34	494	531	694	491 ± 7
z30c	300	148	34	0.0608	0.92	0.094	1.34	579	590	633	578 ± 8
z31	367	53	31	0.0568	1.15	0.0774	1.34	481	481	485	481 ± 7
z32	510	71	44	0.0574	0.87	0.078	1.34	484	488	507	484 ± 7
z33b	375	44	34	0.0555	0.86	0.0829	1.34	513	499	432	515 ± 7
z33c	349	6	29	0.057	0.95	0.0776	1.34	482	484	492	482 ± 7

Table 4.13: Single-grain stepwise evaporation $^{207}\text{Pb}/^{206}\text{Pb}$ data of the upper metavolcanic rocks of the Hiendelaencina region.

Grain	Step	$(^{204}\text{Pb}/^{206}\text{Pb})_{\text{meas}}$	$(^{207}\text{Pb}/^{206}\text{Pb})_{\text{meas}}$	$(^{207}\text{Pb}/^{206}\text{Pb})_{\text{cor}}$	2σ	Age (Ma)
z1	1	0.000095	0.057097	0.055849	0.16	446 ± 4
	2	0.000096	0.056903	0.055639	0.14	438 ± 3
	3	0.000131	0.06254	0.060791	0.19	632 ± 9
	4	0.000055	0.087353	0.086792	0.6	1356 ± 12
	5	0.000031	0.123167	0.123052	0.3	2001 ± 5
z2	2	0.000046	0.057234	0.056702	0.2	480 ± 4
	3	0.000044	0.057274	0.056771	0.4	483 ± 8
	4	0.000066	0.057731	0.056909	0.4	488 ± 9
	5	0.000073	0.057722	0.056798	0.3	484 ± 6
z3	1	0.000047	0.056651	0.056102	0.4	456 ± 9
	2	0.000084	0.059604	0.058526	0.2	549 ± 5
	3	0.000057	0.058511	0.057823	0.2	523 ± 4
z4	1	0.000218	0.058588	0.055546	0.2	434 ± 5
	2	0.000118	0.058199	0.056619	0.5	477 ± 11
	3	0.000042	0.057149	0.056675	0.4	479 ± 9
	4	0.000071	0.057847	0.056952	0.2	490 ± 5
z1	1	0.000041	0.057459	0.057001	0.9	492 ± 19
	2	0.00003	0.060404	0.060115	0.14	608 ± 3
	3	0.000028	0.065274	0.065028	0.6	775 ± 13
	4	0.000073	0.074735	0.073874	0.14	1038 ± 3
z6	1	0.000211	0.058848	0.055911	0.4	449 ± 9

Table 4.14: Rb-Sr data of a sliced K-feldspar megacryst from the upper metavolcanic rocks of the Hiendelaencina region.

Sample	$^{87}\text{Rb}/^{86}\text{Sr}$	error	$^{87}\text{Sr}/^{86}\text{Sr}$	error
Fto-8	1.66714	1.2	0.72401	0.002
Fto-3	1.714108	1.2	0.724281	0.002
Fto-7	1.768224	1.2	0.724439	0.002
Fto-9	1.86337	1.2	0.725494	0.002
Fto-10	2.208603	1.2	0.727668	0.002

Table 4.15: U-Pb laser ablation (LA) results of the Viana do Bolo metagranite.

Grain	Element concentration (ppm)			Isotope ratios				Ages (Ma)			
	U	Th	Pb	$(^{207}\text{Pb}/^{206}\text{Pb})_{\text{meas}}$	σ (%)	$(^{206}\text{Pb}/^{238}\text{U})_{\text{meas}}$	σ (%)	$^{206}\text{Pb}/^{238}\text{U}$	$^{207}\text{Pb}/^{235}\text{U}$	$^{207}\text{Pb}/^{206}\text{Pb}$	$(^{206}\text{Pb}/^{238}\text{U})_{\text{cor}}$
LA010a	295	51	60	0.13006	2.38	0.19283	1.97	1137	1518	2099	-
LA018a	396	90	39	0.06305	2.38	0.09878	1.97	607	629	710	605 ± 13
LA024a	443	190	36	0.05948	2.38	0.077	1.97	478	497	585	477 ± 10
LA026a	465	449	55	0.06501	2.38	0.10131	1.97	622	656	775	619 ± 13
LA032a	669	43	49	0.05816	2.38	0.07769	1.97	482	492	536	482 ± 10
LA034a	958	51	72	0.05952	2.38	0.07848	1.97	487	505	586	485 ± 10
LA038a	580	130	45	0.0586	2.38	0.07784	1.97	483	495	552	482 ± 10
LA042a	665	38	48	0.05975	2.38	0.0767	1.97	476	497	595	475 ± 10
LA048a	586	92	69	0.09227	2.38	0.11814	2.31	720	932	1473	694 ± 17
LA050a	346	63	27	0.06096	2.38	0.08019	2.31	497	523	638	495 ± 12
LA066a	1208	70	90	0.06338	2.38	0.07735	2.31	480	524	721	476 ± 11
LA070a	800	80	59	0.06682	2.38	0.07578	2.31	471	538	832	465 ± 11
LA072a	1105	109	82	0.06017	2.38	0.07785	2.31	483	506	610	481 ± 11
LA074a	1401	66	111	0.05951	2.38	0.08364	2.31	518	531	586	517 ± 12
LA082a	467	261	55	0.06492	2.38	0.09891	2.31	608	644	772	604 ± 14
LA086a	142	51	49	0.1538	2.38	0.29704	2.31	1677	2018	2389	-
LA088a	834	102	82	0.07225	2.38	0.09857	2.31	606	695	993	597 ± 14

Table 4.16: Single-grain stepwise evaporation $^{207}\text{Pb}/^{206}\text{Pb}$ data of the Viana do Bolo metagranite.

Grain	Step	$(^{204}\text{Pb}/^{206}\text{Pb})_{\text{meas}}$	$(^{207}\text{Pb}/^{206}\text{Pb})_{\text{meas}}$	$(^{207}\text{Pb}/^{206}\text{Pb})_{\text{cor}}$	2σ	Age (Ma)
z10	1	0.00099	0.071097	0.056835	0.29	485 ± 6
	2	0.0001	0.059067	0.05775391	0.17	520 ± 3
z11	1	0.00009	0.0582815	0.05711106	0.29	496 ± 6
	2	0.000099	0.058908	0.05760882	0.4	515 ± 9
	3	0.000117	0.06058	0.05902574	0.3	586 ± 6
	4	0.000022	0.059983	0.05980885	0.4	597 ± 8
	5	0.000014	0.06011	0.06005242	0.8	605 ± 17
	6	0.000011	0.060304	0.06029051	0.6	614 ± 13
z12	1	0.000144	0.06226	0.06032564	0.2	615 ± 4
	2	0.000093	0.06355	0.06235744	0.4	686 ± 8
	3	0.000106	0.064836	0.06346038	0.4	724 ± 8
	4	0.000102	0.065683	0.0643689	0.6	754 ± 12
z14	1	0.000804	0.06732	0.05574436	0.17	442 ± 4
	2	0.000818	0.068705	0.0569469	0.18	489 ± 4
	3	0.000807	0.068533	0.05693467	0.14	489 ± 3
	4	0.00081	0.068596	0.05695437	0.18	490 ± 4
	5	0.000496	0.06518	0.0581182	0.2	534 ± 4
	6	0.000021	0.059115	0.05895295	0.19	565 ± 4
	7	0.0000349	0.0603465	0.0599859	0.47	603 ± 10
z15	1	0.0002877	0.0609035	0.05685662	0.5	486 ± 11
	2	0.000287	0.062176	0.0581493	0.35	535 ± 7
	3	0.0002832	0.0622224	0.05825162	0.4	539 ± 9
	4	0.000296	0.062586	0.05843082	0.43	546 ± 9
z2	1	0.000427	0.06607	0.06002957	0.6	605 ± 12
	2	0.00015513	0.062120658	0.06001846	1.3	604 ± 19
z3	1	0.000281	0.060927	0.0569785	0.13	491 ± 3
	2	0.000316	0.061506	0.05704912	0.22	493 ± 4
	3	0.00031	0.061427	0.05705743	0.14	494 ± 3
	4	0.000288915	0.061248872	0.05718697	0.16	499 ± 4
	5	0.000384	0.064168	0.05873839	0.21	557 ± 4
z4	1	0.000082	0.05892	0.05786886	0.22	525 ± 5
	2	0.000075	0.058991	0.05804219	0.24	531 ± 6
	3	0.000073	0.059085	0.0581657	0.3	536 ± 7
	4	0.00003	0.060276	0.05998646	0.7	603 ± 15
	5	0.00006	0.062046	0.06132653	0.37	651 ± 8
	6	0.000048	0.062518	0.06197421	0.7	673 ± 15
	7	0.00007	0.062877	0.06201542	0.7	675 ± 15
	8	0.000024	0.062416	0.06221974	0.5	682 ± 11
z5	1	0.000357	0.061728	0.05667152	0.42	479 ± 9
	2	0.000295	0.061555	0.0574063	0.36	507 ± 8
	3	0.000248	0.061435	0.05797336	0.28	529 ± 6
	4	0.000325	0.064432	0.05986783	0.58	599 ± 10
	5	0.000188	0.062659	0.06008108	1.2	606 ± 26
z6	1	0.000323	0.061188	0.05662576	0.38	477 ± 8
	2	0.0004	0.062468	0.05678711	0.7	483 ± 15
	3	0.000518	0.064565	0.0571724	1	498 ± 22
	4	0.000438	0.064052	0.05782938	0.4	523 ± 8
	5	0.000778	0.069402	0.05824747	0.51	539 ± 11
	6	0.000515	0.065597	0.05826099	0.6	539 ± 13
z7	1	0.000459	0.061779	0.05522252	0.25	421 ± 5
	2	0.000055506	0.058913424	0.05824843	0.6	539 ± 13

z8	1	0.000361159	0.060940015	0.05581572	0.9	445 ± 20
	2	0.000169	0.059037	0.05671573	0.6	480 ± 13
	3	0.000105	0.059027	0.05764078	0.5	516 ± 11
	4	0.000167448	0.060360697	0.05806949	0.6	532 ± 13
	5	0.000164	0.06047866	0.05823844	0.5	539 ± 11
z9	1	0.00012	0.058605	0.05699785	1.3	491 ± 28

Table 4.17: U-Pb laser ablation (LA) and ion microprobe (IM) results of the San Sebastián metagranite. Analyses references ending in b mean rims and c mean cores.

Grain	Element concentration (ppm)			Isotope ratios				Ages (Ma)			
	U	Th	Pb	$(^{207}\text{Pb}/^{206}\text{Pb})_{\text{meas}}$	σ (%)	$(^{206}\text{Pb}/^{238}\text{U})_{\text{meas}}$	σ (%)	$^{206}\text{Pb}/^{238}\text{U}$	$^{207}\text{Pb}/^{235}\text{U}$	$^{207}\text{Pb}/^{206}\text{Pb}$	$(^{206}\text{Pb}/^{238}\text{U})_{\text{cor}}$
<i>Laser-Ablation ICPMS data</i>											
z25b	1495	74	124	0.0567	0.46	0.0766	0.85	476	476	478	476 ± 4
z26m	171	206	18	0.0577	1.35	0.0745	0.88	463	473	519	462 ± 4
z30b	724	53	61	0.0575	1.03	0.0765	0.83	475	481	510	474 ± 4
z30c	171	38	15	0.0576	1.53	0.0763	0.83	474	481	514	474 ± 4
z30c2	36	86	5	0.0579	3.93	0.0753	1.01	468	478	527	467 ± 6
z36b	796	79	66	0.0572	0.71	0.0755	0.83	469	474	498	469 ± 4
<i>SHRIMP data</i>											
z4.1	636	59	44	0.0562	1.2	0.0749	2.2	465	464	459	465 ± 10
z5.1	326	122	23	0.056	1.7	0.076	2.2	472	469	454	473 ± 10
z7.1	1612	190	115	0.0573	0.7	0.076	2.2	472	477	501	472 ± 10
z2.2	55	97	6	0.0572	4.1	0.0749	1.6	466	471	498	466 ± 9

Table 4.18: Single-grain stepwise evaporation $^{207}\text{Pb}/^{206}\text{Pb}$ data of the San Sebastián metagranite.

Grain	Step	$(^{204}\text{Pb}/^{206}\text{Pb})_{\text{meas}}$	$(^{207}\text{Pb}/^{206}\text{Pb})_{\text{meas}}$	$(^{207}\text{Pb}/^{206}\text{Pb})_{\text{cor}}$	2σ	Age (Ma)
z1	1	0.000072	0.057101	0.0561888	0.5	460 ± 11
z4	1	0.000043	0.068459	0.0680007	0.6	869 ± 13
	2	0.000032	0.078264	0.0779999	0.33	1147 ± 7
	3	0.00003	0.060333	0.0901323	0.2	1428 ± 4
	4	0.000057	0.103189	0.1026552	0.5	1673 ± 9
	5	0.118152	0.000061	0.0117617	0.21	1920 ± 4
	6	0.000028	0.143633	0.143619	0.19	2271 ± 3
	7	0.000027	0.148659	0.148673	0.2	2331 ± 4
	8	0.000031	0.147794	0.147754	0.2	2320 ± 4
z3	1	0.000062	0.056942	0.0561607	0.4	459 ± 9
	2	0.000083	0.057348	0.0563122	0.4	465 ± 9
z9	1	0.000021	0.055909	0.0557379	0.6	442 ± 13
	2	0.000015	0.056753	0.0566718	0.8	479 ± 17

Table 4.19: U-Pb laser ablation (LA) and ion microprobe (IM) results of the Miranda do Douro metagranite. Analyses references ending in b mean rims and c mean cores.

