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4 **Authors:** R. Román, J.A. Benavent-Oltra, J.A. Casquero-
5 Vera, A. Lopatin, A. Cazorla, H. Lyamani, C. Denjean, D.
6 Fuertes, D. Pérez-Ramírez, B. Torres, C. Toledano, O.
7 Dubovik, V.E. Cachorro, A.M. de Frutos, F.J. Olmo, and
8 L. Alados-Arboledas

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23 **Retrieval of aerosol profiles combining sunphotometer**
24 **and ceilometer measurements in GRASP code**

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26 **R. Román^{1,2,3}, J.A. Benavent-Oltra^{1,3}, J.A. Casquero-Vera^{1,3}, A. Lopatin⁴, A.**
27 **Cazorla^{1,3}, H. Lyamani^{1,3}, C. Denjean⁵, D. Fuertes⁴, D. Pérez-Ramírez^{1,3}, B.**
28 **Torres⁴, C. Toledano², O. Dubovik^{4,6}, V.E. Cachorro², A.M. de Frutos², F.J.**
29 **Olmo^{1,3}, and L. Alados-Arboledas^{1,3}**

30 [1]{Department of Applied Physics, University of Granada. 18071, Granada (Spain)}

31 [2]{Grupo de Óptica Atmosférica (GOA), Universidad de Valladolid. Paseo Belén, 7,
32 47011, Valladolid (Spain)}

33 [3]{Andalusian Institute for Earth System Research (IISTA-CEAMA), University of
34 Granada, Autonomous Government of Andalusia. 18006, Granada (Spain)}

35 [4]{GRASP-SAS, Remote sensing developments, LOA / Université Lille-1, Villeneuve
36 d'Ascq, France}

37 [5]{CNRM, Centre National de la Recherche Météorologique (UMR3589, CNRS,
38 Météo-France), Toulouse, France}

39 [6]{Laboratoire d'Optique Atmosphérique, Université de Lille 1, Villeneuve d'Ascq,
40 France}

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50 **Abstract**

51 In this paper we present an approach for the profiling of aerosol microphysical and
52 optical properties combining ceilometer and sun/sky photometer measurements in the
53 GRASP code (General Retrieval of Aerosol and Surface Properties). For this objective,
54 GRASP is used with sun/sky photometer measurements of aerosol optical depth (AOD)
55 and sky radiances, both at four wavelengths and obtained from AErosol RObotic
56 NETwork (AERONET), and ceilometer measurements of range corrected signal (RCS)
57 at 1064 nm. A sensitivity study with synthetic data evidences the capability of the
58 method to retrieve aerosol properties such as size distribution and profiles of volume
59 concentration (VC), especially for coarse particles. Aerosol properties obtained by the
60 mentioned method are compared with airborne in-situ measurements acquired during
61 two flights over Granada (Spain) within the framework of ChArMEx/ADRIMED
62 (Chemistry-Aerosol Mediterranean Experiment/Aerosol Direct Radiative Impact on the
63 regional climate in the MEDiterranean region) 2013 campaign. The retrieved aerosol
64 VC profiles agree well with the airborne measurements, showing a mean bias error
65 (MBE) and a mean absolute bias error (MABE) of $0.3 \mu\text{m}^3/\text{cm}^3$ (12%) and $5.8 \mu\text{m}^3/\text{cm}^3$
66 (25%), respectively. The differences between retrieved VC and airborne in-situ
67 measurements are within the uncertainty of GRASP retrievals. In addition, the retrieved
68 VC at 2500 m a.s.l. is shown and compared with in-situ measurements obtained during
69 summer 2016 at a high-atititude mountain station in the framework of the SLOPE I
70 campaign (Sierra Nevada Lidar AerOsol Profiling Experiment). VC from GRASP
71 presents high correlation ($r=0.91$) with the in-situ measurements, but overestimates
72 them, MBE and MABE being equal to 23% and 43%.

73 **Keywords**

74 GRASP, ceilometer, aerosol, profiling, photometer, aerosol volume concentration.

75 **1 Introduction**

76 Aerosols are a key piece in the Earth climatic system because they can increase the
77 cooling or warming of the Earth surface depending on their properties (Boucher et al.,
78 2013). Hence, columnar and vertical aerosol properties must be appropriately known to
79 better understand their impact in the Earth energy balance and therefore on the Earth
80 climate. Furthermore aerosol profiling is also relevant in the management of aviation
81 traffic (Prata, 2009; Flentje et al., 2010).

82 Column-integrated microphysical and optical aerosol properties are commonly
83 retrieved by sun/sky photometer measurements. This is the case of AERONET
84 (AErosol RObotic NETwork; Holben et al., 1998), that derives aerosol optical depth
85 (AOD) from multiwavelength measurements of direct beam sun irradiance, and uses
86 these AOD values in combination with sky radiances measurements for obtaining
87 aerosol properties such as aerosol size distribution, refractive indices, single scattering
88 albedo (SSA), and phase function (Dubovik and King, 2000; Dubovik et al., 2006).
89 However, this kind of measurements does not provide information about the vertical
90 profile of these aerosol properties.

91 Lidar systems are capable of measuring the atmospheric backscatter profile at
92 several wavelengths. The lidar signals are used for profiling optical and even retrieving
93 microphysical aerosol properties applying different methods. These methods depend on
94 the available lidar signals: elastic range corrected signal (RCS) is useful to provide
95 aerosol backscatter (β) profiles (Klett, 1981, 1985; Fernald, 1984; Sasano, 1984); non-
96 elastic (Raman) signal can be used for obtaining independent range-resolved extinction
97 (α) and backscatter coefficients (Ansmann et al., 1990; Whiteman et al., 1992). Elastic
98 and Raman lidar signals can be combined, usually by the so called $3\beta+2\alpha$ configuration,
99 to obtain profiles of aerosol microphysical properties through different inversion

100 techniques (e.g. Müller et al., 1999; Böckmann, 2001; Veselovskii et al., 2002, 2012;
101 Chemyakin et al., 2016).; many papers being already published for characterizing long-
102 transport of biomass-burning (e.g. Veselovskii et al., 2015; Ortiz-Amezcuca et al., 2017),
103 volcanic aerosol (e.g. Navas-Guzmán et al., 2013), dust (e.g. Granados-Muñoz et al.,
104 2016; Veselovskii et al., 2017) pollution (e.g. Wandinger et al., 2002; Noh et al., 2009;
105 Veselovskii et al., 2013), and arctic haze (Müller et al., 2004). In addition, linear particle
106 depolarization ratio measurements allow the detection and assessment of non-spherical
107 particles such as dust or volcanic aerosol (e.g. Ansmann et al., 2009, 2012; Tesche et
108 al., 2009, 2011; Bravo-Aranda et al., 2013) and allows aerosol typing (e.g. Burton et al.,
109 2012; Gross et al., 2013).

110 EARLINET (European Aerosol Research LIdar NETwork; Pappalardo et al.,
111 2014), founded in 2000 and now part of ACTRIS (Aerosols, Clouds, and Trace gases
112 Research InfraStructure ; www.actris.eu/), does include nowadays 31 lidar stations,
113 most of them operating multiwavelength Raman lidars. However, most Raman
114 measurements are sparse and mostly limited to night-time. To retrieve vertical profiles
115 of aerosol microphysics, several inversion techniques were developed within
116 EARLINET/ACTRIS combining backscattering lidar and collocated AERONET
117 sun/sky photometers such as LIRIC (Lidar Radiometer Inversion Code; Chaikovsky et
118 al., 2008, 2016) and GARRLiC (Generalized Aerosol Retrieval from Radiometer and
119 Lidar Combined data; Lopatin et al., 2013). The LIRIC code uses AERONET column-
120 integrated retrievals plus backscattering lidar signals as inputs to provide vertical-
121 resolved aerosol volume concentration (VC), both at fine and coarse mode. However,
122 GARRLiC uses as inputs measured optical depth and sky radiances and the
123 multiwavelength RCS from lidar to provide vertical-resolved aerosol microphysical and
124 optical properties, both at fine and coarse mode, and also improves the classical

125 AERONET columnar retrievals by providing intensive aerosol properties, like refractive
126 indices or SSA, of fine and coarse modes, separately.

127 The Generalized Retrieval of Aerosol and Surface Properties (GRASP; Dubovik
128 et al., 2014) code uses the heritage of AERONET inversion scheme (e.g. Dubovik and
129 King, 2000; Dubovik et al., 2006) and is a versatile and open-source algorithm capable
130 to obtain optical and microphysical aerosol properties from different sources of
131 measurements (www.grasp-open.com). Recently, aerosol properties have been retrieved
132 by GRASP using, among other information sources, satellite images (Kokhanovsky et
133 al., 2015), polar nephelometer data (Espinosa et al., 2017) and different combinations
134 with sun/sky photometer measurements: only spectral AODs (Torres et al., 2017);
135 spectral AODs, sky radiances and polarized sky radiances (Fedarenka et al., 2016); and
136 spectral AODs and sky camera images (Román et al., 2017a). The incorporation of the
137 GARRLiC scheme in GRASP allows to combine AODs, sky radiances and RCS lidar
138 values to retrieve columnar and vertical-resolved aerosol properties discerning between
139 fine and coarse modes (Lopatin et al., 2013; Bovchaliuk et al., 2016; Benavent-Oltra et
140 al., 2017).

141 Although the combination of lidar and sun/sky photometer measurements using
142 GRASP with the GARRLiC scheme is promising, lidar systems are generally expensive
143 and require supervision, so few stations have the set of measurements required to this
144 end. An alternative to multiwavelength lidar systems could be the use of ceilometers,
145 which were originally designed for studying cloud heights but recent ceilometer models
146 are able to detect aerosol layers at altitudes of up to 10 km. Ceilometers only measure at
147 one wavelength and are less accurate than classic lidars, but they are cheaper and more
148 operative than multiwavelength lidar systems and they also can work continuously
149 unattended. In fact, ceilometers have been previously used to obtain aerosol properties

150 as PM_{2.5} (Li et al., 2017), PM₁₀ (Münkel et al., 2007), aerosol backscatter coefficients
151 (Heese et al., 2010; Wiegner and Geiss, 2012; Wiegner et al., 2014; Madonna et al.,
152 2015) or aerosol hygroscopic growth (Haefelin et al., 2016). Moreover, there are some
153 programs nowadays as E-PROFILE, a program of EUMETNET (EUropean
154 METeorological services NETwork), and the COST Action ES1303 TOPROF
155 (TOwards operational ground based PROFiling with ceilometers, doppler lidars and
156 microwave radiometers for improving weather forecasts) dealing with the
157 harmonization and better characterization of ceilometer measurements and products;
158 and there are also ceilometer networks, like the Iberian CEilometer NETwork
159 (ICENET; Cazorla et al., 2017) among others (e.g., de Haij et al., 2007; Emeis et al.,
160 2011), trying to provide ceilometer measurements in near-real time with devices every
161 100 km. These issues motivate to try to combine ceilometer measurements with sun/sky
162 photometer in order to obtain some vertical aerosol information.

163 The main objective of this work is use for the first time the GRASP code to
164 obtain aerosol vertical profiling of aerosol microphysical properties combining
165 AERONET sun/sky photometer measurements with the monochromatic RCS measured
166 by a ceilometer at 1064 nm. The use of this proposed combination of measurements
167 allows the retrievals of column-integrated aerosol microphysical properties, and we
168 explore the possibility of obtaining vertically-resolved aerosol volume concentration.
169 Another important goal is the quantification of the accuracy and uncertainty of all
170 retrieved parameters through synthetic data and also by comparisons of retrieved
171 parameters versus in-situ measurements

172 This paper is structured as follows: Section 2 describes the used instrumentation
173 during the different measurement field campaigns; Section 3 introduces the GRASP
174 code and the methodology to retrieve the aerosol properties; a sensitivity study with

175 synthetic measurements is developed in Section 4 in order to test the capability of the
176 proposed GRASP scheme. Section 5 shows the main results about the comparison of the
177 obtained aerosol retrievals against in-situ measurements and, finally, the main
178 conclusions are summarized in Section 6.

179

180 **2 Instrumentation and campaigns**

181 **2.1 Instrumentation at Granada station**

182 Most of the instrumentation used in this work is installed on the rooftop of the
183 “Andalusian Institute for Earth System Research” (IISTA-CEAMA) building at
184 Granada, Spain (37.1638° N; 3.6051° W; 680 m a.s.l.). This instrumentation is managed
185 by the Atmospheric Physics Group (“Grupo de Física de la Atmósfera”; GFAT) of
186 University of Granada. Granada is a Spanish city located in the South-Eastern of the
187 Iberian Peninsula, in a natural basin surrounded by –Sierra Nevada Mountains with
188 peaks of up to 3300 m a.s.l., showing a Mediterranean climate (Csa in Köppen
189 classification). The city is medium-size with a population about 235000 inhabitants,
190 which increases up to 530000 including the metropolitan area, and non-industrialized
191 being its main aerosol sources the domestic heating based on fuel oil combustion in
192 winter and the heavy traffic along all year (Lyamani et al., 2010, 2011; Titos et al.,
193 2012, 2014). Columnar aerosol pattern in the area is characterized by higher values in
194 summer mostly associated with Saharan dust arrivals (Pérez-Ramírez et al., 2012;
195 Mandija et al., 2016), while the lowest aerosol loads usually corresponds to the arrivals
196 of Atlantic air-masses that clean the atmosphere (Pérez-Ramírez et al., 2016).

197 A CE318-T sun/sky/lunar (triple) photometer (*Cimel Electronique*) is operative on
198 the mentioned station since March 2016 for providing day and night columnar aerosol

199 optical properties (Barreto et al., 2013, 2016). GFAT also operates different sun/sky
200 photometers (hereafter ‘sunphotometers’) which belong to AERONET and have
201 participated in field campaigns in Spain, Brazil, Colombia and Bolivia, and have
202 allowed continuous operation of the site in Granada since the end of 2004. Both
203 sunphotometer models take measurements of direct beam sun irradiance, which retrieve
204 AOD, and sky radiance at several wavelengths, but only the channels of 440, 675, 870
205 and 1020 nm are chosen in this work because they are available in most AERONET
206 sunphotometers. All sunphotometer data used have been obtained from version 2 of
207 AERONET as level 1.5 data. Level 1.5 data are cloud-screened and have been chosen
208 instead of quality assurance level 2.0 data due to the near-real time availability of these
209 data, which can be used to calculate also other products in near-real time.

210 The mentioned Granada station also includes a “CHM-15k Nimbus” ceilometer
211 (*Lufft manufacturer*), which belongs to ICENET (Cazorla et al., 2017) and is detailed in
212 Román et al. (2017b). This instrument works as a one-wavelength lidar which emits at
213 1064 nm (a pulsed Nd:YAG laser) and measures the backscattered signal by the
214 atmosphere at different heights (up to 15360 m a.g.l.) with 15 m resolution. According
215 to the overlap function provided by the manufacturer, the overlap is 90% complete
216 between 555 and 885 m a.g.l. (Cazorla et al., 2017). The firmware of the instrument
217 directly provides NetCDF files with the RCS at 1064 nm which includes background
218 and overlap corrections. In addition, these files include the cloud base height (CBH)
219 product, which is estimated from ceilometer measurements due to the strong
220 backscattered signal of clouds (Martucci et al., 2010). The data are recorded as time
221 averaged data every 15 seconds. More information about this ceilometer and its
222 products can be found in the Jenoptik CHM15k user manual (Jenoptik, 2013).

223 2.2 ChArMEx/ADRIMED 2013

224 One of the main objectives of the ChArMEx/ADRIMED campaign (Chemistry-
225 Aerosol Mediterranean Experiment/Aerosol Direct Radiative Impact on the regional
226 climate in the MEDiterranean region) during summer 2013 was to conduct an
227 experimental campaign, based on surface and aircraft observations, for creating a rich 3-
228 D database of physical, chemical and optical properties of the main Mediterranean
229 aerosols (Mallet et al., 2016). To this end, 16 flights, ascending or descending in a spiral
230 trajectory during 30 min, were performed over the Mediterranean Basin with the ATR-
231 42 aircraft of SAFIRE (French aircraft service for environmental research;
232 <http://www.safire.fr>) during the period from 14th June to 4th July 2013 (Mallet et al.
233 2016; Denjean et al., 2016). The two flights named F30 and F31 of this campaign were
234 done over Granada city on 16th and 17th June 2013, respectively.