Grain	Element concentration (ppm)			Isotope ratios				Ages (Ma)			
	U	Th	Pb	$(^{207}\text{Pb}/^{206}\text{Pb})_{\text{meas}}$	σ (%)	$(^{206}\text{Pb}/^{238}\text{U})_{\text{meas}}$	σ (%)	$^{206}\text{Pb}/^{238}\text{U}$	$^{207}\text{Pb}/^{235}\text{U}$	$^{207}\text{Pb}/^{206}\text{Pb}$	$(^{206}\text{Pb}/^{238}\text{U})_{\text{cor}}$
<i>Laser-Ablation ICPMS data</i>											
z20	455	73	34	0.066	2.38	0.077	1.97	476	537	804	471 ± 10
z22	2439	889	144	0.067	2.38	0.057	1.97	360	433	847	353 ± 7
z24	362	123	34	0.059	2.38	0.092	1.97	568	570	578	568 ± 12
z25	425	140	55	0.066	2.38	0.127	1.97	772	784	819	770 ± 16
z26	624	200	56	0.059	2.38	0.089	1.97	549	551	559	549 ± 11
z27	309	223	20	0.061	2.38	0.058	1.97	366	405	631	363 ± 8
z30c	577	100	132	0.174	2.38	0.198	1.97	1166	1778	2598	-
z30b	603	47	47	0.057	2.38	0.083	1.97	512	507	482	513 ± 11
z31	409	176	24	0.065	2.38	0.056	1.97	350	413	785	344 ± 7
z33	233	147	16	0.085	2.38	0.059	1.97	367	533	1323	353 ± 8
z35	388	55	31	0.057	2.38	0.083	1.97	513	508	489	513 ± 10
z38c	575	471	36	0.064	2.38	0.054	1.97	338	393	729	334 ± 7
z38b	630	397	35	0.042	2.38	0.053	1.97	331	270	-99	336 ± 7
<i>Ion-microprobe data</i>											
z-9c	1447	68	95	0.0774	0.48	0.0625	1.34	391	519	1133	380 ± 5
z-9b	332	41	23	0.0754	0.86	0.0662	1.34	413	532	1080	403 ± 6
z-10	293	23	25	0.0577	0.99	0.0785	1.34	487	492	517	487 ± 7
z-10b	351	34	29	0.0566	1.13	0.0775	1.34	481	480	476	481 ± 7
z-11	257	26	21	0.0562	1.06	0.0771	1.34	479	476	462	479 ± 7
z-12	346	28	29	0.0567	0.9	0.0778	1.35	483	482	478	483 ± 7
z-13b	1529	44	89	0.0931	0.53	0.0579	1.34	363	564	1489	345 ± 5
z-13c	569	139	61	0.0699	0.59	0.0952	1.34	586	661	924	579 ± 8
z-14	339	71	29	0.0574	0.92	0.0769	1.34	478	483	508	477 ± 7
z-15b	636	75	39	0.0686	0.7	0.0576	1.34	361	442	887	354 ± 5
z-15c	830	204	92	0.06	0.73	0.0978	1.34	601	602	605	601 ± 8
z-16	580	219	70	0.0608	0.68	0.1026	1.35	630	630	632	630 ± 9
z-17	396	42	30	0.0637	0.83	0.0709	1.34	442	492	732	437 ± 6
z-19c	516	77	398	0.2445	0.26	0.5984	1.34	3024	3100	3149	-
z-19b	128	53	94	0.2425	0.74	0.5444	1.34	2802	3000	3137	-

Table 4.20: Single-grain stepwise evaporation $^{207}\text{Pb}/^{206}\text{Pb}$ data of the Antoñita metagranite.

Grain	Step	$(^{204}\text{Pb}/^{206}\text{Pb})_{\text{meas}}$	$(^{207}\text{Pb}/^{206}\text{Pb})_{\text{meas}}$	$(^{207}\text{Pb}/^{206}\text{Pb})_{\text{cor}}$	2σ	Age (Ma)
z1	1	0.000068	0.060095	0.084187	1	579 ± 22
z10	1	0.000012	0.056627	0.0565893	1.2	476 ± 26
	2	0.000075	0.05834	0.0573887	0.14	507 ± 4
	3	0.000075	0.058436	0.0585766	0.32	551 ± 7
	4	0.000071	0.060229	0.0593431	0.5	580 ± 11
	5	0.000062	0.060336	0.0595814	0.28	588 ± 6
	6	0.000049	0.060391	0.0598257	0.4	597 ± 9
z11	1	0.000302	0.058943	0.0546703	0.21	399 ± 5
	2	0.000294	0.059071	0.054917	0.21	409 ± 5
	3	0.000072	0.05723	0.0563183	0.17	465 ± 4
	4	0.000075	0.057273	0.0560318	0.26	465 ± 6
	5	0.000074	0.057373	0.0564326	0.28	469 ± 7
	6	0.000067	0.058651	0.0578175	0.47	523 ± 10
	7	0.000078	0.059429	0.0584381	0.9	546 ± 20
z12	1	0.000166	0.06091	0.0586429	0.5	554 ± 11
z14	1	0.000115	0.0566	0.0550566	0.48	415 ± 10
	2	0.000053	0.057141	0.0565065	0.7	472 ± 15
	3	0.00005	0.05905	0.0584657	0.5	547 ± 11
z15	1	0.00047	0.061913	0.0551959	0.6	420 ± 13
	2	0.000647	0.062187	0.0555173	0.29	433 ± 6
	3	0.000131	0.058439	0.05667	0.21	479 ± 5
	4	0.00013	0.058545	0.0567914	0.18	483 ± 4
	5	0.000117	0.059865	0.0583075	0.16	541 ± 3
z2	1	0.000038	0.056395	0.0559771	1.5	451 ± 33
	2	0.00018	0.059013	0.0565307	0.6	473 ± 13
	3	0.000089	0.05895	0.0577969	0.7	522 ± 15
	4	0.000081	0.061906	0.0608819	1.4	635 ± 30
z4	1	0.000356	0.060216	0.0551602	1.3	419 ± 28
	2	0.000082	0.057463	0.0564061	0.4	468 ± 8
	3	0.000087	0.057508	0.0563782	0.4	467 ± 9
	4	0.000037	0.056707	0.0563046	0.38	463 ± 8
	5	0.000045	0.057042	0.0569524	0.31	473 ± 7
z5	1	0.000047	0.05574	0.0551884	0.4	420 ± 10
	2	0.000017	0.056737	0.0566266	0.15	477 ± 3
	3	0.000018	0.056935	0.0568106	0.26	484 ± 6
	4	0.000019	0.057071	0.0567427	0.6	482 ± 13
z6	1	0.0000332	0.0567942	0.0564469	0.3	470 ± 7
	2	0.000044	0.0567365	0.056232	0.4	462 ± 9
	3	0.000036	0.086829	0.0865335	0.52	1350 ± 10
z7	1	0.000024	0.057011	0.0567993	1.3	484 ± 29
	2	0.000049	0.057606	0.0570314	1.5	493 ± 32
	3	0.000041	0.060191	0.0597413	1.4	594 ± 30
z9	1	0.000086	0.10472	0.1037929	0.6	1693 ± 11
	2	0.000018	0.15779	0.1579451	0.24	2434 ± 4
	3	0.000016	0.180637	0.1808804	0.16	2661 ± 3
	4	0.000024	0.18337	0.1835241	0.4	2685 ± 7

Table 4.21: Rb-Sr data of the Antofñita metagranite.

Sample	$^{87}\text{Rb}/^{86}\text{Sr}$	error	$^{87}\text{Sr}/^{86}\text{Sr}$	error
CT-9	4.839	1.2	0.743294	0.002
CT-10	5.45	1.2	0.74794	0.002
CT-11	6.15	1.2	0.753283	0.002
CT-4	22.086	1.2	0.860977	0.002

9.1.2. Tables of geochronological data of the Schist-Graywacke Complex Domain

Table 4.22: U-Pb ion microprobe results of the Fermoselle metagranite. Analyses references ending in b mean rims and c mean cores.

Grain	Element concentration (ppm)			Isotope ratios				Ages (Ma)			
	U	Th	Pb	$(^{207}\text{Pb}/^{206}\text{Pb})_{\text{meas}}$	σ (%)	$(^{206}\text{Pb}/^{238}\text{U})_{\text{meas}}$	σ (%)	$^{206}\text{Pb}/^{238}\text{U}$	$^{207}\text{Pb}/^{235}\text{U}$	$^{207}\text{Pb}/^{206}\text{Pb}$	$(^{206}\text{Pb}/^{238}\text{U})_{\text{cor}}$
z17b	298	26	25	0.056	1.15	0.0765	3.3	475	472	453	476 ± 15
z22b	410	139	70	0.0691	1.24	0.1438	2.02	866	876	902	865 ± 18
z22c	187	77	36	0.0715	1.42	0.1629	2.55	973	972	971	973 ± 25
z23c	78	27	8	0.0587	2.49	0.092	2.06	568	565	557	568 ± 12
z29c	288	92	26	0.0574	1.16	0.0801	1.35	497	499	508	497 ± 7
z31b	261	60	23	0.0562	1	0.0782	1.6	486	481	459	486 ± 8
z31c	715	196	54	0.0559	0.71	0.0653	1.35	408	414	449	407 ± 6
z32b	520	40	44	0.0571	0.57	0.0775	1.36	481	484	495	481 ± 7
z33b	386	31	33	0.0561	0.67	0.0787	1.35	488	482	455	489 ± 7
z33c	446	208	43	0.0571	0.77	0.0789	1.43	489	490	494	489 ± 7

Table 4.23: U-Pb SIMS results of the Vitigudino metagranite. Analyses references ending in b mean rims and c mean cores.

Grain	Element concentration (ppm)			Isotope ratios				Ages (Ma)			
	U	Th	Pb	$(^{207}\text{Pb}/^{206}\text{Pb})_{\text{meas}}$	σ (%)	$(^{206}\text{Pb}/^{238}\text{U})_{\text{meas}}$	σ (%)	$^{206}\text{Pb}/^{238}\text{U}$	$^{207}\text{Pb}/^{235}\text{U}$	$^{207}\text{Pb}/^{206}\text{Pb}$	$(^{206}\text{Pb}/^{238}\text{U})_{\text{cor}}$
z1c	289	65	24	0.0553	1.73	0.0764	1.64	475	466	425	476 ± 8
z7c	944	448	102	0.0596	0.8	0.0923	1.35	569	573	590	569 ± 8
z10b	322	27	27	0.0569	1.17	0.0767	1.5	476	479	489	476 ± 7
z10c	162	119	19	0.0585	1.63	0.0955	1.36	588	580	550	589 ± 8
z15c	141	102	18	0.0614	2.02	0.105	1.35	643	646	654	643 ± 9
z1	510	48	44	0.0588	0.91	0.0798	1.47	495	507	561	494 ± 7
z2	369	58	39	0.075	0.97	0.1712	1.63	1019	1035	1069	-
z3	598	49	45	0.0747	0.74	0.1702	1.48	1013	1028	1059	-
z4	316	40	27	0.0592	0.84	0.0791	1.5	491	506	574	489 ± 7
z5	510	42	46	0.064	0.68	0.0968	1.47	596	627	743	593 ± 9
z6	436	39	37	0.0562	1.04	0.0788	1.48	489	484	462	489 ± 7
z7	722	76	68	0.0582	0.53	0.0859	1.47	531	532	536	531 ± 8
z8	641	49	56	0.058	0.68	0.0803	1.47	498	504	529	497 ± 7
z9	396	52	35	0.0575	0.76	0.0798	1.49	495	498	512	495 ± 7
z10	199	37	17	0.0581	1.03	0.0795	1.47	493	500	533	493 ± 7
z11	762	59	67	0.0582	0.52	0.0801	1.47	497	504	538	496 ± 7
z12	359	72	29	0.0621	0.91	0.0969	1.48	596	614	679	595 ± 9
z13	633	274	72	0.0594	0.56	0.0957	1.47	589	587	580	589 ± 9
z14	175	66	19	0.06	1.23	0.0922	1.47	569	576	603	568 ± 9

Table 4.24: U-Pb ion microprobe results of the Ledesma metagranite. Analyses references ending in b mean rims and c mean cores.

Grain	Element concentration (ppm)			Isotope ratios				Ages (Ma)			
	U	Th	Pb	$(^{207}\text{Pb}/^{206}\text{Pb})_{\text{meas}}$	σ (%)	$(^{206}\text{Pb}/^{238}\text{U})_{\text{meas}}$	σ (%)	$^{206}\text{Pb}/^{238}\text{U}$	$^{207}\text{Pb}/^{235}\text{U}$	$^{207}\text{Pb}/^{206}\text{Pb}$	$(^{206}\text{Pb}/^{238}\text{U})_{\text{cor}}$
z17b	443	29	36	0.0595	1.26	0.0774	1.95	481	499	584	479 ± 9
z18c	525	606	81	0.0624	0.61	0.1064	3.29	652	660	688	651 ± 21
z19b	512	32	40	0.0573	1.08	0.0737	1.95	459	466	504	458 ± 9
z22b	496	4	40	0.0574	1.12	0.0764	1.95	475	480	508	474 ± 9
z22c	39	20	11	0.0784	2.85	0.2247	2.56	1306	1251	1156	-
z29b	374	38	31	0.0583	1.05	0.0764	1.35	474	486	542	473 ± 7
z29c	484	117	51	0.0589	1.02	0.0943	1.61	581	578	565	582 ± 9
z29m	131	37	13	0.06	1.8	0.0883	1.35	546	557	605	545 ± 8
z30b	481	42	39	0.0584	1.25	0.0763	1.36	474	486	546	473 ± 7
z30c	502	432	100	0.0703	0.59	0.1565	1.36	937	934	936	938 ± 13
z31c	686	340	66	0.0562	0.61	0.0804	1.36	499	492	459	499 ± 7
z33c	439	200	53	0.0615	0.82	0.1019	1.35	625	633	658	625 ± 9
z34c	800	189	91	0.0607	0.61	0.1004	1.38	617	620	629	617 ± 9

Table 4.25: U-Pb LA-ICPMS, SIMS and SHRIMP results of the San Pelayo metagranite. Analyses references ending in b mean rims and c mean cores.

Grain	Element concentration (ppm)			Isotope ratios				Ages (Ma)			
	U	Th	Pb	$(^{207}\text{Pb}/^{206}\text{Pb})_{\text{meas}}$	σ (%)	$(^{206}\text{Pb}/^{238}\text{U})_{\text{meas}}$	σ (%)	$^{206}\text{Pb}/^{238}\text{U}$	$^{207}\text{Pb}/^{235}\text{U}$	$^{207}\text{Pb}/^{206}\text{Pb}$	$(^{206}\text{Pb}/^{238}\text{U})_{\text{cor}}$
<i>Laser-Ablation ICPMS data</i>											
z49	209	95	23	0.062	2.38	0.104	1.97	639	646	670	639 ± 13
z51	186	183	23	0.059	2.38	0.104	1.97	640	623	562	642 ± 13
z52	100	31	31	0.118	2.38	0.298	1.97	1683	1791	1919	-
z53	733	343	52	0.059	2.38	0.07	1.97	435	455	559	433 ± 9
z54	228	329	27	0.059	2.38	0.094	1.97	578	578	579	578 ± 12
z55	332	56	104	0.122	2.38	0.311	1.97	1745	1855	1980	-
z56	333	79	24	0.057	2.38	0.075	1.97	467	472	496	466 ± 10
z57	114	156	14	0.063	2.38	0.096	1.97	589	615	712	586 ± 12
<i>SIMS data</i>											
z27	39	24	5	0.0596	2.26	0.0951	1.34	585	586	587.7	585 ± 9
z28b	776	18	65	0.0568	0.56	0.0791	1.34	491	490	484.9	491 ± 7
z28c	307	72	41	0.0638	0.7	0.1176	1.35	717	721	734.3	717 ± 10
z29	166	51	29	0.071	0.79	0.1523	1.34	914	927	958.3	912 ± 12
z30	196	94	25	0.062	0.96	0.1047	1.39	642	649	672.3	641 ± 9
z31	680	29	58	0.0562	0.71	0.0803	1.34	498	491	459.1	498 ± 7
z32	687	176	64	0.0565	0.61	0.0822	1.34	509	503	471.3	510 ± 7
z33	133	167	19	0.0587	1.19	0.097	1.36	597	589	557.1	598 ± 8
z33b	103	91	14	0.059	1.82	0.0976	1.37	600	594	567.9	601 ± 9
<i>SHRIMP data</i>											
z1.1	2191	21	161	0.0601	0.6	0.0796	2.2	491	486	463.1	492 ± 11
z1.2	274	218	107	0.1153	0.8	0.3347	3.1	1861	1872	1883.1	-
z2.1	601	25	43	0.0571	1.2	0.0778	2.2	482	480	465.9	483 ± 11
z2.2	180	70	17	0.0595	2	0.0944	2.3	580	568	519.7	581 ± 14
z3.1	238	118	27	0.0622	1.6	0.1094	2.2	668	651	593.5	669 ± 15
z4.1	1816	42	134	0.0564	0.7	0.0805	2.2	498	488	438.9	499 ± 11
z4.2	1568	171	132	0.0606	0.7	0.0902	2.3	556	556	555.3	556 ± 13
z5.1	586	58	43	0.0579	1.3	0.0776	2.2	481	483	492.9	481 ± 11
z5.2	514	15	38	0.0556	1.3	0.0799	2.2	495	476	388.9	496 ± 11
z6.1	1141	169	357	0.1228	0.3	0.3108	2.2	1744	1861	1993.5	-
z6.2	1316	33	96	0.057	0.8	0.0798	2.2	494	484	437.9	495 ± 11
z7.1	493	142	251	0.1713	0.7	0.4719	2.3	2491	2534	2567.5	-
z7.2	956	47	65	0.0631	1	0.0724	2.2	447	453	479.9	447 ± 10
z8.1	502	23	36	0.0564	1.3	0.0788	2.2	489	483	456.3	489 ± 11
z8.2	158	60	17	0.0591	1.9	0.109	2.3	666	635	528.5	669 ± 16
z10.1	1720	12	109	0.0558	0.8	0.0699	2.2	435	434	424.9	435 ± 10
z10.2	379	130	36	0.0584	1.3	0.095	2.3	584	567	500.3	586 ± 13
z11.1	878	69	80	0.06	0.8	0.0981	2.2	603	600	587.9	603 ± 13

Table 4.26: U-Pb LA-ICPMS and SIMS results of the Castellanos metagranite. Analyses references ending in b mean rims and c mean cores.

Grain	Element concentration (ppm)			Isotope ratios				Ages (Ma)			
	U	Th	Pb	$(^{207}\text{Pb}/^{206}\text{Pb})_{\text{meas}}$	σ (%)	$(^{206}\text{Pb}/^{238}\text{U})_{\text{meas}}$	σ (%)	$^{206}\text{Pb}/^{238}\text{U}$	$^{207}\text{Pb}/^{235}\text{U}$	$^{207}\text{Pb}/^{206}\text{Pb}$	$(^{206}\text{Pb}/^{238}\text{U})_{\text{cor}}$
<i>Laser-Ablation ICPMS data</i>											
z5c	491	167	46	0.06	1.2	0.092	2.2	569	578	614	568 ± 12
z6b	862	3	60	0.056	1.2	0.075	2.2	469	465	445	469 ± 10
z7b	586	202	58	0.062	1.2	0.097	2.2	596	610	661	594 ± 13
z7c	589	234	61	0.062	1.2	0.099	2.2	607	620	668	606 ± 13
z8	677	34	51	0.059	1.2	0.081	2.2	500	509	552	499 ± 11
z10b	544	258	45	0.056	1.2	0.08	2.2	495	485	435	496 ± 11
z10c	444	176	37	0.057	1.2	0.081	2.2	500	498	487	501 ± 11
z12c	301	136	32	0.06	1.2	0.102	2.2	627	618	587	627 ± 14
z15c	365	76	28	0.06	1.2	0.079	2.2	489	508	593	488 ± 11
z16	383	243	46	0.06	1.2	0.108	2.2	663	649	600	664 ± 14
z41a	238	68	23	0.061	1.2	0.088	2.2	543	558	621	542 ± 12
z41a	186	64	18	0.059	1.2	0.09	2.2	557	560	574	556 ± 12
z42a	375	73	32	0.057	1.2	0.081	2.2	504	502	492	504 ± 11
z42b	256	63	22	0.056	1.2	0.082	2.2	507	500	467	508 ± 11
z45	391	348	50	0.061	1.2	0.1	2.2	612	614	622	612 ± 13
z46a	543	59	44	0.057	1.2	0.081	2.2	501	499	490	501 ± 11
z46b	645	40	52	0.059	1.2	0.08	2.2	496	507	559	495 ± 11
z47	271	89	37	0.064	1.2	0.124	2.2	751	747	734	752 ± 16
z48	412	143	46	0.06	1.2	0.103	2.2	630	623	597	631 ± 14
<i>SIMS data</i>											
z-11c	198	120	76	0.1127	0.52	0.2967	1.4	1675	1752	1844	-
z-11b	982	81	99	0.0587	0.82	0.0896	1.41	553	553	555	553 ± 8
z-12	966	21	79	0.057	0.83	0.0777	1.34	482	484	492	482 ± 7
z-13	305	110	32	0.0602	1.16	0.091	1.34	561	571	609	560 ± 8
z-14	29	73	33	0.2467	0.64	0.6006	1.45	3032	3112	3164	-
z-15	2403	888	274	0.0601	0.3	0.0977	1.34	601	602	606	601 ± 8
z-16	190	46	45	0.1377	1.06	0.1929	1.69	1137	1563	2198	-
z-17	485	214	58	0.06	0.63	0.0997	1.34	612	611	604	613 ± 8
z-18	1040	62	91	0.0569	0.54	0.0815	1.34	505	502	489	505 ± 7
z-19	320	16	27	0.057	1.05	0.0801	1.34	496	496	492	497 ± 7
z-20	380	200	38	0.0565	0.99	0.0815	1.34	505	499	471	505 ± 7
z-21b	895	3	77	0.0571	0.53	0.081	1.34	502	501	497	502 ± 7
z-21a	1281	57	99	0.0578	0.88	0.0794	1.34	492	498	523	492 ± 7
z-22	461	221	59	0.0614	0.63	0.1048	1.34	643	645	654	643 ± 9
z-23	516	143	52	0.0579	0.76	0.0873	1.34	540	537	528	540 ± 7
z-24a	913	3	84	0.0573	0.59	0.0872	1.34	539	532	502	540 ± 7
z-24b	221	44	21	0.0583	1.06	0.0851	1.34	527	530	542	526 ± 7
z-25	386	262	55	0.0618	0.73	0.1117	1.34	683	679	668	683 ± 9

Table 4.27: U-Pb LA-ICPMS and SIMS results of the Bercimuelle metagranite. Analyses references ending in b mean rims and c mean cores.

Grain	Element concentration (ppm)			Isotope ratios				Ages (Ma)			
	U	Th	Pb	$(^{207}\text{Pb}/^{206}\text{Pb})_{\text{meas}}$	σ (%)	$(^{206}\text{Pb}/^{238}\text{U})_{\text{meas}}$	σ (%)	$^{206}\text{Pb}/^{238}\text{U}$	$^{207}\text{Pb}/^{235}\text{U}$	$^{207}\text{Pb}/^{206}\text{Pb}$	$(^{206}\text{Pb}/^{238}\text{U})_{\text{cor}}$
<i>Laser-ablation-ICPMS data</i>											
z1b	510	35	66	0.107	1.7	0.135	2.5	814	1100	1720	774 ± 20
z1c	299	79	132	0.174	1.7	0.417	2.5	2245	2430	2588	-
z2b	1585	67	124	0.057	1.7	0.088	2.5	542	524	443	544 ± 14
z2c	1169	533	136	0.065	1.7	0.113	2.5	689	702	741	688 ± 17
z3b	1397	434	112	0.06	1.7	0.083	2.5	511	517.7	548	510 ± 13
z5b	256	73	26	0.06	1.7	0.105	2.5	640	622.1	558	642 ± 16
z5c	442	113	43	0.06	1.7	0.102	2.5	626	612	562	627 ± 16
z6b	594	37	43	0.058	1.7	0.082	2.5	505	501	482	506 ± 13
z6c	423	39	31	0.055	1.7	0.078	2.5	482	458.9	343	485 ± 12
z7b	420	38	31	0.054	1.7	0.078	2.5	484	452.6	297	487 ± 12
z7c	790	163	59	0.057	1.7	0.08	2.5	494	486.3	451	494 ± 12
z8b	163	54	19	0.063	1.7	0.119	2.5	723	706.9	656	725 ± 18
z8c	331	159	42	0.066	1.7	0.122	2.5	741	742.3	746	741 ± 19
z9b	109	17	35	0.121	1.7	0.334	2.5	1854	1900.3	1952	-
z9c	116	19	36	0.121	1.7	0.322	2.5	1796	1867.7	1948	-
z20a	443	18	36	0.054	1.7	0.078	2.5	485	457	318	487 ± 12
z20b	533	195	57	0.06	1.7	0.099	2.5	607	592	537	608 ± 15
z21b	546	162	56	0.062	1.7	0.097	2.5	594	599	617	593 ± 15
z22b	1058	284	99	0.054	1.7	0.078	2.5	483	457	327	485 ± 12
z22c	1587	260	152	0.059	1.7	0.089	2.5	548	540	507	549 ± 14
z23b	647	76	57	0.058	1.7	0.081	2.5	499	493	463	499 ± 13
z23c	261	82	27	0.06	1.7	0.097	2.5	595	587	558	596 ± 15
z25c	622	29	64	0.061	1.7	0.105	2.5	641	632	602	642 ± 16
z26b	436	63	38	0.054	1.7	0.078	2.5	481	449	291	484 ± 12
z26bb	391	52	35	0.061	1.7	0.089	2.5	547	554	584	546 ± 14
z27b	389	48	66	0.122	1.7	0.157	2.5	940	1303	1965	880 ± 23
z27c	435	81	105	0.144	1.7	0.217	2.5	1262	1685	2261	-
z28b	708	113	54	0.059	1.7	0.077	2.5	479	483	501	478 ± 12
z28bb	605	66	47	0.058	1.7	0.078	2.5	481	482	488	481 ± 12
z28c	610	94	49	0.059	1.7	0.079	2.5	487	492	514	486 ± 12
z29b	615	45	209	0.173	1.7	0.309	2.5	1734	2148	2571	-
z29bb	903	54	467	0.198	1.7	0.466	2.5	2462	2650	2796	-
z29c	269	58	136	0.185	1.7	0.444	2.5	2364	2542	2686	-
z30b	525	166	55	0.059	1.7	0.102	2.5	623	599	508	626 ± 16
z30bb	432	23	39	0.058	1.7	0.096	2.5	592	567.9	471	595 ± 15
z30c	436	378	59	0.06	1.7	0.113	2.5	691	661	562	694 ± 17
z31c	523	84	45	0.057	1.7	0.079	2.5	488	475.8	420	489 ± 12

z31b	879	73	76	0.055	1.7	0.091	2.5	560	525	372	564 ± 14
z33b	496	122	42	0.057	1.7	0.089	2.5	548	526	435	550 ± 14
z33bb	609	124	52	0.056	1.7	0.087	2.5	538	515	413	540 ± 14
z33c	739	247	65	0.056	1.7	0.078	2.5	484	471	407	486 ± 12
z34a	248	77	22	0.058	1.7	0.079	2.5	489	488	480	489 ± 12
z36b	462	58	39	0.056	1.7	0.078	2.5	484	468	389	485 ± 12
z36bb	574	201	52	0.058	1.7	0.088	2.5	541	525.6	460	542 ± 14
z36c	477	95	41	0.057	1.7	0.088	2.5	542	523.1	440	544 ± 14
z38a	548	186	58	0.059	1.7	0.103	2.5	630	606.4	518	633 ± 16
z38b	461	137	49	0.06	1.7	0.104	2.5	634	613.8	540	636 ± 16
<i>Ion-microprobe data</i>											
z2	1500	584	184	0.0603	0.36	0.1044	1.34	640	634	613	641 ± 8
z6	459	27	38	0.057	0.78	0.0781	1.34	485	486	492	485 ± 7
z7	638	93	56	0.0573	0.68	0.0793	1.34	492	494	504	492 ± 7
z8a	300	10	25	0.0573	0.96	0.0783	1.34	486	489	501	486 ± 7
z8b	482	8	40	0.0569	0.76	0.0789	1.34	489	489	489	489 ± 7
z9a	72	24	11	0.0652	1.45	0.1259	1.34	765	769	781	764 ± 11
z9b	184	109	29	0.0642	0.95	0.1248	1.35	758	756	750	758 ± 10
z10	268	11	22	0.0572	1.01	0.0781	1.34	485	488	500	485 ± 9

Table 4.28: U-Pb LA-ICPMS and SIMS results of the La Cañada metagranite. Analyses references ending in b mean rims and c mean cores.