235 In both flights the ATR-42 airplane was equipped with different in-situ
236 instrumentation, being used in this work the measurements of fine and coarse aerosol
237 concentrations. For the aerosol concentration measurements in the submicron range: an
238 UHSAS (Ultra-High Sensitivity Aerosol Spectrometer; *Droplet Measurement*
239 *Technologies*) and a SMPS (Scanning Mobility Particle Sizer) with an accuracy of 10%
240 (Cai et al., 2008) and 5% (Wiedensohler et al., 2012), respectively. For coarse particles
241 the optical size distributions was measured by a FSSP-300 (a wing-mounted Forward
242 Scattering Spectrometer Probe, model 300 from *Particle Measuring Systems*) and by the
243 in-cabin GRIMM OPC (sky-optical particle counter; model 1.129 from *Grimm Technik*)
244 in the diameter nominal size ranges of 0.28-20 μm and 0.25-32 μm , respectively. FSSP-
245 300 and GRIMM have an accuracy of 30% (Baumgardner et al., 1992) and 10%
246 (Denjean et al., 2016), respectively. Finally, the profiles of the total aerosol VC (for
247 radius ranging between 0.05 and 15 μm) have been obtained with a resolution of 100 m

248 as in Benavent-Oltra et al. (2017): combining all the measurements of aerosol number
249 size distributions (SMPS, UHSAS, FSSP-300 and GRIMM OPC) and assuming that
250 aerosol particles are spherical (Denjean et al., 2016).

251 **2.3 SLOPE I**

252 The SLOPE I campaign (Sierra Nevada Lidar AerOsol Profiling Experiment) was
253 designed in order to measure relevant data for testing different retrieval schemes of
254 aerosol microphysical and optical vertical-profiles from remote sensing observations.
255 The campaign, developed during summer 2016, combined active and passive remote
256 sensing of the vertical column with in-situ measurements at several levels in the
257 northwestern slope of Sierra Nevada mountain range (Spain). In this framework, a new
258 measurement (SNS: Sierra Nevada Station) was set up in a high-altitude site at Sierra
259 Nevada (37.0958° N; 3.3869° W; 2500 m a.s.l.). This new station is 20 km far from
260 IISTA-CEAMA in horizontal distance and it was equipped with aerosol in-situ
261 instrumentation since May 2016, providing 24-hour aerosol in-situ measurements such
262 as scattering, absorption and extinction coefficients.

263 The in-situ aerosol volume concentration at SNS has been calculated combining
264 SMPS (model 3938 from *TSI Inc.*) and APS (Aerodynamic Particle Sizer; model 3321
265 from *TSI Inc.*) measurements. This volume concentration has been obtained in the 0.05-
266 10 μm radius range with 5 minutes time resolution. For that, Q-value = 1 is assumed for
267 conversion from aerodynamic (APS) to mobility size distribution (Sorribas et al., 2015).

268

269 **3 GRASP retrieval**

270 **3.1 Inputs**

271 **3.1.1 Sun/sky photometer data**

272 CE318 sunphotometers are configured to take a sequence of sky radiance
273 measurements in the almucantar plane (zenith angle equal to solar zenith angle, SZA)
274 for several air masses. AERONET provides the sky radiance usually at the next
275 almucantar azimuth angles (relative to sun): 2°, 2.5°, 3°, 3.5°, 4°, 5°, 6°, 7°, 8°, 10°, 12°,
276 14°, 16°, 18°, 20°, 25°, 30°, 35°, 40°, 45°, 50°, 60°, 70°, 80°, 90°, 100°, 120°, 140°, 160°
277 and 180°. These angles are scanned clockwise and counter clockwise giving two
278 measurements for each angle of symmetric points with respect to the sun position. In
279 this work the sky radiance has been averaged between both points. The azimuth angles
280 below 3.5° are rejected following the same criteria than the version 1 level 1.5 of
281 AERONET (Holben et al., 2006). The angles showing differences above 20% between
282 both almucantar branches are assumed as cloud contaminated and are also discarded as
283 in level 1.5 of AERONET version 2 (Holben et al., 2006). The azimuth at 180° does not
284 have a symmetric point which makes difficult its cloud-screening, and hence this angle
285 is also rejected. These criteria provide, in the most favourable case, 26 sky radiance
286 values at the four channels at 440 nm, 675 nm, 870 nm and 1020 nm.

287 After cloud-screening, the scattering angle criterion of Holben et al. (2006) for
288 AERONET (version 2 level 1.5) is applied. This criterion considers that sky radiance
289 distribution for each wavelength is representative if there is at least one measurement in
290 four regions identified by the scattering angle: $\geq 3.2^\circ$ to 6° ; $\geq 6^\circ$ to 30° ; $\geq 30^\circ$ to 80° ; and
291 $\geq 80^\circ$. In this work the scattering angle of 80° has been replaced in these bins by 78° in
292 order to use almucantars with SZA up to 40° .

293 The GRASP retrievals are done for each available cloud-screened almucantar if
294 it satisfies: (1) the number of sky radiance points at each wavelength is higher or equal
295 than 10 (as in AERONET version 2 level 1.5); (2) at each wavelength there is at least
296 one radiance value at the four mentioned bins, and (3) the closest AOD (level 1.5), also
297 used in the retrieval, is within ± 16 min of almucantar measurement for the four
298 wavelengths. Sky radiance data used as input in GRASP is previously normalized using
299 the “2000 ASTM Standard Extraterrestrial Spectrum Reference E-490-00”
300 (<http://rredc.nrel.gov/solar/spectra/am0>), again the same than in AERONET version 2
301 aerosol inversions. In order to include the filter response of the photometer, the
302 extraterrestrial spectrum is convoluted for each channel by a 10 nm width square filter
303 (similar to the real filters) centred in the effective wavelength of the real photometer
304 filters.

305 3.1.2 Ceilometer

306 For each almucantar dataset the correlative ceilometer RCS values measured
307 without clouds (CBH provided by the instrument is null) are averaged in a ± 15 min
308 window centred around the almucantar time. A minimum of 5 RCS cloud-free profiles
309 is imposed for calculating the average and for consequently running GRASP. This
310 requirement of at least 5 profiles is not too restrictive working with averaged 15-
311 seconds profiles and could provide averaged profiles too noisy when the number of used
312 profiles in the averaging is closer to (and above) 5, but most of them will be only taken
313 into account up to low altitudes due to the used iterative method to reject noisy points
314 that is explained below; this threshold may be increased in future works, but now it
315 permits to obtain more retrievals. The time averaged RCS is vertically smoothed by a
316 moving average of ± 105 m window in order to reduce noise, and later it is normalized at
317 60 log-spaced bins at different heights, as in Lopatin et al. (2013), being the minimum

318 of these heights (z_{min}) equal to 250 m a.s.l. since the ceilometer shows frequently very
 319 noisy signal below this height due to the overlap correction. The maximum height (z_{max})
 320 selected for the 60 log-spaced bins is 7000 m a.s.l. since aerosol layers are rarely
 321 detected above this height and the ceilometer signal is usually too noisy, due to the low
 322 power of ceilometer's laser. The RCS at these 60 log-spaced bins is normalized by
 323 dividing the average of RCS in each logarithmic height interval by the integrated RCS
 324 between z_{min} and z_{max} according to the following equation:

$$325 \quad NRCS_h = \frac{\frac{1}{N} \sum_{z=h_1}^{z=h_N} RCS_z}{\int_{z_{min}}^{z_{max}} RCS_z dz} \quad (1)$$

326 where $NRCS_h$ is the normalized RCS at the h-bin (h ranges from 1 to 60), N is
 327 the number of available RCS values in the height interval given by the h-bin, and $h_1, h_2,$
 328 ... and h_N represents the N heights of the available RCS that are inside the h-bin.

329 Due to the background correction and the noisy signal at high altitudes, the
 330 smooth and normalization process occasionally provides negative values of normalized
 331 RCS, which cannot be processed by GRASP due to the lack of physical sense. An
 332 iterative method has been applied to solve this issue: if any normalized RCS value is
 333 negative then the 60 log-spaced bins and normalized values are recalculated considering
 334 the maximum height 100 m below the last; this loop with z_{max} decreasing 100 m per
 335 iteration stops when all values of normalized RCS are positive.

336 3.1.3 BRDF data

337 A part of measured sky radiance has its source in the light reflected by the Earth
 338 surface; therefore, the Bidirectional Reflectance Distribution Function (BRDF) is used
 339 to take into account this phenomenon. The BRDF is introduced in GRASP through the
 340 BRDF parameters of the Li–Ross model (Ross, 1981; Li and Strahler, 1992). GRASP is
 341 capable to calculate BRDF parameters from satellite images (Dubovik et al., 2014) but

342 the BRDF parameters used for this work are obtained from the V005 Collection
343 MCD43C1 product (V005 MODIS Terra+Aqua BRDF/Albedo 16-Day L3 0.05Deg
344 CMG) of MODIS (MODerate-resolution Imaging Spectroradiometer) with a spatial
345 resolution of 0.05° (Schaff et al., 2011). This product is produced every 8 days with 16
346 days of acquisitions at seven narrow bands, which central wavelengths are 470, 555,
347 659, 858, 1240, 1640 and 2130 nm. The available MCD43C1 data at the Granada
348 coordinates from 2000 to 2014 have been averaged obtaining a table of BRDF
349 parameters every 8 days for one representative year. The BRDF parameters used in a
350 particular GRASP retrieval are obtained from the mentioned table taken into account the
351 date and linearly interpolating the central wavelengths of MCD43C1 product to 675,
352 870, 1020 and 1064 nm and extrapolating to 440 nm.

353 **3.2 Inversion strategy, constraints and products**

354 GRASP includes two independent modules, the first is the forward model based
355 on radiative transfer and aerosol model, which is capable to generate the radiative
356 measurements for a given aerosol scenario (Dubovik et al., 2014). This forward model
357 is used in Section 4 to simulate synthetic data for different aerosol scenarios. The
358 second module corresponds to the numerical inversion, which includes general
359 mathematical operations, based on multi-term least square method (Dubovik and King,
360 2000), not related to the particular physical nature of the inverted data (Dubovik et al.,
361 2014). This module, combined with the forward module, allows flexible and rigorous
362 inversions of the various combinations of the independent multi-source measurements.
363 Detailed description about how GRASP and its modules work using sunphotometer and
364 RCS data was given by Lopatin et al. (2013), who explained the GARRLiC algorithm
365 which nowadays is part of GRASP code.

366 The use of sunphotometer and ceilometer data proposed in this work cannot
367 discern between different aerosol modes in the vertical because the ceilometer provides
368 RCS profiles at only one wavelength. Hence, for the retrieval constraining intensive
369 aerosol properties such as refractive indices, SSA, lidar ratio (LR) or effective radius are
370 assumed equal for fine and coarse mode in the retrieval, which therefore implies that
371 GRASP is not able to provide vertical profiles of these parameters. Column integrated
372 retrieved parameter are aerosol size distribution (22 log-spaced triangle bins from 0.05
373 μm to 15 μm radius as in the operational AERONET retrievals) and fraction of
374 spherical particles (also called sphere fraction). The scheme also provides column-
375 integrated values of real refractive index (RRI), imaginary refractive index (IRI), SSA
376 and LR at 5 wavelengths (440, 675, 870, 1020 and 1064 nm). However, the hypothesis
377 of vertically constant aerosol intensive parameters allows changes in extensive
378 properties and, therefore, vertical profiles of the 60 log-spaced bins of aerosol volume
379 concentration and of extinction, backscatter, absorption and scattering coefficients at the
380 mentioned 5 wavelengths are provided.

381 In the GRASP retrievals we assume: no changes in extensive vertical properties
382 from ground to the z_{min} ; and an exponential decrease in these properties above z_{max} as in
383 Lopatin et al. (2013). GRASP needs an initial aerosol scenario, also known as initial
384 guess (Torres et al., 2017), to initialize each retrieval. The initial guess of each
385 parameter has been assumed the same for all retrievals except for the size distribution,
386 which has been assumed as a trapezoidal distribution proportional to the measured AOD
387 at 440 nm wavelength (AOD_{440}). Finally, GRASP also provides the uncertainty, σ_G , on
388 the retrieved parameters (VC, SSA, etc.), which is calculated from the random and
389 systematic errors estimated by the detailed methodology shown in Sections 2.3 and 2.4
390 of Dubovik et al. (2000). These products obtained by GRASP using the described

391 methodology are labelled in this work as GRASP_{pac}, which sub-index makes reference
392 to the combination of “photometer and ceilometer”. GRASP_{pac} retrievals not showing
393 convergence are rejected. Only 2% of the retrievals obtained in Section 5.2 were
394 discarded by the convergence criteria.

395

396 **4 Retrieval Sensitivity**

397 **4.1 Generation of Synthetic Data**

398 A sensitivity study with synthetic data is done in order to observe the capability of
399 the GRASP_{pac}. To this end, two kinds of aerosol are considered: Smoke and Dust,
400 including different mixtures among them. Smoke and Dust typical size distributions and
401 refractive indices are from Dubovik et al. (2002) for biomass burning in the African
402 savanna (Zambia) and for desert dust at the Arabian Peninsula (Saudi Arabia),
403 respectively. Figure 1 shows the typical size distributions for these two aerosol types
404 (Fig. 1a) and their vertical distribution (Fig. 1b), real refractive index (Fig. 1c), and
405 imaginary refractive index (Fig 1d) according to values reported in the bibliography.

406 The size distribution for each aerosol type (Figure 1a) are assumed as triangle
407 binned and bimodal distribution: the fine mode is log-spaced in 10 radius bins (radius
408 from 0.05 μm to 0.58 μm) and the coarse mode log-spaced in 15 bins (radius from 0.33
409 μm to 15 μm). Fine mode is predominant in Smoke aerosol with residual coarse mode,
410 while for Dust the opposite occurs. We remember that for each scenario both fine and
411 coarse mode have the same refractive indices, RRI being independent on wavelength
412 with values of 1.51 and 1.56 for Smoke and Dust, respectively (Fig. 1c). The IRI is
413 wavelength independent for Smoke, with a value of 0.021, while for Dust it is assumed
414 variable with wavelength varying from 0.003 at 440 nm to 0.001 at 1064 nm (Fig. 1d).
415 The vertical aerosol distribution has been assumed as an exponential decay with altitude

416 for Smoke, while this distribution has been considered as a Gaussian layer centred at
417 2000 m a.g.l. for Dust (Fig. 1b).

418 Different synthetic scenarios are considered consisting to Smoke, Dust and
419 mixtures among them. In the mixtures we assume that fine mode has the intensive
420 properties of Smoke while for coarse mode they are those of Dust. Fine mode of size
421 distribution is proportional to Smoke while that for coarse mode is proportional to Dust.
422 Two different mixtures are considered, Mix-1 that imposes that AOD_{440} is equal for fine
423 and coarse mode, and Mix-2 that imposes AOD at 1064 nm is equal for both modes.
424 The difference between Mix-1 and Mix-2 is the larger volume concentration of fine
425 particles in Mix-2 than in Mix-1. The size distribution and the vertical concentration for
426 these scenarios can be observed in the figures discussed in Section 4.2, labeled as
427 “Original” in Fig. 4 and 5, respectively. From all these scenarios, twelve (4 aerosol
428 types x 3 AODs) synthetic data are computed from different AOD_{440} values: 0.1 (low
429 aerosol load), 0.4 (minimum AOD_{440} used by AERONET to provide quality assured
430 SSA, RRI and IRI in version 2 retrievals) and 1.0 (high aerosol load).