Grain	Element concentration (ppm)			Isotope ratios				Ages (Ma)			
	U	Th	Pb	$(^{207}\text{Pb}/^{206}\text{Pb})_{\text{meas}}$	σ (%)	$(^{206}\text{Pb}/^{238}\text{U})_{\text{meas}}$	σ (%)	$^{206}\text{Pb}/^{238}\text{U}$	$^{207}\text{Pb}/^{235}\text{U}$	$^{207}\text{Pb}/^{206}\text{Pb}$	$(^{206}\text{Pb}/^{238}\text{U})_{\text{cor}}$
<i>Laser-Ablation ICPMS data</i>											
z2	4369	1240	278	0.061	2.77	0.067	3.45	412	416	439	419 ± 14
z3	281	95	20	0.075	9.41	0.072	5.66	436	409	263	438 ± 26
z4	232	281	85	0.121	2.31	0.272	7	1525	1608	1718	-
z6	685	39	36	0.058	3.01	0.056	2.5	350	321	114	352 ± 9
z16	1129	139	64	0.077	15.36	0.056	8.23	338	329	266	338 ± 29
z17	1423	253	88	0.066	4.69	0.061	4.93	377	379	392	377 ± 19
z18	2804	1320	205	0.094	11.57	0.064	11.8	379	378	375	378 ± 43
z27	1533	623	116	0.061	2.51	0.071	2.68	437	405	222	440 ± 12
<i>Ion-microprobe data</i>											
z39b	542	54	45	0.057	0.97	0.0772	1.05	479	481	491	479 ± 5
z39m	1621	609	193	0.0624	0.46	0.1022	1.05	627	641	688	626 ± 7
z41c	562	328	54	0.0582	0.89	0.0788	1.1	489	498	536	489 ± 6
z43b	2698	95	230	0.0569	0.42	0.0802	1.05	497	496	489	497 ± 5
z43c	104	132	14	0.0591	1.84	0.0961	1.06	591	587	572	592 ± 7
z44m	300	296	37	0.0602	1.11	0.0902	1.05	557	568	611	556 ± 6
z45b	998	64	84	0.0559	0.69	0.0787	1.05	488	481	447	489 ± 5
z45c	1504	370	161	0.0594	0.62	0.0942	1.26	581	581	583	581 ± 7
z48c	208	71	22	0.0582	1.74	0.0897	1.08	554	550	536	554 ± 6
z49b	369	32	30	0.0573	1.1	0.0764	1.06	475	480	503	474 ± 5
z49c	339	157	214	0.1818	0.38	0.4819	1.05	2536	2610	2670	-
z50m	433	395	54	0.0598	0.94	0.0933	1.05	575	579	596	575 ± 6
z51b	361	9	31	0.0575	1.11	0.0812	1.06	504	505	511	503 ± 6

Table 4.29: U-Pb SIMS results of La Hoya metagranite. Analyses references ending in b mean rims and c mean cores.

Grain	Element concentration (ppm)			Isotope ratios				Ages (Ma)			
	U	Th	Pb	$(^{207}\text{Pb}/^{206}\text{Pb})_{\text{meas}}$	σ (%)	$(^{206}\text{Pb}/^{238}\text{U})_{\text{meas}}$	σ (%)	$^{206}\text{Pb}/^{238}\text{U}$	$^{207}\text{Pb}/^{235}\text{U}$	$^{207}\text{Pb}/^{206}\text{Pb}$	$(^{206}\text{Pb}/^{238}\text{U})_{\text{cor}}$
z2	215	30	18	0.05695	1.3	0.0768	1.06	477	479	490	477 ± 5
z5	765	55	62	0.05681	0.71	0.0753	1.06	468	471	484	468 ± 5
z8	222	130	21	0.05743	1.3	0.0755	1.05	469	476	508	469 ± 5
z12b	1502	54	101	0.05546	0.54	0.063	1.05	394	399	431	393 ± 4
z14	432	629	51	0.05739	0.87	0.0774	1.05	480	485	507	480 ± 5
z15c	353	532	42	0.05593	1.15	0.0777	1.09	483	477	450	483 ± 6
z16b	547	4	46	0.05712	0.74	0.0789	1.05	490	491	496	489 ± 5
z16c	43	68	8	0.06412	2.25	0.1166	1.05	711	719	746	710 ± 8

Table 4.30: U-Pb SIMS results of the La Estación metagranite. Analyses references ending in b mean rims and c mean cores.

Grain	Element concentration (ppm)			Isotope ratios				Ages (Ma)			
	U	Th	Pb	$(^{207}\text{Pb}/^{206}\text{Pb})_{\text{meas}}$	σ (%)	$(^{206}\text{Pb}/^{238}\text{U})_{\text{meas}}$	σ (%)	$^{206}\text{Pb}/^{238}\text{U}$	$^{207}\text{Pb}/^{235}\text{U}$	$^{207}\text{Pb}/^{206}\text{Pb}$	$(^{206}\text{Pb}/^{238}\text{U})_{\text{cor}}$
z24b	2306	105	197	0.0565	0.4	0.08	0.91	496	492	472	496 ± 4
z26b	2458	23	205	0.0571	0.38	0.0789	0.92	489	491	496	489 ± 4
z29b	913	46	75	0.0568	0.61	0.0769	0.95	477	479	484	477 ± 5
z29c	271	128	32	0.0603	0.95	0.0989	0.93	608	610	616	608 ± 6
z30c	583	976	73	0.0564	0.76	0.079	0.92	490	486	466	490 ± 5
z30m	497	25	41	0.056	0.86	0.0774	0.92	480	475	451	481 ± 5
z32b	765	10	65	0.0568	0.92	0.0808	0.84	501	498	482	501 ± 4
z32c	414	212	49	0.0599	0.93	0.0972	0.83	598	599	602	598 ± 5
z33b	409	38	35	0.0563	1.16	0.08	0.88	496	491	466	497 ± 5
z33c	631	238	115	0.0693	0.55	0.154	0.83	924	919	909	924 ± 8
z34b	586	51	50	0.0566	0.89	0.0795	0.84	493	490	475	493 ± 4
z35c	254	200	26	0.0567	1.34	0.0779	0.84	484	483	481	484 ± 4
z36b	360	43	31	0.0575	1.08	0.0798	0.84	495	498	510	495 ± 4
z36c	95	44	12	0.0607	1.76	0.1054	0.89	646	642	630	646 ± 6

Table 4.31: U-Pb SIMS results of the La Estación II metagranite. Analyses references ending in b mean rims and c mean cores.

Grain	Element concentration (ppm)			Isotope ratios				Ages (Ma)			
	U	Th	Pb	$(^{207}\text{Pb}/^{206}\text{Pb})_{\text{meas}}$	σ (%)	$(^{206}\text{Pb}/^{238}\text{U})_{\text{meas}}$	σ (%)	$^{206}\text{Pb}/^{238}\text{U}$	$^{207}\text{Pb}/^{235}\text{U}$	$^{207}\text{Pb}/^{206}\text{Pb}$	$(^{206}\text{Pb}/^{238}\text{U})_{\text{cor}}$
z39b	1150	51	96	0.0567	0.55	0.0777	1.06	482	482	478	482 ± 5
z40b	3626	93	288	0.0584	0.3	0.0764	1.05	475	487	545	474 ± 5
z47b	518	68	44	0.0576	0.99	0.0774	0.92	481	487	515	480 ± 5
z48c	352	322	34	0.0573	1.24	0.0734	0.92	457	465	503	456 ± 5
z49b	734	72	60	0.0558	1.11	0.0751	0.93	467	463	443	467 ± 5
z49c	669	557	67	0.0571	0.86	0.077	0.93	479	481	494	478 ± 5

Table 4.32: U-Pb SIMS results of the melanocratic metagranites of Vegas de Matute. Analyses references ending in b mean rims and c mean cores.

Grain	Element concentration (ppm)			Isotope ratios				Ages (Ma)			
	U	Th	Pb	$(^{207}\text{Pb}/^{206}\text{Pb})_{\text{meas}}$	σ (%)	$(^{206}\text{Pb}/^{238}\text{U})_{\text{meas}}$	σ (%)	$^{206}\text{Pb}/^{238}\text{U}$	$^{207}\text{Pb}/^{235}\text{U}$	$^{207}\text{Pb}/^{206}\text{Pb}$	$(^{206}\text{Pb}/^{238}\text{U})_{\text{cor}}$
z48b	244	11	20	0.0568	1.86	0.0757	1.2	470	473	482	470 ± 6
z48c	191	95	20	0.0587	2.62	0.0911	1.2	562	561	556	562 ± 8
z49b	481	65	39	0.0579	1.27	0.0757	1.2	470	480	526	469 ± 6
z50b	400	45	34	0.0561	1.11	0.0792	1.2	491	485	455	492 ± 6
z50c	92	212	26	0.0734	1.35	0.1697	1.3	1011	1015	1024	-
z52b	713	38	69	0.059	0.85	0.0906	1.25	559	561	568	559 ± 7
z52c	201	100	23	0.0597	1.34	0.0973	1.2	599	598	594	599 ± 8
z53b	479	70	45	0.0576	0.74	0.0856	1.42	530	527	515	530 ± 8
z53c	407	206	45	0.0589	0.95	0.0929	1.2	573	571	563	573 ± 7
z54c	17	18	9	0.121	1.48	0.3484	1.3	1927	1948	1971	-
z55c	787	161	72	0.0564	0.44	0.0816	1.36	505	499	469	506 ± 7
z57c	369	343	39	0.0565	0.67	0.0772	1.21	479	478	475	479 ± 6
z58b	325	29	27	0.0571	0.71	0.077	1.24	478	481	496	478 ± 6
z59b	224	15	19	0.0565	0.87	0.0766	1.22	476	475	473	476 ± 6
z59c	383	166	41	0.0599	1.28	0.0884	1.22	546	557	602	545 ± 7
z60	371	26	30	0.0565	0.78	0.0757	1.2	470	471	473	470 ± 6
z61b	298	23	24	0.0571	0.8	0.0757	1.2	470	475	497	470 ± 6
z61c	94	39	10	0.0593	1.51	0.0922	1.2	569	571	579	569 ± 7
z63b	292	24	24	0.0574	0.82	0.077	1.35	478	483	506	478 ± 7
z63c	107	83	14	0.0615	1.14	0.0992	1.22	610	620	656	609 ± 8

Table 4.33: U-Pb SIMS results of the leucocratic metagranite of Vegas de Matute. Analyses references ending in b mean rims and c mean cores.

Grain	Element concentration (ppm)			Isotope ratios				Ages (Ma)			
	U	Th	Pb	$(^{207}\text{Pb}/^{206}\text{Pb})_{\text{meas}}$	σ (%)	$(^{206}\text{Pb}/^{238}\text{U})_{\text{meas}}$	σ (%)	$^{206}\text{Pb}/^{238}\text{U}$	$^{207}\text{Pb}/^{235}\text{U}$	$^{207}\text{Pb}/^{206}\text{Pb}$	$(^{206}\text{Pb}/^{238}\text{U})_{\text{cor}}$
z45b	271	32	24	0.0581	1.74	0.0813	2.05	504	509	533	503 ± 11
z44b	985	89	85	0.0573	1.27	0.081	2.05	502	502	503	502 ± 10
z33c	223	277	25	0.0559	1.38	0.0767	1.96	477	472	448	477 ± 10
z35c	126	21	11	0.058	1.78	0.0766	1.95	476	485	530	475 ± 10
z1b	331	114	28	0.0796	1.14	0.0789	1.48	477	490	548	476 ± 7
z2c	297	41	20	0.0702	0.94	0.0595	1.47	365	366	370	365 ± 6
z3b	335	80	26	0.0931	0.7	0.0623	1.47	371	372	378	371 ± 6
z4b	3098	110	275	0.0602	0.7	0.0777	1.47	480	478	470	480 ± 7
z4c	564	152	51	0.1044	1.28	0.07	1.48	411	418	454	409 ± 7
z6c	844	84	71	0.0788	0.67	0.0715	1.47	433	431	424	432 ± 7
z8c	1352	77	112	0.0597	0.45	0.0765	1.48	473	471	462	473 ± 7
z9b	1315	68	112	0.0592	0.39	0.0788	1.47	487	482	457	488 ± 7
z10c	1552	365	158	0.0823	0.29	0.0788	1.47	475	484	529	473 ± 7
z10b	1051	169	101	0.0758	0.81	0.0802	1.49	486	490	504	486 ± 7
z11b	2060	140	174	0.1192	1.14	0.0648	1.48	372	366	326	372 ± 6
z12c	312	44	26	0.0593	0.98	0.0794	1.49	490	479	425	491 ± 7
z15c	456	88	39	0.0593	0.8	0.0786	1.48	486	484	472	486 ± 7
z20c	380	54	32	0.0581	0.84	0.0776	1.48	481	479	470	481 ± 7

Table 4.34: U-Pb SIMS results of the Morcuera metagranite. Analyses references ending in b mean rims and c mean cores.

Grain	Element concentration (ppm)			Isotope ratios				Ages (Ma)			
	U	Th	Pb	$(^{207}\text{Pb}/^{206}\text{Pb})_{\text{meas}}$	σ (%)	$(^{206}\text{Pb}/^{238}\text{U})_{\text{meas}}$	σ (%)	$^{206}\text{Pb}/^{238}\text{U}$	$^{207}\text{Pb}/^{235}\text{U}$	$^{207}\text{Pb}/^{206}\text{Pb}$	$(^{206}\text{Pb}/^{238}\text{U})_{\text{cor}}$
z42b	243	51	21	0.0583	1.28	0.0785	0.84	487	497	543	486 ± 4
z42c	153	80	20	0.061	1.37	0.1064	0.83	652	649	638	652 ± 6
z48b	259	32	22	0.0553	1.2	0.0794	0.84	492	481	426	494 ± 4
z48c	179	162	25	0.0621	1.41	0.1035	0.83	635	644	679	634 ± 6
z50c	177	96	20	0.0583	1.31	0.0927	0.84	571	566	542	572 ± 5
z53b	603	62	49	0.0571	0.74	0.0753	0.83	468	472	494	468 ± 4
z54b	463	46	40	0.0569	0.87	0.0786	0.83	488	488	489	488 ± 4
z54c	205	47	28	0.0652	1.19	0.1193	0.83	726	740	782	725 ± 7
z55	370	30	31	0.0575	1.01	0.0777	0.84	483	488	512	482 ± 4
z57	296	40	25	0.0575	1.07	0.0767	0.84	477	483	512	476 ± 4
z58b	321	30	29	0.0577	1.34	0.0833	0.85	516	517	520	516 ± 5
z59c	483	165	58	0.0613	0.71	0.1029	0.84	632	635	648	631 ± 5

Table 4.35: U-Pb SIMS results of the Buitrago de Lozoya metagranite. Analyses references ending in b mean rims and c mean cores.

Grain	Element concentration (ppm)			Isotope ratios				Ages (Ma)			
	U	Th	Pb	$(^{207}\text{Pb}/^{206}\text{Pb})_{\text{meas}}$	σ (%)	$(^{206}\text{Pb}/^{238}\text{U})_{\text{meas}}$	σ (%)	$^{206}\text{Pb}/^{238}\text{U}$	$^{207}\text{Pb}/^{235}\text{U}$	$^{207}\text{Pb}/^{206}\text{Pb}$	$(^{206}\text{Pb}/^{238}\text{U})_{\text{cor}}$
z10b	506	14	42	0.0574	1.4	0.0774	1.98	481	486	509	480 ± 10
z11b	778	109	74	0.0593	0.82	0.0868	1.98	537	545	578	536 ± 11
z11c	513	111	50	0.0604	1.14	0.0882	1.95	545	560	619	544 ± 11
z13c	88	71	10	0.0598	2.05	0.0912	2.03	563	569	596	562 ± 12
z14b	297	7	24	0.0572	1.09	0.0754	1.63	468	474	499	468 ± 8
z16c	119	62	71	0.18	0.51	0.4518	1.63	2403	2541	2653	-
z17c	120	170	25	0.0667	1.04	0.1389	1.65	838	835	827	839 ± 14
z23b	450	37	38	0.0566	1.09	0.0787	1.63	488	486	475	488 ± 8
z23c	339	283	45	0.0607	0.72	0.1019	1.66	625	626	628	625 ± 10
z27b	372	34	32	0.0562	0.81	0.0786	1.63	488	483	462	488 ± 8
z27c	254	103	32	0.0621	0.78	0.1065	1.64	652	658	678	652 ± 11

Table 4.36: U-Pb SIMS results of the La Berzosa metavolcanic rocks. Analyses references ending in b mean rims and c mean cores.

Grain	Element concentration (ppm)			Isotope ratios				Ages (Ma)			
	U	Th	Pb	$(^{207}\text{Pb}/^{206}\text{Pb})_{\text{meas}}$	σ (%)	$(^{206}\text{Pb}/^{238}\text{U})_{\text{meas}}$	σ (%)	$^{206}\text{Pb}/^{238}\text{U}$	$^{207}\text{Pb}/^{235}\text{U}$	$^{207}\text{Pb}/^{206}\text{Pb}$	$(^{206}\text{Pb}/^{238}\text{U})_{\text{cor}}$
z1b	333	46	28	0.0561	1.4	0.0765	2	475	472	457	475 ± 10
z2	197	106	18	0.0596	2.02	0.0777	1.95	483	502	591	481 ± 10
z4b	311	84	26	0.0562	1.97	0.0731	1.98	455	456	461	455 ± 9
z4b2	127	59	11	0.0571	1.14	0.0733	0.68	456	463	495	456 ± 3
z6c	64	45	9	0.0634	2.62	0.1129	2	690	697	721	689 ± 15
z8c	143	40	86	0.1832	0.61	0.4758	0.66	2509	2606	2682	-
z9b	334	29	31	0.062	0.93	0.0869	1.96	537	564	674	535 ± 10
z11c	240	179	33	0.0611	1.15	0.1035	1.95	635	636	642	635 ± 12
z14c	90	60	42	0.1206	0.94	0.3494	1.96	1932	1948	1965	-
z15c	307	119	35	0.06	0.96	0.0963	1.95	593	595	604	593 ± 11
z15b	207	21	17	0.0641	1.25	0.0744	1.98	463	513	746	458 ± 9
z16c	331	283	42	0.0594	1.14	0.098	1.68	603	598	580	603 ± 10
z18b	556	293	67	0.0624	0.79	0.1	1.7	615	631	690	613 ± 10
z18c	367	154	44	0.0634	0.74	0.1017	1.64	625	646	723	622 ± 10
z19b	607	50	47	0.0642	0.81	0.0726	1.65	452	504	748	447 ± 7

Table 4.37: U-Pb SIMS results of the El Cardoso metavolcanic rocks. Analyses references ending in b mean rims and c mean cores.

Grain	Element concentration (ppm)			Isotope ratios				Ages (Ma)			
	U	Th	Pb	$(^{207}\text{Pb}/^{206}\text{Pb})_{\text{meas}}$	σ (%)	$(^{206}\text{Pb}/^{238}\text{U})_{\text{meas}}$	σ (%)	$^{206}\text{Pb}/^{238}\text{U}$	$^{207}\text{Pb}/^{235}\text{U}$	$^{207}\text{Pb}/^{206}\text{Pb}$	$(^{206}\text{Pb}/^{238}\text{U})_{\text{cor}}$
z1c	177	117	22	0.0622	1.8	0.0994	1.52	610	618	646	609 ± 10
z4b	235	20	20	0.0583	1.22	0.0787	1.49	487	485	473	488 ± 7
z4c	530	224	197	0.1192	0.48	0.3024	1.47	1703	1812	1941	-
z5c	345	408	33	0.0623	1.07	0.0782	1.49	483	488	514	482 ± 7
z7c	307	922	20	0.0755	0.93	0.0526	1.47	323	340	461	321 ± 5
z8c	587	21	50	0.0573	0.64	0.0791	1.48	491	492	496	491 ± 7
z10c	1380	1224	146	0.0576	0.51	0.0796	1.49	494	496	509	493 ± 7
z11c	713	631	69	0.0584	0.62	0.0776	1.48	481	482	485	481 ± 7
z15c	1848	1379	191	0.0566	0.43	0.0784	1.49	486	482	462	487 ± 7
z16c	59	17	10	0.0693	1.31	0.144	1.54	866	868	873	866 ± 14
z16b	497	180	62	0.0638	0.55	0.1078	1.49	659	669	703	658 ± 10
z24b	318	134	34	0.0637	0.86	0.0978	3.36	601	630	733	599 ± 19
z26b	289	19	25	0.0559	1.06	0.0791	3.31	491	483	448	492 ± 16
z30b	188	120	16	0.0578	1.35	0.0748	3.29	465	476	528	464 ± 15
z32b	268	31	22	0.0574	1.2	0.0761	3.29	473	479	507	472 ± 15

Table 4.38: U-Pb SIMS results of the Riaza metavolcanic rocks. Analyses references ending in b mean rims and c mean cores.

Grain	Element concentration (ppm)			Isotope ratios				Ages (Ma)			
	U	Th	Pb	$(^{207}\text{Pb}/^{206}\text{Pb})_{\text{meas}}$	σ (%)	$(^{206}\text{Pb}/^{238}\text{U})_{\text{meas}}$	σ (%)	$^{206}\text{Pb}/^{238}\text{U}$	$^{207}\text{Pb}/^{235}\text{U}$	$^{207}\text{Pb}/^{206}\text{Pb}$	$(^{206}\text{Pb}/^{238}\text{U})_{\text{cor}}$
z21c	975	49	366	0.1214	0.54	0.3303	1.49	1840	1905	1976	-
z22c	221	273	30	0.0614	1.01	0.0936	1.47	576	580	596	575 ± 9
z23c	1068	316	136	0.078	0.38	0.1073	1.47	646	664	724	643 ± 9
z23b	493	35	40	0.0584	0.83	0.0765	1.47	473	464	420	474 ± 7
z24c	365	251	45	0.0834	0.63	0.0915	1.47	548	554	575	547 ± 8
z25b	504	69	40	0.0656	0.86	0.0766	1.49	471	471	473	470 ± 7
z32c	295	39	24	0.063	0.99	0.0769	1.47	473	463	415	474 ± 7
z33c	234	189	31	0.0642	0.87	0.0991	1.48	606	599	575	606 ± 9
z35c	148	116	19	0.0613	1.13	0.0969	1.54	596	600	618	595 ± 9
z35b	440	228	53	0.0599	0.73	0.0976	1.47	600	594	574	600 ± 9
z35b2	415	30	35	0.0567	0.93	0.0784	1.58	486	476	430	487 ± 8
z4c	145	125	21	0.0609	1.29	0.1042	1.99	639	638	637	639 ± 13
z5b	341	15	28	0.0574	1.24	0.0771	1.96	479	484	506	479 ± 9
z5c	189	75	20	0.0576	1.26	0.0881	1.95	544	539	517	545 ± 11
z6b	246	23	20	0.0575	1.12	0.0754	1.95	469	476	510	468 ± 9
z9b	222	73	23	0.0593	0.87	0.0869	0.71	537	545	577	537 ± 4
z13b	121	93	15	0.0591	1.47	0.0964	1.95	594	589	569	594 ± 12

Table 4.39: U-Pb SIMS results of the Polán metagranite. Analyses references ending in b mean rims and c mean cores.

Grain	Element concentration (ppm)			Isotope ratios				Ages (Ma)			
	U	Th	Pb	$(^{207}\text{Pb}/^{206}\text{Pb})_{\text{meas}}$	σ (%)	$(^{206}\text{Pb}/^{238}\text{U})_{\text{meas}}$	σ (%)	$^{206}\text{Pb}/^{238}\text{U}$	$^{207}\text{Pb}/^{235}\text{U}$	$^{207}\text{Pb}/^{206}\text{Pb}$	$(^{206}\text{Pb}/^{238}\text{U})_{\text{cor}}$
z1a	476	297	57	0.0597	0.68	0.0953	1.34	587	588	592	587 ± 8
z1b	639	142	61	0.058	0.66	0.0842	1.34	521	522	528	521 ± 7
z2c	360	30	30	0.056	0.85	0.0774	1.54	480	476	452	481 ± 7
z3a	435	55	35	0.0551	0.99	0.0746	1.35	464	456	414	465 ± 6
z3b	565	54	47	0.0563	0.72	0.076	1.34	472	471	466	472 ± 6
z5	282	28	24	0.0565	0.95	0.0793	1.46	492	489	473	493 ± 7
z6	400	36	34	0.0567	1.17	0.0778	1.34	483	483	481	483 ± 7
z7	382	149	35	0.0566	1.04	0.0775	1.34	481	480	475	481 ± 7
z10	340	36	26	0.056	0.96	0.0689	1.34	430	433	454	429 ± 6

Table 4.40: U-Pb SIMS results of the Mohares metagranite. Analyses references ending in b mean rims and c mean cores.