431 The GRASP forward model is used to compute the synthetic observations
432 (spectral AOD, sky radiances and RCS at 1064 nm) following the conditions described
433 in the flow diagram of Fig. 2 for each of the twelve aerosol scenarios, and varying the
434 SZA by 10° from 40° to 80° in order to test different sets of scattering angles. Note that
435 in all our simulations the ground is assumed as the sea level and the assumed BDRF
436 parameters for these simulations are the climatological values (explained in Section
437 3.1.3) for Granada in summer. Later, using the GRASP forward model the required
438 observations for GRASP_{pac} are computed - AOD and sky radiances (26 values from 3.5°
439 to 160° azimuth angles) at 440, 675, 870 and 1020 nm and RCS (60 heights) at 1064
440 nm.

441 The computed synthetic observations are not representative of real
 442 measurements unless instrument uncertainties are considered, which are ± 0.01 for AOD
 443 and $\pm 5\%$ for sky radiances according to AERONET standards (Holben et al., 1998).
 444 Therefore, next step in the simulation scheme of Fig. 2 is to add uncertainties to the
 445 simulated AOD and sky radiances, which is done by adding random errors generated
 446 from random number that follows a normal distribution with standard deviation equal to
 447 the uncertainties. The addition of noise to the simulated values is done assuming a
 448 constant uncertainty (K) on raw ceilometer signal and, therefore, the uncertainty
 449 $\sigma(RCS)$ varies with the square of the distance (z) and at a level 'z' is given as:

$$450 \quad \sigma(RCS_z) = Kz^2 \quad (2)$$

451 where RCS_z is the range corrected signal at z.

452 The calibration constant for Granada ceilometer obtained by Cazorla et al.
 453 (2017) for molecular (aerosol free) regions presents variations with standard deviation
 454 of approximately 30% (result not published). Thus, the uncertainty of ceilometer RCS
 455 could be assumed as a 30% at the reference height (z_{ref}) where only molecular
 456 backscatter is detected. Then the uncertainty of RCS at z_{ref} can be written as:

$$457 \quad \sigma(RCS_{z=z_{ref}}) = 0.3 * RCS_{z=z_{ref}} \quad (3)$$

458 and combining Eq. (2) and Eq. (3):

$$459 \quad K = \frac{0.3}{z_{ref}^2} RCS_{z=z_{ref}} \quad (4)$$

460 Finally, if Eq. (4) is put in Eq. (2), the uncertainty of RCS at each height can be
 461 expressed as:

$$462 \quad \sigma(RCS_z) = \frac{0.3}{z_{ref}^2} RCS_{z=z_{ref}} * z^2 \quad (5)$$

463 The most frequent value of z_{ref} obtained by the method used in Cazorla et al.
464 (2017) is about 4000 m a.g.l.; therefore, in this work, the uncertainty of ceilometer RCS
465 is calculated by Eq. (5) using 4000 m as z_{ref} .

466 Once RCS uncertainty has been characterized, synthetic RCS is interpolated every
467 15 m, and for each RCS value at 15 m at each scenario, a pseudorandom number
468 normally distributed is generated with a standard deviation equal to the uncertainty of
469 this RCS value, and this random number is added to the previously simulated RCS.

470 As an example, Fig. 3a shows the synthetic RCS after adding uncertainties for the
471 Smoke and Mix-2 scenarios with AOD_{440} equal to 0.4. In addition, a particular example
472 of measured RCS in Granada (dust case with $AOD_{440} = 0.21$) is included to illustrate the
473 capabilities of our scheme to generate synthetic RCS with uncertainties. The iterative
474 method to skip negative values in the measured RCS is applied to the initial values
475 (Measured-Initial) and it is observed as the final signal avoids negative values
476 (Measured-Final). As can be observed the profiles look noisier at higher heights both
477 for the synthetic and measured profiles. In fact, the shape of the added noise to the
478 synthetic profiles is very similar to the one observed in the real measurements, which
479 indicates that the obtained synthetic signal can be considered as realistic. The noise is
480 higher for Smoke likely because for this scenario molecular zone is not completely well
481 represented by the assumed z_{ref} equal to 4000 m a.g.l. Figure 3b shows the RCS of
482 Figure 3a normalized to the 60 heights required as input in GRASP. It can be
483 appreciated that noise is reduced by the averaging of RCS in log-scaled bins.

484 **4.2 Analyses of retrieved parameters**

485 As the diagram of Fig. 2 shows, once the noisy synthetic observations are
486 obtained for each aerosol scenario and SZA value, these data are used as input in

487 GRASP as explained in Section 3. The differences, Δ_{fit} , between the synthetic
 488 observations used as input in GRASP_{pac} and the observations generated by the retrieved
 489 aerosol scenario are calculated to quantify the fitness of each GRASP_{pac} retrieval (see
 490 Fig. 2). Δ_{fit} is defined as:

$$491 \quad \Delta_{fit}(k, n) = O_r(k, n) - O_i(k, n) \quad (6)$$

492 and in percentage as:

$$493 \quad \Delta_{fit}(k, n)(\%) = 100\% \frac{O_r(k, n) - O_i(k, n)}{O_i(k, n)} \quad (7)$$

494 where O represents an observation; the sub-index i and r indicated if the observation is
 495 an input or a value obtained from the retrieved aerosol scenario, respectively (see Fig.
 496 2); k determines the kind of observation (AOD, sky radiances or RCS) and n is the
 497 number of this kind of observation. The fitness of the retrieval can be quantified for
 498 each k -kind observation by the mean (MBE; mean bias error) and standard deviation
 499 (STD) of Δ_{fit} using all n available observations for the k -kind. MBE represents the
 500 accuracy between O_r and O_i , while STD indicates their precision. Following this
 501 method, MBE and STD for AOD (sub-index aod), sky radiance (sub-index rad) and
 502 RCS (sub-index rsc) are calculated for all retrievals and they are shown in Table 1.
 503 MBE_{aod} and STD_{aod} are shown in absolute values while MBE and STD for sky radiance
 504 and RCS are in percentage. Scattering angle interval is also added in Table 1, reaching
 505 bigger angles when SZA increases. Table 1 reveals that MBE_{aod}, MBE_{rsc}, STD_{aod}, and
 506 STD_{rsc} are usually larger for retrievals with AOD₄₄₀=0.1; MBE_{rad} is usually within $\pm 1\%$
 507 and STD_{rad} around 3%. In general, the fitness estimation does not show a clear
 508 dependence on aerosol type, SZA or AOD, which could indicate that differences in
 509 these values for different cases are mainly caused by the noise in the synthetic
 510 measurements since it is random.

511 Several aerosol GRASP_{pac} products are obtained for each retrieval, but this work
512 is mainly focus on columnar size distribution and especially on aerosol VC profiles.
513 Figure 4 shows, for different aerosol types and loads, all the retrieved size distributions
514 for various SZA values. We remind that errors were added to input optical data. The
515 original size distributions are also included. In general, the retrieved size distributions
516 look qualitatively similar to the original ones, especially for the coarse mode, for all
517 aerosol scenarios. Discrepancies on fine mode are more evident especially at low
518 AODs. Worse agreement is expected for small SZA values since the scattering angle
519 range is shorter, however it is not observed. The differences between the original and
520 retrieved size distributions are mostly related with Δ_{fit} . For example, the retrieved size
521 distribution for Mix-1 type with AOD₄₄₀=0.4 differs more from the original at SZA
522 equal to 60° than for the other angles; it should be caused by a worse fit between the
523 inputs and the retrieved observations as it can be observed in Table 1, where MBE_{rad}
524 and STD_{rad} reach their highest values (2.3% and 7.0%, respectively) for all retrievals
525 with AOD₄₄₀=0.4. It can also be appreciated in the Mix-1 type with AOD₄₄₀=1.0 and
526 SZA of 80°.

527 Figure 5 shows the VC profiles for the same data than in Fig. 4. These profiles
528 show a good agreement with the original ones when coarse mode predominates as can
529 be observed for Dust and Mix-1 cases. The larger differences between retrieved and
530 reference profiles are found for Smoke, being particularly noisy for heights above 2 km.
531 This worse agreement for Smoke could be due to the use of RCS at 1064 nm, this
532 wavelength being less sensitive to the fine particles like those prevailing in Smoke. The
533 original Mix-2 profiles present two intense aerosol layers: dust around 2 km and smoke
534 below 1 km; GRASP_{pac} method is able to detect both aerosol layers, although it shows
535 discrepancies compared with the reference. This can be explained by the limited

536 information of using RCS at only one wavelength. To quantify all the differences we
 537 defined Δ_{vc} , as the difference between the retrieved and original VC profiles (see Fig.2)
 538 given by:

$$539 \quad \Delta_{vc}(a, SZA, z) = VC_r(k, SZA, z) - VC_o(a, SZA, z) \quad (8)$$

540 and in percentage as:

$$541 \quad \Delta_{vc}(a, SZA, z)(\%) = 100\% \frac{VC_r(k, SZA, z) - VC_o(a, SZA, z)}{VC_o(a, SZA, z)} \quad (9)$$

542 where VC_r and VC_o represents the retrieved and original VC values, respectively (see
 543 Fig. 2); a determines the aerosol scenario (aerosol type and AOD₄₄₀) and z being one of
 544 the 60 bins of the retrieved VC profiles.

545 Table 2 shows the MBE and STD calculated as the mean and standard deviation,
 546 respectively, of the 60 Δ_{vc} values (Eq. (8) and (9)) of each profile. The Δ_{vc} values with
 547 VC_o below $1 \mu\text{m}^3/\text{cm}^3$ have been discarded in the MBE and STD calculation since they
 548 could provide extreme differences in percentage. The results of Table 2 are showed for
 549 each of the 12 different aerosol scenarios and for different SZA. MBE and STD of
 550 Table 2 do not show any dependence with SZA. The best agreements (minima MBE
 551 and STD) are found for Dust and Mix-1 scenarios, where coarse mode is predominant.
 552 In general, unsigned MBE increases with AOD₄₄₀ while the precision of GRASP_{pac},
 553 given by STD, decreases in percentage with AOD₄₄₀. As a general result, for all
 554 scenarios together GRASP_{pac} systematically underestimates VC showing a MBE of -
 555 5.9% and with an uncertainty, which is given by STD, of 21%. The lowest uncertainties
 556 of GRASP_{pac} are for Dust aerosol (~14%) with bias close to zero, while the highest
 557 uncertainties are for the Smoke type (~28%).

558 In order to observe if the obtained differences between the original VC and the
 559 retrieved by GRASP_{pac} are within σ_G (the estimation of retrieval uncertainty provided by

560 GRASP_{pac}), the percentage of unsigned Δ_{vc} values (Eq. (8)) that are below σ_G and $2\sigma_G$
561 have been calculated and named as $\Delta_{vc}<\sigma_G$ and $\Delta_{vc}<2\sigma_G$, respectively. If $\Delta_{vc}<\sigma_G$ and
562 $\Delta_{vc}<2\sigma_G$ are similar to 68% and 95%, respectively, σ_G will represent the uncertainty in a
563 good way indicating that Δ_{vc} is similar to a normal distribution with a standard deviation
564 equal to σ_G . Table 3 shows the obtained results for each scenario shown in Table 2.
565 $\Delta_{vc}<\sigma_G$ and $\Delta_{vc}<2\sigma_G$ do not show any dependence on SZA or AOD₄₄₀. Mix-2 aerosol
566 scenario presents the $\Delta_{vc}<\sigma_G$ and $\Delta_{vc}<2\sigma_G$ values closer to 68% and 95%; Dust and
567 Mix-1 show even higher values. Smoke aerosol shows the lowest values when all SZA
568 and AOD₄₄₀ values are taken into account, but it is mainly caused by various individual
569 cases with SZA=60° and AOD₄₄₀=0.1 or SZA=40° and AOD₄₄₀=1.0. For the
570 combination of all the different aerosol scenarios, $\Delta_{vc}<\sigma_G$ is 74% and $\Delta_{vc}<2\sigma_G$ is 91%,
571 which are close values to the expected 68% and 95%, and therefore we can conclude
572 that GRASP_{pac} reproduces well the VC profiles within the margins given by the
573 uncertainty associated with the numerical inversion.

574 For backscatter and extinction coefficients at 1064 nm and column integrated
575 intensive properties such as complex refractive index, SSA and LR we also did the same
576 computations (not shown) as in Table 2 and Table 3. Combining all the data of the
577 different aerosol scenarios MBE are -11% and -5% and STD equal to 31% and 21% for
578 backscatter and extinction profiles, respectively. For the backscatter coefficient, MBE
579 presents the largest values for Smoke and Mix-2, while Dust and Mix-2 show the largest
580 STD values. In the case of the extinction coefficient, Dust and Mix-1 present the lowest
581 STD (13% and 15%) and MBE (3% and 2%) values. Regarding the retrieved column-
582 integrated SSA, considering the five wavelengths together, the retrieved SSA fits better
583 the original values when AOD₄₄₀ increases, MBE being equal to -0.02, 0.01 and 0.00
584 and STD equal to 0.08, 0.05, and 0.02 for AOD₄₄₀ of 0.10, 0.4 and 1.0, respectively for

585 all aerosol types and SZA values. The retrieved SSA also agrees better as SZA
586 increases, indicating the importance of large scattering angles in this property as
587 expected (Dubovik et al., 2000), but this dependence is only clear for $AOD_{440}=0.4$ and
588 1.0. Similar dependence on AOD_{440} , but not on SZA, appears for the retrieved LR.
589 These LR retrievals agree with the references when all scenarios are considered together
590 (MBE and STD are 10% and 29%). This agreement is found particularly for the Smoke
591 aerosol cases. MBE and STD are reduced to 1% and 26% when only cases with
592 $AOD_{440}=0.4$ are selected. Finally, for RRI and IRI, good agreements with the reference
593 values are found for high AOD_{440} . Our last computations reveal that the differences
594 between retrieved properties and the original ones are within σ_G , the obtained results
595 indicate that σ_G of backscatter and extinction is representative of the real uncertainty for
596 all AOD_{440} and SZA values. On the other hand, for SSA and LR the percentage of
597 differences below σ_G is lower than the expected and showing an increase with AOD_{440} .

598

599 **5 Results from inversion of real observations**

600 **5.1 Airborne comparison**

601 Figure 6 shows the ceilometer RCS for the period 16-17th June 2013 where flights
602 over Granada were done within the ChArMEx/ADRIMED field campaign. The largest
603 RCS are observed below ~ 2 km a.s.l. that usually corresponds to aerosol in the
604 planetary boundary layer (PBL). During this period, the study region was affected by
605 Saharan dust outbreaks with transport of dust particles (Benavent-Oltra et al., 2017).
606 The presence of long-range transported aerosol is clearly observed in Fig. 6 with
607 significant signal up to 5 km a.s.l., approximately. Decoupled aerosol layers appeared
608 from the 16th June evening to 17th morning, with aerosol entrainment in the PBL also
609 observed, which is typically observed during Saharan dust arrivals at the study station

610 (Bravo-Aranda et al., 2015). Signal decreases are observed from 17th morning,
611 particularly strong at low levels, and explained by advection of clean air-masses at these
612 levels. However, a high-altitude layer remained at 3-5 km a.s.l.. The averaged
613 (\pm standard deviation) daytime AOD at 440 nm and Angström Exponent (AE; in this
614 work calculated only with the AOD at 440 and 870 nm) were 0.26 ± 0.01 and 0.35 ± 0.04
615 (63 data), respectively, for 16th June and 0.20 ± 0.04 and 0.44 ± 0.04 for 17th June (19
616 data); the low AE values indicate the presence of coarse particles. Five-day back-
617 trajectories analyses using HYSPLIT model (Stein et al., 2015) (not shown) point out
618 that the air masses came at Granada from the Saharan desert, which agrees with the
619 presence of coarse particles as Saharan mineral dust.

620 Airplane spirals near the study region were done at 14:15-14:45 UTC (denoted
621 as F30) and at 07:15-07:45 UTC (F31). F31 trajectory (similar to F30) is shown in
622 Benavent-Oltra et al. (2017) and it shows that airborne measurements were done around
623 20 km far from Granada station. The time of both flights are marked in Figure 6 with
624 black vertical lines while the closest GRASP_{pac} retrieval to each flight is indicated by
625 two green vertical lines, with AOD₄₄₀ of 0.27 and 0.21, respectively. The time
626 difference of 2 hours between F30 flight and the closest GRASP_{pac} retrieval is because
627 limitations in SZA (at the exact time of the flight SZA was very small and become
628 larger than 40° from 16:22 UTC). However, stable AOD measurements suggest not big
629 aerosol variations during this 2 hour period.

630 Figure 7 shows the column-integrated size distribution, SSA, and refractive
631 indices obtained by GRASP_{pac} and these provided by AERONET (level 1.5).
632 Comparisons of size distributions reveal that they are very similar between both
633 methodologies, being the differences within the GRASP_{pac} uncertainties. The size
634 distributions also indicate the predominance of the coarse mode as expected for Saharan

635 dust outbreaks (Valenzuela et al., 2012), and both retrievals point out a positive shift of
636 the coarse mode concentration in the morning of 17th June. It is corroborated by the
637 effective radius of the coarse mode given by GRASP_{pac}, which varied from 1.93 μm
638 (Fig. 6a) to 2.22 μm (Fig. 6b). For SSA, Fig. 7c and 7d reveal that values are very
639 similar between GRASP_{pac} and AERONET, and both retrievals show a spectral
640 dependence typical of mineral dust (Dubovik et al., 2002). RRI from AERONET is
641 slightly higher in both cases than from GRASP_{pac}, but both retrievals show wavelength
642 independence and a weak decrease from 16th to 17th June. Finally, for IRI again both
643 AERONET and GRASP_{pac} show similar patterns, typical for dust (Dubovik et al.,
644 2002), and differences between methodologies are within the uncertainties. All these
645 results point out that the column-integrated products from GRASP_{pac} are in accordance
646 with the ones provided by AERONET, at least in the analysed cases.

647 Figure 8 shows vertically-resolved values of particle VC from GRASP_{pac}
648 ($VC_{GRASP_{pac}}$) and the values obtained by airborne measurements ($VC_{Airborne}$). Generally
649 both methodologies present very similar profiles for the two cases. For the flight F30,
650 only one layer is observed with a slight and constant decrease up to 4.5 km
651 approximately, while for F31 three different layers are observed. Most of the differences
652 are within the GRASP_{pac} uncertainty, however, disagreements are found between
653 retrievals and airplane measurements for altitudes below 1.5 km, which can be
654 explained because of the orography and air-traffic restriction that did not allow the
655 flight to perform spiral exactly above the station. This reasoning agrees with the largest
656 aerosol VC values at the lowest layer observed by GRASP retrievals, which can be
657 associated with pollution from the city.

658 To quantify the differences between GRASP_{pac} and airborne profiles, the VC
659 from GRASP_{pac} has been interpolated to the available heights of the airborne

660 measurements. Point-by-point intercomparison between GRASP retrievals and airborne
661 measurements are done. Linear interpolations of $GRASP_{pac}$ are done too for the same
662 altitude than airborne measurements. Cases with very low aerosol load ($VC < 5 \mu m^3/cm^3$)
663 and measurements below 1.25 km a.s.l. (large disagreements in aerosol sampled
664 between both techniques) are rejected in this comparison. Figure 9 shows particle VC
665 obtained by GRASP versus airborne values. The correlation between both
666 methodologies is high (correlation coefficient, r , higher than 0.80), and slightly better
667 for F30 flight. However, the slope of the least square fit indicates that $GRASP_{pac}$
668 underestimates the highest airborne measurements and the abscissa intercept points out
669 that $GRASP_{pac}$ overestimates the lowest values. In addition, the differences, ΔVC ,
670 between VC values from $GRASP_{pac}$ and airborne have been calculated as follows:

$$671 \quad \Delta VC = VC_{GRASP_{pac}} - VC_{Airborne} \quad (10)$$

672 and in percentage as:

$$673 \quad \Delta VC(\%) = 100\% \frac{VC_{GRASP_{pac}} - VC_{Airborne}}{VC_{Airborne}} \quad (11)$$

674 The histograms of ΔVC (Eq. (10)) are shown in Fig. 9d, 9e and 9f for F30, F31
675 and both flights, respectively. These graphs indicate that VC from $GRASP_{pac}$ agrees
676 better with airborne measurements for F30 flight, being the 37% of the absolute ΔVC
677 values below $2.5 \mu m^3/cm^3$ and 89% below $7.5 \mu m^3/cm^3$. The ΔVC distribution for F31
678 flight presents higher values but it is similar to a normal distribution, 61% of ΔVC
679 absolute data being lower than $7.5 \mu m^3/cm^3$; this percentage rises up to 75% when both
680 flights are taken into account. Table 4 shows the mean (MBE), mean of the absolute
681 values (MABE) and standard deviation (STD) of ΔVC (Eq. (10) and (11)) for these
682 three cases of Fig. 9. $GRASP_{pac}$ slightly overestimates the $VC_{Airborne}$ values, showing
683 MBE values of 10.5% and 12.9% for F30 and F31 flights, respectively; however, the

684 absolute MBE is close to $0 \mu\text{m}^3/\text{cm}^3$. Assuming airborne measurements as a reference,
685 the accuracy, given by MBE, of VC from GRASP_{pac} is below 12% when both flights are
686 taken into account. Regarding MABE, F31 flight shows values around the double of
687 that obtained from F30, which indicates that ΔVC differences are much higher in the
688 F31 case, as STD confirms. The precision of GRASP_{pac} using airborne measurements as
689 a reference can be represented by STD, which presents a low value of 18.5% in the F30
690 case, but this value for F31 rises up to 70.8% due to the vertical shift of the lowest layer
691 observed in Fig.8b. The STD for both flights together is 51.4%, but this value is still
692 strongly affected by the differences in F31 flight for low heights. Finally, for both
693 flights together, the percentage of ΔVC values which are below the uncertainty given by
694 GRASP_{pac} is 67.6%; this percentage is 94.4% when the double of the uncertainty is
695 considered. These values are close to 68% and 95%, which points out that the
696 uncertainty estimation provided by GRASP_{pac} is representative of the real uncertainty of
697 the retrieved VC.

698 **5.2 High altitude station comparison**

699 In-situ VC measurements during SLOPE I field campaign at the Sierra Nevada
700 station (VC_{SNS}) are used for evaluating retrieved values by GRASP_{pac} at the same
701 altitude. In-situ measurements measured total particle VC in the range 0.05-10 μm and
702 the GRASP_{pac} retrieved values are integrated in the same range. From retrieved VC
703 profiles, linear interpolations are done to have data at 2500 m a.s.l., which is the altitude
704 of Sierra Nevada station. Figure 10 shows the temporal evolution of in-situ and
705 retrieved VC values for the entire period. While measurements of VC_{SNS} were
706 continuous (24 hours per day), retrieved GRASP_{pac} values are only available during
707 some daytime points every day. The lack of VC_{SNS} data during some short periods were

708 caused by instrumental failures. From Figure 10 can be observed that both measured
709 and retrieved values follow the same temporal evolution, with minimum values
710 associated with clean atmosphere and extreme values associated mostly to Saharan dust
711 arrivals. In fact, the largest values at Sierra Nevada were registered during the morning
712 of 21st July, with in-situ measurements up to $269 \mu\text{m}^3/\text{cm}^3$ and retrieved GRASP_{pac}
713 values from 279 to $364 \mu\text{m}^3/\text{cm}^3$, and were associated with a strong Saharan dust
714 episode that started on 20th July 2016.

715 Figure 11a shows a normalized number density plot of retrieved values by
716 GRASP_{pac} versus in-situ measurements (VC_{SNS}). Selected in-situ measurements are
717 averaged during a time period of ± 15 min from the retrieval time. Most of the VC
718 values on Figure 11a are below $20 \mu\text{m}^3/\text{cm}^3$, being 71% for VC_{SNS} . The linear fit reveals
719 an overestimation of VC from GRASP_{pac} to the VC_{SNS} values around 50%. The data for
720 26th August can be partially responsible of this overestimation with $\text{VC}_{\text{GRASP}_{pac}}$ values
721 $\sim 150 \mu\text{m}^3/\text{cm}^3$ while VC_{SNS} is $\sim 50 \mu\text{m}^3/\text{cm}^3$. These larger differences could be in part
722 due to real differences in the aerosol over the Granada vertical and the aerosol at Sierra
723 Nevada, since SNS could be affected by local effects and sources.

724 The correlation between $\text{VC}_{\text{GRASP}_{pac}}$ and VC_{SNS} is high, being r equal to 0.91; this
725 correlation coefficient is higher than the obtained between the ground measured AOD_{440}
726 and VC_{SNS} , which is 0.79, and the correlation between the retrieved column-integrated
727 VC and VC_{SNS} , which is 0.80. This result points out that the addition of ceilometer
728 signal to the aerosol retrieval improves the capacity to estimate the aerosol vertical
729 concentration.

730 As in Section 5.1, the differences ΔVC between $\text{VC}_{\text{GRASP}_{pac}}$ and the in-situ
731 measurements, in this case VC_{SNS} , have been calculated. ΔVC can be expressed as:

732
$$\Delta VC = VC_{GRASP_{pac}} - VC_{SNS} \quad (12)$$

733 and in percentage as:

734
$$\Delta VC(\%) = 100\% \frac{VC_{GRASP_{pac}} - VC_{SNS}}{VC_{SNS}} \quad (13)$$

735 Figure 11b shows the ΔVC (Eq. (12)) distribution. This frequency histogram is
 736 similar to a normal distribution, the maximum being centred close to 0; however it is
 737 skewed to positive values. 38%, 73% and 87% of $VC_{GRASP_{pac}}$ shows absolute ΔVC
 738 differences lower than 2.5, 7.5 and 12.5 $\mu\text{m}^3/\text{cm}^3$, respectively.

739 Table 5 shows mean values and standard deviations of the differences ΔVC , from
 740 Eq. (12) and (13), for different VC_{SNS} ranges. The percentages of data when ΔVC is
 741 lower than the numerical uncertainty in the inversion, σ_G , are also included. From Table
 742 5 when all ranges of VC are considered mean differences and standard deviations are
 743 31% and 94%, both strongly affected by the low values of VC_{SNS} . In fact, MBE and
 744 STD are 64% and 169%, respectively, for VC_{SNS} values only below 5 $\mu\text{m}^3/\text{cm}^3$.
 745 However, if only data with VC_{SNS} above 5 $\mu\text{m}^3/\text{cm}^3$ are selected (493 in total), mean
 746 difference and standard deviations are reduced to 23% and 59%, respectively. In
 747 general, MBE increases with VC_{SNS} ranging from 10% to 60% if VC_{SNS} below 5
 748 $\mu\text{m}^3/\text{cm}^3$ is not considered. MABE presents values around 40-50% for VC_{SNS} between 5
 749 and 100 $\mu\text{m}^3/\text{cm}^3$. STD varies from 34-64%, showing the lowest values for highest
 750 concentrations. Regarding the ΔVC differences within the $GRASP_{pac}$ uncertainty
 751 estimation, Table 5 shows values below that expected, which indicates that the VC
 752 uncertainty estimation provided by $GRASP_{pac}$ could be not representative of the real
 753 uncertainty in this case. However, the obtained results could be affected by different
 754 factors, independent of $GRASP_{pac}$, which yield a worse agreement than in the airborne
 755 comparison of Section 5.1. In this section the aerosol properties in the free vertical

756 atmosphere over Granada have been assumed equal to the properties at the surface on
757 Sierra Nevada ground station, which could be affected by other aerosol sources and
758 atmospheric conditions. Moreover, the instrumental uncertainty on VC_{SNS} could be also
759 partially responsible of the observed differences.

760 Figure 12 shows the differences ΔVC of Eq. (12) as function of VC_{SNS} (Fig. 12a),
761 AOD_{440} (Fig. 12b), Angström Exponent (Fig. 12c) and sphere fraction (Fig. 12d).
762 Generally it is observed that ΔVC increases with VC_{SNS} , however, some high ΔVC
763 values appear for moderate VC_{SNS} values which correspond to the mentioned case of
764 26th August. ΔVC also increases with AOD_{440} , however high ΔVC values do not appear
765 for moderate AOD_{440} and even low ΔVC values can be observed for high AOD_{440} . ΔVC
766 does not show any clear dependence on AE and sphere fraction, except the highest ΔVC
767 values for the lowest values of AE and sphere fraction, which mainly corresponds to
768 dust particles during the mentioned strong dust episode of 20th -21st July 2016 (see Fig.
769 10).

770 Finally, a case of study based on the dust episode of 20th -21st July 2016 has been
771 analysed as an illustration. Figure 13a and 13b show the retrieved VC profiles and the
772 measured VC_{SNS} at Sierra Nevada on the afternoon of 20th July and on the morning of
773 21st July, respectively. The AOD_{440} from Granada was 0.85 and 0.83 for the Fig. 13a
774 and 13b, respectively, which indicates very similar aerosol load. It indicates that in
775 columnar terms, both cases are similar, but if ceilometer measurements are added to the
776 retrieval, the vertical distribution can be discerned; this is the case in Fig. 13, where the
777 GRASP_{pac} retrieval indicates that VC at SNS increased by about four times from 20th to
778 21st July, which was also appreciated in the measurements of VC_{SNS} . Then, thanks to
779 ceilometer addition, it is known that the dust episode came 20th July in a strong layer
780 located between 3.5-4.0 km a.s.l.. This layer went down providing extreme values at

781 SNS height in the morning of 21st July, but also high dust concentrations in lower
782 heights, which did not happen in the evening of 20th July.

783

784 **6 Conclusions**

785 In this work we have explored the use of collocated sun/sky photometer and
786 ceilometer measurements in the General Retrieval of Aerosol and Surface Properties
787 (GRASP) code to retrieve column-integrated and vertically-resolved optical and
788 microphysical aerosol properties such as backscatter and extinction coefficients and
789 volume concentration, among others. The capability to combining such set of
790 measurements and using them in GRASP has been studied through different sets of
791 simulations for typical dust and biomass-burning aerosol located at different altitudes as
792 well as mixtures of both. In general, the proposed GRASP retrievals reproduce better
793 aerosol properties for coarse particles, likely due to the operational long wavelength of
794 the ceilometer at 1064 nm, and for high aerosol optical depth values. The results of the
795 simulations have demonstrated good agreements for column-integrated size
796 distributions and optical parameters such as complex refractive indices and single
797 scattering albedo. For vertically-resolved aerosol properties, volume concentration
798 presents an accuracy of -6% and an uncertainty of 21%; this accuracy is -11% and 5%
799 for backscatter and extinction profiles at 1064 nm, being the uncertainty 31% and 21%,
800 respectively. The mentioned analysis concludes that the uncertainty of these GRASP
801 retrievals is representative of the real uncertainty of the retrieved parameters, except for
802 column single scattering albedo and lidar ratio where the uncertainty given by GRASP
803 is only representative when aerosol optical depth increases.

804 Two case studies from mid-June 2013 documented during the
805 ChArMEx/ADRIMED field campaign have allowed the comparison of retrieved

806 vertical profiles versus airborne in-situ measurements. The aerosol volume
807 concentration obtained by GRASP presents high correlation with the measured one
808 during the two flights. Differences in this concentration between GRASP retrievals and
809 airborne measurements present a mean value below 12% and a standard deviation
810 around 51%. All these differences are within the uncertainty estimations provided by
811 the GRASP code. Moreover, comparisons of the column-integrated retrieved parameters
812 by the proposed scheme for GRASP versus AERONET retrievals have been done
813 showing a good agreement between both techniques (differences were within
814 uncertainties).