Grain	Element concentration (ppm)			Isotope ratios				Ages (Ma)			
	U	Th	Pb	$(^{207}\text{Pb}/^{206}\text{Pb})_{\text{meas}}$	σ (%)	$(^{206}\text{Pb}/^{238}\text{U})_{\text{meas}}$	σ (%)	$^{206}\text{Pb}/^{238}\text{U}$	$^{207}\text{Pb}/^{235}\text{U}$	$^{207}\text{Pb}/^{206}\text{Pb}$	$(^{206}\text{Pb}/^{238}\text{U})_{\text{cor}}$
z11	643	822	73	0.0574	0.64	0.0795	1.34	493	496	509	493 ± 7
z14	314	49	27	0.0573	0.94	0.0771	1.34	479	483	503	478 ± 7
z16c	912	73	78	0.0557	0.61	0.0794	1.34	493	483	439	494 ± 7
z17	396	75	35	0.0564	0.84	0.0794	1.34	492	488	467	493 ± 7

9.2. Tables of the geochemical data

9.2.1. The Ollo de Sapo Domain

Table 5.3: Major elements and trace elements of the metavolcanic rocks of the Ollo de Sapo Domain.

Sample	T1	T2	T3	T4	T5	T6	T7	T8	T9	T10
Locality	Vivero		Mañón				Puebla de Trives			
SiO ₂	66.53	66.47	64.94	66.8	66.33	66.7	66.19	67.86	66.8	66.29
TiO ₂	0.61	0.74	0.71	0.56	0.61	0.62	0.59	0.51	0.55	0.61
Al ₂ O ₃	16.13	15.74	16.18	15.8	15.43	15.25	16.47	16.27	16.28	15.79
FeO _t	3.61	3.68	4.09	3.3	3.35	3.53	3.52	2.85	3.12	3.73
MgO	1.63	1.91	1.89	1.7	1.73	1.79	1.66	1.63	1.47	1.82
MnO	0.04	0.05	0.05	0.04	0.05	0.05	0.05	0.04	0.04	0.06
CaO	0.37	1.16	0.67	1.38	1.4	1.23	1.05	0.69	1.22	1.48
Na ₂ O	2.73	2.42	2.64	2.86	2.74	2.85	2.87	2.13	3.01	2.75
K ₂ O	4.95	4.09	4.21	4.18	4.36	4.31	5.12	5.26	4.53	4.18
P ₂ O ₅	0.232	0.211	0.229	0.228	0.232	0.234	0.231	0.215	0.216	0.219
Li	73	67	64	56	62	59	89	57	50	76
Rb	170	160	141	143	145	152	186	197	160	160
Cs	7.05	7.81	4.90	6.49	5.87	6.74	9.12	4.97	4.15	24.0
Be	2.40	2.85	2.67	2.22	2.87	2.68	2.91	2.93	2.46	2.34
Sr	46.1	147	63.4	157	141	138	144	112	152	131
Ba	851	655	693	763	800	766	908	820	892	646
Sc	10.3	12.9	12.3	9.82	10.5	10.3	9.59	8.78	8.53	11.1
V	67.2	82.2	78.7	66.3	69.0	67	58.6	57.5	54.6	70.1
Cr	53.8	63.7	59.7	49.5	57.2	53.5	46.8	51.8	45.5	56.0
Co	25.1	57.5	26.5	29.4	27.9	27.4	40.1	32.0	34.2	32.6
Ni	20.0	16.5	27.3	17.1	22.7	20.5	17.2	15.7	10.6	21.5
Cu	16.0	11.5	25.9	12.6	21	20	10.9	1.4	10.2	12.9
Zn	44.4	66.6	65.9	46.8	48.8	52.0	78.3	45.8	54.4	80.3
Ga	21.2	20.6	21.2	20.8	21.1	20.5	21.7	23.8	20.9	21.5
Y	25.7	20.5	22.2	27.4	29.7	34.6	18.5	15.0	28.9	38.3
Nb	10.9	12.2	11.481	10.3	11.1	10.9	11.1	11	10.5	11.3
Ta	0.92	1.12	1.00	0.89	1.01	0.99	1.01	1.10	0.99	1.07
Zr	209	225	213	183	197	191	196	187	189	193
Hf	6.30	7.40	6.30	5.60	6.00	6.00	6.30	5.90	6.30	6.20
Mo	0.70	0.41	0.78	0.89	0.80	0.73	0.38	0.21	0.22	0.37
Sn	5.13	6.05	4.08	4.6	4.47	4.47	5.52	6.81	5.74	8.78
Tl	0.97	0.83	0.80	0.76	0.75	0.83	1.13	1.00	0.88	0.91
Pb	23.2	34.5	22.1	18.8	23.2	20.7	29.5	17.3	25.3	20.4
U	3.84	2.80	3.61	3.12	3.45	3.5	2.93	3.52	3.26	3.73
Th	13.0	13.8	13.6	9.83	11.8	11.6	14.0	14.2	12.8	12.4
La	28.2	42.1	28.4	27.6	30.6	30.7	31.8	32.5	35.9	33.2
Ce	60.5	86.8	72.9	65.4	65.6	66.5	68.4	67.6	74.3	67.8
Pr	7.66	10.6	7.53	7.16	7.97	7.82	8.16	8.39	9.01	8.37
Nd	29.0	40.9	29.3	27.4	30.5	30	30.5	30.8	33.9	32.8
Sm	6.33	8.58	6.09	5.92	6.60	6.50	6.90	6.52	7.22	7.12
Eu	1.13	1.48	1.13	1.04	1.16	1.13	1.24	1.06	1.39	1.17
Gd	5.60	6.67	5.22	5.4	5.53	5.71	5.26	5.33	6.57	6.78
Tb	0.89	0.92	0.82	0.86	0.88	0.97	0.80	0.73	1.08	1.08
Dy	5.24	4.50	4.62	5.04	5.10	5.85	3.94	3.56	5.78	6.43
Ho	0.91	0.76	0.9	1.03	1.02	1.19	0.70	0.58	1.04	1.34
Er	2.32	1.72	2.3	2.65	2.61	3.22	1.59	1.26	2.45	3.50
Tm	0.34	0.23	0.35	0.39	0.40	0.48	0.21	0.16	0.32	0.50
Yb	1.99	1.42	2.3	2.29	2.46	2.80	1.26	0.89	1.75	2.98
Lu	0.31	0.20	0.32	0.35	0.36	0.39	0.20	0.13	0.23	0.42

Table 5.3: Major elements and trace elements of the metavolcanic rocks of the Ollo de Sapo Domain (Continuation).

Sample	T12	T13	T15	T16	T17	T18	T19	T20	T21	T22
Locality	Sanabria						Villadepera		Hiendelaencina	
SiO ₂	58.70	66.65	67.05	67.50	69.06	66.84	70.67	68.02	68.06	72.24
TiO ₂	1.05	0.69	0.52	0.54	0.61	0.5	0.54	0.68	0.5	0.38
Al ₂ O ₃	18.07	16.52	15.62	15.7	16.4	15.56	14.44	15.34	15.76	14.87
FeO _t	6.24	3.45	2.93	2.99	2.38	3.19	3.45	3.73	3.32	2.59
MgO	3.33	1.53	1.53	1.49	1.47	1.51	0.92	1.44	1.49	1.15
MnO	0.09	0.02	0.04	0.04	0.02	0.04	0.04	0.05	0.04	0.24
CaO	1.45	0.05	0.98	0.91	0.18	1.48	1.08	1.28	1.38	0.99
Na ₂ O	2.82	0.37	2.82	2.96	0.1	2.75	3.35	2.17	2.73	2.83
K ₂ O	3.17	6.68	5.17	4.68	6.28	4.55	4.08	4.39	4.18	3.10
P ₂ O ₅	0.21	0.19	0.21	0.22	0.19	0.22	0.19	0.25	0.18	0.15
Li	52	57	48	57	77	54	263	122	42	26
Rb	92.7	290	186	162	228	160	206	174	159	126
Cs	4.09	8.69	5.32	5.93	5.54	5.27	50.7	20.6	4.86	2.78
Be	4.23	2.72	2.44	2.62	2.86	2.66	2.30	2.91	2.22	2.71
Sr	207	42	113	166	16	167	125	129	136	168
Ba	1063	1098	931	909	656	707	753	757	712	863
Sc	20.0	10.9	8.21	8.37	9.50	7.40	8.77	9.38	7.13	6.39
V	129	67.2	51.2	52.8	57.1	48.6	53.0	63.3	61.3	49.4
Cr	106	62.2	43.4	42.7	41.8	46.7	41.8	46.7	57.1	47.7
Co	28.8	27.4	21.3	30.5	19.2	33.1	72.9	60.4	128	105
Ni	35.2	12.8	15.7	15.6	12.9	15.0	16.2	22.1	16.4	10.5
Cu	27.7	15.1	10.1	14.7	1.6	14.6	16.6	15.4	6.7	11.5
Zn	119	70.4	46.8	46.6	37	52.7	78.1	74.7	41.5	40.0
Ga	25.0	23.4	19.5	20.9	24.4	19.0	18.4	22.0	20.0	18.1
Y	36.2	40.9	34.0	33.6	34	32.5	17.8	21.6	28.2	19.0
Nb	14.3	13.3	10.0	10.3	12.2	9.23	10.0	11.3	10.5	8.4
Ta	1.06	1.29	0.93	0.96	1.08	0.87	0.98	1.17	1.14	0.90
Zr	290	242	191	165	209	179	182	217	162	134
Hf	8.50	7.70	6.00	5.20	6.60	5.50	5.00	5.00	5.19	3.60
Mo	0.95	0.30	0.73	0.69	0.66	0.64	0.22	0.63	0.56	0.40
Sn	2.03	7.19	5.53	4.95	6.29	4.88	3.23	4.6	4.07	2.91
Tl	0.52	1.74	1.06	0.97	1.08	0.94	1.16	0.97	0.92	0.62
Pb	22.9	18.9	26.4	29.4	6.56	28.1	48.3	18.9	22.1	21.7
U	2.15	4.56	3.72	2.86	3.94	3.60	2.73	3.12	3.42	3.22
Th	16.7	16.8	13.6	14.2	13.9	11.8	12.1	12.5	11.2	10.3
La	55.8	40.4	25.7	34.7	23.4	28	30.4	35.2	26.8	31.3
Ce	111	84.6	54.7	74.4	50.6	58.0	63.6	74.3	58.7	50.6
Pr	13.8	10.6	6.51	8.82	6.74	6.93	7.73	8.99	7.04	7.84
Nd	50.8	41.2	25.0	31.8	26.5	28.5	29.7	35.4	26.5	29.0
Sm	10.2	8.93	5.76	6.82	6.27	6.29	6.11	7.07	5.70	6.01
Eu	2.10	1.55	1.15	1.31	0.99	1.20	1.04	1.16	1.22	1.46
Gd	9.00	8.19	5.68	6.15	5.93	5.79	4.93	6.02	5.07	5.26
Tb	1.34	1.32	0.91	1.03	0.97	0.91	0.73	0.83	0.86	0.75
Dy	7.14	7.28	5.43	5.86	6.29	5.29	3.70	4.24	5.05	3.83
Ho	1.31	1.44	1.14	1.14	1.24	1.07	0.64	0.76	1.01	0.71
Er	3.28	3.39	2.96	3.12	3.25	2.95	1.49	1.76	2.55	1.76
Tm	0.48	0.48	0.46	0.45	0.48	0.43	0.21	0.25	0.37	0.23
Yb	2.95	2.89	2.87	2.61	2.83	2.52	1.15	1.5	2.14	1.44
Lu	0.46	0.46	0.40	0.41	0.42	0.37	0.17	0.21	0.33	0.22

Table 5.4: Major elements and trace elements of the metagranites of the Ollo de Sapo Domain.