815 Data acquired during the SLOPE I field campaign (summer 2016) at the high
816 mountain Sierra Nevada station, located at 2500 m a.s.l, were used to evaluate the
817 retrieved aerosol volume concentration at a certain altitude. The in-situ volume
818 concentration at a mountain station in Sierra Nevada correlates better ($r=0.91$) with the
819 aerosol volume concentration obtained by GRASP at 2500 m a.s.l. than other variables
820 like aerosol optical depth at Granada (ground station). Discarding the lowest
821 concentration values, the mean differences between retrieved and the measured volume
822 concentrations are of 23% with a standard deviation of 59%, which means that GRASP
823 frequently overestimates the in-situ measurements at Sierra Nevada. However, part of
824 these differences could be caused by uncertainties in the in-situ measurements and
825 assumptions, and in the fact that the aerosol over Granada (where ceilometer monitoring
826 was performed) could not be the same than the aerosol on Sierra Nevada, which could
827 be affected by local dynamic and atmospheric effects, and also to local aerosol sources
828 at the high mountains.

829 Overall, the obtained results indicate that the combination of sun/sky photometer
830 and ceilometer measurements and their use as inputs in GRASP provides reliable

831 products if the uncertainties are considered. Nevertheless, the experimental data
832 obtained were mostly representative of dust and clean conditions, and more evaluations
833 are required for very polluted environment and intense biomass-burning. Therefore, as
834 outlook, the method could be applied in different places, using networks like ICENET,
835 and in long time series in order to characterize the regional and temporal changes on
836 vertical aerosol extensive properties.

837

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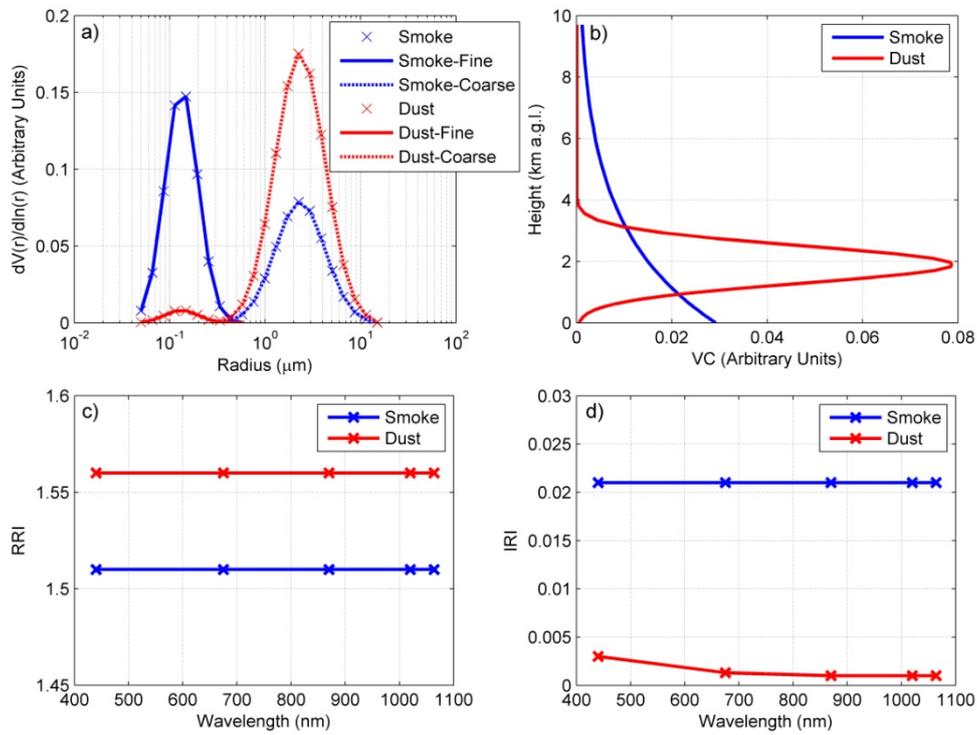
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1218 **Figures**



1219
 1220 Figure 1: Microphysical and optical properties of the two aerosol models (Smoke and
 1221 Dust) used to obtain synthetic data: size distribution (panel a); vertical volume
 1222 concentration, VC, (panel b); real (panel c) and imaginary (panel d) refractive indices,
 1223 RRI and IRI, respectively.

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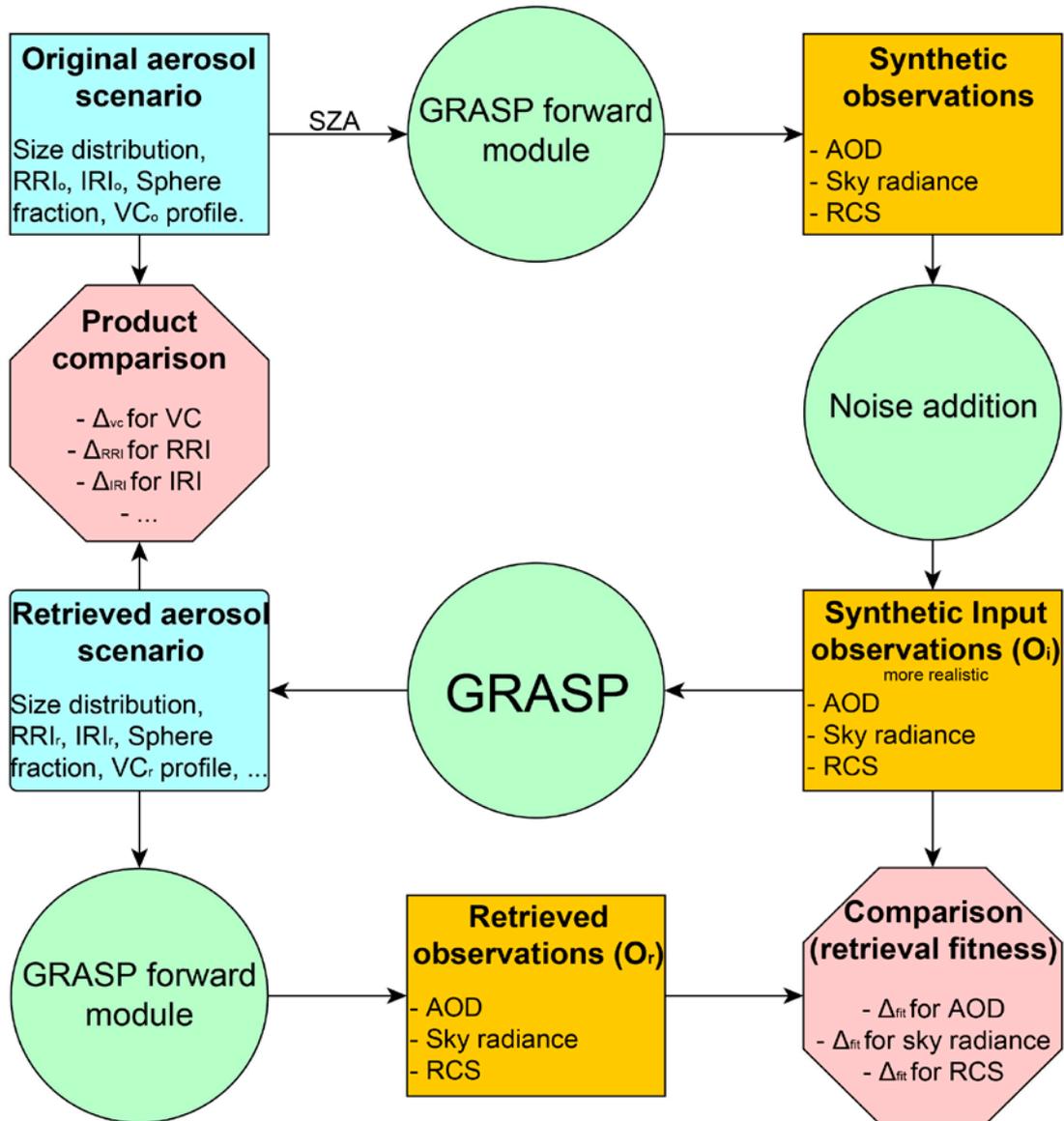
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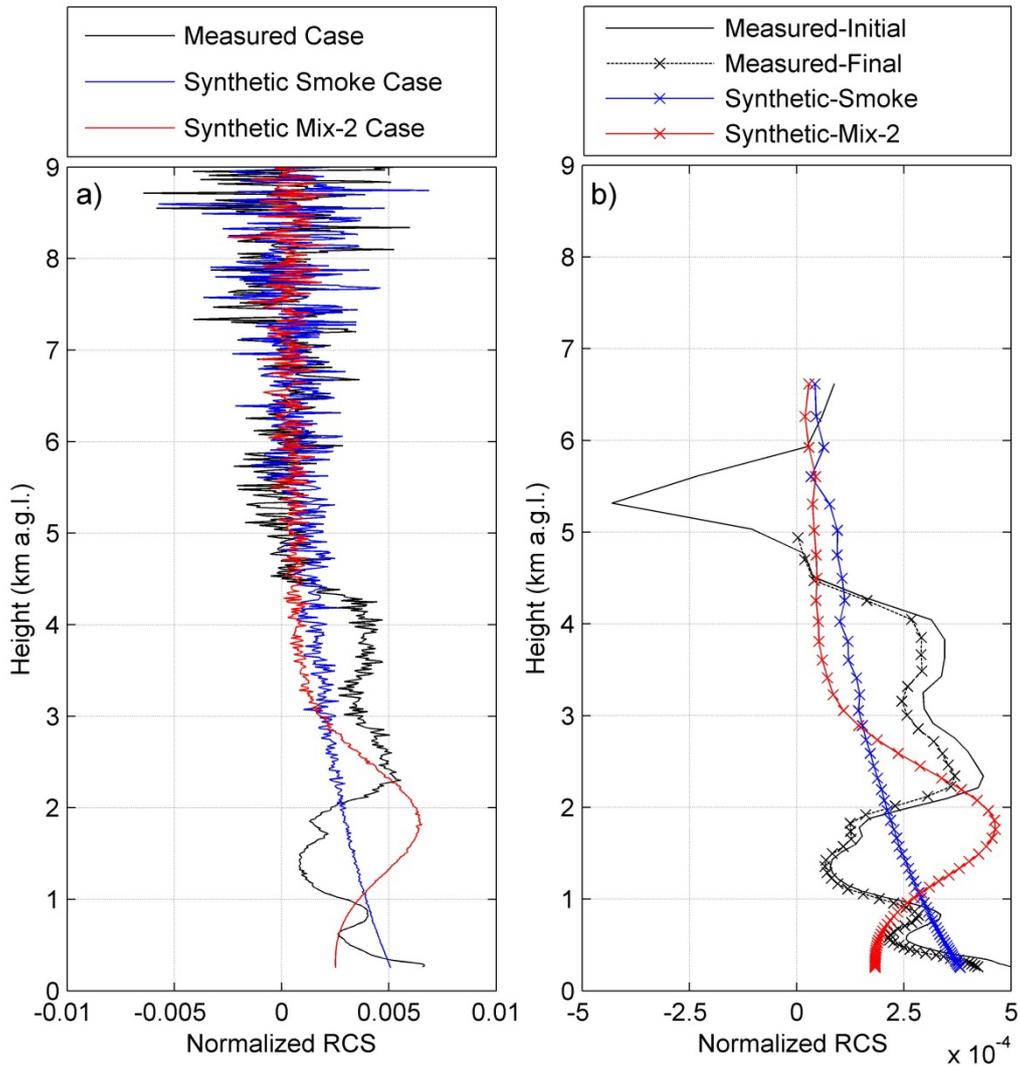
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1231 Figure 2: Flow diagram about retrieval sensitivity study with synthetic data.

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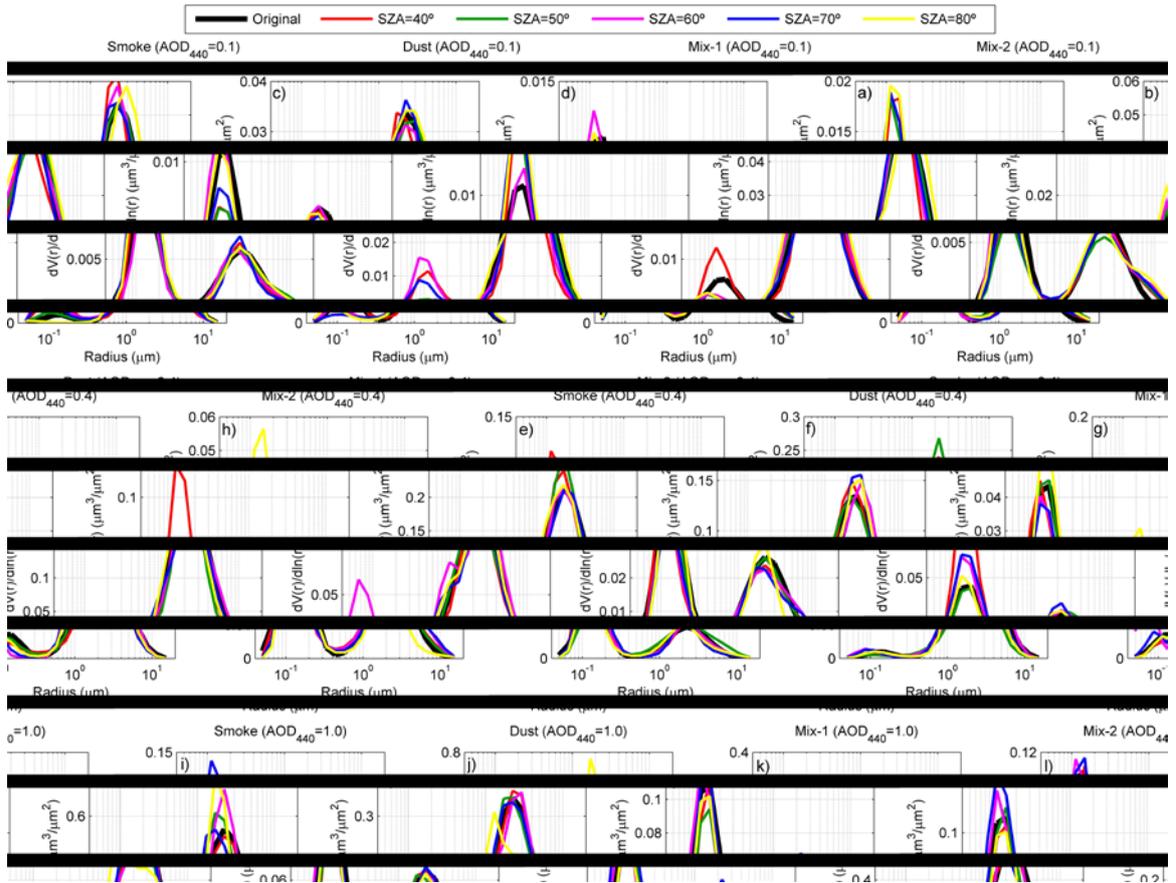
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1234 Figure 3: Range corrected signals (RCS) at 1064 nm, from 250 m to 9000 m every 15
 1235 m, normalized by the sum of all data (panel a) for three cases: half hour average of
 1236 measured ceilometer signal on 17 June 2013, 07:40 UTC ($AOD_{440} \sim 0.21$) (black line);
 1237 synthetic and noisy signal of Smoke with AOD_{440} equal to 0.4 (blue line); and synthetic
 1238 and noisy signal of “Mix-2” with AOD_{440} equal to 0.4 (red line). Panel b shows the RCS
 1239 of panel a, but normalized to 60 log-spaced points following the criteria used for

GRASP_{pac}.

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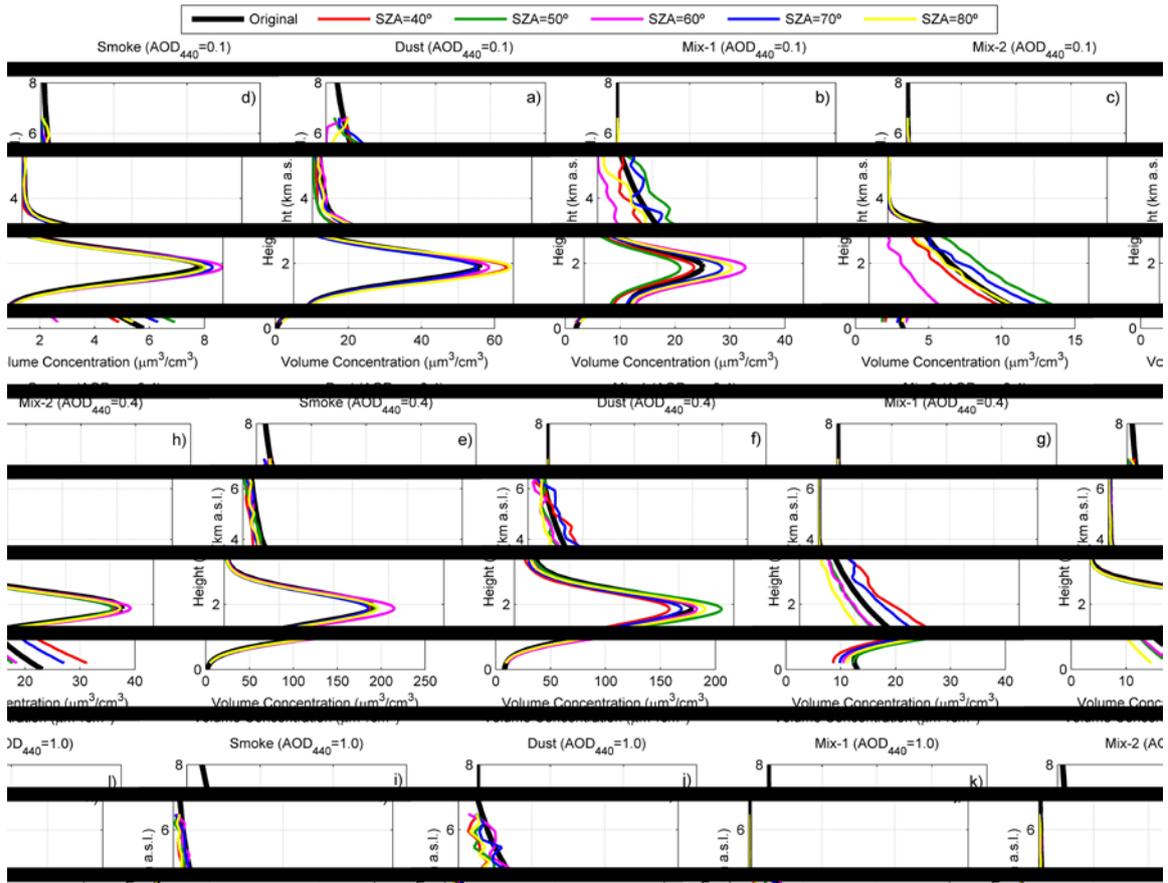


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1244 Figure 4: Original aerosol size distribution as retrieved by GRASP_{pac} for different
 1245 aerosol types (Smoke, Dust, Mix-1 and Mix-2) and loads ($AOD_{440}=0.1, 0.4$ and 1.0),
 1246 and at different solar zenith angles (SZAs) from 40° to 80° .

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1250 Figure 5: Original aerosol volume concentration (VC) vertical profile as retrieved by
 1251 GRASP_{pac} for different aerosol types (Smoke, Dust, Mix-1 and Mix-2) and loads
 1252 ($AOD_{440}=0.1, 0.4$ and 1.0), and at different solar zenith angles (SZA) from 40° to 80° .

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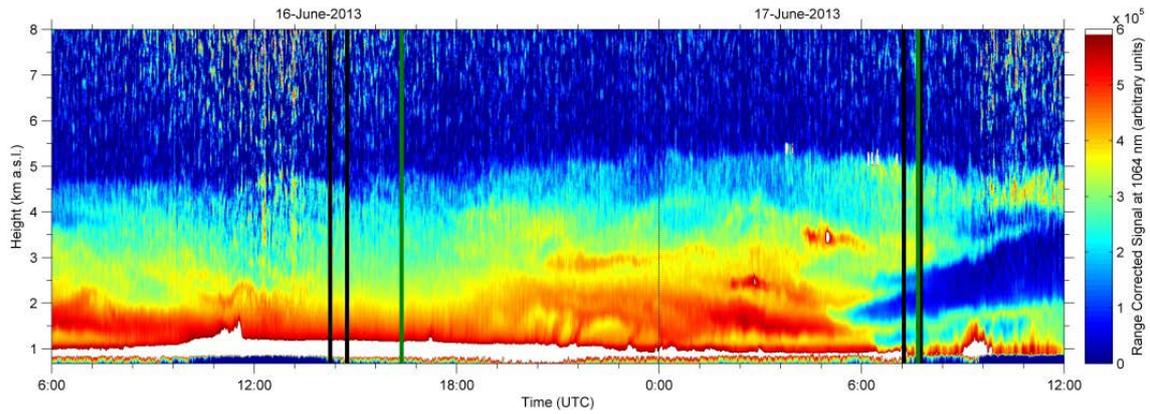
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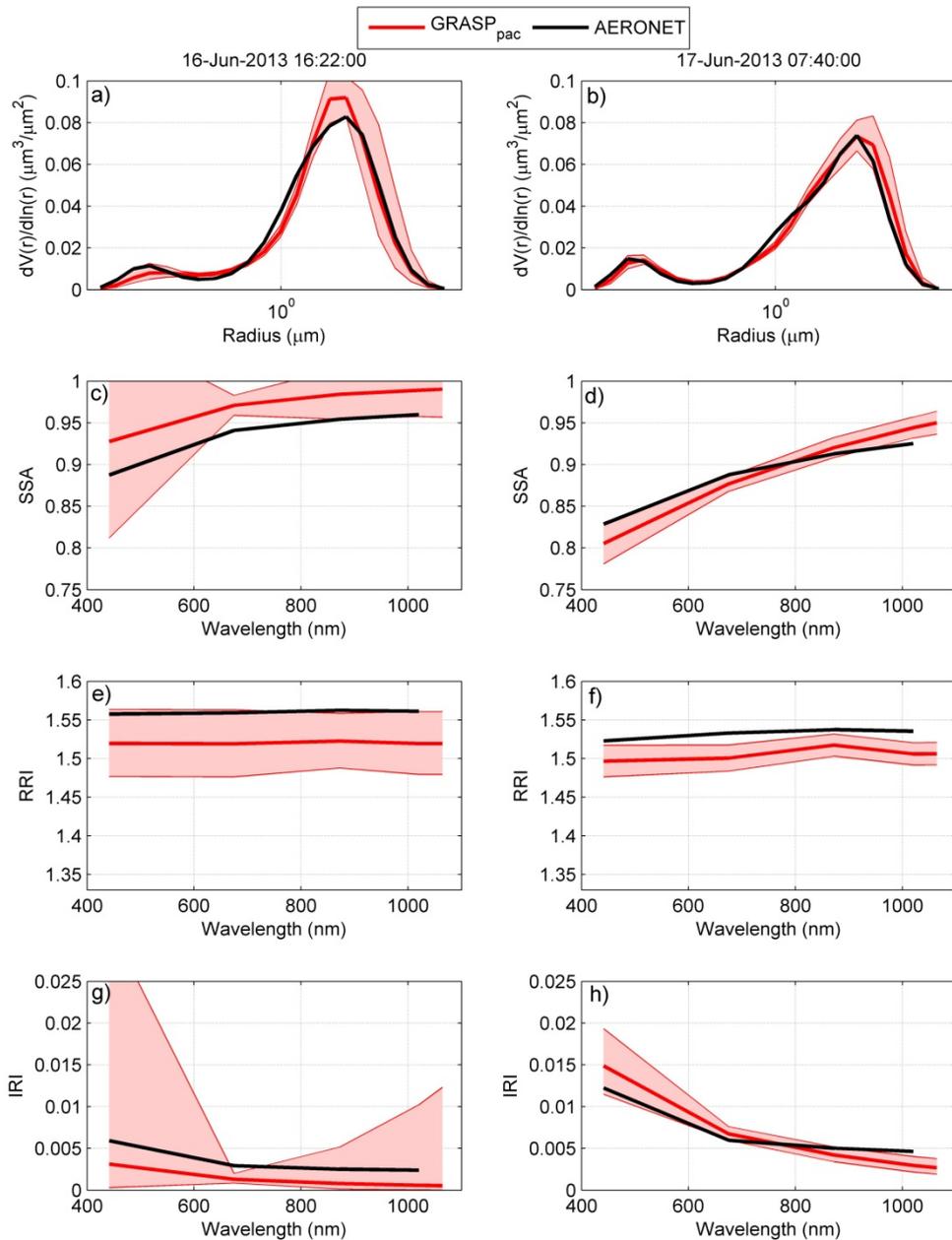
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1265 Figure 6: Ceilometer range corrected signal at 1064 nm as a function of height
 1266 from 16th, 6 UTC, to 17th June, 12 UTC, 2013. White colour represents all values above
 1267 6E5 arbitrary units. The times between vertical black lines corresponds to the F30 and
 1268 F31 flights. Green vertical lines corresponds in time with the sky radiance and AOD
 1269 measurements (sun photometer) nearest to the flights.

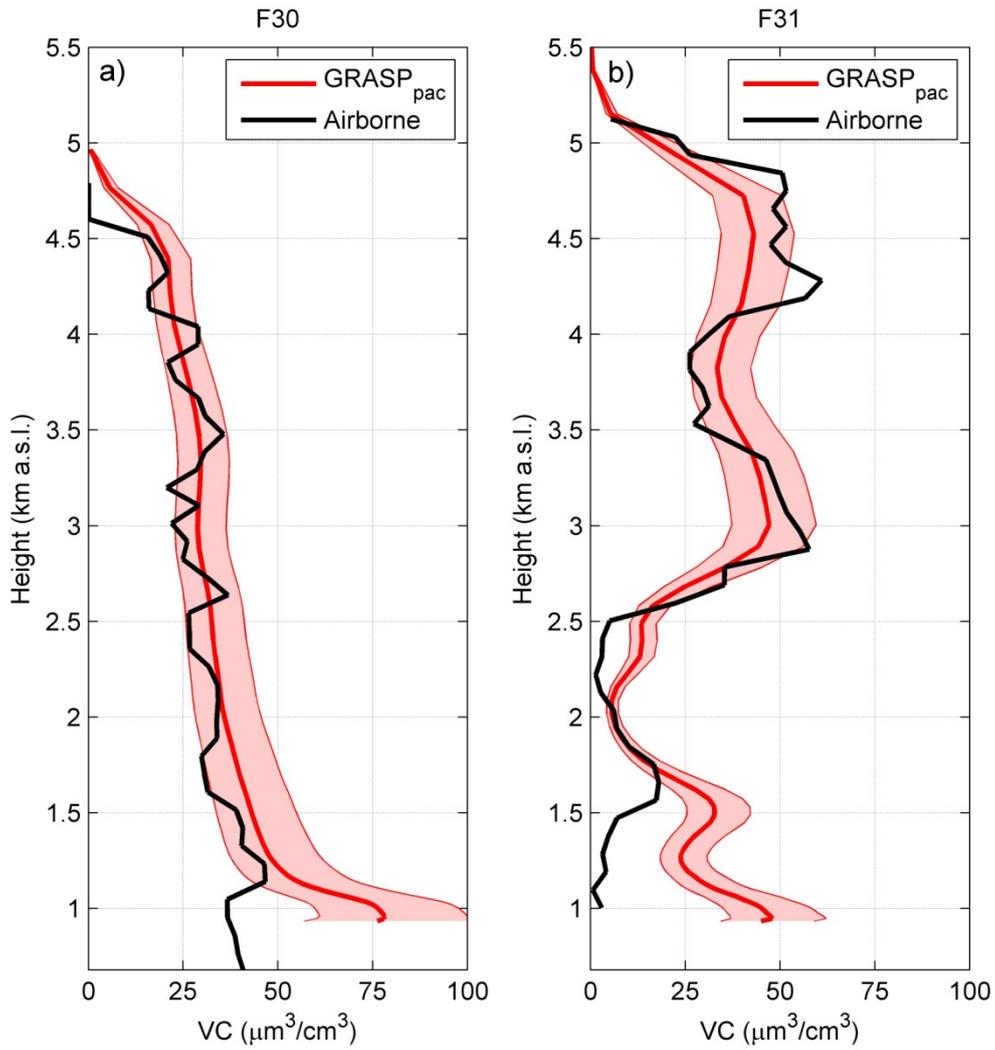


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1271 Figure 7: Columnar size distribution (panels a and b), single scattering albedo (SSA;
 1272 panels c and d), real refractive index (RRI, panels e and f) and imaginary refractive
 1273 index (IRI; panels g and h) obtained by AERONET (black line) and GRASP
 1274 at 16 June 2013 16:22 UTC (left panels) and 17 June 2013 07:40 UTC (right panels).

1275

Shadow band represents uncertainty in the GRASP_{pac} retrieval.



1276

1277 Figure 8: Profiles of aerosol volume concentration (VC) obtained by airborne
 1278 instrumentation (black line) and GRASP_{pac} (red line) at the flights F30 (panel a) and
 1279 F31 (panel b). Shadow band represents uncertainty in the GRASP_{pac} retrieval.

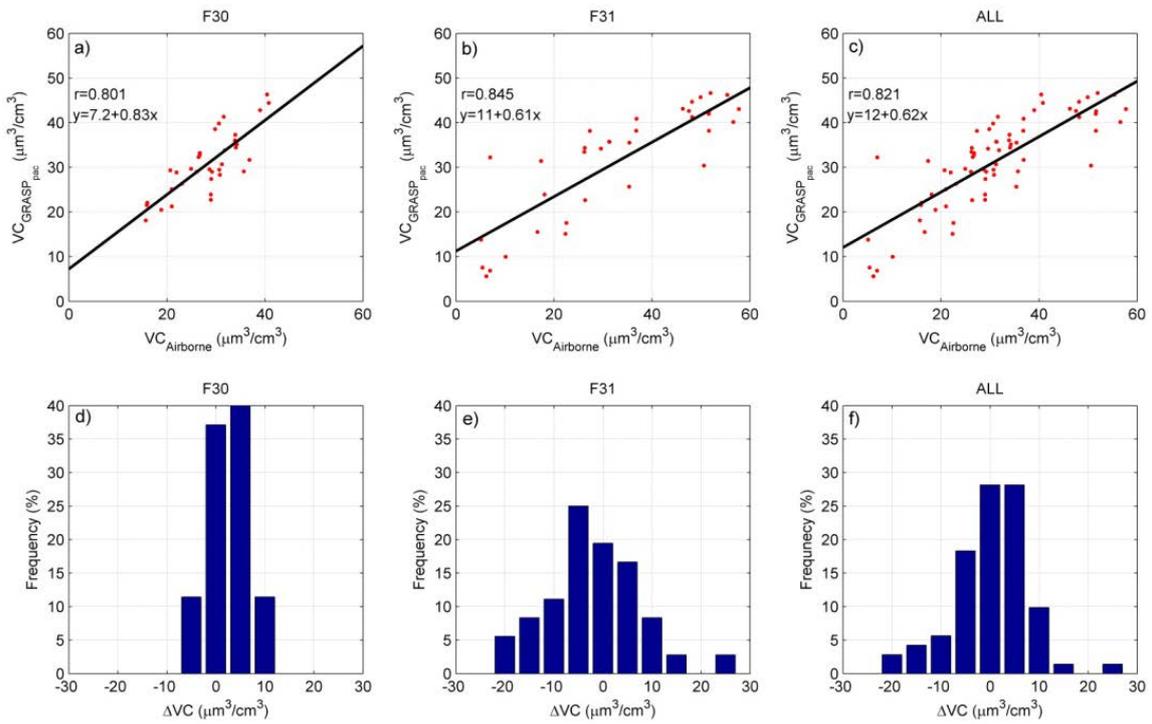
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1286 Figure 9: Aerosol volume concentration (VC) retrieved by GRASP_{pac} as a function of
 1287 the airborne measurements for the flights F30 (panel a), F31 (panel b) and all (panel c).