Sample	T24	T25	T26	T27	T28	T29	T31	T33	T34	T35
Locality	Viana do Bolo		San Sebastián		Miranda do Douro		Antoñita			
SiO ₂	66.24	73.41	74.8	74.83	65.54	64.62	76	70.25	71.46	72.37
TiO ₂	0.79	0.29	0.23	0.22	0.67	0.62	0.09	0.43	0.32	0.35
Al ₂ O ₃	14.83	13.72	12.29	12.74	15.79	16.07	12.88	15.07	14.45	13.92
FeO _t	4.71	1.45	1.26	1.26	4.25	3.83	1.14	2.61	2.1	2.36
MgO	2.49	0.45	0.25	0.25	2.62	2.23	0.12	0.89	0.59	0.77
MnO	0.08	0.03	0.02	0.02	0.1	0.1	0.01	0.04	0.03	0.03
CaO	1.75	0.74	0.84	0.96	2.88	2.93	0.41	0.91	0.8	1.08
Na ₂ O	2.78	2.92	3.12	3.05	3.1	3.21	2.29	2.55	2.74	3.14
K ₂ O	3.43	5.25	5.75	5.54	3.36	4.03	5.83	5.39	5.27	4.28
P ₂ O ₅	0.25	0.17	0.19	0.18	0.35	0.35	0.152	0.196	0.174	0.173
Li	187	111	34	40	159	102	15	35	21	51
Rb	139	218	189	194	217	241	207	217	189	182
Cs	8.93	22.7	3.81	4.29	15.9	14.5	5.72	6.77	3.97	9.23
Be	2.51	5.06	3.76	4.09	5.01	5.91	0.92	2.26	2.06	2.36
Sr	169	53.4	103	135	307	339	25.9	123	101	82.9
Ba	665	359	227	240	799	904	59.4	724	615	459
Sc	17.4	3.78	2.16	2.07	9.96	9.89	1.77	6.00	4.21	3.96
V	116	16.6	10.4	10.3	66.5	65.7	9.15	43.0	30.9	32.4
Cr	83.7	5.39	13.4	12.2	107	83.0	20.6	46.0	32.4	35.1
Co	41.0	79.3	58.6	54.5	73.2	66.2	35.2	37.8	25.8	34.2
Ni	34.0	2.68	1.99	2.03	40.8	33.8	0.16	9.80	7.30	8.04
Cu	23.4	6.10	0.39	0.88	18.1	14.6	3.09	6.47	10.6	2.44
Zn	94.8	39.3	18.2	20.7	84.3	77.0	27.3	49.5	43.7	40.8
Ga	21.9	19	21.6	22.7	23.7	24	15.1	20.9	17.2	18
Y	18.3	18.4	21.1	21.2	22.1	24.0	13.5	18.1	25.0	16.1
Nb	14.3	9.52	20.4	21.4	12.9	11.8	6.65	10.6	7.89	7.95
Ta	1.22	1.47	1.56	1.82	1.36	1.23	0.91	1.10	0.90	0.59
Zr	203	136	151	152	248	232	68.4	194	136	151
Hf	5.70	4.30	5.10	5.20	5.40	5.70	2.63	6.05	4.31	4.97
Mo	1.00	0.15	2.36	3.57	0.73	0.78	0.11	0.37	0.32	0.36
Sn	5.36	14.3	1.80	1.86	6.89	7.45	7.49	8.43	6.20	7.14
Tl	0.79	1.31	0.77	0.75	1.54	1.47	1.14	1.2	1.1	1.03
Pb	18.3	20.6	13.6	14.1	32.0	40.8	25.1	23.4	23.9	22.0
U	2.40	2.56	3.57	6.08	3.07	8.29	2.22	3.96	3.44	3.21
Th	11.0	10.6	13.2	13.8	35.8	33.5	6.86	12.1	9.82	10.1
La	34.5	20.0	23.5	22.2	82.9	83.2	7.96	27.5	22.7	22.2
Ce	71.1	43.8	48.1	45.1	161	161	18.5	59.2	48.3	46.7
Pr	8.74	5.27	5.57	5.22	18.4	18.5	2.26	7.04	5.86	5.79
Nd	33.5	19.5	19.2	18.2	65.8	65.7	7.97	27.4	22.4	22.1
Sm	7.03	4.76	4.45	4.19	9.87	10.0	2.06	6.04	5.05	4.9
Eu	1.25	0.51	0.50	0.49	1.50	1.52	0.18	0.99	0.94	0.89
Gd	6.10	4.00	4.63	4.43	6.44	6.80	2.32	4.97	4.81	4.34
Tb	0.83	0.64	0.70	0.71	0.87	0.89	0.41	0.75	0.75	0.67
Dy	4.15	3.66	3.99	3.94	4.09	4.34	2.49	4.03	4.54	3.68
Ho	0.72	0.66	0.71	0.71	0.75	0.78	0.50	0.7	0.91	0.64
Er	1.64	1.62	1.69	1.68	1.88	2.09	1.40	1.60	2.30	1.50
Tm	0.21	0.23	0.23	0.23	0.27	0.28	0.21	0.22	0.34	0.19
Yb	1.23	1.4	1.33	1.39	1.47	1.74	1.34	1.30	2.01	1.18
Lu	0.17	0.21	0.17	0.19	0.22	0.24	0.20	0.20	0.30	0.17

9.2.2. The Schist-Graywacke Complex Domain

Table 5.5: Major elements and trace elements of the metagranites of the Tormes Dome.

Sample	T36	T37	T38	T39	T40	T41	T42
Locality	Vitigudino		Fermoselle			Ledesma	
SiO ₂	74.67	72.37	75.64	74.57	73.41	77	68.83
TiO ₂	0.23	0.24	0.16	0.27	0.12	0.1	0.56
Al ₂ O ₃	13.51	14.79	13.82	13.7	15.04	12.63	15.62
FeO _t	2.04	2.39	1.74	2.35	1.17	0.88	4.05
MgO	0.52	0.79	0.68	0.61	0.36	0.21	1.47
MnO	0.03	0.05	0.02	0.03	0.01	0.01	0.07
CaO	0.68	0.82	0.66	0.63	0.35	0.59	1.08
Na ₂ O	2.34	2.38	1.94	2.28	1.94	2.16	1.7
K ₂ O	4.82	4.44	4.06	4.07	6.27	5.49	3.77
P ₂ O ₅	0.17	0.16	0.173	0.14	0.11	0.129	0.192
Li	63	82	32	71	73	30	86
Rb	237	230	172	253	242	218	188
Cs	8.43	13.1	4.59	14.2	10.4	8.70	11.2
Be	0.56	0.78	0.75	2.23	2.03	2.73	3.47
Sr	46.3	30.8	54.8	54.1	43.6	51.4	108
Ba	202	120	208	188	224	166	504
Sc	4.74	5.30	2.31	5.23	4.51	2.02	9.00
V	19.3	20.9	11.1	19.3	13.8	6.00	60.1
Cr	11.1	19.5	5.01	17.7	13.1	1.01	40.3
Co	38.5	93.8	47.5	113	83.2	39.6	48.2
Ni	6.37	9.34	3.68	11.8	6.94	2.11	22.2
Cu	5.17	1.89	2.55	1.96	1.50	5.06	16.1
Zn	66.1	61.3	18.6	49.4	33.6	37.3	70.0
Ga	18.0	20.6	13.4	20.1	17.7	14.0	21.8
Y	11.4	9.57	11.5	12.0	8.35	7.52	17.4
Nb	9.42	10.9	3.89	8.34	5.89	5.86	11.9
Ta	1.07	1.42	0.39	1.31	0.71	0.52	1.06
Zr	104	114	86.5	119	63.3	73.6	198
Hf	3.60	2.34	3.10	4.50	2.50	2.70	6.20
Mo	0.08	0	0.09	0.49	0.00	0.10	0.51
Sn	4.95	5.24	3.68	5.22	5.92	3.79	5.44
Tl	1.33	1.33	1.00	1.50	1.44	1.21	1.06
Pb	23.0	18.7	18.3	21.7	21.7	33.1	20.0
U	3.89	5.15	1.86	4.95	2.20	2.41	3.94
Th	9.74	13.0	8.39	12.9	8.98	5.74	16.6
La	15.4	18.9	13.4	20.5	13.6	7.27	38.2
Ce	33.2	40.0	28.6	42.9	29.8	15.4	79.0
Pr	4.11	4.91	3.48	5.06	3.56	1.85	9.40
Nd	14.4	18.5	12.5	19.3	12.8	6.42	36.1
Sm	3.52	4.15	3.11	4.06	3.06	1.67	7.48
Eu	0.39	0.28	0.46	0.49	0.38	0.34	1.01
Gd	2.79	3.25	2.75	3.61	2.51	1.51	5.83
Tb	0.43	0.53	0.40	0.55	0.37	0.24	0.80
Dy	2.10	2.29	2.30	2.55	1.89	1.40	4.15
Ho	0.37	0.35	0.41	0.47	0.33	0.28	0.63
Er	0.89	0.87	0.95	1.04	0.80	0.76	1.38
Tm	0.13	0.11	0.13	0.14	0.11	0.13	0.18
Yb	0.81	0.69	0.81	0.86	0.78	0.86	0.96
Lu	0.12	0.09	0.12	0.14	0.13	0.15	0.13

Table 5.6: Major elements and trace elements of the metagranites of the Sierra de Gredos.

Sample	T43	T44	T45	T46	T47	T48	T49
Locality	San Pelayo	Castellanos		Bercimuelle			
SiO ₂	71.80	70.33	68.54	69.98	68.54	70.19	69.71
TiO ₂	0.38	0.36	0.59	0.46	0.39	0.42	0.45
Al ₂ O ₃	14.37	15.35	15.24	15.35	16.02	15.02	15.29
FeO _t	2.19	2.16	3.5	2.97	2.56	2.86	3.14
MgO	0.93	0.69	1.47	1.13	1.02	1.06	1.14
MnO	0.04	0.03	0.05	0.05	0.04	0.05	0.05
CaO	0.89	0.79	1.07	0.87	0.5	0.89	0.93
Na ₂ O	2.93	3.99	2.66	2.68	2.96	2.85	2.75
K ₂ O	4	4.7	4.26	4.42	5.18	4.54	4.43
P ₂ O ₅	0.23	0.2	0.15	0.22	0.26	0.25	0.23
Li	134	100	137	115	55	93	107
Rb	229	151	196	241	248	235	217
Cs	22.1	17.4	14.5	17.8	8.5	10.2	10.7
Be	1.32	1.55	2.62	2.00	1.74	1.89	1.78
Sr	76.9	78.1	116	83.2	75.2	84.2	91.5
Ba	341	421	570	429	479	414	424
Sc	7.97	4.47	10.7	7.26	7.49	7.77	7.55
V	39.0	29.7	67.2	44.9	41.3	46.3	48.7
Cr	31.8	25.9	51.4	36.0	33.9	41.3	39.5
Co	74.5	94.5	72.8	79.1	76.8	80.7	76.6
Ni	7.64	5.41	11.5	14.2	11.6	14.0	20.1
Cu	16.6	22	10.6	12.9	14.8	13.0	22.8
Zn	51.2	49.9	73.1	64.5	58.4	70.6	65.5
Ga	19.0	17.6	21.2	19.4	20.4	19.0	19.8
Y	14.6	9.64	14.9	13.7	14.1	13.2	13.0
Nb	10.2	8.48	11.5	10.8	11.2	11.2	10.9
Ta	1.32	0.86	1.25	1.44	1.90	1.31	1.20
Zr	122	134	191	142	136	131	138
Hf	3.90	4.20	4.90	4.50	4.40	4.20	4.30
Mo	0.32	0.35	0.77	0.43	0.25	0.59	0.52
Sn	13.2	2.38	6.38	7.05	15.9	4.53	5.25
Tl	1.30	0.87	1.15	1.52	1.40	1.42	1.28
Pb	21.0	28.4	25.0	23.7	24.7	25.5	24.9
U	4.22	3.14	3.91	7.15	4.25	6.11	7.89
Th	9.39	8.27	17.0	10.9	10.0	11.7	11.2
La	20.9	20.4	37.6	25.9	22.9	26.7	26.5
Ce	44.3	45.1	75.8	54.0	47.9	56.7	55.7
Pr	5.22	5.46	9.06	6.79	5.99	7.09	6.91
Nd	19.4	20.9	33.5	25.7	22.6	27.5	26.7
Sm	4.02	4.78	6.74	5.29	4.85	5.74	5.53
Eu	0.64	0.75	1.04	0.79	0.81	0.8	0.84
Gd	3.35	3.32	5.23	4.26	4.14	4.56	4.42
Tb	0.52	0.42	0.73	0.59	0.58	0.65	0.61
Dy	2.84	2.06	3.40	2.79	2.96	2.99	2.78
Ho	0.49	0.35	0.55	0.43	0.51	0.46	0.47
Er	1.22	0.88	1.26	1.01	1.30	1.06	1.10
Tm	0.17	0.12	0.17	0.14	0.18	0.15	0.15
Yb	1.01	0.73	1.03	0.84	1.09	0.94	0.87
Lu	0.15	0.11	0.15	0.12	0.16	0.13	0.13

Table 5.7: Major elements and trace elements of the metagranites of the Sierra de Guadarrama.