1288 Histograms of the differences between the VC retrieved by GRASP and the VC from
 1289 airborne (ΔVC from Eq. (10)) for the flights F30 (panel d), F31 (panel e) and all (panel
 1290 f).

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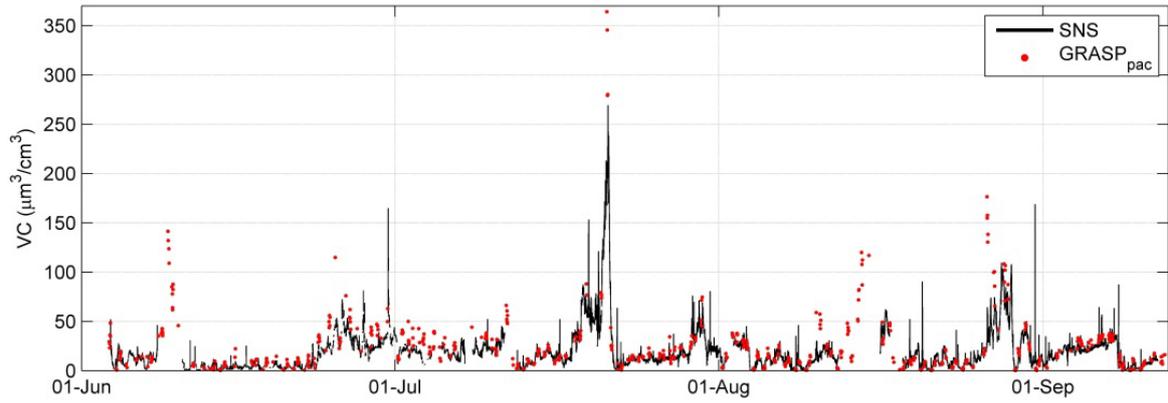
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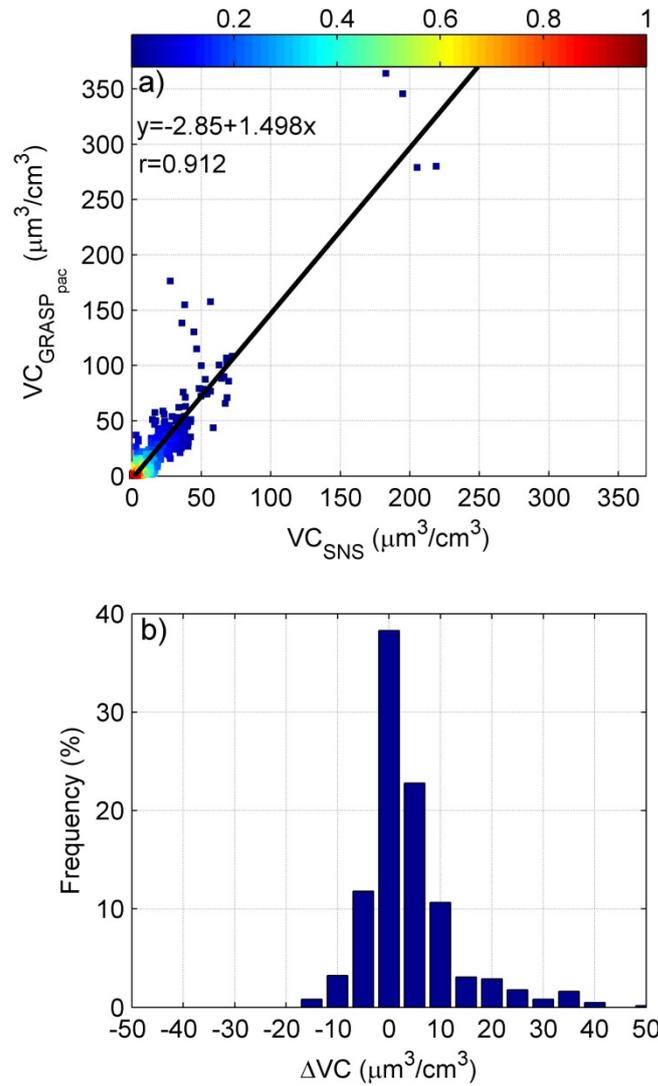


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1300 Figure 10: Temporal evolution of the aerosol volume concentration (VC) measured at
1301 the Sierra Nevada Station (SNS) and the retrieved by GRASP_{pac} at the same altitude.

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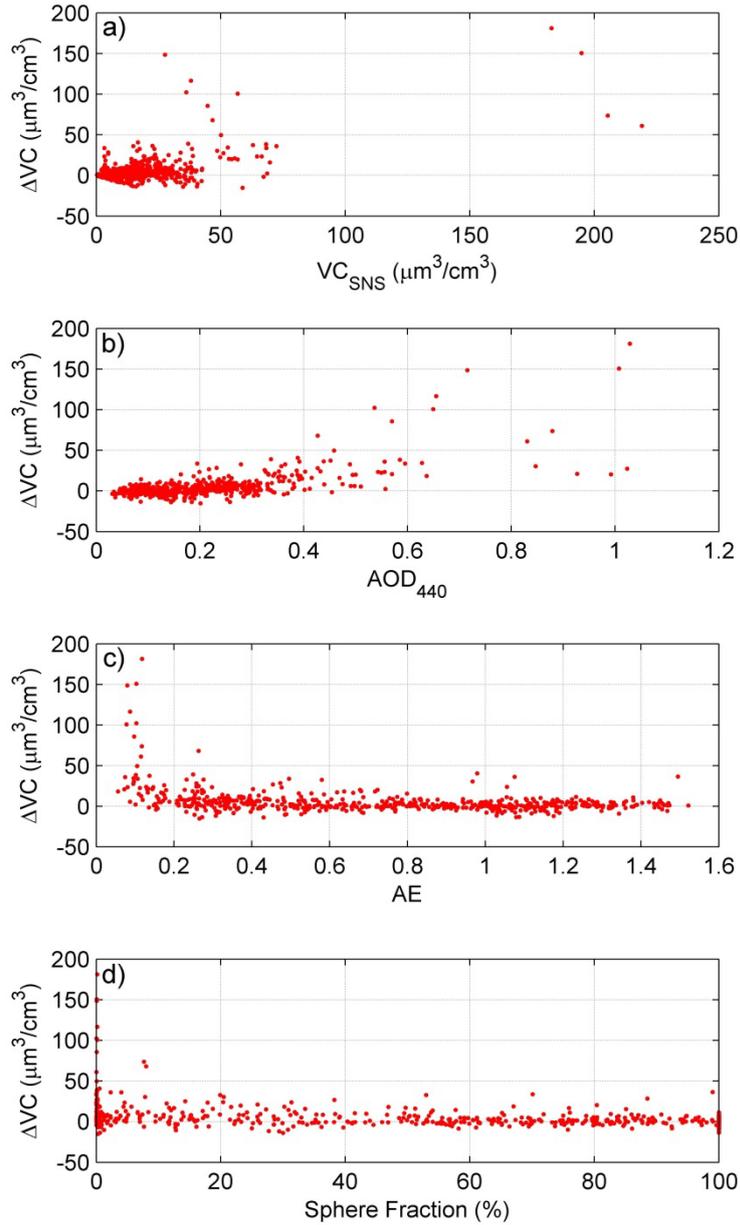
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1305 Figure 11: Aerosol volume concentration (VC) retrieved by GRASP_{pac} at the Sierra
 1306 Nevada Station (SNS) altitude as a function of the VC directly measured at SNS (panel
 1307 a). Colour of points represents the relative density of the points. Histograms of the
 1308 differences (ΔVC from Eq. (12)) between the VC retrieved by GRASP_{pac} at SNS
 1309 altitude and the VC directly measured at SNS (panel b).

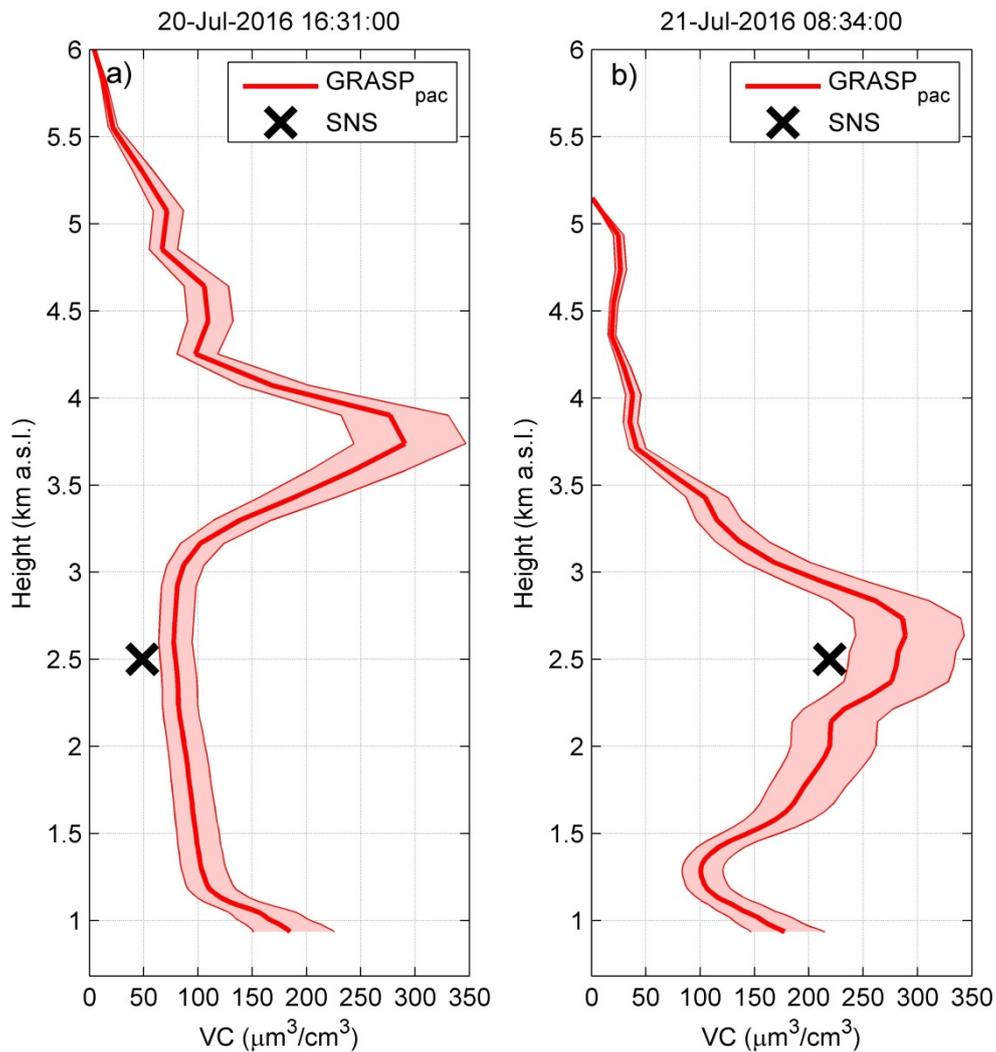
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1312 Figure 12: Differences between the aerosol volume concentration (VC) retrieved by
 1313 GRASP_{pac} at Sierra Nevada Station (SNS) altitude and the VC directly measured at SNS
 1314 (ΔVC from Eq. (12)) as a function of the VC at SNS (panel a), aerosol optical depth at
 1315 440 nm (panel b), Angström Exponent (panel c) and sphere fraction (panel d).

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1318 Figure 13: Profiles of aerosol volume concentration (VC) retrieved by GRASP_{pac} at 20
 1319 July 2016, 18:12 UTC (panel a) and 21 July 2016, 09:32 UTC (panel b). VC measured
 1320 at Sierra Nevada Station (SNS) is marked by a black cross. Shadow band represents
 1321 uncertainty in the GRASP_{pac} retrieval.

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1328 **Tables**

1329 Table 1: Mean bias error (MBE) of the GRASP_{pac} retrievals (Δ_{fit} from Eq. (6) and (7)) of
 1330 AOD, sky radiance, and lidar range-corrected signal (aod, rad, and rcs respective index)
 1331 under different aerosol scenarios. Standard deviation (STD) of Δ_{fit} is in parenthesis.

Aerosol Type*	SZA (°)	Scattering Angle Range (°)	AOD ₄₄₀ =0.1			AOD ₄₄₀ =0.4			AOD ₄₄₀ =1.0		
			MBE _{aod} (x1000)	MBE _{rad} (%)	MBE _{rcs} (%)	MBE _{aod} (x1000)	MBE _{rad} (%)	MBE _{rcs} (%)	MBE _{aod} (x1000)	MBE _{rad} (%)	MBE _{rcs} (%)
Smoke	40	2.3-78.6	5.3 (8.2)	-0.5 (3.0)	0.0 (1.3)	1.8 (3.0)	-0.1 (3.1)	0.2 (4.7)	-0.3 (0.4)	-0.2 (2.8)	0.1 (1.3)
	50	2.7-98.0	16.7 (19.7)	-0.3 (2.7)	0.4 (5.3)	-1.5 (3.0)	0.2 (2.9)	0.0 (0.4)	1.0 (1.5)	-0.5 (3.8)	0.2 (2.3)
	60	3.0-117.1	0.5 (0.7)	-0.1 (3.1)	0.1 (3.1)	0.4 (0.4)	-0.1 (2.8)	0.1 (1.6)	13.5 (17.4)	-1.7 (3.6)	-0.1 (2.7)
	70	3.3-135.5	3.8 (5.8)	-0.6 (3.0)	0.0 (2.0)	-1.1 (2.0)	-0.5 (3.6)	0.2 (3.4)	0.4 (0.2)	-0.1 (3.2)	0.1 (0.9)
	80	3.5-151.8	2.6 (3.7)	-0.6 (2.8)	0.1 (2.3)	1.7 (3.5)	-1.3 (3.0)	0.2 (0.7)	-0.1 (0.8)	-0.3 (3.3)	0.2 (2.5)
Dust	40	2.3-78.6	16.7 (30.4)	-2.9 (4.0)	7.2 (14.7)	-5.5 (8.6)	-1.1 (2.9)	0.5 (4.1)	9.4 (7.7)	-3.2 (4.0)	0.3 (2.6)
	50	2.7-98.0	3.2 (4.2)	-0.8 (3.0)	1.9 (5.4)	-5.1 (7.4)	0.8 (4.4)	0.1 (4.2)	11.6 (11.3)	1.5 (4.5)	1.3 (4.8)
	60	3.0-117.1	16.5 (31.3)	-0.2 (3.2)	0.7 (5.0)	-2.9 (4.7)	0.0 (3.3)	0.0 (2.3)	-3.0 (4.9)	1.1 (4.0)	0.5 (1.9)
	70	3.3-135.5	7.5 (14.3)	-0.6 (3.0)	0.3 (3.7)	-3.7 (6.5)	0.7 (3.4)	0.0 (1.7)	8.7 (6.2)	0.3 (3.6)	0.2 (2.9)
	80	3.5-151.8	-4.0 (7.5)	0.7 (3.2)	0.3 (1.1)	7.6 (10.3)	-0.3 (2.7)	0.2 (2.0)	3.6 (4.0)	0.2 (3.4)	0.2 (1.9)
Mix-1	40	2.3-78.6	4.1 (6.3)	-0.5 (2.8)	0.9 (3.1)	-3.0 (4.0)	-0.5 (3.4)	0.4 (1.7)	-5.9 (10.4)	0.1 (2.9)	-0.2 (1.9)
	50	2.7-98.0	-7.6 (13.5)	0.1 (3.6)	0.9 (2.3)	11.2 (16.7)	0.1 (3.3)	0.3 (6.8)	-0.4 (0.8)	0.2 (3.6)	-0.1 (2.8)
	60	3.0-117.1	-5.6 (7.6)	2.4 (5.6)	8.4 (11.8)	-2.1 (4.5)	2.3 (7.0)	0.0 (2.0)	-1.0 (1.3)	0.6 (3.1)	-0.2 (1.5)
	70	3.3-135.5	-2.8 (4.3)	0.4 (3.4)	1.2 (4.1)	-8.4 (16.4)	2.1 (4.7)	0.6 (2.7)	1.9 (3.0)	0.5 (3.4)	0.1 (1.1)
	80	3.5-151.8	-2.8 (8.0)	6.4 (8.0)	8.1 (11.1)	-2.2 (3.0)	1.7 (3.8)	-0.2 (1.6)	9.3 (12.5)	1.5 (4.6)	0.4 (3.9)
Mix-2	40	2.3-78.6	-3.3 (6.7)	-1.8 (3.5)	1.4 (5.1)	2.4 (3.7)	1.3 (5.1)	0.8 (8.3)	-0.1 (0.2)	-0.2 (3.3)	0.0 (1.1)
	50	2.7-98.0	-3.8 (6.0)	-2.0 (3.2)	1.1 (6.0)	0.7 (1.1)	-0.4 (3.1)	0.1 (2.0)	0.4 (1.2)	0.0 (2.9)	0.1 (5.4)
	60	3.0-117.1	1.1 (2.5)	-0.2 (3.4)	0.6 (3.6)	1.6 (2.6)	1.4 (4.9)	0.4 (3.2)	0.0 (0.4)	1.0 (3.5)	0.0 (0.9)
	70	3.3-135.5	1.9 (4.9)	-3.5 (4.4)	2.1 (9.7)	-0.4 (1.0)	-0.1 (2.6)	0.0 (1.0)	-0.8 (0.5)	0.1 (3.3)	0.0 (0.3)
	80	3.5-151.8	-0.7 (0.9)	-0.1 (2.5)	0.1 (2.6)	0.0 (1.1)	0.0 (3.5)	0.1 (1.4)	0.1 (0.1)	0.1 (2.4)	0.0 (1.4)

1332 *See section 4.1 for the aerosol models description.