Sample	T51	T52	T53	T54	T56	T57	T58	T59	T60	T61	T63	T64
Locality	Navas del Marqués	Santa Maria de la Alameda	La Hoya	La Estación		La Estación II	Vegas de Matute (Melanocratic)		Vegas de Matute (Leucocratic)	La Morcuera	Buitrago de Lozoya	Prádena del Rincón
SiO ₂	72.30	71.05	74.23	69.46	73.72	71.80	62.67	64.41	75.45	70.16	69.61	75.59
TiO ₂	0.27	0.51	0.24	0.39	0.22	0.41	0.91	0.75	0.08	0.54	0.45	0.13
Al ₂ O ₃	14.77	14.13	13.12	15.83	13.54	13.71	16.67	16.79	13.64	14.58	15.29	13.32
FeO _t	2.11	3.42	1.63	3.19	1.82	2.85	6.47	4.91	1.34	3.44	3.22	1.4
MgO	0.87	0.76	0.39	1.35	0.59	0.65	2.44	2.14	0.09	1.00	1.19	0.23
MnO	0.03	0.04	0.02	0.0	0.02	0.0	0.10	0.07	0.03	0.04	0.04	0.02
CaO	0.63	1.24	0.91	1.06	0.53	0.98	1.48	1.46	0.34	0.97	1.07	0.46
Na ₂ O	2.25	2.64	2.31	2.33	1.98	2.29	1.98	2.12	2.8	2.32	2.23	2.53
K ₂ O	4.96	4.78	5.48	4.12	5.46	5.68	3.55	3.66	4.99	4.7	4.92	5.01
P ₂ O ₅	0.16	0.12	0.08	0.16	0.14	0.13	0.20	0.20	0.25	0.26	0.17	0.17
Li	70	79	34	101	38	76	68	94	164	89.7	77.5	64.0
Rb	213	222	165	191	184	248	111	129	511	158	185	270
Cs	10.1	11.7	3.71	7.24	3.71	7.44	5.82	5.42	10.8	10.3	8.76	5.72
Be	2.41	1.13	1.52	2.36	1.14	2.53	2.76	3.36	0.40	3.60	0.78	1.03
Sr	81.1	83.1	142	117	93.9	72.0	183	189	9.66	92.8	120	55.6
Ba	383	398	501	485	472	405	947	906	18.8	520	549	171
Sc	4.19	6.24	2.69	5.36	4.04	5.17	17.6	11.1	3.21	6.01	7.51	2.99
V	22.7	33.6	15.6	37.8	20.1	26.9	107	81.6	5.26	37.6	49.0	7.77
Cr	9.92	12.2	2.30	25.9	9.09	6.31	79.22	54.9	1.12	16.3	30.6	1.03
Co	78.0	153	104	89.8	83.7	104	81.8	59.7	95.5	95.0	69.2	81.4
Ni	18.5	5.10	17.8	10.7	134	3.19	33.6	23.9	0.31	6.23	31.66	1.47
Cu	6.30	3.57	1.86	14.0	17.6	5.71	36.5	25.5	2.38	12.9	13.2	2.18
Zn	33.7	58.8	29.3	59.9	37.7	47.6	105.5	91.3	23.3	70.3	65.2	35.4
Ga	19.6	21.9	16.0	21.7	16.8	20.6	23	22.7	24.0	20.7	19.9	18.1
Y	13.0	25.2	16.5	21.5	18.5	24.0	29	27.2	10.1	49.5	16.3	15.6
Nb	7.11	21.4	9.24	8.82	6.28	19.6	13.6	12.9	12.2	15.4	9.60	7.16
Ta	0.98	1.64	0.64	0.76	0.69	1.76	1.16	0.97	1.65	1.76	0.85	0.90
Zr	108	326	198	156	101	278	287	246	54.1	298	164	69.6
Hf	3.50	8.20	6.40	4.90	3.30	7.80	8.60	7.30	2.40	8.40	4.90	2.60
Mo	0.28	0.48	0.25	0.46	0.12	0.37	1.24	1.12	0.17	0.29	0.59	0.24
Sn	5.88	3.40	2.80	3.34	5.07	5.66	3.26	3.44	7.12	8.1	4.8	13.4
Tl	1.10	1.10	0.74	0.98	1.03	1.23	0.62	0.80	2.47	0.82	0.98	1.46
Pb	27.8	24.8	38.7	22.9	28.6	23.4	24.0	23.3	11.7	23.7	28.4	19.9
U	3.34	3.37	2.68	2.48	2.53	3.40	2.9	2.26	2.35	2.86	3.30	10.0
Th	10.0	28.3	23.8	12.4	8.4	24.5	16.7	16.3	8.10	23.8	12.7	7.00
La	19.4	57.1	40.5	31.0	15.9	46.8	48.3	46.3	4.90	47.9	29.7	8.18
Ce	40.3	118.3	83.6	62.7	33.3	95.2	100	94.7	12.5	101	62.8	19.1
Pr	4.85	13.4	9.56	7.54	4.08	10.9	12.2	11.4	1.54	12.44	7.54	2.29
Nd	18.0	47.5	32.9	28.2	15.5	39.1	46.7	43.6	5.31	46.04	28.66	8.59
Sm	3.93	9.49	6.60	6.09	3.48	8.10	9.44	9.09	1.48	10.0	5.80	2.26
Eu	0.66	0.98	1.14	0.91	0.65	0.77	1.46	1.50	0.03	0.81	0.83	0.21
Gd	3.47	7.97	5.85	4.86	3.16	6.83	8.38	7.70	1.22	8.82	5.15	2.25
Tb	0.55	1.16	0.83	0.74	0.53	1.04	1.13	1.08	0.26	1.39	0.72	0.42
Dy	2.82	6.0	4.07	4.15	3.25	5.22	5.96	5.67	1.71	8.62	3.60	2.75
Ho	0.49	0.92	0.60	0.71	0.64	0.86	1.04	0.90	0.31	1.61	0.60	0.55
Er	1.19	1.92	1.10	1.62	1.79	2.07	2.44	2.00	0.82	4.00	1.31	1.52
Tm	0.16	0.23	0.14	0.22	0.28	0.27	0.35	0.23	0.13	0.53	0.17	0.21
Yb	0.97	1.41	0.84	1.37	1.70	1.44	1.96	1.31	0.79	3.15	1.02	1.19
Lu	0.13	0.19	0.12	0.20	0.25	0.19	0.27	0.17	0.11	0.40	0.16	0.18

Table 5.8: Major elements and trace elements of the metavolcanic rocks of the Sierra de Guadarrama (Continuation).

Sample	T66	T67	T68	T69	T70	T71
Locality	La Berzosa		El Cardoso			Riaza
SiO ₂	72.51	76.37	66.58	65.72	65.1	72.75
TiO ₂	0.35	0.05	0.80	0.79	0.79	0.29
Al ₂ O ₃	14.7	13.01	15.96	16.48	16.24	13.9
FeO _t	2.54	1.2	4.97	5.05	5.14	2.28
MgO	0.64	0.09	1.84	1.9	2.21	0.74
MnO	0.03	0.06	0.07	0.06	0.07	0.03
CaO	0.67	0.40	1.56	1.66	1.25	0.75
Na ₂ O	2.04	2.14	2.32	2.36	2.48	2.23
K ₂ O	4.5	5.94	3.74	3.87	3.95	4.64
P ₂ O ₅	0.16	0.17	0.20	0.20	0.20	0.15
Li	55.2	19.28	58.7	58.7	55	40.4
Rb	147	245	154	155	145	229
Cs	7.28	4.21	5.47	5.11	4.42	11.4
Be	2.64	1.32	2.35	2.36	2.13	1.74
Sr	159	65.0	237	223	168	73.7
Ba	688	776	925	832	773	250
Sc	14.4	4.52	13.3	12.9	13.7	4.45
V	99.1	17.7	85.8	84.6	91.3	28.4
Cr	80.6	15.9	61.1	58.0	69.2	15.8
Co	79.3	84.6	42.4	40.0	32.4	88.4
Ni	24.3	5.22	42.6	25.1	28.9	7.89
Cu	16.9	4.14	24.7	20.1	24.4	2.49
Zn	80.4	27.1	93.0	93.0	87.9	43.9
Ga	21.8	17.1	22.5	22.7	22.4	18.9
Y	14.3	12.8	26.8	30.4	19.0	15.2
Nb	12.3	6.35	12.6	12.7	12.0	7.63
Ta	1.11	1.19	1.01	1.00	1.04	0.97
Zr	170	59.0	227	225	219	116
Hf	5.20	2.40	6.40	6.30	6.10	3.70
Mo	0.37	0.25	0.80	1.22	0.92	0.34
Sn	5.88	13.0	4.64	4.75	4.44	7.83
Tl	0.90	1.35	0.86	0.91	0.83	1.26
Pb	21.6	27.8	22.3	18.5	19.7	19.0
U	3.59	5.82	3.51	3.41	3.39	4.04
Th	13.4	8.33	13.0	12.8	12.3	9.00
La	35.7	14.3	46.5	42.5	37.0	18.7
Ce	75.0	29.5	94.2	86.4	76.1	39.6
Pr	9.22	3.60	11.6	10.5	9.09	4.76
Nd	35.2	13.6	43.4	40.0	34.9	18.2
Sm	6.41	2.98	9.25	8.44	7.28	3.84
Eu	1.36	0.58	1.29	1.26	1.18	0.51
Gd	5.76	2.58	7.29	7.81	6.06	3.34
Tb	0.85	0.45	1.13	1.11	0.85	0.51
Dy	3.38	2.32	6.20	6.24	4.03	2.97
Ho	0.63	0.45	1.03	1.04	0.61	0.51
Er	1.28	1.22	2.56	2.57	1.48	1.24
Tm	0.16	0.21	0.33	0.35	0.21	0.18
Yb	0.99	1.15	2.09	2.27	1.34	1.18
Lu	0.14	0.19	0.33	0.30	0.23	0.18

Table 5.9: Major elements and trace elements of the metagranite of the Anatectic Complex of Toledo.

Sample	T74
Locality	Mohares
SiO ₂	72.91
TiO ₂	0.68
Al ₂ O ₃	12.50
FeO _t	3.60
MgO	0.94
MnO	0.03
CaO	1.44
Na ₂ O	2.29
K ₂ O	3.59
P ₂ O ₅	0.19
Li	46
Rb	185
Cs	2.46
Be	0.91
Sr	81.9
Ba	478
Sc	6.36
V	47.6
Cr	15.0
Co	99.8
Ni	8.57
Cu	12.0
Zn	76.3
Ga	21.2
Y	35.3
Nb	27.0
Ta	1.39
Zr	406
Hf	12.8
Mo	1.47
Sn	0.79
Tl	1.00
Pb	21.8
U	4.27
Th	44.5
La	77.4
Ce	158
Pr	18.4
Nd	66.9
Sm	13.6
Eu	1.21
Gd	11.1
Tb	1.64
Dy	8.39
Ho	1.44
Er	3.38
Tm	0.45
Yb	2.83
Lu	0.42

9.3. Table of Sr and Nd isotopic data

Table 5.10: Sr and Nd isotopes of the: MROS, metavolcanic rocks of the Ollo de Sapo Domain; MOS, metagranites of the Ollo de Sapo Domain; MTD, metagranites of the Tormes Dome; MGR, metagranites of the Sierra de Gredos; MGU, metagranites and metavolcanic rocks of the Sierra de Guadarrama

Sample	Group	$(^{87}\text{Sr}/^{86}\text{Sr})_{\text{now}}$	$(^{143}\text{Nd}/^{144}\text{Nd})_{\text{now}}$	$(^{87}\text{Sr}/^{86}\text{Sr})_{490\text{Ma}}$	$(\epsilon\text{Sr})_{490\text{Ma}}$	$(^{143}\text{Nd}/^{144}\text{Nd})_{490\text{Ma}}$	$(\epsilon\text{Nd})_{490\text{Ma}}$	$\text{Nd}_{(\text{CHUR})}$ (Ga)	$T_{(\text{DM})}$ (Ga)
T-2	MROS	0.72983	0.512172	0.70789	56	0.511765	-4.7	1.0	1.7
T-4	MROS	0.73008	0.512208	0.7117	111	0.511789	-4.2	1.0	1.7
T-5	MROS	0.73142	0.51217	0.7106	95	0.511750	-5.0	1.1	1.8
T-6	MROS	0.73230	0.512203	0.71014	88	0.511786	-4.3	1.0	1.7
T-7	MROS	0.73521	0.512258	0.70904	73	0.51184	-3.3	0.9	1.6
T-8	MROS	0.74630	0.512206	0.71104	101	0.511795	-4.1	1.0	1.7
T-10	MROS	0.73389	0.512183	0.70920	75	0.511762	-4.8	1.1	1.8
T-12	MROS	0.71901	0.512177	0.70997	86	0.511787	-4.3	0.9	1.6
T-15	MROS	0.74444	0.512159	0.70930	76	0.511712	-5.8	1.3	2.0
T-16	MROS	0.73305	0.512126	0.7133	133	0.511710	-5.8	1.2	1.9
T-18	MROS	0.73018	0.512148	0.71085	98	0.511720	-5.6	1.2	1.9
T-19	MROS	0.73857	0.512199	0.70719	46	0.511785	-4.3	1.0	1.7
T-20	MROS	0.73302	0.512174	0.70668	39	0.511755	-5	1.1	1.8
T-21	MROS	0.73272	0.512219	0.71067	96	0.511800	-4.0	1.0	1.7
T-22	MROS	0.72429	0.512214	0.71002	87	0.511807	-3.9	0.9	1.6
T-24	MOS	0.72535	0.512209	0.70876	69	0.511802	-4.0	0.9	1.7
T-25	MOS	0.74058	0.512565	0.70353	-5.6	0.512115	2.1	0.2	1.2
T-26	MOS	0.73273	0.512463	0.70371	-3.0	0.51202	0.20	0.5	1.4
T-27	MOS	0.7999	0.512258	0.7173	190	0.511785	-4.3	1.2	2.0
T-28	MOS	0.71892	0.512136	0.70502	16	0.511832	-3.4	0.8	1.3
T-29	MOS	0.7179	0.512232	0.70424	4.5	0.511918	-1.7	0.6	1.2
T-31	MOS	0.86098	0.512407	0.70905	73	0.511884	-2.4	1.0	2.2
T-33	MOS	0.74329	0.512268	0.70963	81	0.511840	-3.3	0.9	1.7
T-34	MOS	0.7479	0.512244	0.71004	87	0.51181	-3.9	1.0	1.8
T-35	MOS	0.75328	0.512307	0.71055	94	0.511872	-2.6	0.8	1.6
T-37	MTD	0.84616	0.512278	0.71500	157	0.511834	-3.4	0.9	1.8
T-39	MTD	0.79199	0.512152	0.70683	41	0.511707	-5.9	1.3	2.0
T-45	MGR	0.77014	0.512208	0.70678	41	0.511778	-4.5	1.0	1.8
T-46	MGR	0.76471	0.512264	0.71133	105	0.511832	-3.4	0.9	1.7
T-48	MGR	0.74529	0.512276	0.70837	63	0.511844	-3.2	0.9	1.7
T-51	MGU	0.76120	0.512209	0.70616	32	0.511781	-4.4	1.0	1.8
T-53	MGU	0.74091	0.512246	0.71663	180.5	0.511850	-3.1	0.8	1.5
T-56	MGU	0.75353	0.512231	0.71134	105	0.511788	-4.3	1.1	1.9
T-59	MGU	0.72419	0.512212	0.71011	88	0.511816	-3.7	0.9	1.6
T-61	MGU	0.74710	0.512268	0.71122	104	0.511839	-3.3	0.9	1.7
T-63	MGU	0.74376	0.512188	0.71125	104	0.511779	-4.4	1.0	1.7
T-64	MGU	0.81822	0.512169	0.71519	160	0.511607	-7.8	3.3	3.8
T-66	MGU	0.72812	0.512179	0.71031	91	0.511790	-4.2	0.9	1.6
T-67	MGU	0.78285	0.512191	0.71393	142	0.511747	-5.1	1.2	1.9
T-68	MGU	0.72393	0.512185	0.71051	94	0.511788	-4.3	0.9	1.6
T-70	MGU	0.72758	0.512166	0.70954	80	0.511760	-4.8	1.0	1.7
T-71	MGU	0.7746	0.512212	0.70961	81	0.511782	-4.4	1.0	1.8