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1336 Table 2: MBE and STD from the differences between the VC retrieved by GRASP_{pac}
1337 and the original VC (Δ_{vc} from Eq. (8) and (9)) under different aerosol scenarios and
1338 SZA values. Original VC values below $1 \mu\text{m}^3/\text{cm}^3$ have not been taken into account in
1339 the calculations. MBE and STD are given in % in parenthesis.

Aerosol Type*	SZA (°)	AOD ₄₄₀ =0.1		AOD ₄₄₀ =0.4		AOD ₄₄₀ =1.0		All	
		MBE ($\mu\text{m}^3/\text{cm}^3$)	STD ($\mu\text{m}^3/\text{cm}^3$)	MBE ($\mu\text{m}^3/\text{cm}^3$)	STD ($\mu\text{m}^3/\text{cm}^3$)	MBE ($\mu\text{m}^3/\text{cm}^3$)	STD ($\mu\text{m}^3/\text{cm}^3$)	MBE ($\mu\text{m}^3/\text{cm}^3$)	STD ($\mu\text{m}^3/\text{cm}^3$)
Smoke	40	-0.5 (-15.5)	0.1 (8.4)	6.0 (39.2)	3.2 (17.1)	-12.7 (-38.2)	4.8 (6.5)	-2.5 (-4.5)	8.5 (35.0)
	50	1.2 (33.6)	0.3 (5.9)	-2.5 (-18.6)	1.0 (5.7)	-8.7 (-26.4)	3.2 (6.9)	-3.5 (-4.9)	4.5 (27.1)
	60	-2.0 (-57.9)	0.5 (9.4)	-2.2 (-18.7)	0.7 (11.0)	-0.5 (-1.8)	0.8 (9.0)	-1.6 (-25.2)	1.0 (25.2)
	70	0.6 (14.9)	0.3 (9.8)	3.6 (24.8)	1.8 (13.3)	0.6 (-0.7)	1.4 (8.4)	1.6 (12.9)	2.0 (15.1)
	80	-0.1 (-6.1)	0.2 (11.8)	-4.6 (-33.6)	1.8 (5.8)	-8.9 (-27.5)	3.1 (6.9)	-4.7 (-22.9)	4.2 (14.4)
	All	-0.2 (-6.2)	1.1 (32.3)	0.0 (-1.4)	4.4 (30.4)	-6.0 (-18.9)	6.0 (16.8)	-2.1 (-8.9)	5.2 (28.2)
Dust	40	0.3 (-1.5)	1.7 (13.1)	-2.0 (-2.7)	6.7 (12.4)	-3.1 (-1.4)	14.0 (11.3)	-1.6 (-1.9)	9.2 (12.2)
	50	0.5 (3.6)	1.6 (10.6)	-3.1 (-5.2)	7.2 (13.8)	-9.0 (-5.7)	16.2 (11.6)	-4.0 (-2.6)	11.1 (12.7)
	60	2.2 (11.5)	2.7 (10.9)	3.6 (5.4)	6.9 (12.7)	-1.9 (-1.5)	14.3 (12.5)	1.3 (5.0)	9.6 (13.1)
	70	1.3 (5.8)	2.1 (11.0)	-0.5 (2.1)	6.5 (12.1)	-10.1 (-7.2)	15.5 (11.4)	-3.2 (0.1)	11.1 (12.7)
	80	1.1 (15.5)	1.7 (18.2)	-0.3 (-1.0)	5.7 (11.8)	-20.1 (-11.8)	25.0 (10.9)	-6.6 (0.6)	17.8 (17.8)
	All	1.1 (7.0)	2.1 (14.3)	-0.5 (-0.3)	6.9 (13.0)	-8.8 (-5.5)	18.5 (12.1)	-2.8 (0.3)	12.4 (14.1)
Mix-1	40	1.4 (1.8)	2.1 (18.1)	-1.1 (-8.9)	3.4 (14.0)	0.9 (-6.1)	7.1 (13.9)	0.3 (-4.8)	4.9 (15.8)
	50	-0.1 (-1.5)	0.9 (8.2)	0.1 (-6.8)	3.3 (12.5)	-0.7 (-7.0)	7.9 (12.5)	-0.3 (-5.3)	5.1 (11.6)
	60	0.6 (2.1)	1.1 (9.3)	6.1 (18.6)	7.3 (17.9)	-5.0 (-9.7)	9.7 (12.1)	0.5 (3.7)	8.6 (18.2)
	70	-0.2 (-3.1)	0.9 (8.6)	-1.1 (-1.1)	3.7 (13.1)	-3.0 (-10.6)	8.8 (14.2)	-1.5 (-5.1)	5.8 (13.0)
	80	2.0 (14.9)	2.1 (8.7)	1.0 (-2.7)	3.5 (13.0)	-2.3 (-10.7)	6.0 (14.8)	0.1 (-0.4)	4.7 (16.3)
	All	0.7 (2.9)	1.8 (12.8)	1.0 (-0.2)	5.2 (17.2)	-2.0 (-8.8)	8.2 (13.6)	-0.2 (-2.4)	6.0 (15.6)
Mix-2	40	-0.8 (-24.5)	0.2 (13.2)	-3.1 (-28.2)	1.1 (14.1)	-6.7 (-23.7)	2.7 (11.3)	-3.7 (-25.6)	3.0 (13.0)
	50	-1.2 (-32.2)	0.2 (10.9)	1.2 (-1.2)	2.0 (21.4)	-4.6 (-18.0)	1.9 (12.1)	-1.6 (-16.1)	3.0 (20.1)
	60	1.1 (21.9)	0.9 (12.3)	-0.9 (-12.5)	1.0 (12.8)	0.4 (-3.2)	3.3 (12.6)	0.2 (0.8)	2.3 (18.9)
	70	0.3 (2.7)	0.6 (10.2)	-1.9 (-18.5)	0.8 (12.4)	0.2 (-4.2)	3.3 (12.7)	-0.5 (-7.3)	2.3 (14.7)
	80	0.7 (11.1)	0.7 (12.6)	-0.2 (-8.6)	1.3 (14.8)	-10.1 (-32.0)	4.8 (9.5)	-3.5 (-11.2)	5.8 (21.4)
	All	0.0 (-4.2)	1.1 (23.9)	-1.0 (-13.8)	2.0 (17.8)	-4.2 (-16.2)	5.2 (16.1)	-1.8 (-11.9)	3.8 (19.9)
All	40	0.1 (-10.1)	1.6 (17.1)	0.0 (0.0)	5.4 (29.1)	-5.5 (-18.0)	9.4 (18.3)	-1.9 (-9.4)	7.0 (23.5)
	50	0.1 (1.9)	1.3 (25.4)	-1.0 (-8.1)	4.3 (15.8)	-5.6 (-14.6)	9.4 (13.8)	-2.3 (-7.3)	6.6 (19.8)
	60	0.4 (-7.3)	2.2 (33.6)	1.6 (-2.2)	6.0 (20.3)	-1.7 (-4.1)	8.7 (12.0)	0.1 (-4.4)	6.5 (23.2)
	70	0.5 (5.3)	1.3 (11.9)	0.0 (1.8)	4.3 (20.2)	-2.8 (-5.6)	9.6 (12.4)	-0.8 (0.3)	6.4 (16.1)
	80	0.9 (8.4)	1.6 (15.9)	-1.1 (-12.0)	4.0 (17.7)	-10.0 (-20.8)	13.9 (14.4)	-3.6 (-8.9)	9.8 (20.0)
	All	0.4 (-0.4)	1.6 (23.3)	-0.1 (-4.1)	4.9 (21.7)	-5.1 (-12.6)	10.7 (15.8)	-1.7 (-5.9)	7.5 (21.0)

1340 *See section 4.1 for the aerosol models description.

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1345 Table 3: Percentage of differences between the VC retrieved by GRASP_{pac} and the
1346 original VC (Δ_{vc} from Eq. (8)) that is below the uncertainty, σ_G , of VC given by
1347 GRASP_{pac}, for different aerosol scenarios and SZA values. The same percentage but for
1348 differences below $2\sigma_G$ is also shown.

Aerosol Type*	SZA (°)	AOD ₄₄₀ =0.1		AOD ₄₄₀ =0.4		AOD ₄₄₀ =1.0		All	
		$\Delta_{vc}<\sigma_G$ (%)	$\Delta_{vc}<2\sigma_G$ (%)						
Smoke	40	95.0	100.0	8.3	98.3	1.7	1.7	35.0	66.7
	50	90.0	100.0	88.3	98.3	1.7	85.0	60.0	94.4
	60	1.7	1.7	81.7	93.3	95.0	98.3	59.4	64.4
	70	96.7	100.0	66.7	98.3	93.3	100.0	85.6	99.4
	80	91.7	95.0	3.3	68.3	0.0	85.0	31.7	82.8
	All	75.0	79.3	49.7	91.3	38.3	74.0	54.3	81.6
Dust	40	93.3	100.0	91.7	100.0	93.3	100.0	92.8	100.0
	50	90.0	100.0	80.0	91.7	86.7	98.3	85.6	96.7
	60	86.7	90.0	83.3	95.0	86.7	96.7	85.6	93.9
	70	78.3	85.0	80.0	95.0	85.0	96.7	81.1	92.2
	80	63.3	83.3	86.7	98.3	80.0	90.0	76.7	90.6
	All	82.3	91.7	84.3	96.0	86.3	96.3	84.3	94.7
Mix-1	40	81.7	83.3	85.0	95.0	85.0	98.3	83.9	92.2
	50	100.0	100.0	90.0	100.0	81.7	96.7	90.6	98.9
	60	100.0	100.0	91.7	98.3	80.0	96.7	90.6	98.3
	70	91.7	100.0	88.3	100.0	75.0	93.3	85.0	97.8
	80	98.3	100.0	88.3	100.0	81.7	98.3	89.4	99.4
	All	94.3	96.7	88.7	98.7	80.7	96.7	87.9	97.3
Mix-2	40	55.0	95.0	38.3	80.0	36.7	80.0	43.3	85.0
	50	40.0	90.0	91.7	93.3	61.7	93.3	64.4	92.2
	60	78.3	95.0	88.3	95.0	95.0	98.3	87.2	96.1
	70	95.0	95.0	66.7	86.7	96.7	98.3	86.1	93.3
	80	96.7	98.3	83.3	91.7	1.7	58.3	60.6	82.8
	All	73.0	94.7	73.7	89.3	58.3	85.7	68.3	89.9
All	40	81.3	94.6	55.8	93.3	54.2	70.0	63.7	86.0
	50	80.0	97.5	87.5	95.8	57.9	93.3	75.1	95.6
	60	66.7	71.7	86.3	95.4	89.2	97.5	80.7	88.2
	70	90.4	95.0	75.4	95.0	87.5	97.1	84.4	95.7
	80	87.5	94.2	65.4	89.6	40.8	82.9	64.6	88.9
	All	81.2	90.6	74.1	93.8	65.9	88.2	73.7	90.9

1349 *See section 4.1 for the aerosol models description.

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1354 Table 4: Statistical estimators MBE, MABE and STD from ΔVC (Eq. (10)) for the
 1355 comparison of VC retrieved by GRASP_{pac} and the airborne measured for the F30, F31
 1356 and both flights together. Values within parentheses are in % (from Eq. (11)).

Flight	N	MBE ($\mu\text{m}^3/\text{cm}^3$)	MABE ($\mu\text{m}^3/\text{cm}^3$)	STD ($\mu\text{m}^3/\text{cm}^3$)	$\Delta VC < \sigma_G$ (%)	$\Delta VC < 2\sigma_G$ (%)
F30	35	2.5 (10.5)	4.1 (15.7)	4.4 (16.5)	77.1	100
F31	36	-1.8 (12.9)	7.5 (33.9)	9.5 (70.8)	58.3	88.9
All	71	0.3 (11.7)	5.8 (24.9)	7.7 (51.4)	67.6	94.4

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1380 Table 5: Statistical estimators for the comparison of VC retrieved by GRASP_{pac} and the
 1381 measured by in-situ instrumentation at SNS (ΔVC from Eq. (12)) along SLOPE I
 1382 campaign for different VC_{SNS} intervals. Values within parentheses are in % (from Eq.
 1383 (13)).

VC _{SNS} range	N	MBE ($\mu\text{m}^3/\text{cm}^3$)	MABE ($\mu\text{m}^3/\text{cm}^3$)	STD ($\mu\text{m}^3/\text{cm}^3$)	$\Delta VC < \sigma_G$ (%)	$\Delta VC < 2\sigma_G$ (%)
0-Max. $\mu\text{m}^3/\text{cm}^3$	619	5.5 (31.1)	7.7 (56.0)	16.6 (94.0)	37.6	68.2
5-Max. $\mu\text{m}^3/\text{cm}^3$	493	6.3 (22.7)	8.9 (43.2)	18.2 (58.9)	40.0	72.0
0-5 $\mu\text{m}^3/\text{cm}^3$	126	2.3 (64.0)	3.2 (106.4)	5.9 (169.4)	28.6	53.2
5-10 $\mu\text{m}^3/\text{cm}^3$	132	0.7 (10.0)	3.6 (50.0)	4.3 (59.4)	27.3	59.8
10-20 $\mu\text{m}^3/\text{cm}^3$	184	3.1 (19.8)	5.9 (39.9)	8.1 (53.1)	45.1	75.0
20-30 $\mu\text{m}^3/\text{cm}^3$	97	8.8 (36.4)	9.9 (40.7)	16.7 (63.4)	45.4	80.4
30-50 $\mu\text{m}^3/\text{cm}^3$	58	11.2 (28.5)	15.1 (39.5)	25.6 (65.7)	55.2	79.3
50-100 $\mu\text{m}^3/\text{cm}^3$	18	27.1 (46.1)	29.0 (49.2)	24.1 (42.9)	11.1	66.7
100-Max. $\mu\text{m}^3/\text{cm}^3$	4	116.7 (60.1)	116.7 (60.1)	58.6 (33.9)	0.0	50.0

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